

# PC2232 Physics for Electrical Engineers: Tutorial 8

## Question 1: Stirling's approximation

[Griffiths 5.27] Stirling's approximation, given in pg 7 of Lecture 11, is

$$\ln z! \simeq z \ln z - z, \quad \text{for } z \gg 1. \quad (1)$$

This relation, being an approximation, is of course not expected to be 100% accurate.

- (a) Find the percentage error in Stirling approximation for  $z = 10$ .
- (b) What is the smallest integer  $z$  such that the error is less than 5%?

## Question 2: Two-dimensional density of states

Consider spin-1/2 electrons in a two-dimensional box with the length of sides  $L_x$  and  $L_y$ . Calculate the density of states,  $\rho(E)$  for this case.

## Question 3: Bose-Einstein distribution

The Fermi-Dirac distribution was obtained for *fermions*. Here, calculate the corresponding distribution for *bosons*, where the multiplicity  $W$  for bosons is given by

$$W(N_1, \dots, N_N) = \prod_{n=1}^{\infty} \frac{(N_n + d_n - 1)!}{N_n! (d_n - 1)!}, \quad (2)$$

under the constraints of number and energy conservation. *Hint:* use the method of Lagrange multipliers similar to the fermion case:

$$G = \ln W + \alpha \left( N - \sum_n N_n \right) + \beta \left( E - \sum_n N_n E_n \right). \quad (3)$$

## Question 4: Blackbody radiation

According to Plank's hypothesis, each photon of light carries an energy given by  $E = \hbar\omega$ . You are given that for photons, the number of states per  $k$ -space volume is the same as for electrons,

$$C(k_x, k_y, k_z) = \frac{2V}{\pi^3}, \quad (4)$$

where  $V$  is the real-space volume of the box. Show that, for an infinitesimal interval of energy between  $E$  and  $E + dE$ , there are  $d_k$  states that can carry this energy, where  $d_k$  is given by

$$d_k = \frac{V k^2 dk}{\pi^2}. \quad (5)$$

Show that the energy density for photons in the box, defined by

$$\frac{N\hbar\omega}{V} = \rho(\omega) d\omega = \frac{\hbar\omega^3 d\omega}{\pi^2 c^3 (e^{\hbar\omega/k_B T} - 1)}. \quad (6)$$

This is the well-known spectrum describing blackbody radiation.

**Question 5: Wien's Law**

- (a) Use Eq. (6) to determine the energy density in the *wavelength* range  $d\lambda$ . *Hint*: Set  $\rho(\omega)d\omega = \bar{\rho}(\lambda)d\lambda$ .
- (b) Find the *peak intensity wavelength*,  $\lambda = \lambda_{\max}$  for  $\bar{\rho}(\lambda)$ . Express  $\lambda_{\max}$  in terms of temperature and the other physical constants. *Hint*: You may need the numerical solution to the following transcendental equation:

$$(5 - x) e^x = 5, \quad \rightarrow \quad x \simeq 4.965 \dots \quad (7)$$