Table 1: Key units in Physical Chemistry

N_{Av}:	6.02214×10^{23}	mol^{-1}		
1 amu:	1.6605×10^{-27}	kg		
k_{B}:	1.38065×10^{-23}	$\mathrm{J\ K}^{-1}$	8.61734×10^{-5}	$\mathrm{eV\ K}^{-1}$
R:	8.314472	$\mathrm{J\ K}^{-1}\ \mathrm{mol}^{-1}$	8.2057×10^{-2}	$\mathrm{l\ atm\ mol}^{-1}\ \mathrm{K}^{-1}$
σ_{SB}:	5.6704×10^{-8}	$\mathrm{J\ s}^{-1}\ \mathrm{m}^{-2}\ \mathrm{K}^{-4}$		
c:	2.99792458×10^{8}	$\mathrm{m\ s}^{-1}$		
h:	6.62607×10^{-34}	J s	4.13566×10^{-15}	eV s
$\hbar$:	1.05457×10^{-34}	J s	6.58212×10^{-16}	eV s
hc:	1239.8	eV nm		
e:	1.60218×10^{-19}	C		
m_e :	$9.10938215 \times 10^{-31}$	kg	1: 0.5109989	$\mathrm{MeV\ c}^{-2}$
ϵ_0:	8.85419×10^{-12}	$\mathrm{C}^2\ \mathrm{J}^{-1}\ \mathrm{m}^{-1}$	5.52635×10^{-3}	e^2 $\mathrm{Å}^{-1}\ \mathrm{eV}^{-1}$
$e^2/4\pi\epsilon_0$:	2.30708×10^{-28}	J m	14.39964	eV Å
a_0:	0.529177×10^{-10}	m	0.529177	Å
E_{H}:	1	Ha	27.212	eV

1 The Classical Foundations

1.1 Lecture 0: Introduction

1. Burning lighter
2. Foundations of Physical Chemistry
   (a) Quantum mechanics
   (b) Statistical mechanics
   (c) Thermodynamics, kinetics, spectroscopy
   (d) Physical and chemical properties of matter

1.2 Lecture 1: Basic statistics

1. Discrete probability distributions—Coin flip
   (a) Example of Bernoulli trial, 2^n possible outcomes from n flips
   (b) Number of ways to get i heads in n flips, ${}_nC_i = n!/i!(n-i)!$
   (c) Probability of i heads $P_i \propto {}_nC_i$
   (d) *Normalized* probability, $\tilde{P}_i = P_i / \sum_i P_i = {}_nC_i/2^n$
   (e) Expectation value $\langle i \rangle = \sum_i i\tilde{P}_i$
2. Continuous distributions—temperature
   (a) Probability density $P(x)$ has units $1/x$
   (b) Normalized $\tilde{P}(x) = P(x)/\int P(x)dx$
   (c) (Unitless) probability $a < x < b = \int_a^b \tilde{P}(x)dx$

Table 2: Energy conversions and correspondences

	J	eV	Hartree	kJ mol^{-1}	cm^{-1}
1 J =	1	6.2415×10^{18}	2.2937×10^{17}	6.0221×10^{20}	5.0340×10^{22}
1 eV =	1.6022×10^{-19}	1	0.036748	96.485	8065.5
1 Ha =	4.3598×10^{-18}	27.212	1	2625.6	219474.6
1 kJ mol^{-1} =	1.6605×10^{-21}	0.010364	3.8087×10^{-4}	1	83.5935
1 cm^{-1} =	1.986410^{-23}	1.23984×10^{-4}	4.55623×10^{-6}	0.011963	1

(d) Expectation value $\langle f(x) \rangle = \int f(x)\tilde{P}(x)dx$

(e) Mean $= \langle x \rangle$

(f) Mean squared $= \langle x^2 \rangle$

(g) Variance $\sigma^2 = \langle x^2 \rangle - \langle x \rangle^2$

(h) Standard deviation $\Delta x = \sigma$

3. Boltzmann distribution

   (a) $P(E) \propto e^{-E/k_BT}$, in some sense the definition of temperature

   (b) Energy and its units

   (c) Absolute temperature and its units

   (d) k_BT as an energy scale, 0.026 eV at 298 K

   (e) Gravity example

      i. $E(h) = mgh$, linear, continuous energy spectrum

      ii. molecule vs car in a gravitational field

      iii. Barometric law for gases, $P = P_0 e^{-mgh/k_BT}$

   (f) Kinetic energy in 1-D example

      i. $KE = \frac{1}{2}mv_x^2$

      ii. $P_{1D}(v_x) = \left(\frac{m}{2\pi k_BT}\right)^{1/2} \exp\left(-\frac{m|v_x|^2}{2k_BT}\right)$

