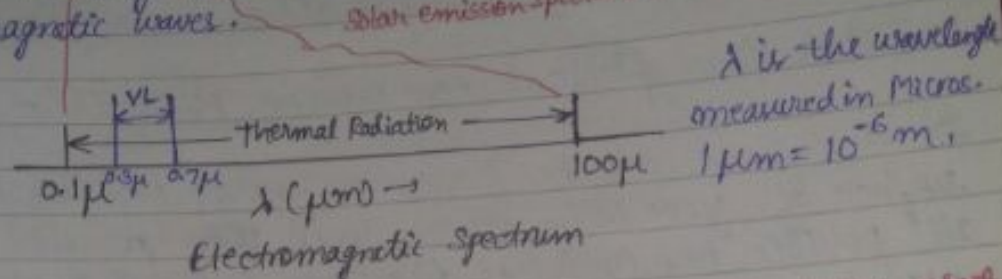


RADIATION

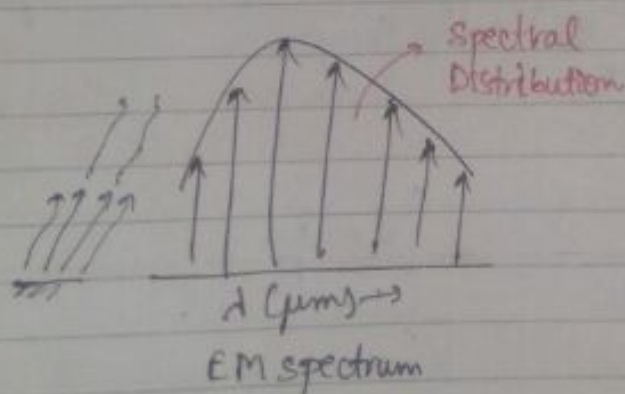
→ All bodies at all temperatures emit thermal radiation except the body at 0K or -273.15°C in the form of electromagnetic waves.



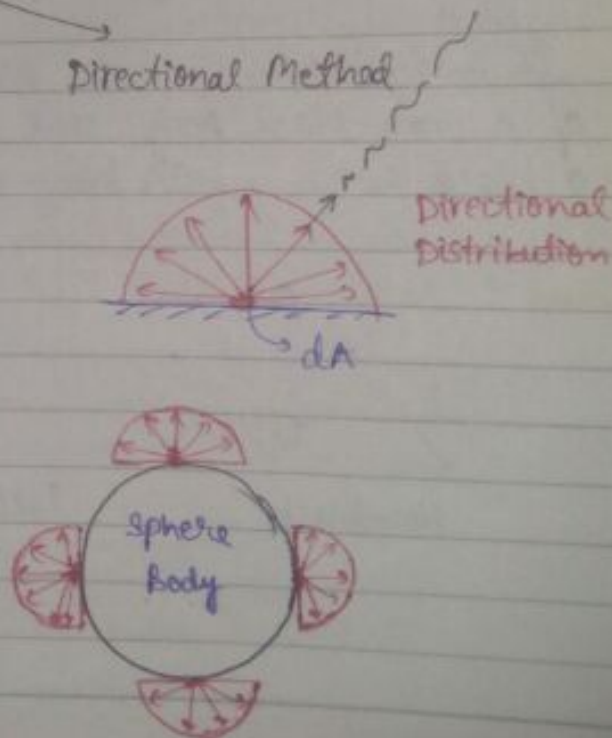
* All high temp. bodies like SUN, not only emits very great amount of thermal radiation but also ~~the~~ most of the radiation occurs at short wavelengths.

Thermal Radiation

Wavelength Module



Directional Method



* Any body at any temp will emit thermal radiation in all hemispherical directions w/ differential area considered on the body and at all probable wavelengths on the electromagnetic spectrum (but in the range of 0.1 to $100 \mu\text{m}$).

Basic Definitions of Radiation HT -

1- Total Hemispherical emissive power (E) -

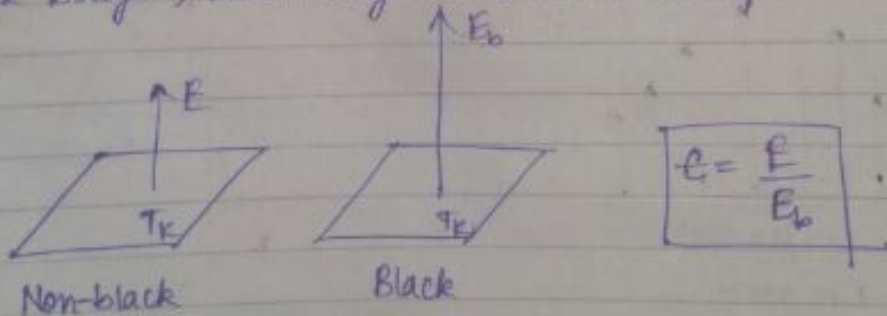
It is defined as the radiation energy emitted from the surface of your body per unit time per unit area in all possible hemispherical directions integrated over all the wavelengths.

