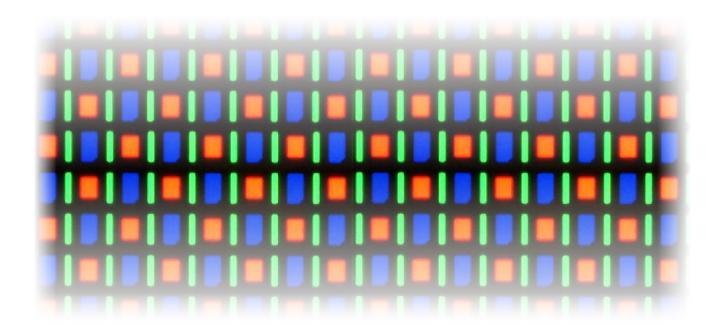


Thermally Activated Delayed Flourescence (TADF)

Thermally Activated Delayed Flourescence (TADF)

Given nowadays popularity of smartphones and many other potential applications, OLED materials and organic electronics are a very active field of research and industrial developments.



The following tutorial illustrates how advanced electroluminescence phenomena of OLED materials can be simulated and studied. After a more general introduction to the topic, the approach presented below is largely based on the following recent publications:

P.K. Samanta, D. Kim, V. Coropceanu, J.-L. Brédas *Up-Conversion Intersystem Crossing Rates in Organic Emitters for Thermally Activated Delayed Fluorescence: Impact of the Nature of Singlet vs Triplet Excited States*, J. Am. Chem. Soc. 139, 4042-4051 (2017).

A broader overview on the topic can be found for example in

Y.Olivier, B. Yurash, L. Muccioli, G. D'Avino, O. Mikhnenko, J. C. Sancho-García, C. Adachi, T.-Q. Nguyen, D. Beljonne *Nature of the singlet and triplet excitations mediating thermally activated delayed fluorescence*, Phys. Rev. Materials 1, 075602 (2017).

Note

This tutorial illustrates the modeling of TADF on the example of a realistic dye molecule. Expect several hours of computation time for each single-point evaluation.

General Remarks on Modelling OLED Emitters

Electronic Structure of OLED Materials

OLED devices typically consist of a thin layer of a host material which is doped with an organic (or metalorganic) emitter material. Applying an electric voltage across such semiconductor layers results in free electrons and electron holes, so-called charge carriers, travelling through the material. When these charge carriers recombine, strongly bound electron-hole pairs, so-called excitons are formed, which are entirely localized on a single emitter molecule and thereby correspond to electronic excitations.

As free electrons (and holes) are equally likely to be created with spin up or spin down configurations, 75% of the formed excitons emerge in a triplet state (total spin S=1) while 25% are found in a singlet state (S=0).

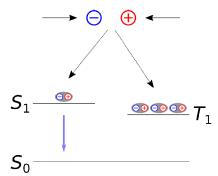


Fig. 11 Mechanism of first generation OLED emitters involving a flourescent decay path with a 25% maximum theoretical quantum yield.

This spin statistics negatively affects the quantum efficiency of an OLED material, hence on the proportion of electron-hole pairs that successfully undergo a radiative decay in to the ground state S_0 and emit a photon. A direct de-excitation from T_1 to S_0 is spin-forbidden in a non-relativistic theory. In the absence of any spin-orbit coupling effects in the dye, a radiative $T_1 \to S_0$ emission becomes very unlikely and light is only be emitted from a flourescent $S_1 \to S_0$ transition.

To address this problem a second generation of OLED emitters was developed involving OLED dyes containing heavy element atoms. Electrons in the valence shells of such atoms (e.g. Iridium)

https://www.scm.com/doc/Tutorials/Advanced/ADF ...

 $T_1 o S_0$ decay and emit light.

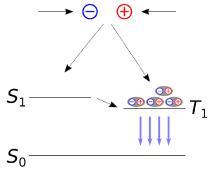


Fig. 12 Second generation OLED dyes rely on intersystem crossing to convert all excitons into a triplet state from which phosphorescent emissions occur.

While SOC-based emitters can reach quantum efficiencies of nearly 100%, the presence of long-living triplet states can cause chemical degradations. This tendency increases with the energy difference to the ground state (and thus with the frequency of the emitted light). As a result no stable, long-lasting OLED materials to deep blue light have been found as of yet. Apart from that, SOC-based dyes involve relatively rare and expensive heavy elements which significantly increases their costs.

