$\begin{array}{c} {\rm Let's\ Chat\ About}\\ Quantum\ Mechanics \end{array}$

Andrew Zhen Ning Yong

Department of Physics & Astronomy
University of Edinburgh

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1 Preamble

First, something of an explanation and a disclaimer regarding this document.

This is my attempt at writing an informal discussion based on the content in the PHYS09053/09051 Quantum Mechanics course by C. Stock and A. Huxley. Needless to say, at best, this document should be read *in complement with* the lecture notes provided, rather than instead of.

One final thing before we start: I am prone to careless mistakes, especially where signs and numerical factors are concerned. So, if you notice anything suspicious with the expressions detailed below, please send me an email or send a pull request to my git repository.

Now let's chat about Quantum Mechanics.

2 An Electron and Its Proton Pal

2.1 Spin-Orbit Interaction

To be written up

3 An Amber of Electrons

In the above, we have described the going-ons of a hydrogen atom. Already with just a single proton and electron, we discover some pretty interesting phenomena that a lonely electron would not experience, eg the spin-orbit interaction. However, there is a whole periodic table of atoms, and all but save the hydrogen has multiple electrons. This leads us nicely to the physics of identical particles.

In quantum mechanics, the term 'identical' (or indistinguishable) has a specific meaning. That is, for a system with identical particles (eg an amber¹ of electrons), the measurable observables (eg the energy levels) must be invariant under the exchange of labels assigned to the particles. Let's put this in practice.

First, I will quote the version of Pauli's Exclusion Principle from [1] that is relevant to us:

A system containing several electrons must be described by an antisymmetric total eigenfunction (wavefunction).

Since electrons have spin, we have two possible combinations that make an overall antisymmetric wavefunction. Let ψ_i and χ_i be the *i*'th electron's spatial and spin wavefunction respectively, we have:

1. symmetric in space, antisymmetric in spin:

$$\Psi_A(x_1, x_2; s_1, s_2) = \frac{1}{2} \left(\psi_1(x_1) \psi_2(x_2) + \psi_1(x_2) \psi_2(x_1) \right) \left(\chi_1(s_1) \chi_2(s_2) - \chi_1(s_2) \chi_2(s_1) \right). \tag{3.1}$$

 ${f 2}.$ antisymmetric in space, symmetric in spin:

$$\Psi_A(x_1, x_2; s_1, s_2) = \frac{1}{\sqrt{2}} \left(\psi_1(x_1) \psi_2(x_2) - \psi_1(x_2) \psi_2(x_1) \right) \chi_1(s_1) \chi_2(s_1), \tag{3.2}$$

$$\Psi_A(x_1, x_2; s_1, s_2) = \frac{1}{2} \left(\psi_1(x_1) \psi_2(x_2) - \psi_1(x_2) \psi_2(x_1) \right) \left(\chi_1(s_1) \chi_2(s_2) + \chi_1(s_2) \chi_2(s_1) \right), \tag{3.3}$$

¹not actual collective noun

$$\Psi_A(x_1, x_2; s_1, s_2) = \frac{1}{\sqrt{2}} \left(\psi_1(x_1) \psi_2(x_2) - \psi_1(x_2) \psi_2(x_1) \right) \chi_1(s_2) \chi_2(s_2). \tag{3.4}$$

In the notation above, Equation 3.1 is what we call a *singlet* state and Equations 3.2-3.4 are called the *triplet* states. The subscript 'A' stands for antisymmetric. That is, it has the following properties

$$\Psi_A(x_1, x_2; s_1, s_2) = -\Psi_A(x_2, x_1; s_2, s_1). \tag{3.5}$$

However, as mentioned in the second paragraph, the observables are invariant under the exchange of the labels $\{x_i, s_i\}$. I encourage you to verify this by swapping the labels for a simple obserable, eg the probability density $|\Psi_A|^2$.

