Bayesian Statistical Methods for Astronomy

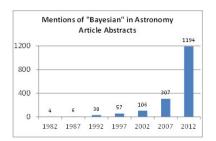
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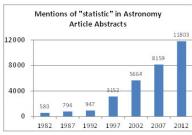
Statistics Section, Imperial College London

Stat4Astro — Autrans, France — October 2017

Bayesian Renaissance in Astronomy

The use of Statistical Methods in general and Bayesian Methods in particular is growing exponentially in Astronomy.





Source: http://magazine.amstat.org/blog/2013/12/01/science-policy-intel/

Why Use Bayesian Methods?

Advantages of Bayesian methods:

- Directly model complexities of sources and instruments.
- Allows science-driven modeling. (Not just predictive modeling.)
- Combine multiple information sources and/or data streams.
- Allow hierarchical or multi-level structures in data/models.
- Bayesian methods have clear mathematical foundations and can be used to derive principled statistical methods.
- Sophisticated computational methods available.

Challenges:

 Require us to specify "prior distributions" on unknown model parameters.

Outline of Topics

- BACKGROUND: Motivation; modern Bayesian tools; comparisons with likelihood methods; evaluating an estimator.
- BASIC MODELS: Poisson, binomial, and normal models; conjugate, informative, non-informative, and Jeffries prior distributions; summarizing posterior inference; the posterior as an average of the prior and data; nuisance parameters.
- MODEL FITTING: (Markov chain) Monte Carlo Methods, convergence detection, data augmentation
- MIERARCHICAL MODELS: Random-effects models and shrinkage; Multilevel models; Examples: selection effects, spectral and image analysis in high-energy astrophysics.
- MODEL CHECKING, SELECTION, AND IMPROVEMENT: Posterior predictive checks, Bayes factors, comparisons with significance tests and p-values.

Outline

- Foundations of Bayesian Data Analysis
 - Probability
 - Bayesian Analysis of Standard Poisson Model
 - Building Blocks of Modern Bayesian Analyses
- Further Topics with Univariate Parameter Models
 - Bayesian Analysis of Standard Binomial Model
 - Transformations
 - Prior Distributions
 - Final comments

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Probability

Bayesian Analysis of Standard Poisson Mode Building Blocks of Modern Bayesian Analyses

Defining Probability

What do we mean by:

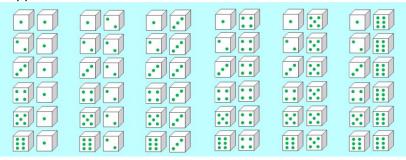
- Pr(Roll two dice and get doubles) =
- Pr(Rain today) =
- Pr(You will get at least a 75 in this class) =

How should we define "probability"?

Probability Bayesian Analysis of Standard Poisson Mode

Rolling Dice

Suppose we roll two dice:



• Let S be the set of possible outcomes.

Mathematical Definition of Probability

Definition

(Kolmogorov Axioms) A <u>probability function</u> is a function such that

- i) $Pr(A) \ge 0$, for all subsets of S.
- ii) Pr(S) = 1.
- iii) For any pair of disjoint subsets, A_1 and A_2 , of S, $Pr(A_1 \text{ or } A_2) = Pr(A_1) + Pr(A_2)^a$

But what does this this mean in real applications? How do we interpret a probability?

^a(Countable additivity) More generally, if $A_1, A_2, ...$ are pairwise disjoint subsets of S then $\Pr\left(\bigcup_{i=1}^{\infty} A_i\right) = \sum_{i=1}^{\infty} \Pr(A_i)$.

Defining Probability

What do we mean by:

- Pr(Roll two dice and get doubles) =
- Pr(Rain today) =
- Pr (You will get at least a 75 in this class) =

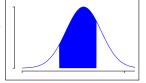
How should we define "probability"?

- Frequency-based definition.
- Subjective definition.
- Advantages and Difficulties of each.
- Is there a right or a wrong definition?