These disadvantages motivated the development of third generation OLED materials: Dye molecules in which the states S_1 and T_1 are close in energy and (to some degree) vibrationally coupled, can exhibit a phenomenon known as thermally activated delayed flourescence (TADF). Using emitter molecules specifically tailored to maximize TADF, this effect can be exploited to avoid the degradation problems of SOC-based emitters. In addition, TADF occurs in purely organic dyes so that these third generation OLED emitters do not require expensive heavy metal elements.

TADF involves a reverse intersystem crossing (RISC) process, in which S_1 states are populated from T_1 levels. This is followed again by a normal flourescent transition $S_1 \to S_0$ and allows for light emissions with nearly 100% quantum yields.

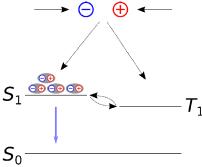


Fig. 13 The thermally assisted reverse intersystem crossing mechanism in third generation OLED

SCM

RISC competes with other processes such as the $S_1 \to T_1$ intersystem crossing. However, if non-radiative paths are negligible, $k_{\rm RISC}$ represents the rate-determining factor for TADF.

TADF represents a rather complex process and $k_{\rm RISC}$ has been shown to be strongly affected by various individual properties of a given dye material:

- singlet-triplet gap $\Delta E_{\mathrm{ST}} = E(S_1) E(T_1)$
- spin-orbit interactions
- vibrational couplings between different electronic states
- solvent effects

Each of these aspects will be addressed in the following.

Computational Description of TADF 1: Electronic Structure

A small singlet-triplet gap $\Delta E_{\rm ST}=E(S_1)-E(T_1)$ is essential for the RISC process. Indeed, the reverse intersystem crossing rate $k_{\rm RISC}$ is very sensitive regarding $\Delta E_{\rm ST}$, so that this quantity has to be predicted very precisely. Because of the required accurate electronic structure descriptions of excited states and their tiny energy differences, computational studies of TADF processes can be quite challenging for practical applications. As TADF dyes typically consist of hundreds of atoms, time-dependent density functional theory (TDDFT) often is the only computationally viable approach still accurate enough for modelling TADF processes.

To perform such TDDFT calculations, a suitable DFT approximation (i.e. exchange-correlation functional) needs to be chosen. Indeed, this is nontrivial question in itself. TDDFT results based on standard density functional approximations are inherently plagued by the self-interaction error. For the charge-transfer (CT) excitations typically found in OLED emitters such self-interaction effects cause spurious artifacts in the energies and electronic structure of excited states. While hybrid DFT methods ameliorate self-interaction errors, the excitation energies and electronic coupling parameters of TADF emitters are very sensitive regarding the amount of exact exchange incorporated in a given hybrid DFT approximation.

In range-separated hybrid (RSH) DFT methods the amount of exact exchange increases with the electron-electron distance. In many RSH approximations this used to restore the correct long-range asymptotic behavior of the corresponding potential. Their proper asymptotic behavior renders RSH methods potentially more accurate for the description of electronic excitations. The range-separation of RSH methods is controlled by and additional parameter ω , which corresponds to the inverse distance of the midpoint of the switching function between short-and long-range terms. The parameter ω can be tuned for a given emitter material to improve the



Excited States Geometry Optimizations

Following Samanta et al., we use TDDFT calculations in the Tamm-Dancoff approximation (TDA) to describe the TADF process in 1,2,3,5-tetrakis(carbazol-9-yl)-4,6-dicyanobenzene (4CzIPN), a typical TADF emitter.

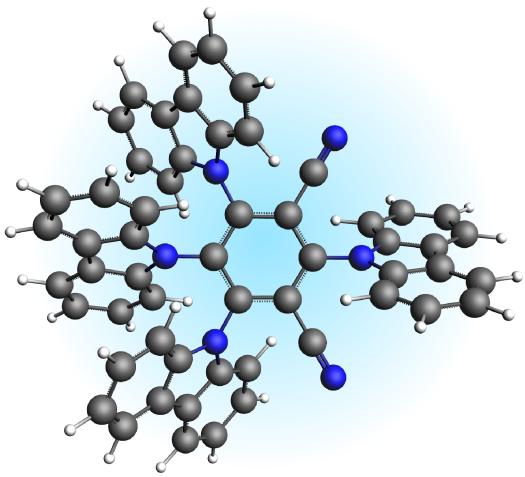


Fig. 14 1,2,3,5-tetrakis(carbazol-9-yl)-4,6-dicyanobenzene (4CzIPN)

Thereby we use the range-separated LCY-PBE functional, which employs a Yukawa potential to realize the switching function between long- and short-range terms. The sharpness of the Yukawa potential is controlled by the parameter γ , for which a value of 0.22 is used in the following.