$$E \rightarrow \text{Joule/sec-m}^2 \rightarrow \text{W/m}^2.$$

$$E = f(T) \quad \text{Kelvin}$$

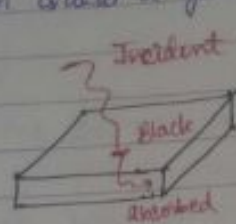
2- Total Emmissivity (ϵ) -

It is defined as the ratio b/w total hemispherical emissive power of a non-black body and total hemispherical emissive power of a black body, both being at the same temp.



Black body - It is the body which absorbs all the thermal radiation incident or falling upon the body.

- Body is absorbing the radiation and on the same time emitting.
- Radiation is only dependent on temp. of the body.
- Good emitter is also a good absorber (applicable even for non-black body)



Black body

- Perfect absorber
- Ideal emitter
- Diffusive



A narrower tiny cavity in a hollow spherical container is a black body.

- Examples -
- (i) A small hole in a furnace wall is black body.
 - (ii) In radiation analysis, sun is also treated as black body.

Note: A thermally black body absorbing all the incident thermal radiation may not appear black in colour to human eye.

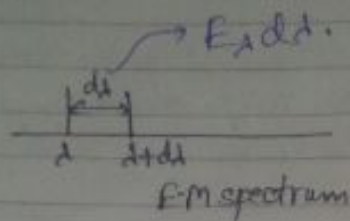
Ex: eg; ice and snow are thermally black.

* For any body,

$\epsilon \leq 1$. → For any body

$\epsilon_b = 1$. → For black body

3- Monochromatic (or Spectral) hemispherical emissive power (E_λ) -



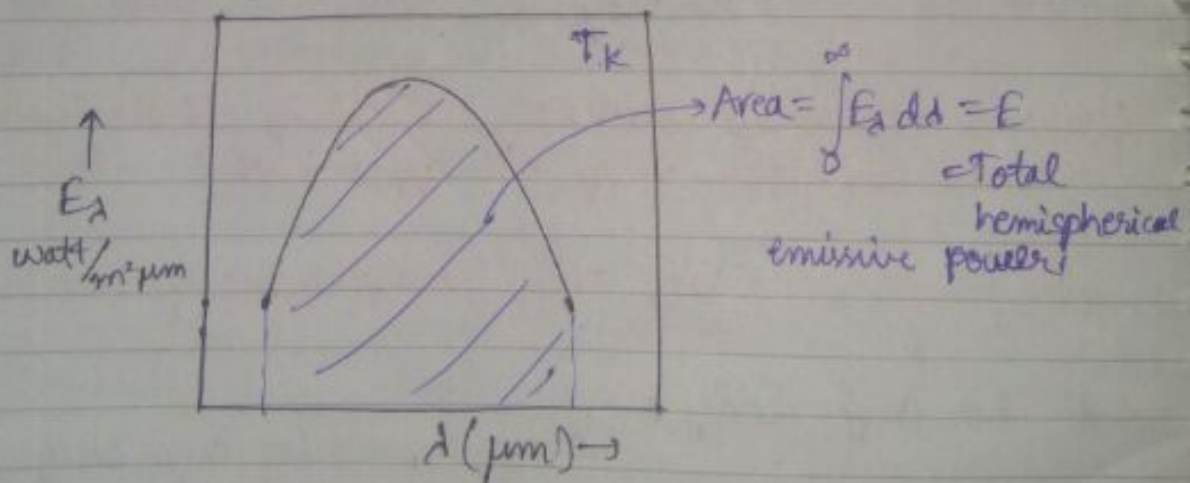
$d\lambda \rightarrow$ Differentially small increment in wavelength (λ)

E_λ at a particular wavelength λ is defined as the quantity which when multiplied by $d\lambda$ shall give the radiation energy emitted from the surface of the body per unit time per unit area in the wavelength region λ to $\lambda + d\lambda$.

$$E_\lambda \rightarrow \text{Joule/sec-m}^2\mu\text{m} \rightarrow \text{Watt/m}^2\mu\text{m}.$$