The Calculus of Probability

I assume you are familiar with:

- Probability density and mass functions, e.g.,
 - $\Pr(a < X < b) = \int_a^b p_X(x) dx$ or $\Pr(a \le X \le b) = \sum_{x=a}^b p_X(x)$
 - $\bullet \int_{-\infty}^{\infty} p(x) dx = 1$
- Joint probability functions, e.g.,
 - $Pr(a < X < b \text{ and } Y > c) = \int_a^b \int_c^\infty p_{XY}(x, y) dy dx$
 - $p_X(x) = \int_{-\infty}^{\infty} p_{XY}(x, y) dy$
- Conditional probability functions, e.g.,



When it is clear from context, we omit the subscripts: $p(x) = p_X(x)$.

Bayes Theorem

Bayes Theorem allows us to reverse a conditional probability:

Theorem

Bayes Theorem:

$$p_Y(y|x) = \frac{p_X(x|y)p_Y(y)}{p_X(x)} \propto p_X(x|y)p_Y(y)$$

 Bayes Theorem follows from applying the definition of conditional probability twice:

$$p_Y(y|x) = \frac{p_{XY}(x,y)}{p_X(x)} = \frac{p_X(x|y)p_Y(y)}{p_X(x)} \propto p_X(x|y)p_Y(y)$$

• The denominator does not depend on y and thus can be viewed as a normalizing constant. Advantage?

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A Poisson Model

Consider a Poisson model for a photon counting detector.

Simplest case: single-bin detector

$$Y \stackrel{\text{dist}}{\sim} \mathsf{Poisson}(\lambda_{\mathcal{S}}\tau).$$

(τ is the observation time in seconds and λ_S is expected counts/sec.)

• The sampling distribution is the probability function of data:

$$p_{Y}(y|\lambda_{S}) = \frac{e^{-\lambda_{S}\tau}(\lambda_{S}\tau)^{y}}{v!}.$$

Definition

The <u>likelihood function</u> is the sampling distribution viewed as a function of the parameter. Constant factors may be omitted.

The <u>maximum likelihood estimator</u> (MLE) is the value of the parameter that maximizes the likelihood.

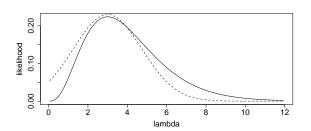
Likelihood for Poisson Model

Likelihood Function: For a single-bin detector,

$$\mathsf{likelihood}(\lambda_{\mathcal{S}}) = \frac{e^{-\lambda_{\mathcal{S}^{\mathcal{T}}}}(\lambda_{\mathcal{S}^{\mathcal{T}}})^{\mathcal{Y}}}{\mathcal{Y}!}$$

$$\mathsf{loglikelihood}(\lambda_{\mathcal{S}}) = -\lambda_{\mathcal{S}}\tau \! + \! y \, \mathsf{log}(\lambda_{\mathcal{S}})$$

Maximum Likelihood Estimation: Suppose y = 3 with $\tau = 1$



The likelihood and its normal approximation.

MLE:
$$\hat{\lambda}_{S} = \frac{y}{\tau}$$

Can estimate λ_S and its error bars.

Data-Appropriate Models and Methods

- Many methods based on χ^2 or Gaussian assumptions.
- Bayesian/Likelihood methods easily incorporate more appropriate distributions.
- E.g., for count data, we use a Poisson likelihood:

$$\chi^2$$
 fitting: $-\sum_{\text{bins}} \frac{(y_i - \lambda_i)^2}{\sigma_i^2}$

Gaussian Loglikelihood:
$$-\sum_{\text{bins}} \sigma_i - \sum_{\text{bins}} \frac{(y_i - \lambda_i)^2}{\sigma_i^2}$$

Poisson Loglikelihood:
$$-\sum_{\text{bins}} \lambda_i + \sum_{\text{bins}} y_i \log \lambda_i$$

A Prior Distribution for Poisson Model

Definition

The <u>prior distribution</u> quantifies knowledge regarding parameters obtained prior to the current observation.