Note

The value γ = 0.22 is the result of the RSH-tuning prescription of Karolewski et al. for the 4CzIPN molecule. Thereby the difference between the ionization potential and the HOMO energy is simultaneously minimized for a neutral and anionic system. Please follow the advanced tutorial tuning the range separation for range separated hybrids. Note, that γ is



To model the TADF process of 4CzIPN properly, the geometry of the system needs to be relaxed for all three relevant electronic states, S_0 , S_1 , and T_1 , respectively. We can do this from GUI with ADF with the following settings:

Download the 4CzIPN structure & here

Start ADFinput

Load the structure with File → Import Coordinates...

In the **Main** panel:

Task: → Geometry Optimization

XC-Functional \rightarrow Range Separated \rightarrow LCY PBE

Click on the ... button left of the XC-Functional field

Enter Gamma: 0.22

Return to Main panel

Relativity → Scalar

Basis set → TZP

Frozen core → None

Numerical quality → Good

Click on Model panel and select Solvation

Solvation method: → COSMO

COSMO solvent → Toluene

Save the input File \rightarrow Save as... under e.g. 4CzIPN GSopt.adf

A few more additional options are needed to define the excited states geometry optimizations. In the case of the S_1 -optimization:

In the window with the previously created *4CzIPN_GSopt.adf* file still open:

Properties → Excitations (UV/Vis), CD

Type of excitations: \rightarrow SingletOnly

TDA: → Yes

Number of excitations: \rightarrow 2

Calculate: → NTOs

Calculate: → Charge transfer descriptors

Properties → Excited State Geometry

Excitation: enter 1A



Analogously for the T_1 -state:

Properties → Excitations (UV/Vis), CD

Type of excitations: \rightarrow TripletOnly

TDA: → Yes

Number of excitations: \rightarrow 2

Calculate: → NTOs

Calculate: → Charge transfer descriptors

Properties → Excited State Geometry

Excitation: enter 1A Spin type → Triplet

Save the input File \rightarrow Save as... under e.g. 4CzIPN T1opt.adf

We are now ready to run all three optimization jobs. Note that, for a system like 4CzIPN a excited states geometry optimization with RSH can take multiple hours. For your convenience, the relaxed structures for the S_0, S_1 , and T_1 states, respectively as well as the corresponding input scripts can be downloaded here:

See also

- Land Table 1 Land Table 1 Land Table 2 Land Tab
- **Land Property of Structure** Landscape Acceleration Landscape Acce
- Land Table 4 Table 4
- **L** 4CzIPN GO GS.run run script ground state optimization
- $4CzIPN_{exGO_{S1.run}}$ run script S_1 optimization
- \angle 4CzIPN_exGO_T1.run run script T_1 optimization

After these optimizations are concluded, the respective minimum energies of the states can be extracted at the end of the logfile:

```
<Feb15-2019> <23:02:54> Converged.
                                         -33.75755925 a.u.
<Feb15-2019> <23:03:11>
                      Bond Energy
<Feb15-2019> <23:03:11> Bond Energy
                                        -918.58992548 eV
<Feb15-2019> <23:03:11> Bond Energy
                                      -21183.19 kcal/mol
<Feb15-2019> <23:03:11> Excited state energy (ground + excitation energy) for: 1A(Singlet-Triplet)
-33.67699976 a.u.
<Feb15-2019> <23:03:11>
                       Excited State Energy:
                                                             -916.39779022 eV
                                                                         kcal/mol
<Feb15-2019> <23:03:11>
                      Excited State Energy:
                                                           -21132.64
<Feb15-2019> <23:03:12> NORMAL TERMINATION
<Feb15-2019> <23:03:14> END
```



$E_{ m S1}-E_{ m S0}$	$E_{ m T1}-E_{ m S0}$	$\Delta E_{ m ST} = E_{ m S1} - E_{ m T1}$
2.55 eV	2.39 eV	0.16 eV

The result is somewhat different from the results ($\Delta E_{\rm ST}$ = 0.01 eV) of the original paper due to the different electronic structure description used therein. Compared to the experimental value of $\Delta E_{\rm ST}$ = 0.1 eV, the difference is however smaller and we will use our computed results of 0.16 eV in the subsequent calculations of this tutorial.