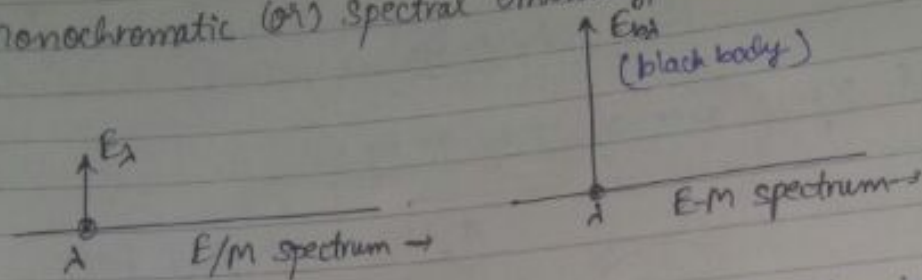
\rightarrow For any given body at a given temperature,

$$E_\lambda = f(\lambda).$$



Total Hemispherical
Emissive power $= E = \int_0^\infty E_\lambda d\lambda \text{ W/m}^2.$

4-Monochromatic (or) Spectral emissivity (ϵ_λ) -



It is defined as the ratio b/w monochromatic hemispherical emissive power of a non-black body and monochromatic hemispherical emissive power of a black body, both being at the same temp. and wavelength.

$$\boxed{\epsilon_\lambda = \frac{E_\lambda}{E_{\lambda b}}} \Rightarrow E_\lambda = \epsilon_\lambda E_{\lambda b}$$

$$E = \frac{E}{E_b} = \frac{\int_0^\infty E_\lambda d\lambda}{\int_0^\infty E_{\lambda b} d\lambda} = \frac{\int_0^\infty \epsilon_\lambda E_{\lambda b} d\lambda}{\int_0^\infty E_{\lambda b} d\lambda}$$

$$E = \frac{\epsilon_\lambda \int_0^\infty E_{\lambda b} d\lambda}{\int_0^\infty E_{\lambda b} d\lambda} \Rightarrow \boxed{E = \epsilon_\lambda}$$

If $\epsilon_\lambda \neq f(\lambda)$ or rather remaining constant.

Gray Body -
 Such body whose monochromatic E_λ is independent of wavelength (λ) or rather remaining constant is known as gray body or gray surface. i.e. for gray body

$$[E_\lambda = \text{constant}]$$

Physical significance of gray body -

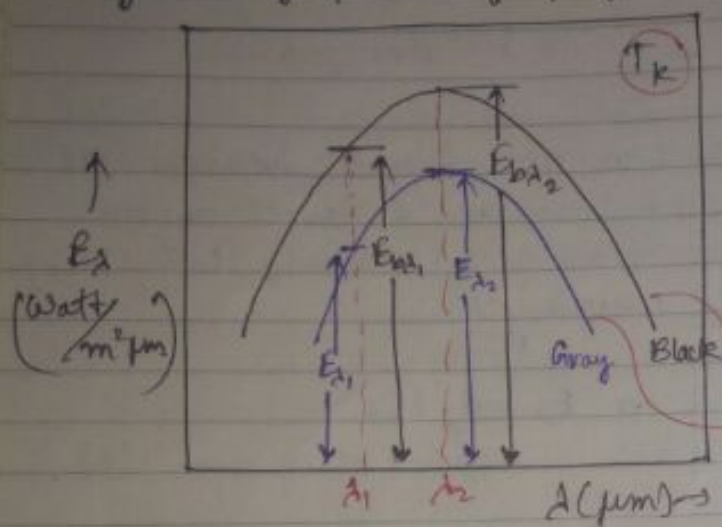


Fig. (i)

For gray body,
 $E_\lambda = C \cdot f(\lambda)$
 i.e. $E_{\lambda_1} = E_{\lambda_2}$

$$\Rightarrow \frac{E_{\lambda_1}}{E_{\lambda_1}} = \frac{E_{\lambda_2}}{E_{\lambda_2}}$$

Both the curves are in ~~asynchronous~~ ^{synchronous},
 (i.e. varying in synchronous)