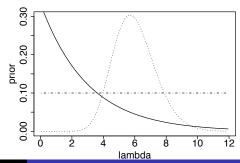
The gamma distribution is a flexible family of prior dist'ns:

$$p(\lambda_{\mathcal{S}}) = \frac{\beta^{\alpha}}{\Gamma(\alpha)} \lambda_{\mathcal{S}}^{\alpha-1} e^{-\beta \lambda_{\mathcal{S}}}$$

for $\lambda_S > 0$.

•
$$E(\lambda_S) = \alpha/\beta$$

•
$$Var(\lambda_S) = \alpha/\beta^2$$



The Posterior Distribution for Poisson Model

Definition

So:

The <u>posterior distribution</u> quantifies combined knowledge for parameters obtained prior to and with the current observation.

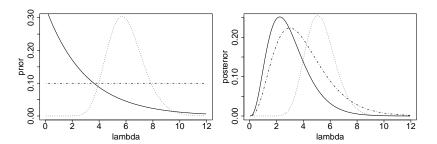
Bayes Theorem and the Posterior Distribution:

$$\begin{array}{lcl} \rho(\lambda_S|y) & = & \rho(y|\lambda_S)\rho(\lambda_S)/\rho(y) \\ \text{posterior}(\lambda_S|y) & \propto & \text{likelihood}(\lambda_S|y) \times \rho(\lambda_S) \\ & \propto & \frac{(\lambda_S\tau)^y e^{-\lambda_S\tau}}{y!} \times \frac{\beta^\alpha}{\Gamma(\alpha)} \lambda_S^{\alpha-1} e^{-\beta\lambda_S} \\ & \propto & \lambda_S^y e^{-\lambda_S\tau} \times \lambda_S^{\alpha-1} e^{-\beta\lambda_S} \\ & \propto & \lambda_S^{y+\alpha+1} e^{-(\tau+\beta)\lambda_S} \end{array}$$

 $\lambda_{S}|\mathbf{y} \sim \mathsf{GAMMA}(\mathbf{y} + \alpha, \beta + \tau)$

The Posterior Distribution for Poisson Model

The posterior dist'n combines past and current information:



Bayesian analyses rely on probability theory.

Summary: Bayesian Analysis of Poisson Model

Definition

If the prior and the posterior distributions are of the same family, the prior dist'n is called that likelihood's conjugate prior distribution.

$$\begin{array}{c} \text{If } Y|\lambda_{\mathcal{S}} \overset{\text{dist}}{\sim} \mathsf{POISSON}(\lambda_{\mathcal{S}}\tau) \text{ and } \lambda_{\mathcal{S}} \overset{\text{dist}}{\sim} \mathsf{GAMMA}(\alpha,\beta) \\ \text{then } \lambda_{\mathcal{S}}|Y \overset{\text{dist}}{\sim} \mathsf{GAMMA}(y+\alpha,\tau+\beta). \end{array}$$

- Conjugate prior distributions simplify computation!
- Using formulae for the Gamma distribution:
 - A Bayesian estimator of λ_S : $E(\lambda_S|y) = \frac{y+\alpha}{\tau+\beta}$
 - A Bayesian error bar: $\sqrt{\text{Var}(\lambda_S|Y)} = \frac{\sqrt{y+\alpha}}{\tau+\beta}$

"Prior Data"

Compare the MLE and the posterior expectation of λ_S :

$$\mathsf{MLE}(\lambda_{\mathcal{S}}) = \frac{\mathsf{y}}{\tau} \qquad \mathsf{E}(\lambda_{\mathcal{S}}|\mathsf{y}) = \frac{\mathsf{y} + \alpha}{\tau + \beta}$$

- The prior distribution has as much influence as α observed events in an exposure of β seconds.
- We can use this formulation of the prior in terms of "prior data" to
 - meaningfully specify the prior distribution for λ_S and
 - limit the influence of the prior distribution.

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Model Specification

- The first step in a Bayesian analysis is specifying the statistical model
- This consists of specification of
 - the prior distribution
 - the likelihood function
- Both of these involves subjective choices
 - Comprehensive description can be overly complex.
 - Parsimony: simple w/out compromising scientific objectives.
 - What is a model?
 - What do we model? Or consider fixed?
 (E.g., calibration, preprocessing, selection, etc.)

