

QFT for Cosmology, Achim Kempf, Lecture 7

Note Title

The driven harmonic oscillator cont'd:

D. Energy eigenstates

* Recall $\hat{H}(t) = \omega(a^+(t)a(t) + \frac{1}{2}) - \frac{1}{\sqrt{2\omega}}(a^+(t)+a(t))J(t)$

$$= \begin{cases} \omega(a_{in}^+ a_{in} + \frac{1}{2}) & \text{for } t < 0 \\ \text{something} & \text{for } 0 \leq t \leq T \\ \omega(a_{out}^+ a_{out} + \frac{1}{2}) & \text{for } T < t \end{cases}$$

(Q: How come that $\hat{H}(t=0) \neq \hat{H}(t>T)$?
 (A: We use the Heisenberg picture!)

Here, $a_{in} := a(0)$, $a_{out} := a(T)$ and $a_{out} = a_{in} + J_0$.

with: $J_0 := \frac{i}{\sqrt{2\omega}} \int_0^T J(t') e^{i\omega t'} dt'$

* For $t < 0$, we diagonalized the Hamiltonian

$$\hat{H}(t) = \omega(a_{in}^+ a_{in} + \frac{1}{2}) = \hat{H}_{in} = \text{const.}$$

by using $[a_{in}, a_{in}^+] = 1$ to construct its eigenvectors:

$$\hat{H}_{in} |n_{in}\rangle = E_n^{(in)} |n_{in}\rangle$$

Namely:

$$E_n^{(in)} = \omega(n + \frac{1}{2}), n = 0, 1, 2, 3 \dots$$

$$|n_{in}\rangle := \frac{1}{\sqrt{n!}} (a_{in}^+)^n |0_{in}\rangle$$

Note: The set $\{|n_{in}\rangle\}$ is a Hilbert basis of the Hilbert space \mathcal{K} .

* By $t > T$, the Hamiltonian has become a different operator:

$$\hat{H}(t) = \omega(a_{\text{out}}^+ a_{\text{out}} + \frac{1}{2}) = \hat{H}_{t>T} = \text{const.}$$

What are its eigenvectors $|n_{\text{out}}\rangle$ and eigenvalues E_m^{out} ?

Observation:

We have:

$$[a_{\text{out}}, a_{\text{out}}^+] = 1$$

\Rightarrow we can construct the eigenbasis of $H_{t>T}$

with the same method as the eigenbasis of H_{free} :

* There is a unique vector $|0_{\text{out}}\rangle \in \mathcal{H}$ obeying:

$$a_{\text{out}} |0_{\text{out}}\rangle = 0$$

* We define the set of vectors $\{|n_{\text{out}}\rangle\}$:

$$|n_{\text{out}}\rangle := \frac{1}{\sqrt{n!}} (a_{\text{out}}^+)^n |0_{\text{out}}\rangle$$

* Proposition:

$$\hat{H}_{t>T} |n_{\text{out}}\rangle = E_n^{(\text{out})} |n_{\text{out}}\rangle \text{ with } E_n^{(\text{out})} = \omega(n + \frac{1}{2}) = E_n^{(\text{in})}$$

The operators \hat{H}_{free} and $\hat{H}_{t>T}$ are different and have different eigenvectors: $|n_{\text{in}}\rangle$ and $|n_{\text{out}}\rangle$. Why are the eigenvalues the same? They both describe a free oscillator of frequency ω .

* Proposition:

The set $\{|n_{\text{out}}\rangle\}$ is a ON Hilbert basis of the Hilbert space \mathcal{H} .

How are the two bases related?

* Recall: Both, $\{|n_{\text{in}}\rangle\}$ and $\{|n_{\text{out}}\rangle\}$ are ON bases of \mathcal{H} .

\Rightarrow Each basis vector $|n_{\text{in}}\rangle$ is a linear combination of the basis vectors $\{|n_{\text{out}}\rangle\}$ and vice versa.

