

Inertial and gravity waves

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Waves vs Modes

- Waves are travelling, don't care about boundaries till they reflect. They are 'local' in nature.
- Modes are aware of boundaries and are 'global' in nature.

Types of waves in planetary fluid dynamics

Type of wave	Restoring force(s)
Acoustic (p -modes)	$-\nabla p$
Inertial	$-\nabla p - 2\boldsymbol{\Omega} \times \mathbf{u}$
Surface/Internal gravity (f -/ g -modes)	$\rho' \mathbf{g}$
Inertia-gravity or gravito-inertial	$-\nabla p - 2\boldsymbol{\Omega} \times \mathbf{u} + \rho' \mathbf{g}$
Alfvén	$\mathbf{B} \cdot \nabla \mathbf{B}$
Magnetoacoustic	$-\nabla p + \mathbf{B} \cdot \nabla \mathbf{B}$
Magneto-Coriolis (MC)	$-\nabla p - 2\boldsymbol{\Omega} \times \mathbf{u} + \mathbf{B} \cdot \nabla \mathbf{B}$
Magnetic, Archimedean, Coriolis (MAC)	$-2\boldsymbol{\Omega} \times \mathbf{u} + \rho' \mathbf{g} + \mathbf{B} \cdot \nabla \mathbf{B}$

Internal gravity oscillations



Internal gravity oscillations



Equation of motion of blob:

$$\begin{aligned}\rho_0 \frac{d^2}{dt^2} \Delta z &= \rho_0(z + \Delta z)g - \rho_0(z)g \\ &= \left(\rho_0(z) + \frac{d\rho_0}{dz} \Delta z \right) g - \rho_0(z)g\end{aligned}\tag{1}$$

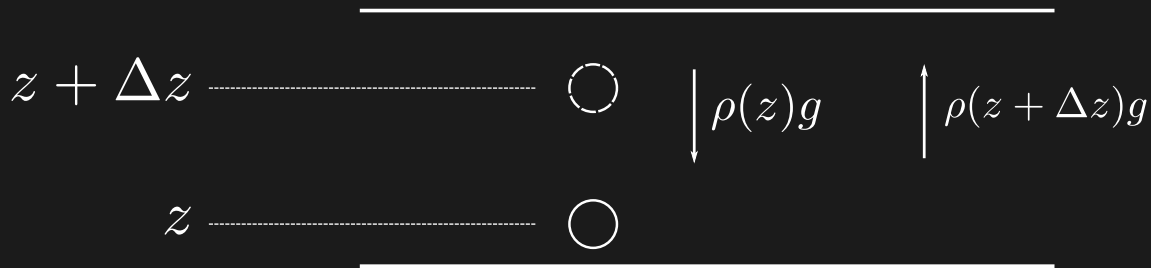
Internal gravity oscillations



$$\frac{d^2}{dt^2}\Delta z + N^2\Delta z = 0 \quad (2)$$

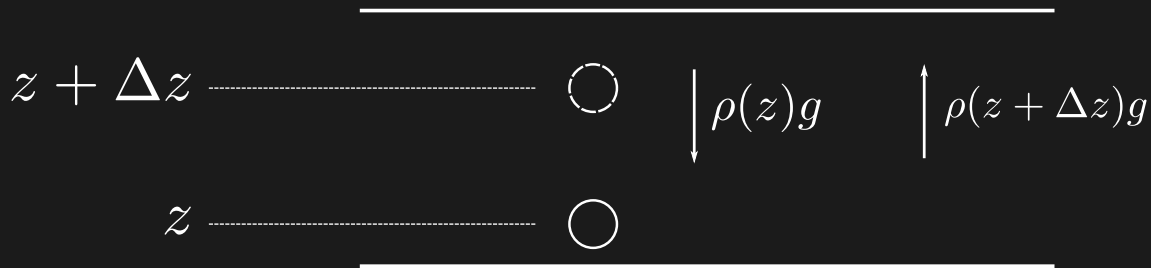
$$N = \sqrt{-\frac{g}{\rho_0} \frac{d\rho_0}{dz}} \quad (3)$$

Internal gravity oscillations



What happens when $N^2 > 0$ and vice versa?