All models are wrong, but some are useful.

—George Box

Multilevel (and Hierarchical) Models

Example: Background contamination in a single bin detector

- Contaminated source counts: $y = y_S + y_B$
- Background counts: x
- Background exposure is 24 times source exposure.

A Poisson Multi-Level Model:

```
LEVEL 1: y|y_B, \lambda_S \stackrel{\text{dist}}{\sim} \text{Poisson}(\lambda_S) + y_B,
```

LEVEL 2:
$$y_B|\lambda_B \stackrel{\text{dist}}{\sim} \operatorname{Pois}(\lambda_B)$$
 and $x|\lambda_B \stackrel{\text{dist}}{\sim} \operatorname{Pois}(\lambda_B \cdot 24)$,

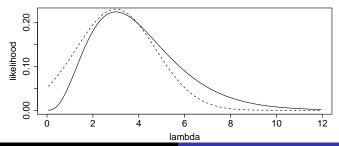
LEVEL 3: specify a prior distribution for
$$\lambda_B$$
, λ_S .

Each level of the model specifies a dist'n given unobserved quantities whose dist'ns are given in lower levels.

Bayesian Statistical Summaries

- The full statistical summary: the posterior distribution.
- ② But researchers would like summaries: A parameter estimate: The posterior mean. An error bar: The posterior standard deviation.

But is the enough??



Posterior Intervals or Regions

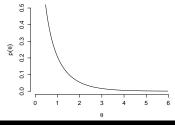
For non-Gaussian posterior distins, we find L and U so that

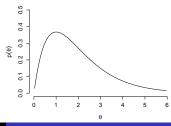
$$Pr(L < \theta < U|y) = \int_{L}^{U} p(\theta|y) d\theta = 68\% \text{ or } 95\% \text{ or } \dots$$

or more generally, Θ so that

$$\mathsf{Pr}(\theta \in \Theta|y) = \int_{\theta \in \Theta} p(\theta|y) d\theta = 68\% \ \mathsf{or} \ 95\% \ \mathsf{or} \ \dots$$

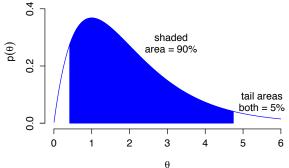
But the choice is not unique! Are there optimal choices?





Choice of Posterior Intervals

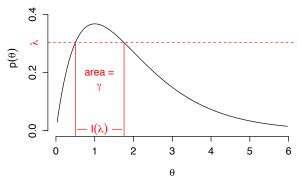
The Equal-Tailed Interval



- The simplest interval to compute (e.g., via Monte Carlo).
- Preserved under monotonic transformations.
 - E.g., If (L_{θ}, U_{θ}) is a 95% equal-tailed interval for θ , then $\Big(\log(L_{\theta}), \log(U_{\theta})\Big)$ is a 95% equal-tailed interval for $\log(\theta)$

Choice of Posterior Intervals (con't)

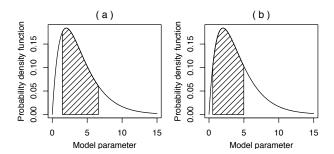
The Highest Posterior Density (HPD) Interval



- As λ decrease, probability (γ) of interval $(I(\lambda))$ increases.
- HPD interval is shortest interval of a given probability.

Choice of Posterior Intervals (con't)

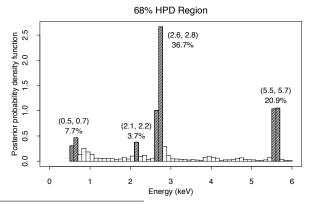
Equal-tailed and HPD intervals for a skewed gamma dist'n:



The difference is more pronounced for more extreme distributions!