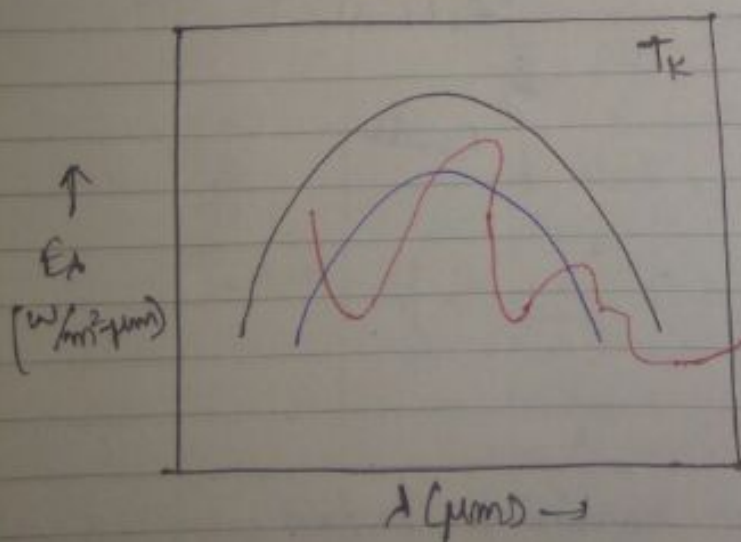


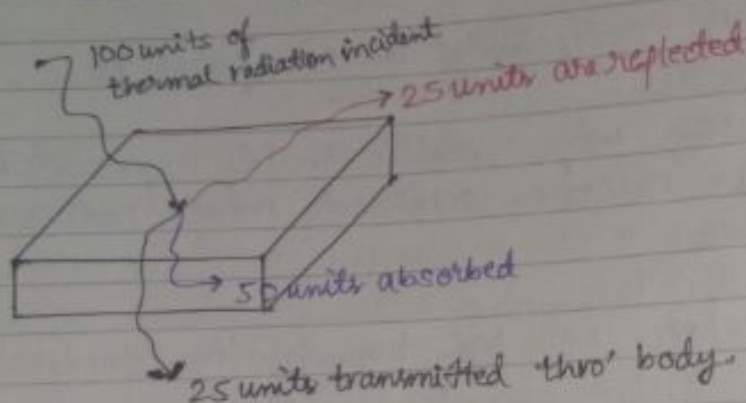
Fig. (ii)

Non-gray surface
 $E_\lambda = f(\lambda)$

Fig 1.1

— The ratio b/w the area under bottom curve on λ -axis & the area under the top curve on λ -axis will be equal to total emissivity of gray body which happens to be equal to its monochromatic emissivity.

Absorptivity (α), Reflectivity (ρ), Transmissivity (τ)



Absorptivity (α) = $\frac{50}{100} = 0.5$ = Fraction of radiation energy incident upon a surface which is absorbed by it.

Reflectivity (ρ) = $\frac{25}{100} = 0.25$ = Fraction of radiation energy incident upon a surface which is reflected by it.

Transmissivity (τ) = $\frac{25}{100} = 0.25$ = Fraction of radiation energy incident upon a surface which is reflected & transmitted thro' it.

* All these effects are taking into place altogether

∴ For any surface, $\alpha + \rho + \tau = 1$.

For opaque surface, which doesn't transmit any energy, $\tau = 0$.

∴ For opaque surface, $\alpha + \rho = 1$.

→ For black body, which absorbs all radiation energy incident

ie. $\alpha_b = 1$,
and $\epsilon_b = 1$.

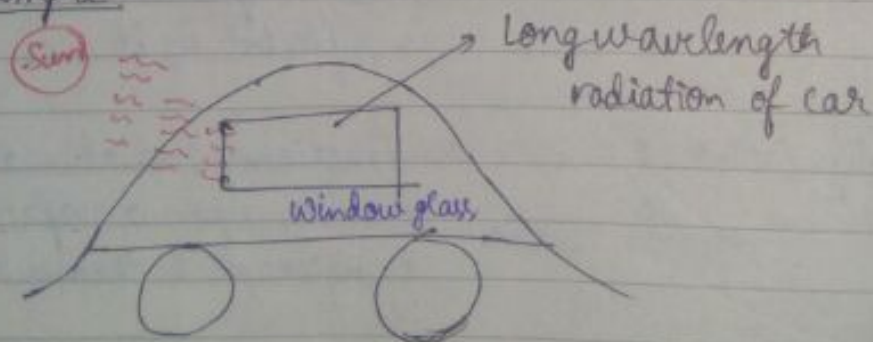
* Metals have high reflectivity as compared to non-metals.

→ This is the reason why metallic shields are provided in the furnaces as radiation screens to reduce radiation heat exchange.

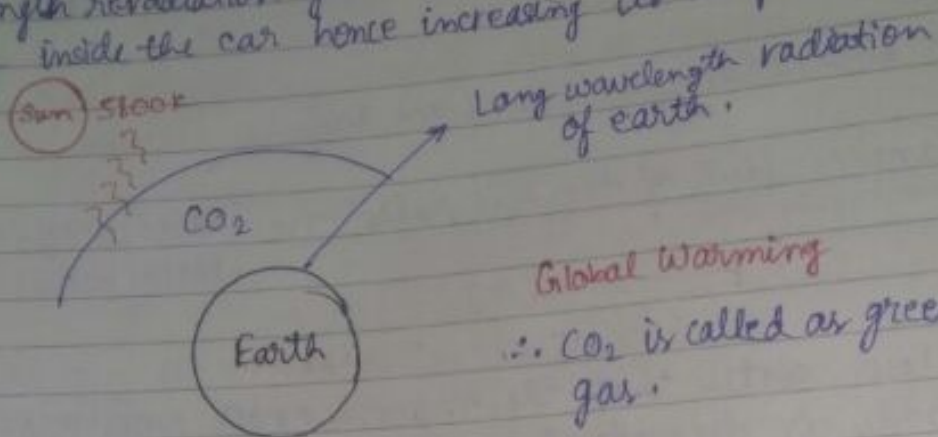
* Gases like O_2 , N_2 , etc. have high transmissivity (transparent to thermal radiation).