* Therefore, in particular:

There must exist coefficients $\Lambda_n \in \mathbb{C}$ so that:

$$|0_{\text{in}}\rangle = \sum_n \Lambda_n |n_{\text{out}}\rangle$$

↳ "Bogoliubov Transformation"

* Meaning of the Λ_n ?

▢ Recall: The system's state is frozen in state $|y\rangle = |0_{\text{in}}\rangle$.

▢ Assume we measure at a time $t > T$ the energy,
i.e., we measure

$$\hat{A}(t) = \omega(a_{\text{out}}^+ a_{\text{out}} + \frac{1}{2})$$

▢ What is the probability amplitude for finding the energy eigenvalue E_n ?

▢ Clearly:

$$\text{prob.amp.}(|n_{\text{out}}\rangle \text{ at } t > T) = \langle n_{\text{out}} | y \rangle$$

$$\text{i.e.:} \quad \text{prob.}(|n_{\text{out}}\rangle \text{ at } t > T) = |K_{n_{\text{out}}} |y\rangle|^2$$

Q Calculate:

$$\langle n_{\text{out}} | j(x) \rangle = \langle n_{\text{out}} | 0_{\text{in}} \rangle$$

$$= \langle n_{\text{out}} | \sum_m \Lambda_m | m_{\text{out}} \rangle$$

$$= \Lambda_n$$

\Rightarrow If the oscillator started in its ground state, then

at time $t > T$ the probability for finding the oscillator in its n^{th} excited state is given by:

$$\text{prob.} (|n_{\text{out}}\rangle \text{ at } t > T) = |\Lambda_n|^2$$

Remark: In QFT, this will be the prob. for finding n particles after charges and currents $j(x, t)$ excited the vacuum.

Calculation of Λ_n :

Proposition: $\Lambda_n = e^{-\frac{i}{2}|\mathcal{J}_0|^2} \frac{1}{\sqrt{n!}} \mathcal{J}_0^n$

Proof: The claim is that $|0_{\text{in}}\rangle = \sum_n e^{-\frac{i}{2}|\mathcal{J}_0|^2} \frac{1}{\sqrt{n!}} \mathcal{J}_0^n |n_{\text{out}}\rangle$.

We need to check that indeed: $a_{\text{in}} |0_{\text{in}}\rangle = 0$

Using $a_{\text{out}} = a_{\text{in}} + \mathcal{J}_0$, we need to check: $(a_{\text{out}} - \mathcal{J}_0) |0_{\text{in}}\rangle = 0$

Indeed:

$$(a_{\text{out}} - \mathcal{J}_0) \sum_n e^{-\frac{i}{2}|\mathcal{J}_0|^2} \frac{1}{\sqrt{n!}} \mathcal{J}_0^n |n_{\text{out}}\rangle \stackrel{\frac{1}{n!}}{=}$$

$$= e^{-\frac{i}{2}|\mathcal{J}_0|^2} (a_{\text{out}} - \mathcal{J}_0) \sum_n \mathcal{J}_0^n \underbrace{\frac{1}{\sqrt{n!}}}_{\frac{1}{\sqrt{n!}}} \underbrace{(a_{\text{out}}^+)^n}_{(a_{\text{out}}^+)^n} |0_{\text{out}}\rangle$$

$$= e^{-\frac{1}{2}|\beta_0|^2} (a_{out} - \beta_0) e^{\beta_0 a_{out}^\dagger} |0_{out}\rangle$$

$$= e^{-\frac{1}{2}|\beta_0|^2} \left(a_{out} e^{\beta_0 a_{out}^\dagger} - \beta_0 e^{\beta_0 a_{out}^\dagger} \right) |0_{out}\rangle$$

using $AB = [A, B] + BA$

$$= e^{-\frac{1}{2}|\beta_0|^2} \left(\underbrace{[a_{out}, e^{\beta_0 a_{out}^\dagger}]} + e^{\beta_0 a_{out}^\dagger} a_{out} - \beta_0 e^{\beta_0 a_{out}^\dagger} \right) |0_{out}\rangle$$

$$\stackrel{(*)}{=} e^{-\frac{1}{2}|\beta_0|^2} \left((\beta_0 - \beta_0) e^{\beta_0 a_{out}^\dagger} + e^{\beta_0 a_{out}^\dagger} a_{out} \right) |0_{out}\rangle = 0 \quad \checkmark$$