Choice of Posterior Intervals (con't)

For a multimodal posterior, HPD may not be an interval! 1



¹See Park, van Dyk, and Siemiginowska (2008). Searching for Narrow Emission Lines in X-ray Spectra: Computation and Methods. *ApJ*, **688**, 807–825.

Benefits of Mathematical Foundation

EXAMPLE: The Posterior Odds.

$$\frac{p(\theta_1|y)}{p(\theta_2|y)} = \frac{p(y|\theta_1)p(\theta_1)/p(y)}{p(y|\theta_2)p(\theta_2)/p(y)} = \frac{p(y|\theta_1)}{p(y|\theta_2)} \times \frac{p(\theta_1)}{p(\theta_2)}$$

$$= \text{likelihood ratio} \times \text{prior odds}.$$

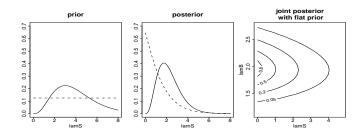
- Used to compare two parameter values of interest.
- @ Geneses of Bayesian methods for model comparison.
- No new methods required, just standard probability calculations.

Nuissance Parameters

Summarizing the posterior distribution:

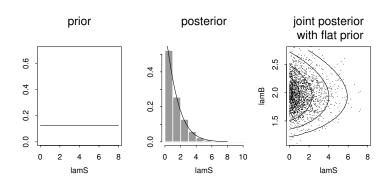
- We can plot the contours of the posterior distribution.
- Plot the marginal distributions of the parameters of interest:

$$p(\lambda_S \mid y, y_B) = \int p(\lambda_S, \lambda_B \mid y, y_B) d\lambda_B$$



Markov Chain Monte Carlo

Exploring the posterior distribution via Monte Carlo.



Easily generalizes to higher dimensions.

Bayesian Data Analysis: The Big Picture



- Statisticians: Model checking and model improvement.
- Scientists: Model comparison and model selection.

But remember....

All models are wrong, but some are useful.

—George Box

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Bayesian Analysis of Standard Binomial Model

EXAMPLE: Hardness Ratios in High Energy Astrophysics²

Let

- $H \sim POISSON(\lambda_H)$ be the observed hard count.
- $S \sim \mathsf{Poisson}(\lambda_S)$ be the observed soft count.
- n = H + S be the total count.

If *H* and *S* are independent,

$$H|n\sim \mathsf{BINOMIAL}\left(n,\pi=rac{\lambda_H}{\lambda_H+\lambda_S}
ight)$$

We will conduct a Bayesian Analysis of this model, treating π as the unknown parameter.

²For more on Bayesian analysis of Hardness Ratios see Park et al. (2006). Hardness Ratios with Poisson Errors: Modeling and Computations. *ApJ*, **652**, 610–628.

Details of Binomial Analysis

Likelihood:

$$p_H(h|\pi) = \frac{n!}{h!(n-h)!} \pi^h (1-\pi)^{n-h} \text{ for } h = 0, 1, \dots, n$$

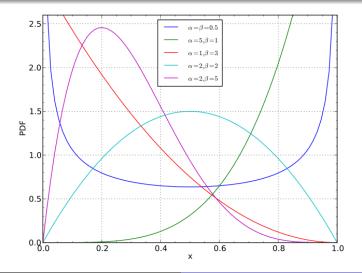
Beta prior distribution:

$$p(\pi) = \frac{\Gamma(\alpha + \beta)}{\Gamma(\alpha)\Gamma(\beta)} \pi^{\alpha - 1} (1 - \pi)^{\beta - 1} \text{ for } 0 < \pi < 1$$

where α and β are hyper parameters, which define prior dist'n.

The beta family is a flexible class of prior distributions on the unit interval.

Beta Distributions: A Flexible Class of Priors



Beta Dist'n is Conjugate to the Binomial

If
$$H|n,\pi \overset{\mathrm{dist}}{\sim} \mathsf{BINOMIAL}(n,\pi)$$
 and $\pi \overset{\mathrm{dist}}{\sim} \mathsf{BETA}(\alpha,\beta)$ then $\pi|H,n \overset{\mathrm{dist}}{\sim} \mathsf{BETA}(h+\alpha,n-h+\beta)$.