* The above radiation properties mentioned change with wavelength of incident thermal radiation, surface roughness of the body and also with its temp.

Example:-



The window glass of a car is very much transparent (i.e. $\epsilon \approx 1$) to the short wavelength solar radiation, but the same window glass almost becomes opaque to the long wavelength re-radiation given by inside of the car thus trapping energy inside the car hence increasing its temp.



\therefore CO₂ is called as green house gas.

→ H₂O (water vapour) is also a green house gas that's why deserts are hotter in day and cold at night.

→ As the surface roughness of the body ↓ by polishing it, the reflectivity (ρ) of the surface ↑. This is the reason why highly polished copper & aluminium sheets having very good reflectivity are generally used in the furnaces to reduce radiation heat exchange.

Laws of Thermal Radiation -

1- Kirchhoff's Law of thermal radiation - The law states that whenever a body is in thermal equilibrium with its surroundings, its emissivity is equal to its absorptivity.

$$\boxed{\alpha = \epsilon}$$

A good absorber is always a good emitter.
eg. Black body ($\alpha = \epsilon = 1$).

→ The law is also supposed to be valid even under non-equilibrium condⁿ.

→ For both black & non-black body.

2- Planck's law of thermal radiation - Valid only for black body.

$$E_{\lambda} = f(\lambda, T).$$

T in Kelvin.

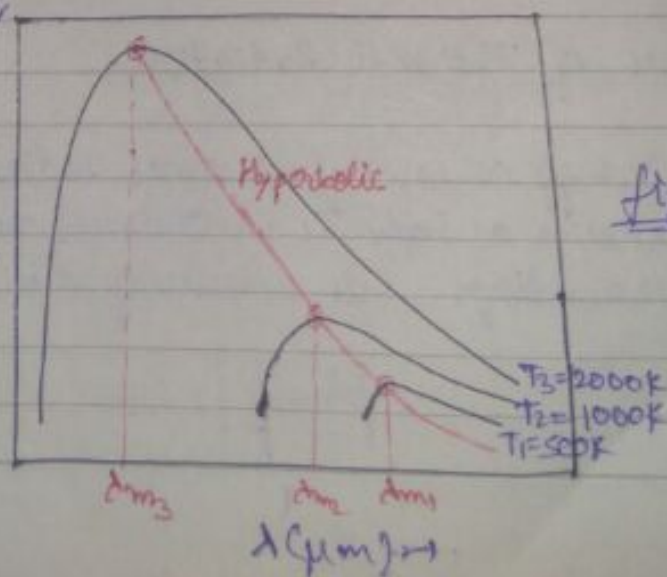
The law states that the monochromatic emissive power of a blackbody is dependent on both absolute temp. of black body and also on wavelength of radiation energy emitted λ (in μm).

$$E_{\lambda} = \frac{2\pi C_1}{\lambda^5 (e^{\frac{C_2}{\lambda T}} - 1)} \text{ Watt/m}^2\mu.$$

C_1 & C_2 are experimental constants.

The above functional relationship among the 3 variables can be graphically plotted as

λ_m = Wavelength at which E_{λ} is max. at a given absolute temp of blackbody.



Note: At a given absolute temp. of a black body, as wavelength λ increases, E_{λ} also increases reaches a maximum & then decreases.

Also as the absolute temp. of black body \uparrow (each time getting doubled) E_{λ} values enormously increase but now most of thermal radiation at higher temp. will be shifted to shorter (smaller) wavelength.

As T increases $\Rightarrow \lambda_m$ decreases.

$$\Rightarrow \lambda_m \propto \frac{1}{T}$$

This law is used in Optical pyrometer.
(very high temp. measurement)

$$\lambda_m T = a = \text{constant} = 2898 \mu\text{m.K.}$$

This equation is called as Wein's Displacement Law.⁽³⁾

\Downarrow
Valid only for black body.