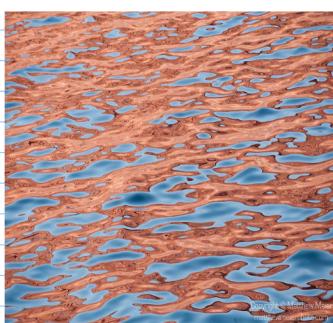
Note: In the last step, (*), we used that: $[a_{out}, e^{\beta_0 a_{out}^\dagger}] = \beta_0 e^{\beta_0 a_{out}^\dagger}$.

Exercise: Show that, more generally, $[\alpha, \alpha^\dagger] = 1$ implies $[\alpha, f(\alpha^\dagger)] = f'(\alpha^\dagger)$ by induction.

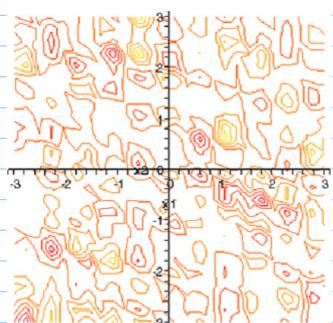
Hint: Show that: $[\alpha, \alpha^\dagger] = 1, [\alpha, \alpha^{\dagger 2}] = 2\alpha^\dagger, [\alpha, \alpha^3] = 3\alpha^2, \dots, [\alpha, \alpha^{\dagger n}] = n(\alpha^\dagger)^{n-1}$

Exercise: Verify that $|0_n\rangle = \sum_n e^{-\frac{1}{2}|\beta_n|^2} \frac{1}{\sqrt{n!}} \beta_n^\dagger |0_{out}\rangle$ obeys $\langle 0_m | 0_n \rangle = 1$.

Apply this strategy to the mode oscillators in QFT:



Making waves...



Making EM waves...

e^-

Recall:

$$\hat{H}(t) = \frac{1}{2} \int_{\mathbb{R}^3} \hat{\pi}^2(x, t) - \hat{\phi}(x, t)(\Delta - m^2) \hat{\phi}(x, t) + J(x, t) \hat{\phi}(x, t) d^3x$$

Example interpretation:

- * $\hat{\phi}(x, t)$ may be viewed as a slightly simplified version of the quantum electromagnetic field.
- * $J(x, t)$ may be viewed as a simplified version of a given classical electric charge and current density function.

Example:

A (Klein-Gordon) charge traveling a path $\tilde{x}^i(t)$:

$$\text{Then: } J(x, t) = q \delta(x - \tilde{x}(t))$$

In- and out periods

Let us consider the case where

$$J(x, t) = 0 \text{ for all } t \notin [0, T]$$

\Rightarrow It suffices to consider the periods $t < 0$ and $t > T$ in both of which $J(x, t) = 0$ (and then to relate the bases).

The free (i.e., undriven) QFT: $(t < 0 \text{ or } t > T)$

$$\hat{H}(t) = \frac{1}{2} \int_{\mathbb{R}^3} \hat{\pi}^2(x, t) - \hat{\phi}(x, t)(\Delta - m^2) \hat{\phi}(x, t) d^3x$$

* We need to solve:

$$\hat{\pi}(x,t) - (\Delta - m^2) \hat{\phi}(x,t) = 0$$

$$[\hat{\phi}(x,t), \hat{\pi}(x',t)] = i \delta^3(x-x')$$

* Fourier transformed,

$$\hat{\phi}_k(t) := \frac{1}{(2\pi)^3 \omega} \int_{\mathbb{R}^3} \hat{\phi}(x,t) e^{-ikx} d^3x$$

we need to solve:

$$\ddot{\hat{\phi}}_k(t) + \underbrace{(k^2 + m^2)}_{\text{Definition: } = \omega_k^2} \hat{\phi}_k(t) = 0 \quad (\text{EoM})$$

$$[\hat{\phi}_k(t), \hat{\pi}_k(t)] = i \delta^3(k+k') \quad (\text{CCRs})$$

* Recall: $\hat{\phi}^+(x,t) = \hat{\phi}(x,t)$ means $\hat{\phi}_k^+(t) = \hat{\phi}_k(t)$.