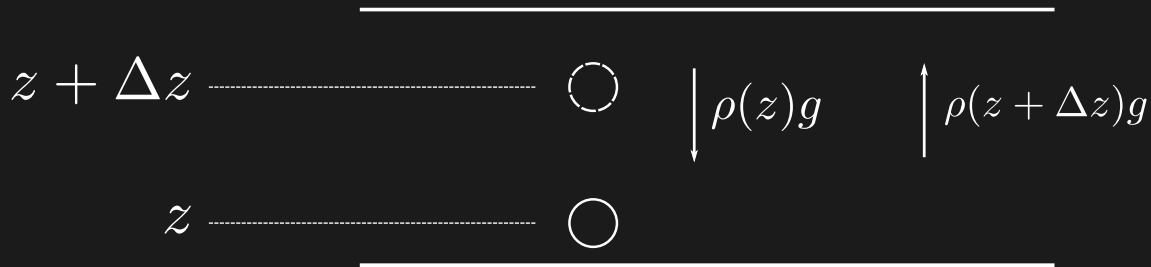
Internal gravity oscillations



What happens when $N^2 > 0$ and vice versa?

- $N^2 > 0$: Internal gravity waves
- $N^2 < 0$: Convective instability

Internal gravity oscillations

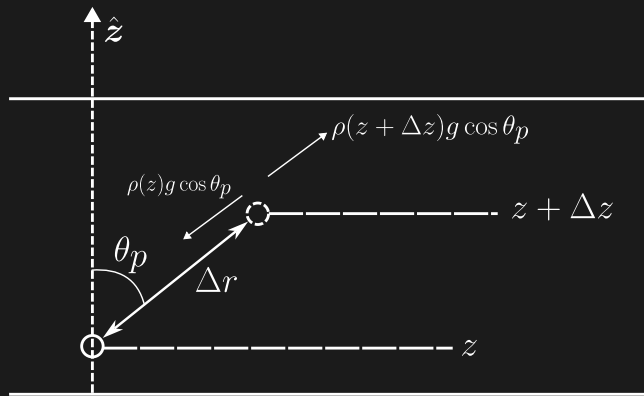


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N is called the Brunt–Väisälä frequency or buoyancy frequency.

At an angle?



At an angle?

$$\begin{aligned}\rho_0 \frac{d^2}{dt^2} \Delta r &= (\rho_0(z + \Delta z)g - \rho_0(z)g) \cos \theta_p \\ &= \left[\left(\rho_0(z) + \frac{d\rho_0}{dz} \Delta z \right) g - \rho_0(z)g \right] \cos \theta_p \\ &= \left(\frac{d\rho_0}{dz} \Delta r \cos \theta_p \right) \cos \theta_p\end{aligned}\tag{4}$$

Frequency of oscillation, $\omega = N \cos \theta_p$

When pressure variations are important

- In previous example, we assumed the parcel did not change its density when moving up.
- This is no longer true if there are significant pressure variations across height.
- Requires a simple change, replace density by potential density ρ_θ

$$N = \sqrt{-\frac{g}{\rho_\theta} \frac{d\rho_\theta}{dz}} = \sqrt{\frac{g}{\theta} \frac{d\theta}{dz}} \quad (5)$$

When pressure variations are important

$$\theta = T \left(\frac{P_o}{P} \right)^{(R/C_p)}$$

Thus,

$$N = \sqrt{\frac{g}{\theta} \frac{d\theta}{dz}} = \sqrt{g \left[\frac{1}{T} \frac{dT}{dz} - \frac{R}{C_p} \frac{1}{P} \frac{dP}{dz} \right]} \quad (6)$$

Schwarzschild criterion

For convection to occur:

$$\begin{aligned} N^2 &< 0 \\ \Rightarrow g \left[\frac{1}{T} \frac{dT}{dz} - \frac{R}{C_p} \frac{1}{P} \frac{dP}{dz} \right] &< 0 \\ \Rightarrow \frac{P}{T} \frac{dT}{dP} &< \frac{R}{C_p} \\ \Rightarrow \frac{d \ln T}{d \ln P} &< \left(\frac{d \ln T}{d \ln P} \right)_{ad} \end{aligned} \tag{7}$$

Internal gravity waves

- Wave involves a coherent oscillation of a domain of parcels/particles
- Need a continuum analysis