The output file provides the charge transfer descriptors according to the prescription of Plasser and Lischka, which are obtained from transition density matrices.

```
0.0002
1: 19/a
              -> 213a
1: 204a
                              0.0002
             -> 207a
1: 205a
             -> 242a
                              0.0002
    NTO weight: 0.9941
      occ: 0.9755 205a:1
                                  -0.1434 188a:1
                                                        -0.1317 193a:1
     virt: -0.9996 206a:1
                                  0.0187 245a:1
                                                        -0.0160 221a:1
    Descriptors (based on Plasser, Lischka, et al.)
      CT AT =
                 0.8969 R HE =
                                        4.6809
    Descriptors (Peach, Tozer, et al.)
     LAMBDA =
                  0.3919
                             R HE =
                                        1.6717
   Q
```

For the T_1 state we find a charge transfer (CT) character of 90% with an average electron-electron hole distance (R_HE) of 4.7 Å, while in the case of S_1 the CT character amounts to 94% with an average of 5 Å between the excited electron and electron hole. Note, that such high CT characters and electron-hole distances are a direct consequence of the design principles of typical OLED emitters featuring well-separated electron donor and acceptor groups. In consequence, these excited states also exhibit a relatively small exchange energy, which contributes most to the small $\Delta E_{\rm ST}$ value computed above from TDDFT excitation energies.

Vertical Absorption

Following Samanta et al. and to analyze the system further, we compute vertical excitations at the ground state geometry to obtain the absorption spectrum of 4CzIPN. Note, that these results do not enter the calculation of the RISC rate but rather serve as an additional assessment of the accuracy of the description employed here. Regarding most aspects this is done analogously to the previous calculations. The main differences are that we now treat the excitations in a spin-orbit description which allows for pure spin-states (i.e. no spin contamination). Furthermore, the response of the solvation model is ignored as the solvent molecules cannot rearrange within the short timespan of a vertical excitation; a better



Open a new instance of ADFinput

Load the previously optimized GS structure with File → Import Coordinates...

In the Main panel:

Task: → Single Point

XC-Functional → Range Separated → LCY PBE

Click on the ... button left of the XC-Functional field

Enter Gamma: 0.22

Return to **Main** panel

Relativity → Scalar

Basis set → TZP

Frozen core → None

Numerical quality → Good

Click on Model panel and select Solvation

Solvation method: → COSMO

COSMO solvent → Toluene

Click on **Details** panel and select **COSMO**

Untick Include response in TDDFT

Click on Properties and select Excitations (UV/Vis), CD

Type of excitations: \rightarrow Spin-Orbit (Perturbative)

TDA: \rightarrow Yes

Number of excitations: 5

Calculate: → NTOs

Calculate: → Charge transfer descriptors

Save the input **File** → **Save as...** under e.g. 4CzIPN_GS_VA.adf

See also

• 🚣 4CzIPN_VA.run run script vertical absorption calculation

After completing this calculation, we open the results with ADFoutput and search for the results with Response Properties → All Spin-Orbit Coupling Excitation Energies:

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Free trial

2:	0.09372	2.55025	0.7248E-09	4.889	Α
3:	0.09372	2.55025	0.1266E-08	2.798	Α
4:	0.10037	2.73109	0.1514E-05	0.2041E-02	Α
5:	0.10037	2.73109	0.2371E-08	1.303	Α
6:	0.10037	2.73109	0.1645E-07	0.1879	Α
7:	0.10164	2.76580	0.8882E-01	0.3392E-07	Α
8:	0.10592	2.88216	0.1760E-09	15.76	Α
9:	0.10592	2.88216	0.1001E-06	0.2772E-01	Α
10:	0.10592	2.88216	0.1068E-06	0.2597E-01	Α
11:	0.10753	2.92604	0.1321E-06	0.2037E-01	Α
12:	0.10753	2.92604	0.3438E-08	0.7829	Α
13:	0.10753	2.92604	0.3785E-08	0.7112	Α
14:	0.11237	3.05776	0.7005E-01	0.3518E-07	Α
15:	0.11321	3.08072	0.7619E-01	0.3187E-07	Α
16:	0.11327	3.08233	0.6733E-08	0.3603	Α
17:	0.11327	3.08233	0.3842E-07	0.6314E-01	Α
18:	0.11327	3.08234	0.4878E-03	0.4972E-05	Α
19:	0.11537	3.13932	0.2435E-02	0.9605E-06	Α
20:	0.12095	3.29119	0.1663E-01	0.1279E-06	Α

Recalling the spin-forbidden transitions appear with near vanishing oscillator strengths in an SOC treatment, we note the following from these results:

- ullet Excitation no. 7 is the lowest with a relevant oscillator strength f, it corresponds to $S_0 o S_1$.
- ullet The $S_0 o S_1$ excitation energy of 2.77 eV compares well with the experiment value, 2.85 eV.
- ullet At the GS geometry the T_1 and the T_2 state lie below $S_1.$

Computational Description of TADF 2: Spin-Orbit Coupling

As discussed above, transitions between S and T states are formally forbidden within a non-relativistic framework. Phosphorescence or intersystem crossings thus become only possible due to the presence of spin-orbit coupling (SOC) effects. Compared to other interactions these couplings are typically small and can therefore be treated as a perturbation of the systems electronic structure mediated by $\hat{H}_{\rm SOC}$, the SOC operator. There exist various different approaches and approximation for $\hat{H}_{\rm SOC}$. See the recent review of Marian for an overview of spin-orbit methods in the context of excited states electronic structures and intersystem crossing rates.