Suppressing the conditioning on *n*,

$$\rho(\pi|h) \propto \rho(h|\pi) \rho(\pi)$$

$$= \frac{n!}{h!(n-h)!} \pi^h (1-\pi)^{n-h} \times \frac{\Gamma(\alpha+\beta)}{\Gamma(\alpha)\Gamma(\beta)} \pi^{\alpha-1} (1-\pi)^{\beta-1}$$

$$\propto \pi^{h+\alpha-1} (1-\pi)^{n-h+\beta-1},$$

which is proportional to a Beta $(h + \alpha, n - h + \beta)$ density.

Beta Dist'n is Conjugate to the Binomial

If
$$H|n, \pi \overset{\text{dist}}{\sim} \mathsf{BINOMIAL}(n, \pi)$$
 and $\pi \overset{\text{dist}}{\sim} \mathsf{BETA}(\alpha, \beta)$ then $\pi|H, n \overset{\text{dist}}{\sim} \mathsf{BETA}(h + \alpha, n - h + \beta)$.

NOTE:

- The posterior distribution is an "average" of the data/likelihood and the prior distribution.
- We can interpret the hyperparameters α and β as "prior hard and soft counts".
- As n increases, choice of prior matters less.
- Point estimate for π :

$$E(\pi|h) = \frac{h+\alpha}{n+\alpha+\beta}$$

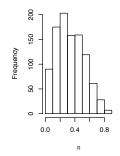
But be cautious of summarizing a dist'n with its mean!

Sample R code

```
# set (flat) prior
> alpha <- 1
> beta <- 1
>
> # set data
> hard <-1
> soft <- 3
>
  # Monte Carlo sample of posterior
> post.sample.pi <- rbeta(1000, hard + alpha, soft +beta)</pre>
>
> estimate <- mean(post.sample.pi)</pre>
> error.bar <- sd(post.sample.pi)</pre>
> lower <- sort(post.sample.pi)[25]</pre>
> upper <-sort(post.sample.pi)[975]</pre>
>
> hist(post.sample.pi, xlab =expression(pi), main="")
```

Sample R output

- > estimate
- 0.3237472
- > error.bar
- 0.1719679
- > lower
- 0.05146435
- > upper
- 0.6926952



Two 95% intervals

- estimate \pm 2× error bars: (-0.02, 0.66)
- equil-tail: (0.05, 0.69)

Why the difference?

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Parameterization of Hardness Ratio

We have formulated our analysis of Hardness ratios in terms of

$$\pi = \frac{\lambda_H}{\lambda_H + \lambda_S}.$$

Other formulations are more common:

simple ratio:
$$\mathcal{R} = \frac{\lambda_{\mathcal{S}}}{\lambda_{\mathcal{H}}} = \frac{1-\pi}{\pi}$$

color: $C = \log_{10}\left(\frac{\lambda_{\mathcal{S}}}{\lambda_{\mathcal{H}}}\right) = \log_{10}(1-\pi) - \log_{10}(\pi)$

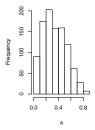
fractional difference: $\mathcal{H}\mathcal{R} = \frac{\lambda_{\mathcal{H}} - \lambda_{\mathcal{S}}}{\lambda_{\mathcal{H}} + \lambda_{\mathcal{S}}} = 2\pi - 1$

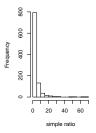
Transformations of scale and/or parameter are common.

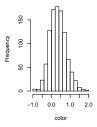
Parameterization of Hardness Ratio

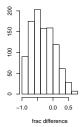
With an MC sample from posterior, transformations are trivial:

- # Monte Carlo sample of posterior of transformed parameters
- > post.sample.ratio <- (1-post.sample.pi)/post.sample.pi
- > post.sample.color <- log10(post.sample.ratio)</pre>
- > post.sample.diff <- 2*post.sample.pi 1



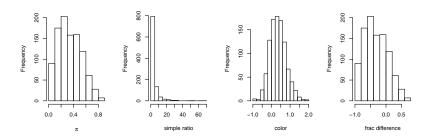






Frequency

Parameterization of Hardness Ratio



- How will the equal tail intervals compare with that for π ?
- How will the HPD intervals compare?
- How will the "estimate \pm 2× error bar" interval compare?
- What transformation is "best" from a stats perspective?

