# CTAPIPE:

experimental framework for CTA event data processing

github.com/cta-observatory/ctapipe

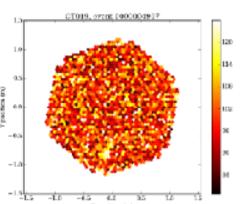
## Goal

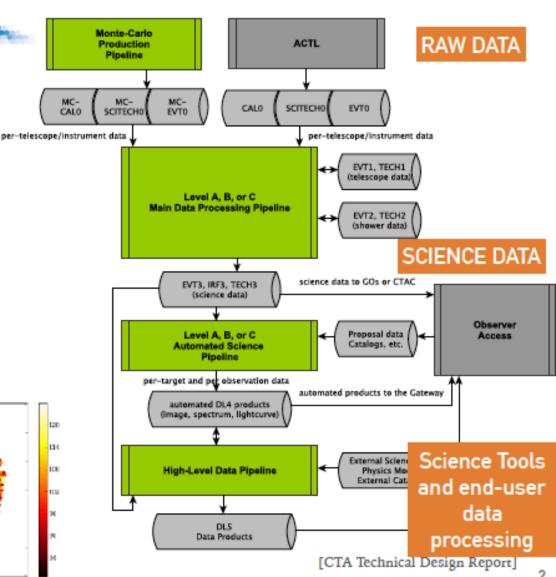
#### Process raw data:

- images of air-showers products by gamma-rays or cosmic rays, seen by an array of telescopes
- Calibrate Camera data
- Process Camera Images
- Reconstruct showers
- Select gamma-like events

#### Produce science data products:

- event lists (photons + bg)
- instrument response tables
- These are then given to endusers to do science





## Building a Framework

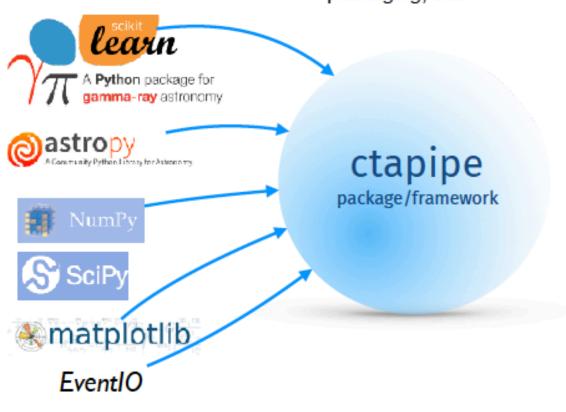
python at all)

### **Bottom-Up approach** Top-Down approach start here Python Python Numba, Cython C/C++ C/C++ start here Most current frameworks did Our approach: start early with python and high-level it this way (if they use

API

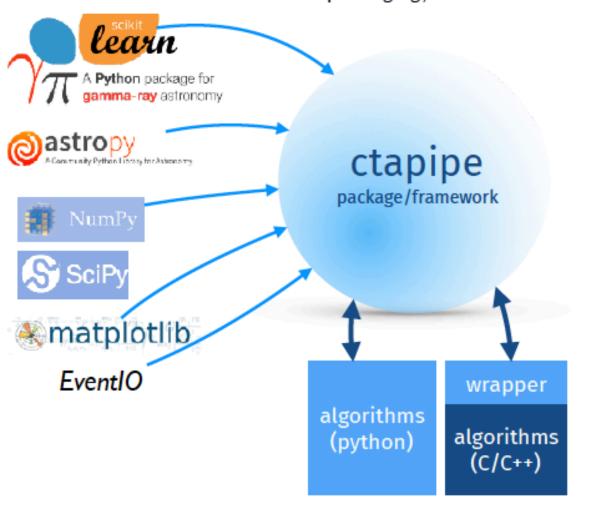
ctapipe will be glue between various components.

Provides common APIs and user interfaces packaging, etc.