\rightarrow Since sun is a blackbody, for solar radiation

$$(\lambda_m)_{\text{solar}} \times T_{\text{solar}} = 2898$$

$$(\lambda_m)_{\text{solar}} = \frac{2898}{5800} \approx \frac{1}{2} \mu\text{m.}$$

4- Stefan's Boltzman's Law-

The total hemispherical emissive power of a black body is directly proportional to the fourth power of the absolute temp. of the black body.

$$E_b \propto T^4$$

$$E_b = \sigma T^4 \text{ W/m}^2.$$

$$\sigma = \text{Stefan-Boltzmann's constant.}$$

$$= 5.67 \times 10^{-8} \text{ W/m}^2 \cdot \text{K}^4.$$

Proof -

$$E_b = \int_0^\infty E_m d\lambda \text{ W/m}^2.$$

$$= \int_0^\infty \frac{2\pi C_1}{\lambda^5 \left(e^{\frac{C_2}{\lambda T}} - 1 \right)} d\lambda.$$

$$= \sigma T^4 \text{ W/m}^2.$$

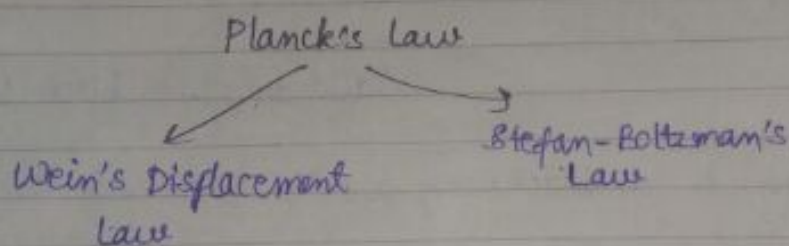


Fig. 15

* Note: The ratio b/w the area under topmost curve on x-axis and the bottom most curve on x-axis will be equal to 4^4 i.e. 256.

→ For a non-black body whose emissivity is ϵ , its total hemispherical emissive power $= E = \epsilon E_b$.

$$= \epsilon \sigma T^4 \text{ W/m}^2.$$

If 'A' is the total surface area of non-black body.

Then, radiation energy emitted from entire non-black body $= EA \text{ watt} = \epsilon \sigma T^4 A \text{ watt}.$

23- $E_1 = \epsilon_1 \sigma T_1^4 = 500$

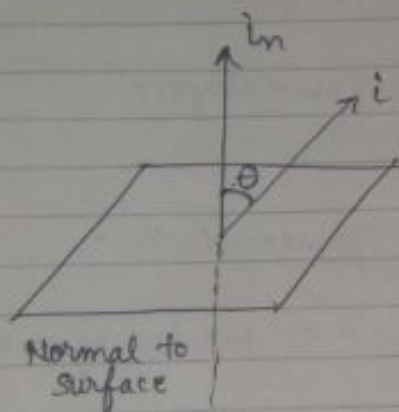
(c) $E_2 = \epsilon_2 \sigma T_2^4 = 1200$

$$\frac{T_1}{T_2} = \left(\frac{5}{12}\right)^{1/4} = 0.802$$

31- ~~$E_{\text{rad}} = T \Delta m T^4$~~ $dm T = \text{Constant}$,
 (b) $0.5 \times 5800^\circ = dm_2 \times 10000$

$$dm_2 = 2.90 \mu\text{m}$$


Lambert's Cosine Law-



$$i = i_n \cos \theta$$

i_n = Normal intensity of radiation.
 i = Intensity of radiation along a dirⁿ making an angle ' θ ' w.r.t. normal direction.

$i \left(\frac{\text{Joule}}{\text{sec-m}^2\text{-steradian}} \right) = \text{W/m}^2\text{-steradian}$



Steradian is the unit of solid angle.

Intensity of radiation (i) along an given dirⁿ is defined as the radiation energy emitted from the surface of the body per unit time per unit area normal to that dirⁿ & per unit solid angle

about that dirⁿ.

$$i = \frac{dE}{d\omega}$$

Watt / m²-steradian

i is a sum of all the radiations of all the wave lengths in one direction.

∴ Total hemispherical emissive power = $E = \int i d\omega$ W/m².

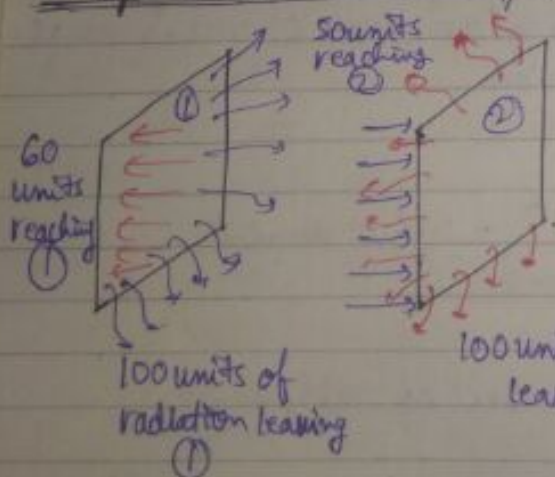
→ A diffuse surface has the same intensity of radiation along all the direction i.e. for diffuse surface i is independent of dirⁿ. eg. blackbody

For blackbody, $E_b = \pi i_b$ W/m².



