

# Using Penalizing Likelihood Method and Nuisance Parameters in Gammapy

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Parameters which are not of intrinsic interest (in contrast to model parameters), but describe any kind of systematic uncertainty.



# Using Penalizing Likelihood Method and Nuisance Parameters in Gammapy

Maximum Likelihood method with an additional penalizing term describing the behavior of the Nuisance Parameters (to avoid over-fitting)



### Applications:

- In general: describing any kind of systematic uncertainty
- Examples:
  - Uncertainty on the effective area (in progress)
  - Uncertainty on the energy reconstruction
  - Uncertainty on the 3D background modeling ....

### **Motivation: Fit Statistics**



 Cash Poisson data with background model

$$C = 2 imes (\mu_{
m sig} + \mu_{
m bkg} - n imes log(\mu_{
m sig} + \mu_{
m bkg}))$$

- N: number of counts, Poisson random variable
- mu\_sig/ mu\_bkg: expected number of counts from the source / bkg
- MapDataset, SpectrumDataset
- Not including background estimation uncertainties

 WStat Poisson data with background measurement

$$W = 2 ig( \mu_{
m sig} + (1 + 1/lpha) \mu_{
m bkg} - n_{
m on} \log \left( \mu_{
m sig} + \mu_{
m bkg} 
ight) - n_{
m off} \log \left( \mu_{
m bkg} / lpha 
ight) ig)$$

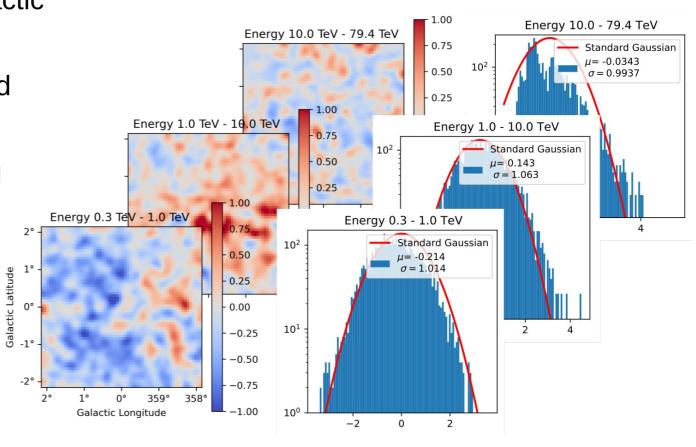
- n\_off/ n\_on: number of counts in OFF region (bkg only) / ON region (bkg + signal)
- Alpha: ratio of ON and OFF acceptances
- MapDatasetOnOff,
   SpectrumDatasetOnOff
- Including background estimation uncertainties

#### **Motivation: Fit Statistics**



- Example: 3D analysis of the Galactic Center
- Residual maps show mismodelled background
- Significance distribution deviating from the Normal distribution

 → Nuisance parameters to describe the systematics due to background mismodelling



# **Penalizing Likelihood Method**



**Nuisance Parameters** 

- Set of Parameters to describe the 'background perturbations'
- Change the background prediction (bin-wise weights)
- Larger Parameters space →
   widening of the Lscan → Larger
   model parameter uncertainties
- N bins → N additional parameters
- Penalizing term to avoid overfitting

$$\begin{aligned} \bar{\mathbf{bg}}_k &= \mathbf{bg}_k \cdot (1 + \vec{\mu}_k) \\ \mathcal{L}_k &= \frac{\mu_k^{n_k}}{n_k!} \exp(-\mu_k) \\ -2 \ln \mathcal{L}_k &= -2 \left[ n_k \ln(\mu_k) - \mu_k \right] \end{aligned}$$

$$\bar{\mathcal{L}}_k = \frac{\mu_k^{n_k}}{n_k!} \exp(-\mu_k) \cdot \exp\left[-\frac{1}{2}\Delta N_k \sum_{l=0}^N (K^{-1})_{kl} \Delta N_l\right]$$
$$-2 \ln \bar{\mathcal{L}}_k = -2 \left[n_k \ln(\mu_k) - \mu_k\right] + \left[\Delta N_k \sum_{l=0}^N (K^{-1})_{kl} \Delta N_l\right]$$