In this case Fermi's golden rule provides an expression for the reversed intersystem crossing rate

$$k_{ ext{RISC}} = rac{2\pi}{\hbar} |V_{ ext{SOC}}|^2 imes
ho_{ ext{FCWD}}$$

which will be used in the following. $|V_{\rm SOC}|^2$ is directly proportional to $k_{\rm RISC}$ and is to be understood as coupling term of the matrix representation of $\hat{H}_{\rm SOC}$. We model the transition from T_1 to S_1 , whereas T_1 actually consists of three individual states the system can assume;



$$\left|V_{\mathrm{SOC}}
ight|^{2}=rac{1}{3}{\left|\left\langle S_{1}\left|\hat{H}_{\mathrm{SOC}}
ight|T_{1}
ight
angle
ight|^{2}}=rac{1}{3}{\sum_{J=-1,0.1}}\!\!\left\langle S_{1}\left|\hat{H}_{\mathrm{SOC}}
ight|T_{1}^{J}
ight
angle \left\langle S_{1}\left|\hat{H}_{\mathrm{SOC}}
ight|T_{1}^{J}
ight
angle ^{st}$$

As $|V_{\rm SOC}|^2$ is strongly affecting the reverse intersystem crossing rate, developing novel TADF emitters with larger SOCME values is an important part of many OLED related research activities.

The other important factor entering the $k_{\rm RISC}$ rate is $\rho_{\rm FCWD}$, the Frank-Condon-weighted density of states, which will be addressed in the corresponding section below.

Calculating Spin-Orbit Couplings

The reverse intersystem crossing process itself occurs on a too short timescale for the systems geometry to readapt during the $T_1 \to S_1$ transition. The factor $|V_{\rm SOC}|^2$ in the above expression for $k_{\rm RISC}$ is therefore computed at the optimized geometry of the T_1 state.

Open a new instance of ADFinput

Load the previously optimized T1 structure with File → Import Coordinates...

In the Main panel:

Task: → Single Point

XC-Functional \rightarrow Range Separated \rightarrow LCY PBE

Click on the ... button left of the XC-Functional field

Enter Gamma: 0.22

Return to Main panel

Relativity → Scalar

Basis set → TZP

Frozen core → None

Numerical quality → Good

Click on Model panel and select Solvation

Solvation method: → COSMO

COSMO solvent → Toluene

Click on **Details** panel and select **COSMO**

Enable Include response in TDDFT

Click on Properties and select Excitations (UV/Vis), CD

Type of excitations: \rightarrow Spin-Orbit (Perturbative)

TDA: → Yes

Number of excitations: 5

Calculate: → NTOs



See also

• **L** 4CzIPN_SOCME.run run script to compute the spin-orbit coupling matrix elements

After the calculation is completed, open the output file with ADFoutput. In the ADFoutput the SOCME results can be found with Response Properties \rightarrow Spin-orbit couplings:

quares of	spin-orbit	coupling	matrix el	
T1	T2	т3	Т4	Т5
0.40	1.07	0.35	0.17	2.17
0.89	0.40	1.67	1.14	0.16
0.05	1.43	0.17	0.29	0.66
0.09	1.72	0.25	0.18	0.88
0.30	0.09	0.61	0.49	0.02
0.00	1.24	0.13	0.27	3.31
1.24	0.00	2.15	2.05	0.21
0.13	2.15	0.00	0.44	0.44
0.27	2.05	0.44	0.00	1.18
3.31	0.21	0.44	1.18	0.00
	quares of the uncou T1 0.40 0.89 0.05 0.09 0.30 0.00 1.24 0.13 0.27	quares of spin-orbit the uncoupled state T1 T2 0.40 1.07 0.89 0.40 0.05 1.43 0.09 1.72 0.30 0.09 0.00 1.24 1.24 0.00 0.13 2.15 0.27 2.05	quares of spin-orbit coupling the uncoupled states) in cm-1 T1 T2 T3 0.40 1.07 0.35 0.89 0.40 1.67 0.05 1.43 0.17 0.09 1.72 0.25 0.30 0.09 0.61 0.00 1.24 0.13 1.24 0.00 2.15 0.13 2.15 0.00 0.27 2.05 0.44	0.40 1.07 0.35 0.17 0.89 0.40 1.67 1.14 0.05 1.43 0.17 0.29 0.09 1.72 0.25 0.18 0.30 0.09 0.61 0.49 0.00 1.24 0.13 0.27 1.24 0.00 2.15 2.05 0.13 2.15 0.00 0.44 0.27 2.05 0.44 0.00