A Diffusive Surface

Shape Factor or View factor or Configuration Factor -



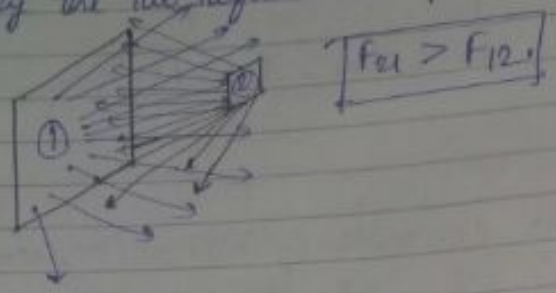
$F_{12} = \frac{60}{100} = 0.6$ = Fraction of radiation energy leaving surface ① that reaches surface ②.

$F_{21} = 0.6$ = Energy leaving ② reaches ①.

In general, F_{mn} = Fraction of radiation energy leaving surface (m) that reaches surface (n).

$$0 \leq F_{mn} \leq 1.$$

The shape factor b/w 2 surfaces is independent of their temp., emissivity, wavelengths of emission but depends only on how closely their sizes, their shapes and how closely the two surfaces are kept w.r.t each other.



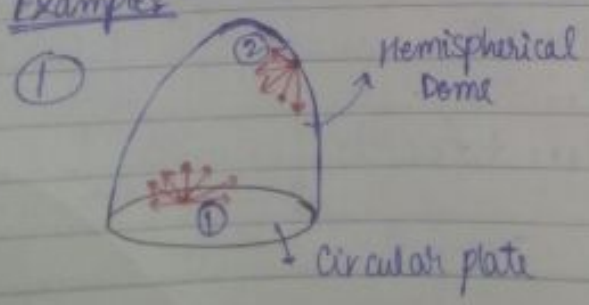
→ $F_{11} = 0$
 ① Flat Surface.

→ $F_{11} > 0$
 ① Concave surface

→ $F_{11} = 0$
 ① Convex surface

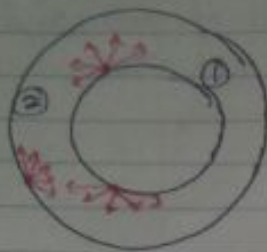
When one body or a surface is completely surrounded by another body as a surface, the shape factor of inner body w.r.t. the outer is equal to one.

Examples -



$$\begin{aligned} F_{12} &= 1. \\ F_{21} &< 1 \\ F_{22} &< 1 \\ F_{11} + F_{22} &= 1. \end{aligned}$$

② Two concentric spherical ~~ep~~ surfaces -



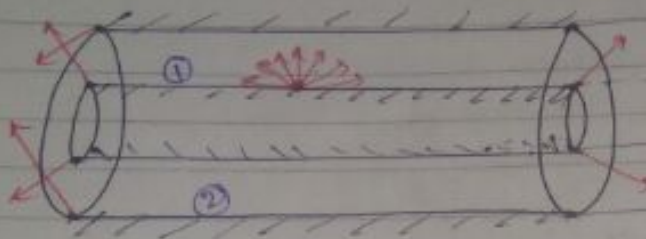
$$F_{12} = 1$$

$$F_{21} < 1$$

$$F_{22} < 1$$

Here also, $F_{21} + F_{22} = 1$.

③ Two infinitely long concentric cylindrical surfaces -



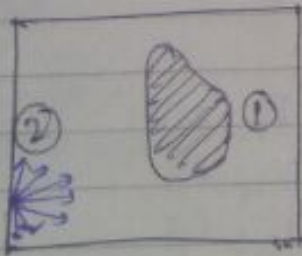
end view

$$F_{12} = 1 \quad (\text{Neglecting end effects}), \text{ as long cylinder}$$

$$F_{21} < 1 \quad (\text{See in end views}).$$

$$F_{22} < 1 \quad F_{21} + F_{22} = 1 \quad (\text{Neglect end effects}).$$

④ A body kept in an enclosure -



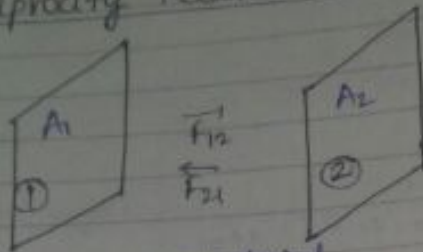
$$F_{12} = 1.$$

$$F_{21} < 1$$

$$F_{22} < 1$$

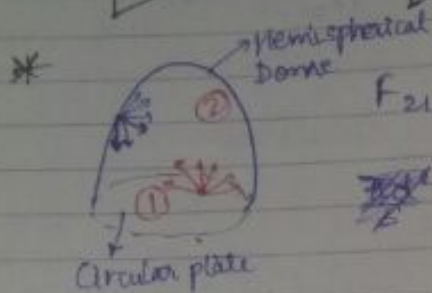
$$F_{21} + F_{22} = 1.$$

Reciprocity Relation b/w shape factors -



$$A_1 F_{12} = A_2 F_{21}$$

Reciprocity theorem.