      iii. Gaussian distribution, mean μ, variance σ^2

$$G(x) = \frac{1}{\sigma\sqrt{2\pi}} \exp\left(-\frac{(x-\mu)^2}{2\sigma^2}\right)$$

      iv. By inspection, $\mu = \langle v_x \rangle = 0$, $\sigma^2 = \langle v_x^2 \rangle = k_BT/m$

      v. Molecule vs car again

   (g) Equipartition – energy freely exchanged between all degrees of freedom

1.3 Lecture 2: Kinetic theory of gases

1. Postulates

   (a) Gas is composed of molecules in constant random, thermal motion

   (b) Molecules only interact by perfectly elastic collisions

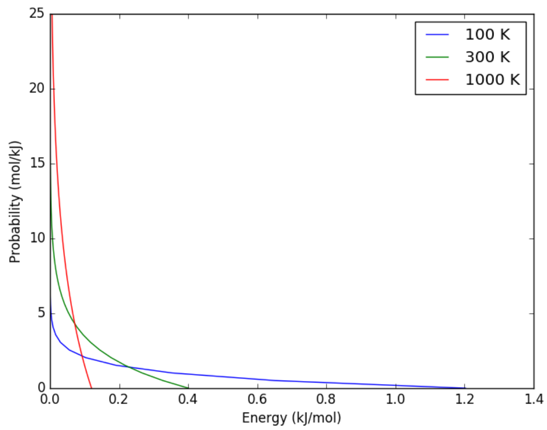

Figure 1: Boltzmann distribution at various temperatures

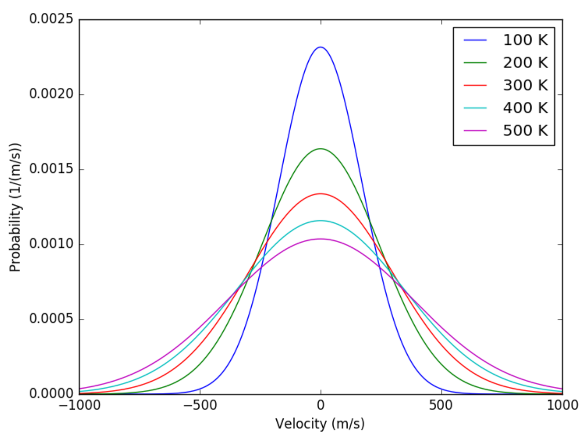

Figure 2: One-dimensional (Gaussian) velocities of N_2 gas

   (c) Volume of molecules is << total volume

2. Maxwell-Boltzmann distribution of molecular speeds
   (a) Speed $v = \sqrt{v_x^2 + v_y^2 + v_z^2}$
   (b) $P_{MB}(v)dv = P_{1D}(v_x)P_{1D}(v_y)P_{1D}(v_z) * \text{degeneracy}(v)dv$
   (c) mean speeds $\propto \sqrt{T}$
   (d) mean energy $U = \frac{3}{2}RT$ and heat capacity $C_v = \frac{3}{2}R$

3. Flux and pressure
   (a) Velocity flux $j(v_x)dv_x = v_x \frac{N}{V} P(v_x)dv_x$, molecules /area /time /v_x
   (b) Wall collisions, J_w, total collisions /area /time
   (c) Momentum exchange, pressure, ideal gas law

4. Collisions and mean free path
   (a) Collision cross section $\sigma = \pi d^2$, size of molecule
   (b) Molecular collisions, z per molecule and z_{AA} per volume
   (c) Mean free path, λ, mean distance between collisions