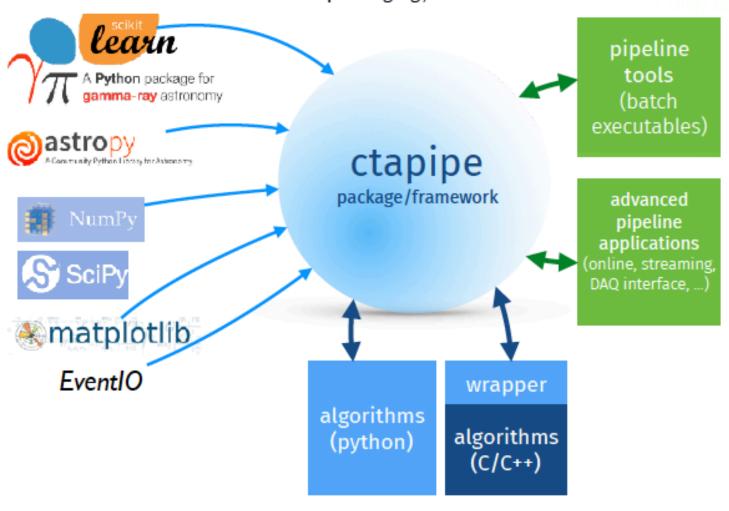
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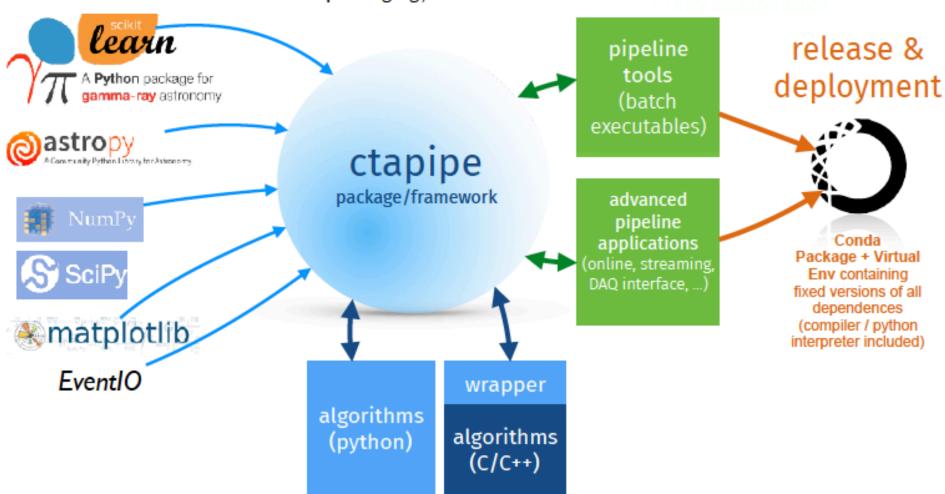


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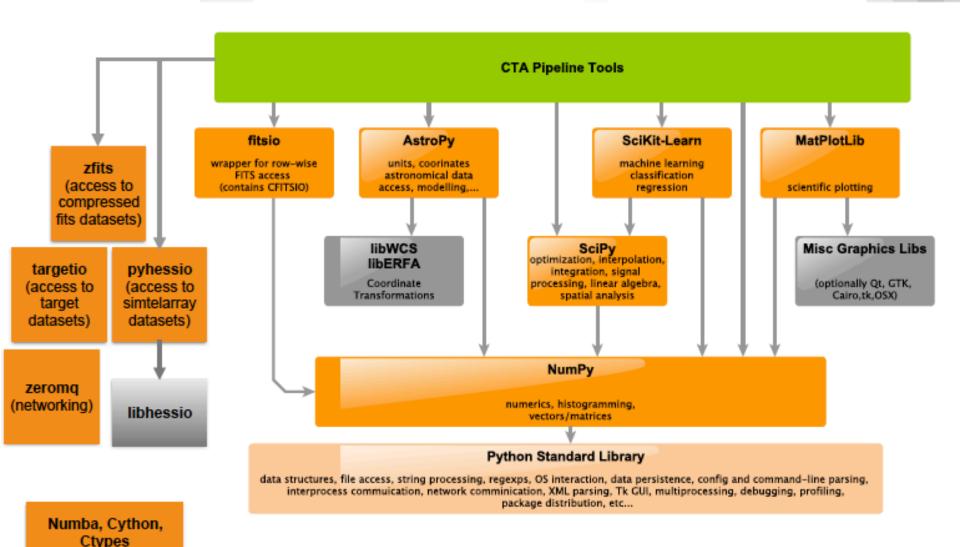


**ctapipe** will be **glue** between various components. Provides common APIs and user interfaces packaging, etc.



## Core dependencies

speed improvements and c-library access



leverage code developed by wide scientific and industrial community!