# **Penalizing Likelihood Method**



Penalty Term

- Gaussian Penalty term
- Correlation matrix K
  - Describing the systematic
  - Matrix-product of the spatial and spectral correlation matrices
  - Gaussian: parameterized by amplitude and length

$$\begin{aligned} \mathbf{b} \mathbf{g}_k &= \mathbf{b} \mathbf{g}_k \cdot (1 + \vec{\mu}_k) \\ \mathcal{L}_k &= \frac{\mu_k^{n_k}}{n_k!} \exp(-\mu_k) \\ -2 \ln \mathcal{L}_k &= -2 \left[ n_k \ln(\mu_k) - \mu_k \right] \end{aligned}$$

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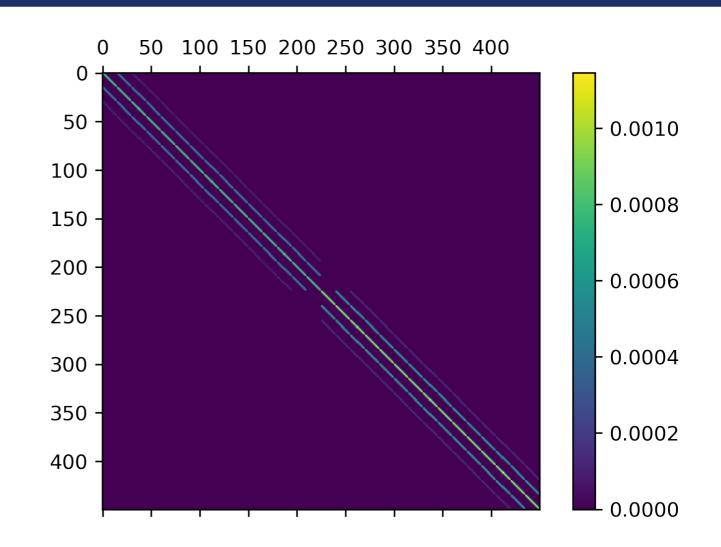
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## **Penalizing Likelihood Method**



Penalty Term

- Gaussian Penalty term
- Correlation matrix K
  - Describing the systematic
  - Matrix-product of the spatial and spectral correlation matrices
  - Gaussian: parameterized by amplitude and length
  - (two analysis methods to determine the values)







Implementation

```
class MapDatasetNuisance(MapDataset):
    """Map dataset for likelihood fitting with nuisance parameters to describe the systematics.
inv_corr_matrix: `~numpy.array`
```

Inverted correlation matrix to describe the correlations between the nuisance parameters. Needs to have same dimension as nuisance parameters.

N\_parameters : `~ gammapy.modeling.Parameters`

Nuisance Parameters to be fitted.

nuisance\_mask: `~ gammapy.map`

Masking the regions where the nuisance parameters are applied to the background.



Reduction of the amount of parameters (computation time): Not one parameter for each bin, but averaging the systematic over multiple neighbor bins (depending on the angular size of the systematics)





Best Fit Nuisance Parameters

Energy 0.7 TeV - 1.0 TeV

5<sup>h</sup>48<sup>m</sup> 36<sup>m</sup> 24<sup>m</sup>

Right Ascension

Implementation – Nuisance Parameters

```
def N map(self):
    """Map of the Nuisance parameters
                                                                                                Energy 0.6 TeV - 0.7 TeV
    Helper function to evaulate the nuisance parameters in the correct geometry.
                                                                                                                    - 10
    Returns
                                                                                            E 22°
    -----
    npred background : `Map`
        Predicted counts from the nuisance parameters.
                                                                                                                    -10
    if self.nuisance mask is not None:
                                                                                                  5<sup>h</sup>48<sup>m</sup>
                                                                                                      36<sup>m</sup>
                                                                                                          24<sup>m</sup>
        from scipy.interpolate import interp2d
                                                                                                    Right Ascension
        N map = Map.from geom(self.geoms['geom down'])
        N map.data[np.where(self.nuisance mask== True)] = self.N parameters.value
        fac = int(self.geoms['geom'].data shape[1] / self.geoms['geom down'].data shape[1])
        N map up = Map.from geom(self.geoms['geom'])
        x = self.geoms['geom down'].to image().get idx()[0][0] * fac + fac/2
        y = self.geoms['geom down'].to image().get idx()[0][0] * fac + fac/2
        x new = self.geoms['geom'].to image().get idx()[0][0]
        y new = self.geoms['geom'].to image().get idx()[0][0]
        for e, z e in enumerate(N map.data):
             f = interp2d(x = x, y = y, z = z e, kind='cubic', fill value = None, bounds error = False)
             N map up.data[e,:,:] = f(x \text{ new}, y \text{ new}).reshape(
                 np.shape(N map up.data[0,:,:]))
    return N map up
```