Solution strategy due to Fock:

* Proceed analogously to the driven oscillator, e.g., during $t < 0$:

□ Introduce new variables:

$$\text{QM: } a(t) := \sqrt{\frac{\omega}{2}} \hat{q}(t) + i \frac{1}{\sqrt{2\omega}} \hat{p}(t)$$

$$\text{QFT: } a_k(t) := \sqrt{\frac{\omega_k}{2}} \hat{\phi}_k(t) + i \frac{1}{\sqrt{2\omega_k}} \hat{\pi}_k(t)$$

□ Equation of motion and CCRs:

$$\text{QM: } \dot{a}(t) = -i\omega a(t) \quad [a(t), a^*(t)] = 1$$

$$\text{QFT: } \dot{a}_k(t) = -i\omega_k a_k(t) \quad [a_k(t), a_k^*(t)] = \delta^3(k-k')$$

□ Remark: Valid only while no force and while ω is constant.

□ Solution, using an initial condition:

$$QM: \quad a(t) = e^{-i\omega t} a_{in}, \quad [a_{in}, a_{in}^+] = 1$$

$$QFT: \quad a_k(t) = e^{-i\omega_k t} a_{in_k}, \quad [a_{in_k}, a_{in_k}^+] = \delta^3(k \cdot k)$$

□ Explicitly \Rightarrow

$$QM: \quad q(t) = \frac{1}{\sqrt{2}} \left(\frac{e^{-i\omega t}}{\sqrt{\omega}} a_{in} + \frac{e^{i\omega t}}{\sqrt{\omega}} a_{in}^+ \right)$$

$$QFT: \quad \hat{\phi}_k(t) = \frac{1}{\sqrt{2}} \left(\frac{e^{-i\omega_k t}}{\sqrt{\omega_k}} a_{in_k} + \frac{e^{i\omega_k t}}{\sqrt{\omega_k}} a_{in_k}^+ \right) \quad (S)$$

Exercise:
verify

$$\left(\text{i.e.: } \hat{\phi}(x, t) = \int \frac{1}{(2\pi)^{3/2}} \frac{1}{\sqrt{2\omega_k}} \left(a_{in_k} e^{-i\omega_k t + ikx} + a_{in_k}^+ e^{i\omega_k t - ikx} \right) dk \right)$$

The Hilbert space of states:

* Analogous to the case of QM, there is a vector, $|0_{in}\rangle \in \mathcal{H}$, which obeys:

$$a_{in_k} |0_{in}\rangle = 0, \quad \text{now for all vectors } k.$$

$$|0_{in}\rangle = \bigotimes_k |0_{in_k}\rangle$$

* The Hamiltonian reads (for $t < 0$):

$$\begin{aligned} \hat{H} &= \frac{1}{2} \int_{\mathbb{R}^3} \hat{p}_k^+ \hat{p}_k^- + \omega^2 \hat{\phi}_k^+ \hat{\phi}_k^- d^3 k \\ &= \int_{\mathbb{R}^3} \omega_k (a_{in_k}^+ a_{in_k}^- + \frac{1}{2} \delta^3(0)) d^3 k \end{aligned}$$

↓ This is called a "Infrared divergence"

In a box: $\hat{H} = L^{-3/2} \sum_k \omega_k (a_{in_k}^+ a_{in_k}^- + \frac{1}{2}) \quad (\text{because } \delta(k, k') \text{ is now } \delta_{k,k'})$

Notice: The divergence $\sum_k L^{-3/2} \omega_k \frac{1}{2} = \infty$ is an "ultraviolet divergence".

After the driving ends, $t > T$:

* One obtains: $a_n(x) = e^{-i\omega_n t} a_{\text{out},n}$ with $a_{\text{out},n} = a_{\text{in},n} + J_{\text{out},n}$

$$J_{\text{out},n} := \frac{i}{\sqrt{2\omega_n}} \int_0^T J_k(t') e^{i\omega_n t'} dt'$$

Here: $J_k(t)$ is the Fourier transform of $J(x, t)$.