Here comes my favourite aspect of identical particles. Consider what happens to the probability density when two identical particles approach each other in the box. There are two scenarios again, which I write with a \pm below:

$$\lim_{x_1, x_2 \to x} |\Psi_A(x_1, x_2; s_1, s_2)|^2 = \frac{1}{2} \lim_{x_1, x_2 \to x} (\psi_1^*(x_1) \psi_2^*(x_2) \pm \psi_1^*(x_2) \psi_2^*(x_1)) \cdot (\psi_1(x_1) \psi_2(x_2) \pm \psi_1(x_2) \psi_2(x_1)),$$

$$= \frac{1}{2} \lim_{x_1, x_2 \to x} (|\psi_1(x_1)|^2 |\psi_2(x_2)|^2 + |\psi_1(x_2)|^2 |\psi_2(x_1)|^2 \pm (\psi_1^*(x_1) \psi_1(x_2) \psi_2^*(x_2) \psi_2(x_1) + \psi_1^*(x_2) \psi_1(x_1) \psi_2^*(x_1) \psi_2(x_2))),$$

$$= \frac{1}{2} \lim_{x_1, x_2 \to x} (|\psi_1(x_1)|^2 |\psi_2(x_2)|^2 + |\psi_1(x_2)|^2 |\psi_2(x_1)|^2 \pm (|\psi_1(x_1)|^2 |\psi_2(x_2)|^2 + |\psi_1(x_2)|^2 |\psi_2(x_1)|^2),$$

$$\pm (|\psi_1(x_1)|^2 |\psi_2(x_2)|^2 + |\psi_1(x_2)|^2 |\psi_2(x_1)|^2),$$

$$(3.6)$$

 2 I hope it is clear the \pm corresponds to the singlet and triplet states respectively. Now, consider the consequence. For the case of the singlet, we have a '+', and so the probability density is enhanced by the cross terms by a factor of 2. In the case of the triplet states, the cross terms reduce the probability density to zero as the electrons approach each other.

The interpretation is clear yet profound. When the spins are anti-aligned (singlet state), there exists an *attractive* force that enhances the probability density - so we can expect to **find electrons** with 'opposite' spins clumping together. On the other hand, if the two electrons have the same alignment, there exists a *repulsive* force such that the probability density vanishes as they coincide. This means, statistically, **electrons with the 'same' spin tend to stay apart**. Remember, we haven't introduced Coulomb interaction at this stage - so this attractive/repulsive force arises from the fact that these are identical particles!

That's enough chat about identical particles for now. In the following, let's look at two most obvious interactions we can think of for a system with two electrons.

3.1 Nuclear-Electron Coulomb Attraction

Simply put: we have a nucleus that has a net positive charge, and we have electrons that are negatively-charged - they will feel a Coulomb attraction from each other. We can quantify this interaction in the following Hamiltonian

²The keen learner may ask if it is reasonable to exclude the spin component in the equation above. Since the probability density appears in an integration over **spatial** variables, we can expect the spins to orthonormalise appropriately.

$$H_0 = \frac{Ze^2}{4\pi\epsilon_0 r_1} + \frac{Ze^2}{4\pi\epsilon_0 r_2},\tag{3.7}$$

where r_i corresponds to the i'th electron's distance from the nucleus.

A moment's glance and you will recognise each term to be just the usual Coulomb potential term. This time, we have Z=2 for a nucleus with two protons. We can write this nucleus-electron interaction as a sum (as opposed to a product) because the Coulomb attraction experienced by each electron is independent of the other. In other words, whatever force electron 1 feels due to the nucleus does not in any way affect what electron 2 experiences.

In physics lingo, we say this interaction is **separable**. That is, the interactions of each electrons can be neatly separated out into a sum(or products). In the following, we will meet an interaction which is not trivially separable.