From the $|\langle S_1|\hat{H}_{
m SOC}|T_1
angle|$ value of 0.4 cm⁻¹ we then obtain $|V_{
m SOC}|^2=$ 8.2·10⁻¹⁰ eV 2 .

Computational Description of TADF 3: Vibrations

Recall the aforementioned expression for the reverse intersystem crossing rate:

$$k_{
m RISC} = rac{2\pi}{\hbar} |V_{
m SOC}|^2 imes
ho_{
m FCWD}$$

While $|V_{\rm SOC}|^2$ can be interpreted as probability for the $T_1 \to S_1$ transition, $\rho_{\rm FCWD}$ describes the thermokinetic barrier associated with this process.

Semiclassical Marcus Theory

Semiclassical Marcus theory (SCMT) provides a simple way to describe this thermokinetic barrier. The potential energy surfaces of both states around their respective minimum structures are thereby described in terms of parabolic functions.

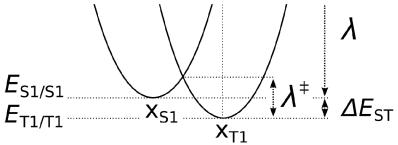


Fig. 15 Potential energy surfaces of the S_1 and T_1 states as quadratic functions centered at x_{S1} and x_{T1} , respectively. In Marcus theory the difference between $E_{S1/S1}$ and $E_{S1/T1}$ is used to estimate the thermokinetic barrier of the $T_1 \to S_1$ transition.

Assuming the same curvature of both states, i.e. $E(T_1)=a\cdot x^2$ and $E(S_1)=a\cdot (x+x_{\rm S1}-x_{\rm T1})^2+\Delta E_{\rm ST}$, one finds the for the $T_1\to S_1$ barrier height as the intersection of both parabolas:

$$\lambda^{\ddagger} = rac{(\Delta E_{
m ST} + \lambda)^2}{4\lambda}$$

Using this barrier height in an Arrhenius-type expression yields then the RISC rate:

$$ho_{ ext{FCWD}}^{ ext{SCMT}} = rac{1}{\sqrt{4\pi\lambda k_{ ext{B}}T}} ext{exp}igg(\!-rac{\left(\Delta E_{ ext{ST}}+\lambda
ight)^2}{4\lambda k_{ ext{B}}T}igg)$$

The reorganization energy λ can thereby be considered as the energy required to bring the system in the S_1 to the minimum structure of the T_1 state. Designating the energies of the S_1 state at $x_{\rm S1}$ and $x_{\rm T1}$ as $E_{\rm S1/S1}$ and $E_{\rm S1/T1}$, respectively, on can use

$$\lambda = E_{
m S1/T1} - E_{
m S1/S1}$$

to calculate k_{RISC} . We can thereby set up the necessary calculations analogously to the vertical absorption calculations:

Open a new instance of ADFinput

Load the previously optimized S1 structures with File → Import Coordinates...

In the Main panel:

Task: → Single Point

XC-Functional \rightarrow Range Separated \rightarrow LCY PBE

Click on the ... button left of the XC-Functional field

Enter Gamma: 0.22

Return to Main panel



FIUZEII CUIE - MUIIE

Numerical quality → Good

Click on Model panel and select Solvation

Solvation method: → COSMO

COSMO solvent → Toluene

Click on **Details** panel and select **COSMO**

Untick Include response in TDDFT

Click on Properties and select Excitations (UV/Vis), CD

Type of excitations: \rightarrow Spin-Orbit (Perturbative)

TDA: → Yes

Number of excitations: 5

Calculate: → NTOs

Calculate: → Charge transfer descriptors

Save the input File \rightarrow Save as... under e.g. $4CzIPN_S1_S1.adf$

Analogously for the S1/T1 calculation we set:

Load the previously optimized T1 structures with File \rightarrow Import Coordinates...