Energy 1.0 TeV - 1.3 TeV

Right Ascension

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Implementation – Nuisance Parameters

```
def npred background(self):
    """Predicted background counts
    The predicted background counts depend on the parameters
    of the `FoVBackgroundModel` defined in the dataset and on the values of the nuisance parameters.
    Returns
    npred background : `Map`
        Predicted counts from the background.
    background = self.background
    if self.background model and background:
        if self. background parameters changed:
            values = self.background model.evaluate geom(geom=self.background.geom)
            if self. background cached is None:
                self. background cached = background * values
            else:
                self. background cached.quantity = (
                    background.quantity * values.value
        if self.N parameters is not None:
            self. background cached.data = self. background cached.data* (1+ self.Nap().data)
        return self. background cached
    else:
        return background
    return background
```





Implementation - Likelihood

$$-2\ln\bar{\mathcal{L}}_k = -2\left[n_k\ln(\mu_k) - \mu_k\right] + \left[\Delta N_k \sum_{l=0}^N (K^{-1})_{kl} \Delta N_l\right]$$

```
def stat_array(self):
    """Likelihood per bin given the current model parameters + Gaussian of N parameters"""
    G = np.matmul(self.inv_corr_matrix ,self.N_parameters.values.ravel() )*self.N_parameters.values.ravel()
    stat = cash(n_on=self.counts.data, mu_on=self.npred().data)
    return G+stat

def stat_sum(self):
    """Total likelihood given the current model parameters + Gaussian of N parameters."""
    counts, npred = self.counts.data.astype(float), self.npred().data
    G = np.matmul(np.matmul(self.inv_corr_matrix ,self.N_parameters.value.ravel() ),self.N_parameters.value.ravel() )
    if self.mask is not None:
        return cash_sum_cython(counts[self.mask.data], npred[self.mask.data]) + G
    else:
        return cash sum cython(counts.ravel(), npred.ravel()) + G
```

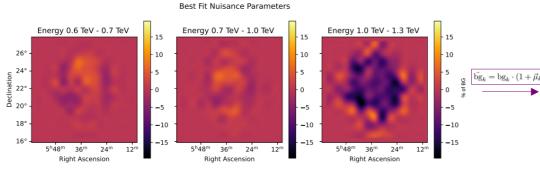
Additional parameters are used in methods like: from\_hdulist, slice\_by\_idx, etc.

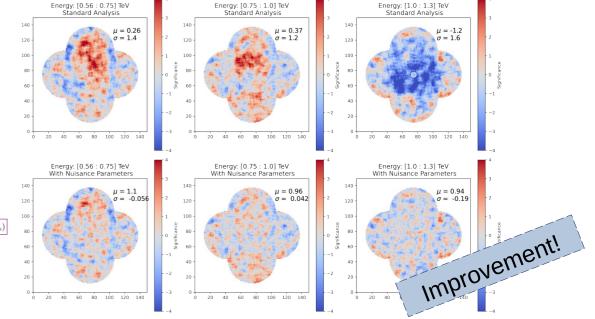
### **Fitting**

# H.E.S.S.

#### Results and Comparison to Standard Method

- Obtain best fit nuisance and model parameters
- Improved Significance map and distribution
- Model parameter errors including statistical and systematic errors



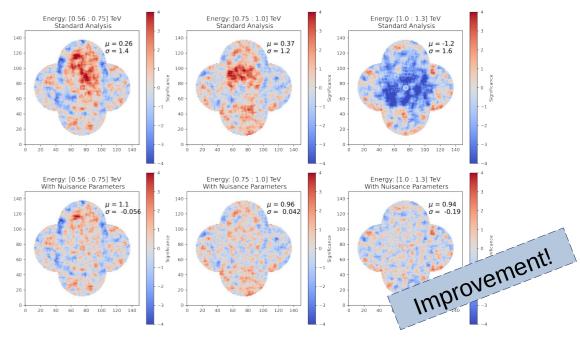


	Amplitude [1e-11 cm-2 s-1 TeV-1]	Spectral Index	Bg Normalisation
Standard Analysis	$3.729 \pm_{stat} 0.025$	$2.557 \pm_{stat} 0.0143$	$0.997 \pm_{stat} 0.002$
With Nuisance Parameters	$3.739 \pm_{stat} 0.025 \pm_{sys} 0.001$	$2.555 \pm_{stat} 0.0143 \pm_{sys} 0.0003$	$1.007 \pm_{stat} 0.002 \pm_{sys} 0.0001$