* Construct the out-basis $\{|n_{\text{out},k}\rangle\}$ from:

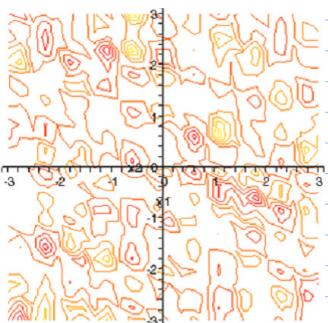
$$a_{\text{out},n} |0_{\text{out}}\rangle = 0$$

→ Can calculate, e.g., $|\langle n_{\text{out},k} | 0_{\text{in}} \rangle|^2$ i.e., the probability for $J(x, t)$ to have created n particles of momentum k .

Recall:



Making waves...



Making EM waves...

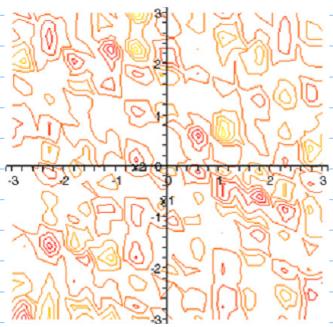


Described by $J(x, t)$.

Upgrade: Give the charge $j(x, t)$ its own dynamics:



A cork with
a spring.



Often, only one atomic transition is of interest.
Then, we can model the atom as a 2-level system.

Described by QM, e.g. atom or qubit

$$H(t) = \frac{1}{2} \int_{\mathbb{R}^3} \hat{\pi}^2(x, t) - \hat{\phi}(x, t)(\Delta - m^2) \hat{\phi}(x, t) + \hat{j}(x, t) \hat{\phi}(x, t) d^3x$$

number-valued, classical
↓

and upgrade it to:

$$H(t) = \frac{1}{2} \int_{\mathbb{R}^3} \mathbf{1} \otimes \left(\hat{\pi}^2(x, t) - \hat{\phi}(x, t)(\Delta - m^2) \hat{\phi}(x, t) \right) + \hat{j}(x, t) \otimes \hat{\phi}(x, t) d^3x$$

with $\hat{j}(x, t) = \hat{Q}(t) \delta(x - \tilde{x}(t))$

↑
an operator acting on the Hilbert space of the atom.

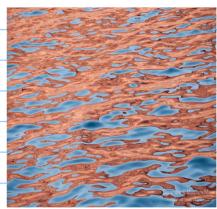
The Hilbert space: $\mathcal{H}_{total} = \mathcal{H}_{atom} \otimes \mathcal{H}_{field}$

Simplified notation:

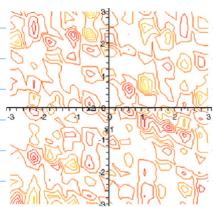
$$H(t) = \frac{1}{2} \int_{\mathbb{R}^3} \hat{\pi}^2(x, t) - \hat{\phi}(x, t)(\Delta - m^2) \hat{\phi}(x, t) + \hat{j}(x, t) \hat{\phi}(x, t) d^3x$$



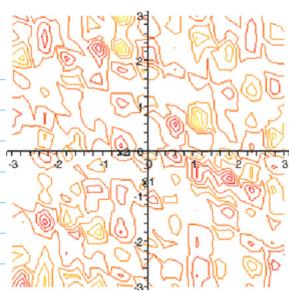
The charged systems can act as emitters and as receivers of waves:



And, quantumly, they can act as emitters and receivers of particles!



Definition: (Unruh, deWitt):
A "particle", such as a photon
is what a "particle detector", such
as an atom, can detect, by
getting excited.



With acceleration:

An accelerated atom's charges can excite the field.

This, in turn can excite the atom: the Unruh effect

→ An accelerated atom may detect particles
even if inertial observers only see the vacuum.

→ Related to gravity via the equivalence principle

→ Related to Hawking radiation.