Repeat the above steps

Save the input **File** \rightarrow **Save as...** under e.g. $4CzIPN_s1_T1.adf$

See also

- \clubsuit 4CzIPN_S1atS1.run run script to compute $E_{
 m S1/S1}$
- riangle 4CzIPN_S1atT1.run run script to compute $E_{
 m S1/T1}$

After running the calculations the required energies can be extracted again with ADFoutput. Note, that the excited states energies are obtained from the ground state energy and the respective excitation energies. These can be retrieved in ADFoutput as *Total Bonding Energy* under **Properties** \rightarrow **Bonding Energy Decomposition** and as the first excitation with a significant oscillator strength under **Response Properties** \rightarrow **All Spin-Orbit Coupling Excitation Energies**, respectively.

For 4CzIPN this results in a value of λ = $E_{\rm S1/T1}$ - $E_{\rm S1/S1}$ = 0.06 eV. This result can now be used along with the previously obtained values $\Delta E_{\rm ST}$ = 0.16 eV and $|V_{\rm SOC}|^2$ = 8.2·10⁻¹⁰ eV² in the above expressions for $\rho_{\rm FCWD}^{\rm SCMT}$ and $k_{\rm RISC}$.

Compared to that, Samanta et al. obtained higher rates for 4CzIPN ($4.9\cdot10^6$ – $1.8\cdot10^7$ / s). This difference can be rationalized by their usage of different DFT approximations, the significantly smaller value of $\Delta E_{\rm ST}$ obtained thereby, and a different handling of the spin-orbit couplings. We also note that Samanta et al. argued that for emitters like 4CzIPN λ can be replaced with an effective value between 0.1 and 0.2 eV. When using λ = 0.1 eV instead of the previously calculated reorganization energy, the RISC rate rises to $k_{\rm RISC}$ = $6.3\cdot10^5$ / s.

Franck-Condon Principle and Marcus-Levich-Jortner Theory

While the semiclassical often provides acceptably accurate results, it relies on some strong assumptions and the real situation turns out more intricate: First, the potential energy surfaces around both minima are described by the entirety of vibrational normal modes. Each normal mode defines a separate harmonic oscillator system which, in a quantum mechanical description, leads to a separate set of vibrational eigenfunctions (quanta). The Franck-Condon principle then states that the $T_1 \to S_1$ transition requires a significant overlap between nuclear wavefunctions of both electronic states.

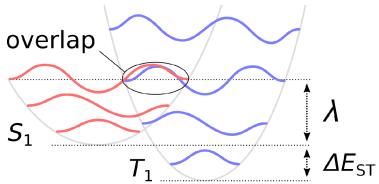


Fig. 16 The Franck-Condon principle describes $T_1 \to S_1$ transition rate in terms of the overlap between the nuclear vibronic wavefunctions of the initial and final state, respectively.

One can distinguish between two limiting cases: the quanta low frequency vibrations from shallow modes and from rearranging solvent molecules (see below) typically form an essentially continuous spectrum. Pairs of overlapping quanta from low frequency modes will therefore occur at energies close to the intersection of the two parabolic potential curves, which is leads to results equivalent to those of semiclassical Marcus theory. Within semiclassical Marcus theory the effect of these modes can be described in terms of a Marcus reorganization energy $\lambda_{\rm M}$. If the transition is dominated by high frequent modes, nuclear tunneling effects and the overlap of individual pairs of quanta are prevalent. The effect of such modes consists mainly in a renormalization of the coupling parameters.

An extension of SCMT, Marcus-Levich-Jortner theory includes includes both of these limiting



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$$ho_{ ext{FCWD}}^{ ext{MLJT}} = rac{1}{\sqrt{4\pi\lambda_{ ext{M}}k_{ ext{B}}T}} \sum_{n=0}^{\infty} \exp(-S_{eff}) rac{S_{eff}^n}{n!} \exp\Biggl(-rac{\left(\Delta E_{ ext{ST}} + n\hbar\omega_{eff} + \lambda_{ ext{M}}
ight)^2}{4\lambda_{ ext{M}}k_{ ext{B}}T} \Biggr)$$

Thereby, S_{eff} denotes the Huang-Rhys factor of this effective mode, which is a measure for the strength of the electron-phonon coupling. S_{eff} and the corresponding frequency ω_{eff} are obtained from the Huang-Rhys factors and frequencies of the individual high frequency normal modes.

$$S_{eff} = \sum_i S_i \qquad \omega_{eff} = \sum_i rac{\omega_i S_i}{S_{eff}}$$

Whereas modes with $\omega_i > 1000 \ {\rm cm}^{\text{-}1}$ are typically included in the summation.