Linearized equations of Boussinesq approximation

$$\nabla \cdot \mathbf{u} = 0 \quad (8)$$

$$\frac{\partial \mathbf{u}}{\partial t} = -\frac{1}{\rho_0} \nabla p - \frac{\rho'}{\rho_0} \hat{\mathbf{z}} \quad (9)$$

$$\frac{\partial \rho'}{\partial t} + w \frac{d\rho_0}{dz} = 0 \quad (10)$$

Useful vector relations

$$\nabla \times \nabla \times \mathbf{A} = \nabla(\nabla \cdot \mathbf{A}) - \nabla^2 \mathbf{A} \quad (11)$$

$$\begin{aligned} \nabla \times \nabla \times (f \hat{\mathbf{z}}) &= (\dots) \hat{\mathbf{x}} + (\dots) \hat{\mathbf{y}} - \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) f \hat{\mathbf{z}} \\ &= (\dots) \hat{\mathbf{x}} + (\dots) \hat{\mathbf{y}} - \nabla_H^2 f \hat{\mathbf{z}} \end{aligned} \quad (12)$$

Linearized equations of Boussinesq approximation

Taking double curl of equation (9) and only considering the z -component,

$$-\frac{\partial}{\partial t} \nabla^2 w = \frac{g}{\rho_0} \nabla_H^2 \rho' \quad (13)$$

Differentiating this wrt time and using equation (10), we get

$$\frac{\partial^2}{\partial t^2} \nabla^2 w = \frac{g}{\rho_0} \frac{dT_0}{dz} \nabla_H^2 w = -N^2 \nabla_H^2 w \quad (14)$$

Dispersion relation

Assuming, $w = \hat{w}e^{i(k_x x + k_y y + k_z z - \omega t)}$, equation (14) becomes,

$$-\omega^2 k^2 = -N^2 k_h^2 \Rightarrow \omega^2 = N^2 \frac{k_h^2}{k^2} \quad (15)$$

where, $k^2 = k_x^2 + k_y^2 + k_z^2$ and $k_h^2 = k_x^2 + k_y^2$. Thus,

$$\omega = \pm N \cos \theta \quad (16)$$

where, θ is the angle to the *horizontal*.

Particle motion

$$\nabla \cdot \mathbf{u} = 0 \Rightarrow \mathbf{k} \cdot \mathbf{u} = 0$$

Thus, particles move perpendicular to the wavevector. This can also be seen in the definitions of θ_p and θ .

Phase and Group velocities

Phase and Group velocities

$$c_{\text{phase}} = \frac{\omega}{k} \frac{\mathbf{k}}{k} = N \frac{k_h}{k^3} \mathbf{k} \quad (17)$$

$$\begin{aligned} c_{\text{group}} = \nabla_{\mathbf{k}} \omega(\mathbf{k}) &= \frac{d\omega}{dk_x} \hat{\mathbf{x}} + \frac{d\omega}{dk_y} \hat{\mathbf{y}} + \frac{d\omega}{dk_z} \hat{\mathbf{z}} \\ &= \frac{N}{k^3} \left[\frac{k_z^2}{k_h} (k_x \hat{\mathbf{x}} + k_y \hat{\mathbf{y}}) - k_z k_h \hat{\mathbf{z}} \right] \end{aligned} \quad (18)$$

$$c_{\text{phase}} \cdot c_{\text{group}} = \frac{N}{k^3} (k_z^2 k_h^2 - k_z^2 k_h^2) = 0$$