## Core Data Structures

#### Container

- a class with metadata per element (name, type default, helpstring)
- nested hierarchies supported
- · conversion to dict, flattened dict
- · schema defined in class definition, attributes locked

#### Component

- Simple class that wraps an algorithm and handles algorithm configuration params
- currently based on traitlets.conflg.Conflgurable

#### Tool

- · command-line application base class
- currently based on traitlets.conflg.Application (may change)

data.mo.tel[1]

stapipe.io.containers.ECCsmaraZwantContainers

gheto\_electres\_image: reference image in pure photoelectrons, with no noise

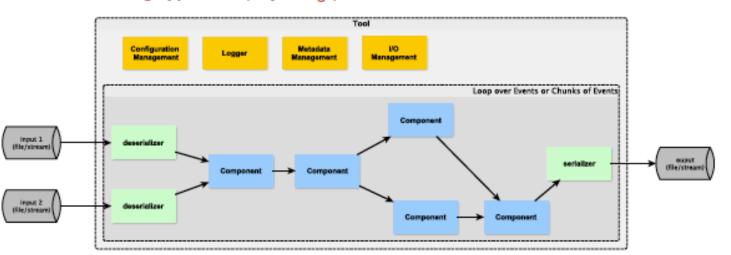
reference\_pulse\_shape: reference pulse shape for each channel

time\_slice: width of time\_slice [ns]

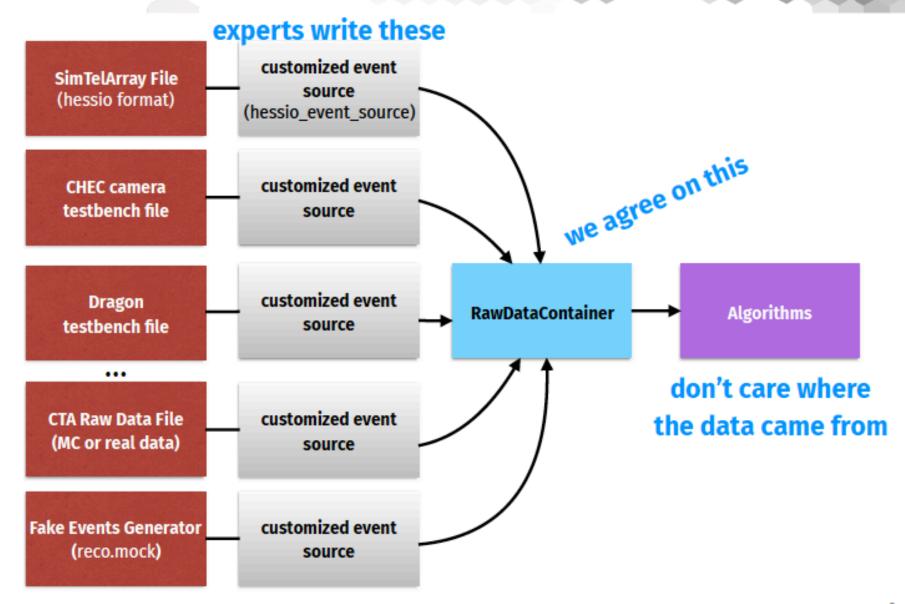
do\_to\_pe: DC/PE calibration arrays from NE file
 pudsetol. pudsetol calibration arrays from NE file
 sainuth\_ray: New asfauth engle (redians from x->E) for the
 telescope

altitude\_row: New asfauth engle (redians) for the telescope
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 altitude\_row: the tracking Aslanth corrected for polating
 servers for the telescope

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 servers for the telescope



### Data access



## io: Monte-Carlo Input



## Working with data is supposed to be simple:

```
from ctapipe.io.hessio import hessio_event_source
source = hessio_event_source("gammas.simtel.gz")

for event in source:
    print(event.trig.tels_with_trigger)
```

complex I/O hidden in python generator

set of hierarchical containers for various data items

> attempt to keep the framework lightweight for algorithm designers (*lesson learned*), while supporting advanced processing techniques