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Interpreting prior distributions

Using hardness ratios for illustration,

POPULATION/FREQUENCY INTERPRETATION: Imagine a population of sources, experiments, or universes from which the current parameter is draw.

"This source is drawn from a population of sources."

- STATE OF KNOWLEDGE: A subjective probability dist'n.
- **3** LACK OF KNOWLEDGE: UNIFORM(0, 1) corresponds to "no prior information". This choice of prior does draw $E(\pi|h)$ toward 1/2, but has relatively large prior variance.

We refer to "subjective" and "objective" Bayesian methods

Objective Bayesian Methods

Definition

A <u>reference prior</u> is a prior distribution than can be used as a matter of course under a given likelihood. That is, once the likelihood is specified the reference prior can be automatically applied.

Reference priors might be formulated to

- minimize the information conveyed by the prior, or
- optimize other statistical properties of estimators.

For example, we may find the prior that maximizes

 $Var(\theta|y)$ (for all y and/or choice of θ ??)

or yields confidence intervals with correct frequency coverage.

Non-informative Prior Distributions

Definition

A non-informative prior is a prior that aims to play a minimal role in the statistical inference.

Common choice: flat or uniform prior over range of parameter.

EXAMPLE: $h \mid \pi \sim \mathsf{BINOMIAL}(n,\pi)$ with $\pi \sim \mathsf{UNIFORM}(0,1)$. What does this choice of prior correspond to for:

simple ratio:
$$\mathcal{R} = \frac{\lambda_{\mathcal{S}}}{\lambda_{\mathcal{H}}} = \frac{1-\pi}{\pi}$$

color: $C = \log_{10}\left(\frac{\lambda_{\mathcal{S}}}{\lambda_{\mathcal{H}}}\right) = \log_{10}(1-\pi) - \log_{10}(\pi)$

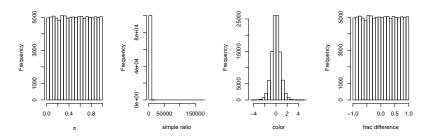
fractional difference: $\mathcal{H}\mathcal{R} = \frac{\lambda_{\mathcal{H}} - \lambda_{\mathcal{S}}}{\lambda_{\mathcal{H}} + \lambda_{\mathcal{S}}} = 2\pi - 1$

The Effect of Transformation on the Prior

R-code for an Monte Carlo study:

```
> prior.sample.pi <- runif(100000,0,1)</pre>
>
 # Monte Carlo sample of prior of transformed parameters
> prior.sample.ratio <- (1-prior.sample.pi)/prior.sample.pi</pre>
> prior.sample.color <- log10(prior.sample.ratio)</pre>
 prior.sample.diff <- 2*prior.sample.pi -1
 # Histograms
> pdf("hr-2.pdf", width=8, height=3)
> par(mfrow=c(1,4))
> hist(prior.sample.pi, xlab =expression(pi), main="")
> hist(prior.sample.ratio, xlab = "simple ratio", main="")
> hist(prior.sample.color, xlab = "color", main="")
> hist(prior.sample.diff, xlab = "frac difference", main="")
> dev.off()
```

Effect of Transformation on the Prior (cont)



- While the idea of a "flat prior dist'n" seem sensible enough, it is completely determined by the choice of parameter.
- Color is a standard normalizing transformation in stats.³
- Why not use flat prior on $\psi = \text{color}$: $p(\psi) \propto 1$ for $-\infty < \psi < \infty$?

³But statisticians call $\ln(\pi/(1-\pi))$ the log odds.

Improper Prior Distributions

Definition

An improper prior distribution is a positive-valued function that is not integrable, but that is used formally as a prior distribution.