Hamiltonian for the Nuclear-Electron Coulomb Interaction

$$H_0 = \frac{Ze^2}{4\pi\epsilon_0 r_1} + \frac{Ze^2}{4\pi\epsilon_0 r_2}. (3.8)$$

3.2 Electron-Electron Coulomb Repulsion

In a similar fashion to §3.1, the electrons can feel the charge of their fellow neighbours and, since they have identical charges, experience a Coulomb repulsion. We can write this interaction as

$$H_1 = \sum_{i>j} \frac{e^2}{4\pi\epsilon_0 |r_i - r_j|}, \quad (\geq 2 \text{ electrons})$$

$$= \frac{e^2}{4\pi\epsilon_0 |r_1 - r_2|} \quad (2 \text{ electrons}).$$
(3.9)

For more than two electrons we have a concise expression in the first line. The summation $\sum_{i>j}$ ensures that we do not double count the number of interaction terms since the Coulomb repulsion experienced by electron i due to electron j is exactly identical to that felt by electron j due to electron i.

Suppose we treat H_1 as a perturbation to the system described by the Hamiltonian in Equation 3.7. Let's see how the energy shifts in the following example cases. In the following, the notation $\langle \alpha_1 | \langle \beta_1 | | \alpha_2 \rangle | \beta_2 \rangle = \langle \alpha_1 | \alpha_2 \rangle \langle \beta_1 | \beta_2 \rangle$, ie the leftmost bra acts with the leftmost ket and so forth.

1. The Ground State

In the ground state, the spin of the electrons must be anti-aligned. So we use the singlet expression as in Equation 3.1. We can quantify the energy shift on the ground state (1s) energy as:

$$\Delta E = \langle 1s, 1s | \hat{H}_{1} | 1s, 1s \rangle,
= \frac{e^{2}}{4\pi\epsilon_{0}} \langle 1s, 1s | \frac{1}{|\hat{r} - \hat{r}'|} | 1s, 1s \rangle,
= \frac{e^{2}}{4\pi\epsilon_{0}} \cdot \frac{1}{2} \int d^{3}r_{1}d^{3}r_{2} (\langle 1s | r_{2} \rangle \langle 1s | r_{1} \rangle + \langle 1s | r_{2} \rangle \langle 1s | r_{1} \rangle) \frac{1}{|r_{1} - r_{2}|} (\langle r_{1} | 1s \rangle \langle r_{2} | 1s \rangle + \langle r_{1} | 1s \rangle \langle r_{2} | 1s \rangle),
= \frac{e^{2}}{4\pi\epsilon_{0}} \cdot \frac{1}{2} \int d^{3}r_{1}d^{3}r_{2} (\psi_{1s}^{*}(r_{2})\psi_{1s}^{*}(r_{1}) + \psi_{1s}^{*}(r_{2})\psi_{1s}^{*}(r_{1})) \frac{1}{|r_{1} - r_{2}|} (\psi_{1s}(r_{1})\psi_{1s}(r_{2}) + \psi_{1s}(r_{1})\psi_{1s}(r_{2})),
= \frac{e^{2}}{4\pi\epsilon_{0}} \cdot 2 \int d^{3}r_{1}d^{3}r_{2} \frac{|\psi_{1s}(r_{1})|^{2}|\psi_{1s}(r_{2})|^{2}}{|r_{1} - r_{2}|},$$
(3.10)

where in the third line I have used the completeness relation, $\mathbb{1} = \int d^3r \, |r\rangle \, \langle r|$ and the factor of 2 in the final line comes from both electron being in the same state n=1, L=0 state.

Now, you might ask: have I forgotten the spin part? To which I will answer: the interaction Hamiltonian, H_1 depends only on the separation distance between the electrons, *i.e* it doesn't care whether the particle has spin or not. Thus, the spin part just normalises to unity, which I encourage you to check.

To evaluate the integral, we use the hydrogen wavefunction with n = 1, L = 0 as we have shown in Problem Sheet 7.

2. First Excited State

In the first excited state, things become more interesting. As an electron is promoted to the n=2 state, it can enter either the s- or the p-shell, where L=0,1 respectively. Since both electrons are in different states, we are now confronted with the option of a singlet or triplet state again since the spins can be either parallel or anti-parallel. In the following, I will denote +/- on the spatial wavefunctions for the singlet and triplet states respectively.