$$F_{21} + F_{22} = 1.$$

$$\cancel{\frac{\pi d^2}{4}} \times F_{12} = \cancel{2\pi R^2} \times F_{21}$$

$$F_{21} = \frac{1}{2}.$$

$$F_{22} = 0.5$$

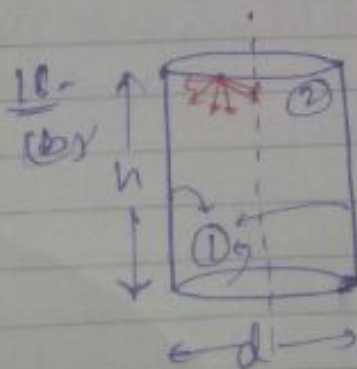
$$34 - F_{12} \times 4\pi d^2 = 4\pi R^2 \times F_{21}$$

$$F_{21} = \frac{4}{9}^{0.44} = 0.44.$$

$$28 - \frac{4\pi d^2}{A_2} \times F_{12} = 6d^2 F_{21}$$

(d)

$$F_{21} = \frac{\pi}{6}$$

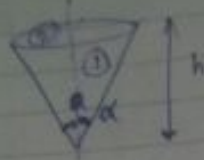


① → Circular Base +
Cylindrical Surface.

$$F_{11} = 1$$

$$F_{12} = \frac{A_2}{A_1} = \frac{d}{d + 4h}$$

19
(b)



$$F_{21} = 1.$$

$$\pi r^2 \times 1 = \frac{1}{2} \pi r l \times F_{12}$$

$$F_{12} = \frac{2}{l}.$$

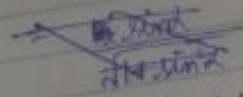
$$\text{Cone height} = \frac{h}{\sin \alpha}$$

$$l = \frac{h}{\cos \alpha}$$

$$r = h \tan \alpha$$

$$\tan \alpha = \frac{r}{h}$$

$$r = h \tan \alpha$$



$$= \frac{K \tan \alpha \times \cos \alpha}{r}$$

$$F_{12} = \sin \alpha$$

20
(b)

$$F_{21} = 0.004.$$

$$0.6 \times 1 = A_{21} \times 0.004$$

$$A_{21} = \frac{0.6 \times 1}{0.004} = 150 \text{ m}^2.$$

$$8a^2 = 150$$

$$a = 5 \text{ m}.$$

Summation Rule among Shape Factors:-

If there are 'n' no. of surfaces involved in any radiation heat exchange then

$$F_{11} + F_{12} + F_{13} + \dots + F_{1n} = 1.$$

$$F_{21} + F_{22} + F_{23} + \dots + F_{2n} = 1.$$

$$F_{m1} + F_{m2} + \dots + F_{mn} = 1.$$

If any particular surface is either flat or convex surface, then its self shape factor becomes zero.

Note: Even now reciprocity is valid ^{rel} b/w any 2 surfaces.
 eg., $A_2 F_{23} = A_3 F_{32}$
 $A_1 F_{1m} = A_m F_{m1}$

2- $F_{11} + F_{12} + F_{13} = 1$
 (d) $F_{12} = 1 - 0.17 = 0.83$

$F_{12} \times \frac{\pi D^2}{4} = \frac{\pi D^2}{4} \times F_{21}$

$F_{21} = 0.2075$

From symmetry of fig.
 $F_{21} = F_{23}$

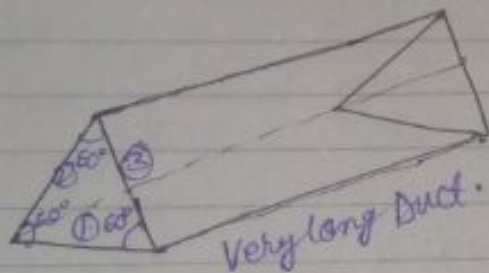
$\therefore F_{23} = 0.2075$

~~$F_{22} = 1 - 0.2075$~~
 ~~$= 0.7925$~~

$F_{22} = 1 - 0.2075 - 0.2075$
 $= 1 - 0.4150$

$F_{22} = 0.585$

*



From symmetry of fig.
 $F_{12} = F_{13}$

Since, duct is very long,
 (Neglecting end effects)

$F_{11} + F_{12} + F_{13} = 1$

$F_{12} = F_{13} = \frac{1}{2}$