Movie
Movie 2

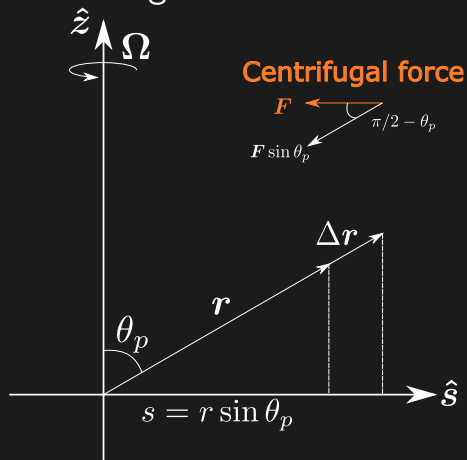
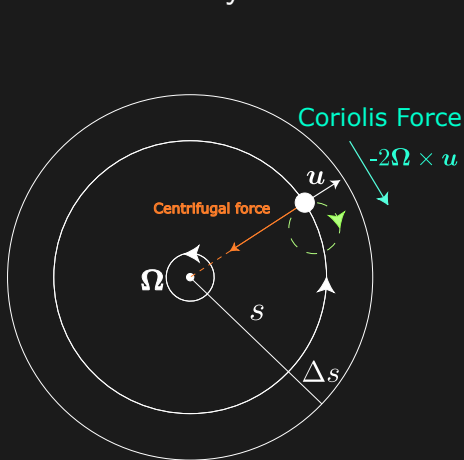
Inertial Oscillations

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A fluid in solid body rotation is stratified in angular momentum.

Inertial Oscillations

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Inertial Oscillations

$$\ddot{\Delta r} + 4\Omega^2 \sin^2 \theta_p \Delta r = 0$$

Oscillation frequency,

$$\omega = \pm 2\Omega \sin \theta_p$$

Inertial waves

Taylor-Proudman theorem:

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Taking curl:

$$\frac{\partial}{\partial t} (\nabla \times \mathbf{u}) = 2\mathbf{\Omega} \frac{\partial \mathbf{u}}{\partial z} \quad (21)$$

Inertial waves

Taking curl again,

$$-\frac{\partial}{\partial t} \nabla^2 \mathbf{u} = 2\Omega \frac{\partial}{\partial z} \nabla \times \mathbf{u}$$

Differentiating wrt times and using equation (21), we get

$$\frac{\partial^2}{\partial t^2} \nabla^2 \mathbf{u} + 4\Omega^2 \frac{\partial^2 \mathbf{u}}{\partial z^2} = 0 \quad (22)$$

Inertial waves

Consider an ansatz, $\mathbf{u}, p \propto e^{i(\mathbf{k} \cdot \mathbf{r} - \omega t)}$. Substituting in the continuity equation,

$$\mathbf{u} \cdot \mathbf{k} = 0 \quad (23)$$

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Substituting in (22), we obtain

$$\omega^2 k^2 - 4\Omega^2 k_z^2 = 0 \Rightarrow \omega = \pm 2\Omega \frac{k_z}{k} = \pm 2\Omega \cos \theta \quad (24)$$

What do these physically mean?

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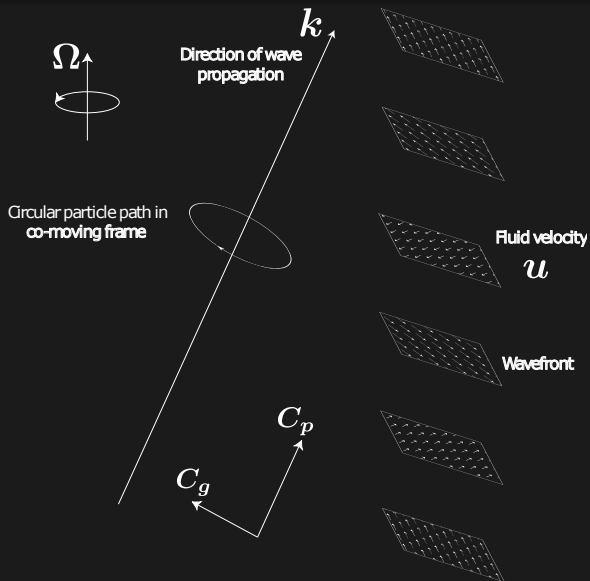
- The velocity perturbations are perpendicular to the wave propagation direction \Rightarrow transverse waves (like on a string).
- The frequency of the waves lie between $\pm 2\Omega$.
- Frequency depends on the angle θ from rotation axis.

Inertial waves

$$c_{\text{phase}} = \frac{\omega(\mathbf{k})}{|\mathbf{k}|} \hat{\mathbf{k}} = \pm \frac{2\Omega \hat{\mathbf{z}} \cdot \hat{\mathbf{k}}}{|\mathbf{k}|} \hat{\mathbf{k}}$$

$$c_{\text{group}} = \nabla_{\mathbf{k}} \omega(\mathbf{k}) = \pm 2\Omega \frac{\hat{\mathbf{k}} \times \hat{\mathbf{z}} \times \hat{\mathbf{k}}}{|\mathbf{k}|}$$

Inertial waves



Inertial waves

Movie

Inertia-gravity waves

What happens when both forces are present?