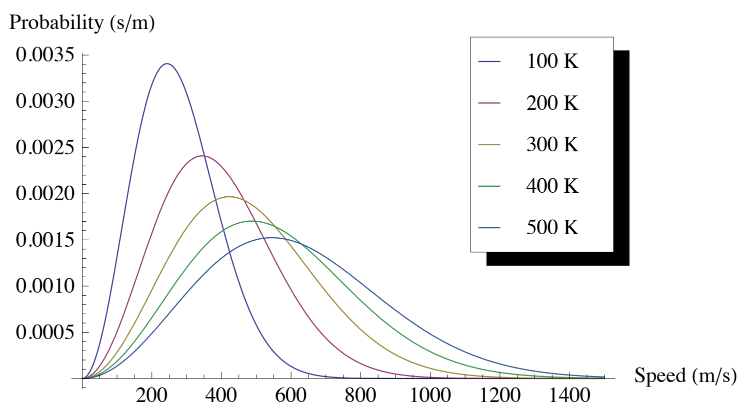

Figure 3: Maxwell-Boltzmann speed distribution of N_2 gas

1.4 Lecture 3: Transport

1. Effusion and Graham's law, effusion rate $\propto MW^{-1/2}$

2. Fick's first law: net flux proportional to concentration gradient
   (a) $j_x = -D\frac{dc}{dx}$
   (b) Self-diffusion constant, $D = \frac{1}{3}\lambda\langle v\rangle$

3. Knudsen diffusion, $D = \frac{1}{3}l\langle v\rangle$

4. Fick's second law: time evolution of concentration gradient

Table 3: Kinetic theory of gases key equations

Boltzmann distribution ($g(E)$: degeneracy of E)	$P(E) = g(E)e^{-E/k_BT}$
Maxwell-Boltzmann distribution	$P_{\text{MB}}(v) = 4\pi v^2 \left(\frac{m}{2\pi k_BT}\right)^{3/2} \exp\left(-\frac{mv^2}{2k_BT}\right)$
Mean and RMS speeds	$\langle v\rangle = \left(\frac{8k_BT}{\pi m}\right)^{1/2} \quad \langle v^2\rangle^{1/2} = \left(\frac{3k_BT}{m}\right)^{1/2}$
Pressure	$\langle P\rangle = \frac{\Delta p}{\Delta t} = m\frac{N}{V}\frac{1}{3}\langle v^2\rangle = \frac{Nk_BT}{V} = \frac{nRT}{V}$
Wall collision frequency	$J_W = \frac{1}{4}\frac{N}{V}\langle v\rangle = \frac{P}{(2\pi mk_BT)^{1/2}}$
Molecular collision frequency	$z = \sqrt{2}\sigma\langle v\rangle\frac{N}{V} = \frac{4\sigma P}{(\pi mk_BT)^{1/2}}$
Total collisions	$z_{AA} = \frac{1}{2}\frac{N}{V}z$
Mean free path	$\lambda = \frac{\langle v\rangle}{z} = \frac{V}{\sqrt{2}\sigma N}$
Graham's effusion law	$\frac{dN}{dt} = \text{Area}\cdot J_w \propto 1/m^{1/2}$
Effusion from a vessel	$P = P_0e^{-t/\tau}, \tau = \frac{V}{A}\left(\frac{2\pi m}{k_BT}\right)^{1/2}$
Self-diffusion constant	$D_{11} = \frac{1}{3}\langle v\rangle\lambda$
Diffusion rate	$\langle x^2\rangle^{1/2} = \sqrt{2Dt} \quad \langle r^2\rangle^{1/2} = \sqrt{6Dt}$
Einstein-Smoluchowski equation	$D_{11} = \frac{\delta^2}{2\tau}$
Stokes-Einstein equation for liquids	$D_{11} = \frac{k_BT}{4\pi\eta r}$ "Slip" boundary
	$D_{\text{Brownian}} = \frac{k_BT}{6\pi\eta r}$ "Stick" boundary

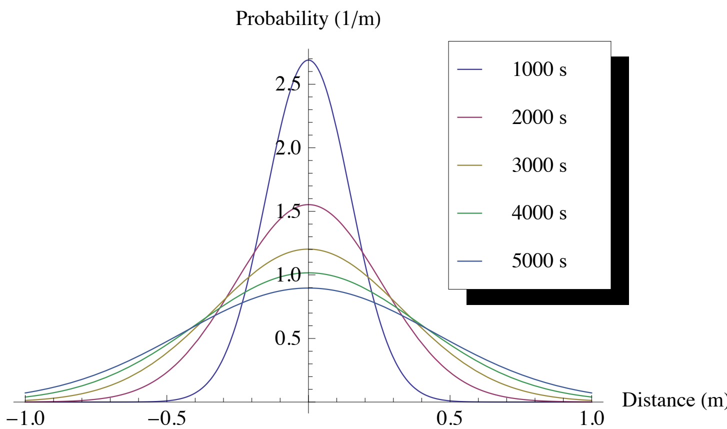

Figure 4: Diffusional spreading, $\sqrt{\langle x^2 \rangle} = \sqrt{2Dt}$

(a) Continuity with no advection: $\frac{\partial c}{\partial t} = -\nabla \cdot \vec{j} + \text{gen}$