### **Camera Calibration**



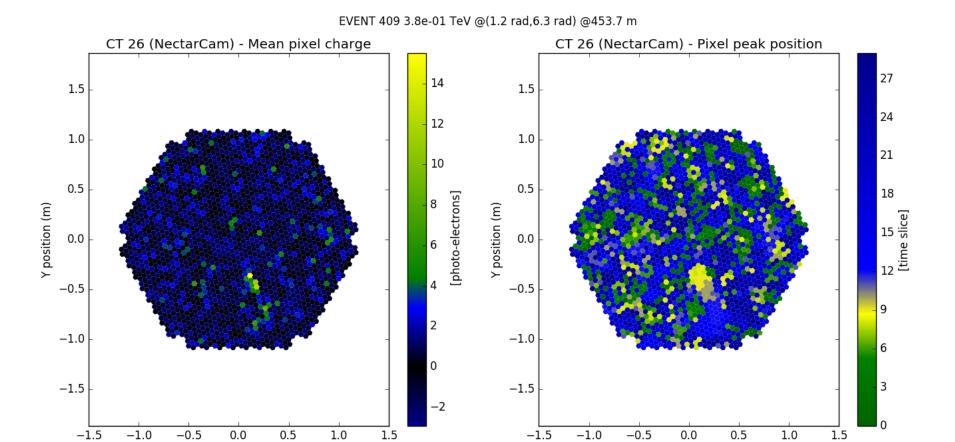
- ctapipe
  - analysis
  - calib
    - array
    - atmo
    - camera
      - tests
        - 🧓 \_\_init\_\_.py
        - able charge\_extractors.py
        - 🧓 dl0.py
        - 🐌 dl1.py
        - 🐌 r1.py
        - README.rst
        - reductors.py

- Split into the 3 data levels
  - R1 Camera/MC pre-calibration into photoelectron (or ADC) samples
  - DL0 Data reduction algorithms before storage
  - DL1 Charge extraction and conversion into photoelectrons, one value per pixel.
- First two calibration stages are not usually done by the offline calibration, except for MC and prototyping. However it is good that the same methods may be available to the offline analysis.

### examples/calibration\_pipeline.py



p examples/calibration\_pipeline.py -D



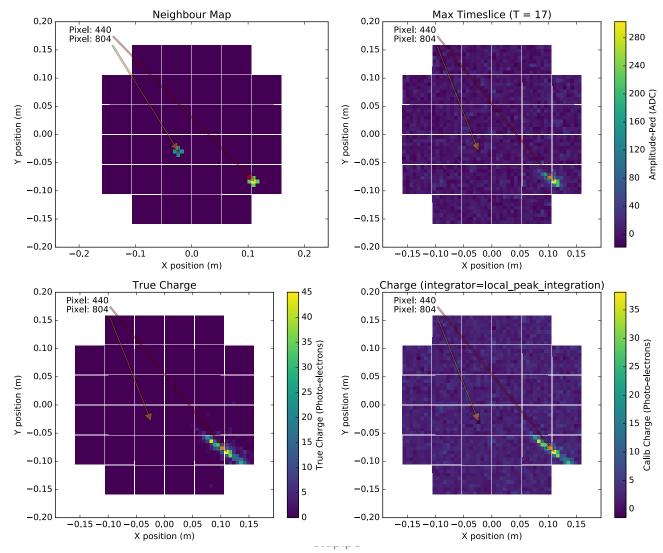
ctapipe MPIK Python Bootcamp 28<sup>th</sup> February 2017

X position (m)

X position (m)

### examples/display\_integrator.py

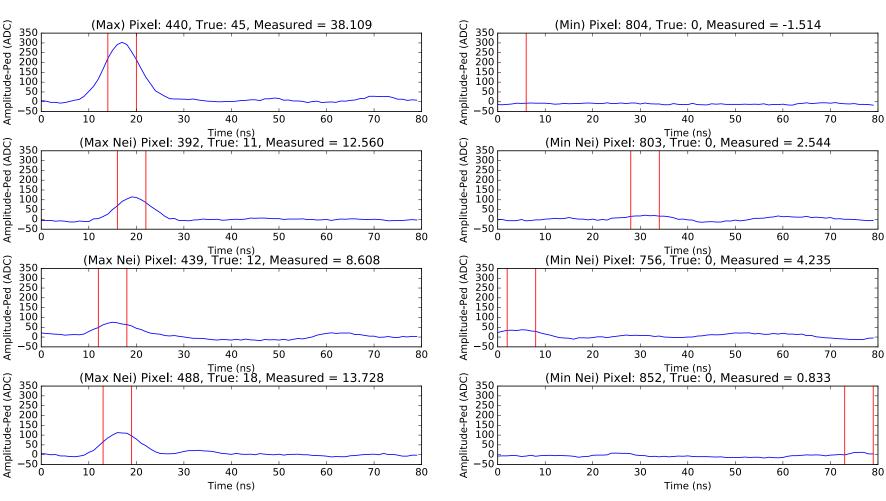




### examples/display\_integrator.py



Integrator = local\_peak\_integration



### **Verification: Charge Resolution**