Samanta et al. computed S_i factors from mode resolved reorganization energies $\lambda_i=k_i\Delta Q_i^2/2$, which in turn results from the frequencies $k_i=\omega_i^2$ and ΔQ_i , the normal mode projection onto the transition path.

For a large emitter molecule like 4CzIPN however, the calculation of the full Hessian matrix turns out computationally too expensive.

Effective Modes and Huang-Rhys Factors from DFTB and FCF

In some cases, the DFTB method can be used as a more approximate but efficient alternative to compute the normal modes. The FCF program of the Amsterdam Modeling Suite can then be used to calculate the S_i factors which then lead to the S_{eff} and ω_{eff} values. The following procedure illustrates how this can be achieved with the Amsterdam Modeling Suite.

Warning

The calculation of Huang-Rhys factors crucially depends on the quality of the normal modes. Their computation from TD-DFTB excited states gradient must be carefully tested to determine their validity for a given dye system.

Open a new instance of ADFinput

Load the previously optimized S_1 structure via File \rightarrow Import Coordinates

In the options panel click $ADF \rightarrow DFTB$

Task: → Geometry Optimization

Followed by: → Frequencies



FIUPEILIES -> EXCITATIONS (OV) VIS

Type of excitations: \rightarrow Singlet

Method: → Davidson

Number of excitations: $\rightarrow 2$

Analogously, for the T_1 state:

Open a new instance of ADFinput

Load the previously optimized T_1 structure via File \rightarrow Import Coordinates

...

Properties → Excitations (UV/Vis)

Type of excitations: \rightarrow Triplet

Method: → Davidson

Number of excitations: \rightarrow 2

See also

- \blacktriangle 4CzIPN_optFreqDFTB_S1.run DFTB optimization and normal modes on S_1
- \blacktriangle 4CzIPN_optFreqDFTB_T1.run DFTB optimization and normal modes on T_1

After completing both calculations, the binary result files dftb.rkf containing the normal modes are found in the respective directories <jobname>.results. Create a new directory and copy these rkf files naming them e.g. $S1_optFreqDFTB.rkf$ and $T1_optFreqDFTB.rkf$. Based on these results the program-tool fcf can be invoked to compute S_{eff} and ω_{eff} values using the following bash script:

```
$ADFBIN/fcf << eor
STATES T1_optFreqDFTB.rkf S1_optFreqDFTB.rkf
QUANTA 0 0
TRANSLATE
ROTATE
eor</pre>
```

Running this script yields the frequencies and Electron-Phonon couplings for both states as shown in the following example:



Frequency cm^-1	Displacement a0 amu^1/2	<pre>Electron-Phonon coupling (dimensionless)</pre>
12.070318	-11.804599	2.642917
13.872303	3.368437	0.808492
19.113716	-2.303894	0.649094
976.643338	-0.001580	0.003183
986.370540	-0.000023	0.000046
1041.616941	-0.001475	0.003069
1075.333767	-0.019010	0.040172
1078.281693	-0.017297	0.036603
3075.114508	0.033650	0.120250
3086, 236558	0.013900	0.049764
3087.310659	0.005134	0.018383

We are interested in the $T_1 \to S_1$ transition and are therefore using the results from the first state. The mode-resolved Huang-Rhys factors S_i are obtained as the square of the above Electron-Phonon coupling parameters. Using the above expressions for S_{eff} and ω_{eff} , we then sum over all modes with $\omega_i >$ 1000 cm⁻¹. This results in S_{eff} = 0.837 and ω_{eff} = 1498 cm⁻¹.

These quantities are then used in the above expression for $\rho_{\mathrm{FCWD}}^{\mathrm{SCMT}}$ (with λ_{M} = λ = 0.06 eV) whereas the summation is stopped at n = 5. With $|V_{\mathrm{SOC}}|^2 = 8.2 \cdot 10^{-10}$ eV 2 this results in a reverse intersystem crossing rate that is essentially identical to the result from semiclassical Marcus theory of k_{RISC} = 2.3·10 5 / s. This supports the finding of Samanta et al. that quantum tunneling effects are negligible in the reverse intersystem crossing process of 4CzIPN and that a classical treatment suffices for its description.

Computational Description of TADF 4: Solvent Effects

The approach presented above included solvent effects implicitly, as the COSMO model was used in the calculations.

Besides examining vibrational modes that simultaneously involve both, the OLED emitter and explicit solvent molecules, and pursuing analogously to the approach for intramolecular vibrations discussed above, many more explicit treatments have been suggested in the literature. Outersphere vibronic contributions can for example be obtained in terms of dielectric response functions (Rühle et al.), polarizable force fields (McMahon and Troisi), or from QM/MM