Inertia-gravity waves

$$\nabla \cdot \mathbf{u} = 0 \quad (25)$$

$$\frac{\partial \mathbf{u}}{\partial t} = -\frac{1}{\rho_0} \nabla p - \frac{\rho'}{\rho_0} g \hat{\mathbf{z}} - 2\Omega \hat{\mathbf{z}} \times \mathbf{u} \quad (26)$$

$$\frac{\partial \rho'}{\partial t} + w \frac{d\rho_0}{dz} = 0 \quad (27)$$

Inertia-gravity waves

Take curl of (26),

$$\frac{\partial}{\partial t} \nabla \times \mathbf{u} = -\nabla \times \frac{\rho'}{\rho_0} g \hat{\mathbf{z}} + 2\Omega \frac{\partial \mathbf{u}}{\partial z} \quad (28)$$

Taking the z -component gives,

$$\frac{\partial \omega_z}{\partial t} = 2\Omega \frac{\partial w}{\partial z} \quad (29)$$

Taking the curl of (28) and taking the z -component gives

$$-\frac{\partial}{\partial t} \nabla^2 w = \frac{g}{\rho_0} \nabla_H^2 \rho' + 2\Omega \frac{\partial \omega_z}{\partial z} \quad (30)$$

Inertia-gravity waves

Differentiating wrt time and using equations (27) and (29), we get

$$\begin{aligned} -\frac{\partial^2}{\partial t^2} \nabla^2 w &= -\frac{g}{\rho_0} \frac{d\rho_0}{dz} \nabla_H^2 w + 4\Omega^2 \frac{\partial^2 w}{\partial z^2} \\ \Rightarrow -\frac{\partial^2}{\partial t^2} \nabla^2 w &= N^2 \nabla_H^2 w + 4\Omega^2 \frac{\partial^2 w}{\partial z^2} \end{aligned} \quad (31)$$

Inertia-gravity wave

Using the wave solution, $w = \hat{w}e^{i(k_x x + k_y y + k_z z - \omega t)}$, we get:

$$\omega^2 k^2 = N^2 k_h^2 + 4\Omega^2 k_z^2 \quad (32)$$

$$\omega^2 = N^2 \left(\frac{k_h}{k} \right)^2 + 4\Omega^2 \left(\frac{k_z}{k} \right)^2 \quad (33)$$

Inertia-gravity wave

When gravity and rotation are misaligned,

$$\omega^2 = N^2 \left(\frac{k_h}{k} \right)^2 + f^2 \left(\frac{k_z}{k} \right)^2 = N^2 \cos^2 \theta + f^2 \sin^2 \theta \quad (34)$$

where, θ = angle wrt horizontal.

This limits the range of ω to $f \leq \omega \leq N$.

Movie

Particle paths

Note that the continuity equation still gives $\mathbf{k} \cdot \mathbf{u} = 0$. Particles go in elliptical paths perpendicular to the wave propagation direction.

Phase and group velocities

$$c_{\text{phase}} = \frac{\omega}{k} \frac{\mathbf{k}}{k} = \frac{\omega}{k^2} \mathbf{k} \quad (35)$$

$$\begin{aligned} c_{\text{group}} = \nabla_{\mathbf{k}} \omega(\mathbf{k}) &= \frac{d\omega}{dk_x} \hat{\mathbf{x}} + \frac{d\omega}{dk_y} \hat{\mathbf{y}} + \frac{d\omega}{dk_z} \hat{\mathbf{z}} \\ &= \frac{N^2 - f^2}{\omega k^2} \left[k_z^2 (k_x \hat{\mathbf{x}} + k_y \hat{\mathbf{y}}) - k_z k_h^2 \hat{\mathbf{z}} \right] \end{aligned} \quad (36)$$

$$c_{\text{phase}} \cdot c_{\text{group}} = \frac{N^2 - f^2}{k^4} \left[k_z^2 k_h^2 - k_z^2 k_h^2 \right] = 0 .$$