# **Summary**



- Binned Maximum Likelihood method in 3D with Gaussian penalizing Prior term
- Nuisance parameters to describe systematic uncertainties of unknown distribution
- Advantages:
  - Improved background description
  - Automatic computation of the systematic error on the model parameters
- Outlook:
  - Testing and quantifying improvement and robustness (Asimov datasets)
  - Application (for now mainly GC)
  - Including other sources of systematics (effective area, energy reconstruction)





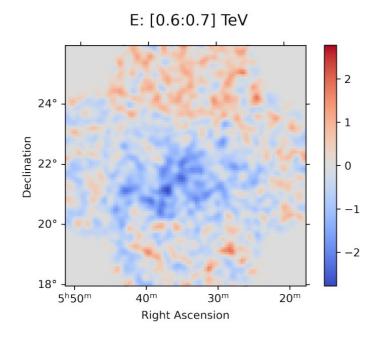




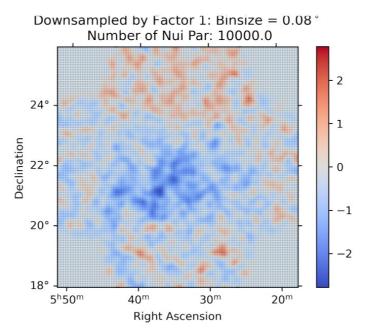
# Backup

#### **Method of Down-sampling**

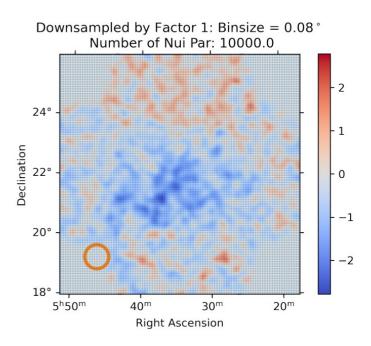




Smoothed spatial residual of the standard analysis



Adding the spatial nuisance parameters in each bin



Visualisation of the Gaussian spatial correlation length of 0.8 deg = 100 strongly correlated nuisance parameters

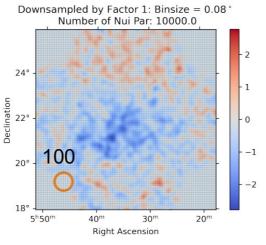
#### **Method of Down-sampling**

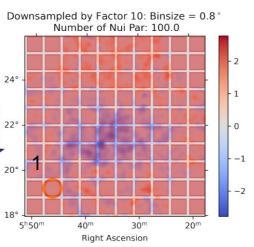


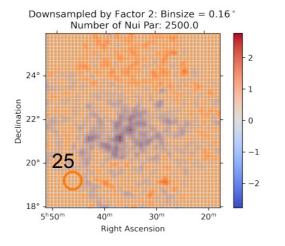
#### Idea:

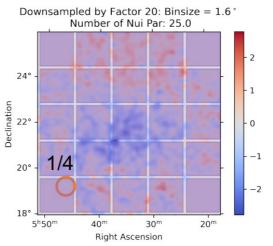
Use one nuisance parameter as average over multiple spatial neighbor bins!
Use different binning for the nuisance parameters than for counts cube

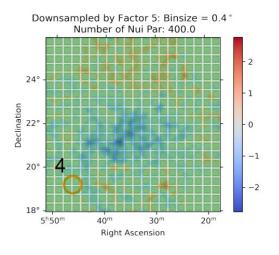
Choose the bin size such that one bin is correlation length

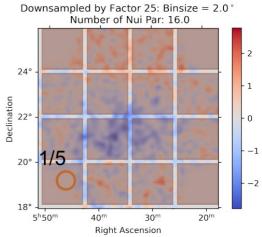












#### **Fourier Trafo to Obtain Correlation Length**



- What is the optimal amplitude σ and correlation length?
- Spatial residual of the standard analysis:
  - Angular spectrum from 2D Fourier transformation

 Compare with angular spectrum of datasets containing only statistical fluctuations