NOTE:

- Because improper priors are not distributions, we can not rely on probability theory alone.
- However, improper priors generally cause no problem so long as we verify that the resulting posterior distribution is a proper distribution.
- If the posterior distribution is not proper, no sensible conclusions can be drawn.

Example of an Improper Prior Distribution

If
$$H|n, \pi \overset{\text{dist}}{\sim} \mathsf{BINOMIAL}(n, \pi)$$
 and $\pi \overset{\text{dist}}{\sim} \mathsf{BETA}(\alpha, \beta)$ then $\pi|H, n \overset{\text{dist}}{\sim} \mathsf{BETA}(h + \alpha, n - h + \beta)$.

The flat improper prior distribution on color:

$$p(\phi) \propto 1$$
 for $-\infty < \phi < \infty$

corresponds to the (improper) distribution on π

$$\pi \sim \textit{Beta}(\alpha = 0, \beta = 0).$$

The posterior distribution, however, is proper so long as

- 0 $h \ge 1$ and
- ② $n h \ge 1$.

Bayesian Analysis of Standard Binomial Model Transformations Prior Distributions Final comments

Jeffrey's Invariance Principle

Question: Can we find an objective rule for generating priors that does not depend on the choice of parameterization?

Definition

Jeffery's invariance principle says that any rule for determining a (non-informative) prior distribution should yield the same result if applied to a transformation of the parameter.

NOTE: Any subjective prior distribution should adhere to Jeffery's invariance principle. (At least in principle.)

Jeffrey's Prior Distribution

In likelihood-based statistics, the Expected Fisher Information is

$$J(\theta) = -\mathrm{E}\left[rac{\mathrm{d}^2\log p(y| heta)}{\mathrm{d}^2 heta} \mid heta
ight]$$

Definition

The Jeffery's prior distribution is

$$p(\theta) \propto \sqrt{J(\theta)}$$

or in higher dimensions,

$$p(\theta) \propto \sqrt{|J(\theta)|}$$
.

Example of Jeffrey's Prior

Example: For the binomial model,

$$\log(p_H(h|\pi)) = h\log(\pi) + (n-h)\log(1-\pi) + \text{ constant }.$$

and the expected Fisher information is

$$-\mathrm{E}\left[-\frac{h}{\pi^2}-\frac{n-h}{(1-\pi)^2}\mid\pi\right]=\frac{n}{\pi(1-\pi)}.$$

So the Jeffrey's Prior is

$$p(\pi) \propto \sqrt{J(\pi)} \propto \pi^{-1/2} (1-\pi)^{-1/2} = \mathsf{BETA}(\alpha = 1/2, \beta = 1/2).$$

This prior is invariant, but is it non-informative??

Bayesian Analysis of Standard Binomial Mode Transformations Prior Distributions Final comments

Outline

- Foundations of Bayesian Data Analysis
 - Probability
 - Bayesian Analysis of Standard Poisson Model
 - Building Blocks of Modern Bayesian Analyses
- Further Topics with Univariate Parameter Models
 - Bayesian Analysis of Standard Binomial Model
 - Transformations
 - Prior Distributions
 - Final comments

Choosing the Prior Distribution

Solance: Any reasonable prior distribution results in exactly the same asymptotic frequency properties as likelihood methods.

Worry: Only if you want to do better than likelihood-based methods in small samples.

Diligence: Nonetheless in practice much effort is put into selecting priors that help us best achieve our objectives.

Advantage: The choice of prior is an additional degree of freedom in methodological development.

Choice of prior can even improve frequency properties!

Subjective vs. Objective Analysis

All *statistical analyses are subjective*. Choices of data, parametric forms, statistical/scientifc models, "what to model".

But Bayesian methods have one more subjective component, the quantification of prior knowledge in through a distribution.

And prior distributions need't be used in subjective manner.

Everything follows from basic probability theory once we have established $p(y|\theta)$ and $p(\theta)$, Compare with likelihood theory.

Asymptotic results and counter intuitive definitions (e.g., for a CI or a p-value) *are not required*.