2.1. Excited Electron in 2s-shell

Both electrons are in the s-shell. So, $l_1 = l_2 = 0 \Rightarrow L = 0$. Again, ignoring the spin component for this Hamiltonian:

$$\Delta E(L=0) = \langle 1s, 2s | \hat{H}_{1} | 1s, 2s \rangle,
= \frac{e^{2}}{4\pi\epsilon_{0}} \cdot \frac{1}{2} \int d^{3}r_{1}d^{3}r_{2} (\langle 1s | r_{2} \rangle \langle 2s | r_{1} \rangle \pm \langle 2s | r_{2} \rangle \langle 1s | r_{1} \rangle) \frac{1}{|r_{1} - r_{2}|} (\langle r_{1} | 1s \rangle \langle r_{2} | 2s \rangle \pm \langle r_{1} | 2s \rangle \langle r_{2} | 1s \rangle),
= \frac{e^{2}}{4\pi\epsilon_{0}} \cdot \frac{1}{2} \int d^{3}r_{1}d^{3}r_{2} (\psi_{1s}^{*}(r_{2})\psi_{2s}^{*}(r_{1}) + \psi_{2s}^{*}(r_{2})\psi_{1s}^{*}(r_{1})) \frac{1}{|r_{1} - r_{2}|} (\psi_{1s}(r_{1})\psi_{2s}(r_{2}) + \psi_{2s}(r_{1})\psi_{1s}(r_{2})),
= \frac{e^{2}}{4\pi\epsilon_{0}} \cdot \frac{1}{2} \int d^{3}r_{1}d^{3}r_{2} \underbrace{\psi_{1s}^{*}(r_{1})|^{2}|\psi_{2s}(r_{2})|^{2}}_{|r_{1} - r_{2}|} \pm \underbrace{\psi_{1s}^{*}(r_{2})\psi_{2s}(r_{2})\psi_{1s}(r_{1})\psi_{2s}^{*}(r_{1})}_{|r_{1} - r_{2}|}. \tag{3.11}$$

The first term, labelled the 'direct term', together with the e^2 out front, gives us the charge density. However, we pick up an additional term as well, which I have labelled the 'exchange term'. Once again, the very fact that we are dealing with identical particles has given rise to an enhancement/suppression of the charge density due to the singlet/triplet state.

So, you see, even though H_1 is 'blind' to the fact that electrons carry spin, the intrinsic quality of spin in electrons helps us to distinguish them when we observe the energy spectrum! 2.2. Excited Electron in 2p-shell

Now, suppose one of the electron is promoted to the 2p shell. So $l_1 = 0, l_2 = 1 \Rightarrow L = 1$. Algebraically, having an electron in the 2p-shell gives an energy shift of the same form, so we skip to the final line:

$$\Delta E(L=1) = \frac{e^2}{4\pi\epsilon_0} \cdot \frac{1}{2} \int d^3r_1 d^3r_2 \underbrace{\frac{\det(r_1)|^2 |\psi_{2p}(r_2)|^2}{|\psi_{1s}(r_1)|^2 |\psi_{2p}(r_2)|^2}}_{\text{direct term}} \underbrace{\psi_{1s}^*(r_2)\psi_{2p}(r_2)\psi_{1s}(r_1)\psi_{2p}^*(r_1)}_{\text{l}_1 - r_2}.$$
 (3.12)

What is the difference? Well, look at the exchange term. For each integration, it is a product of two different wavefunctions, $\psi_{1s}\psi_{2p}$. Since $|\psi_{1s}|^2$ overlaps more with $|\psi_{2s}|^2$ than $|\psi_{2p}|^2$ (see plot below), this convolution(aka product) means that the contribution from the exchange term is greater in the L=0 than L=1.

TODO:Put a plot of radial wavefunction

Hamiltonian for the Electron-Electron Coulomb Interaction

$$H_1 = \sum_{i>j} \frac{e^2}{4\pi\epsilon_0 |r_i - r_j|}. (3.13)$$

References

[1] R. Eisberg and R. Resnick, Quantum Physics of Atoms, Molecules, Solids, Nuclei, and Particles, 2nd Edition. 1985.