(b) One-dimension: $\frac{dc}{dt} = D\frac{d^2c}{dx^2}$

(c) Diffusion has Gaussian probability distribution: $c(x,t)/c_0 = [2\sqrt{\pi Dt}]^{-1}\exp(-x^2/4Dt)$

5. Seeing is believing—Brownian motion
   (a) Seemingly random motion of large particles (“dust”) due to “kicks” from invisible molecules
   (b) Einstein receives Nobel Prize for showing:
      i. Motion follows same Gaussian diffusion behavior
      ii. From steady-state arguments in a field, diffusion constant is ratio of Boltzmann energy, k_BT, to mobility
      iii. Mobility inversely related to viscosity
   (c) Stokes-Einstein equation
   (d) Allows measurement of Avogadro’s number, final proof of kinetic theory
   (e) Similar model for diffusion of liquid molecules, slip boundary
6. Random walk model of diffusion
   (a) Binomial distribution
   (b) Large N and Stirling approximation
   (c) Einstein-Smoluchowski relation

2 Quantum Mechanics: Blurred Lines Between Particles and Waves

2.1 Lecture 4: Duality and demise of classical physics

1. Properties of waves
   (a) traveling waves, $\psi(x,t) = A\sin(kx - \omega t)$, $k = 2\pi/\lambda$, $\omega = 2\pi\nu$
   (b) standing waves, $\psi(x,t) = A\sin(kx)\cos(\omega t)$
   (c) interference, diffraction

   (d) energy proportional to amplitude squared
   (e) Expected energy of a classical oscillator, $\langle \epsilon \rangle_\nu = k_B T$ for all ν

2. Blackbody radiation
   (a) Hohlraum spectrum
   (b) Stefan-Boltzmann law, total irradiance
   (c) Wien's displacement law
   (d) Rayleigh-Jeans and ultraviolet catastrophe
   (e) Planck model
      i. Energy spectrum of oscillators are *quantized*, $\epsilon_\nu = nh\nu$
      ii. Expected energy of a quantized oscillator, $\langle \epsilon \rangle_\nu = h\nu / \left(e^{h\nu/k_B T} - 1 \right)$
      iii. Planck expression for blackbody radiation works!

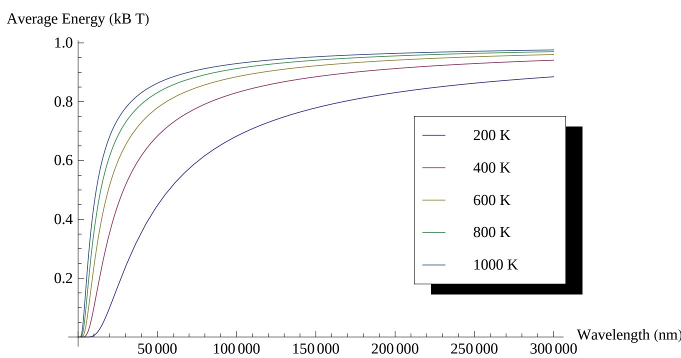

Figure 5: Planck oscillator energy $\langle \epsilon \rangle_\lambda$ vs wavelength, normalized to $k_B T$

3 Statistical Mechanics: The Bridge from the Tiny to the Many

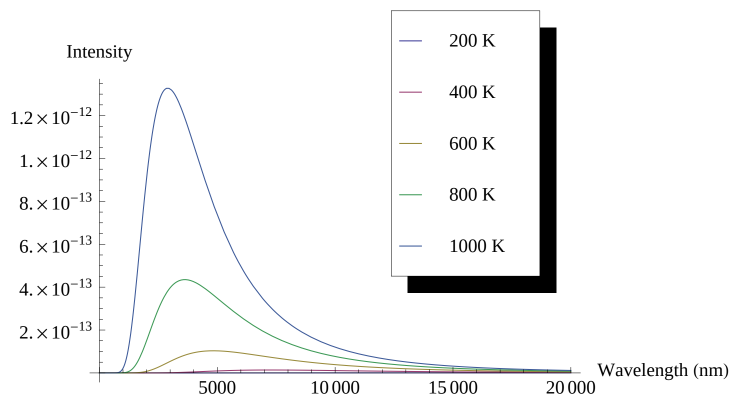

Figure 6: Black body radiation intensity $I(\lambda, T)$ vs wavelength

