

PELEG BAR SAPIR

# PHYSICS SIMULATIONS (LECTURE NOTES)



# 1

## Introduction

### 1.1 *Why are Simulations Used?*

Text

### 1.2 *Python*

Text

### 1.3 *A Bit About Git*

Text

### 1.4 *Some Mathematical Background*

Text

### 1.5 *Harmonic Oscillator*

#### 1.5.1 *Simple Harmonic Oscillator*

Many systems in physics present a simple, periodic (repeating) motion. One such system is a simple mass-less spring connected to a mass  $m$  and allowed to move in a single dimension only. If we ignore the effects of gravity, the only force acting on the mass arises from the spring itself: the more we pull or push the spring, the stronger it will resist to that change. This resistant force is given by

$$F = -kx, \tag{1.1}$$

where  $k$  is the **spring constant**, and  $x$  is the amount by which the spring contracts or expands relative to its rest position  $x_0$ . In Figure 1.1 we show three cases: when the mass is at the springs rest position,  $x_m = x_0$ , the spring applies no force on it. When the mass is displaced by a positive amount  $\Delta x > 0$ , the spring applied a *negative* force on it:  $F = -k\Delta x < 0$  (see note 1.1). And when the mass

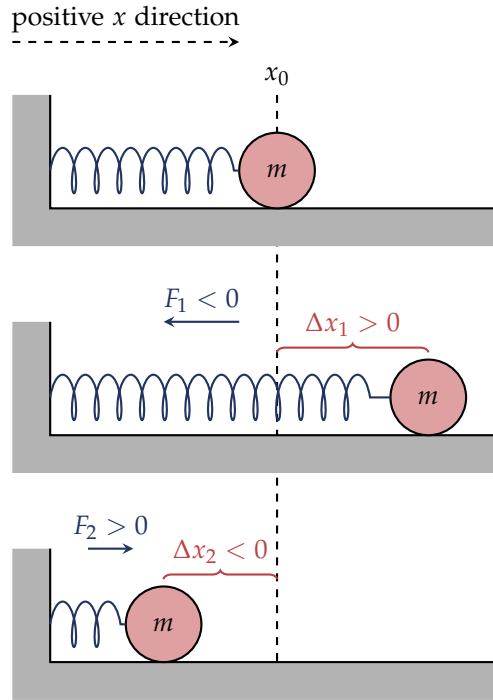
is displaced such that it contracts the spring,  $\Delta x < 0$  and thus the force applied by the spring is positive,  $F = k\Delta x > 0$ .

**Note 1.1 Negative and positive forces**

Recall that in this context, negative force means a force in the negative  $x$  direction.



Figure 1.1: A simple spring-mass system with spring constant  $k$  and a mass  $m$ . The top figure shows the spring at rest - i.e. when the mass is located at position  $x_0$  the spring applies no force on the mass (since  $\Delta x = x_m - x_0 = 0$ ). The middle figure shows the spring being at a *positive* displacement  $\Delta x_1 > 0$ , causing the spring to pull back with a negative force  $F_1 = -k\Delta x_1$ . The bottom picture shows the spring contracting by  $\Delta x_2 < 0$ , causing the spring to apply a positive force  $F_2 = -k\Delta x_2$  on the mass.



Using Newton's second law of motion (REF?) with  $x$  being the displacement from the rest position  $x_0$ , we get the relation

$$a = -kx, \quad (1.2)$$

but since  $a = \frac{dv}{dt} = \frac{d^2x}{dt^2}$ , we can re-write Equation 1.2 as

$$\frac{d^2x}{dt^2} = -kx, \quad (1.3)$$

or even more succinctly as

$$\ddot{x} = -kx. \quad (1.4)$$

Equation 1.4 is one of the simplest possible 2nd-order *ordinary* differential equations. Its solution is a combination of the two basic trigonometric equations:

$$x(t) = c_1 \sin(\alpha t) + c_2 \cos(\beta t), \quad (1.5)$$

where  $c_1$  and  $c_2$  are constants which we can find using the *starting conditions* (see note 1.2), and  $\alpha, \beta$  are parameters of the motion. Since function arguments must be unitless, these two parameters also cause the total quantity inside the trigonometric functions to

be unitless by having units corresponding to  $\left[\frac{1}{\text{time}}\right]$ . For example, if we measure the time in  $[\text{s}]$ , then the units of  $\alpha$  and  $\beta$  are  $[\text{s}^{-1}] = [\text{Hz}]$ .

**Note 1.2 Starting conditions for solving ODEs**

Recall that in order to completely solve an ordinary differential equation of order  $n$  we must have  $n$  starting conditions.



In the case where the initial position of the mass is  $x_0$  and the initial velocity is  $\dot{x}_0 = 0$  we get the simplified solution

$$x(t) = x_0 \cos(\omega t), \quad (1.6)$$

where  $\omega = \sqrt{\frac{k}{m}}$ . The plot of the position  $x(t)$  of the mass is given in Figure 1.2. Note how, when displayed side-by-side using the same time axes, the position plot is “lagging behind” the velocity plot by  $\phi = \frac{\pi}{2}$ : when we start the motion, the velocity is 0 and then starts to increase in the negative direction (i.e. the mass is moving to the left). The position is at its maximum at that time, and decreases from  $x_{\max}$  at  $t = 0$  to  $x = 0$  at  $t = \frac{\pi}{2}$ , while still staying positive. At  $t = \frac{\pi}{2}$  the velocity reaches its maximum negative value,  $\dot{x} = -v_{\max}$  (which is 1 in units of  $\frac{1}{v_{\max}}$ ), and the position becomes negative (since it is to the left of the rest position of the spring).

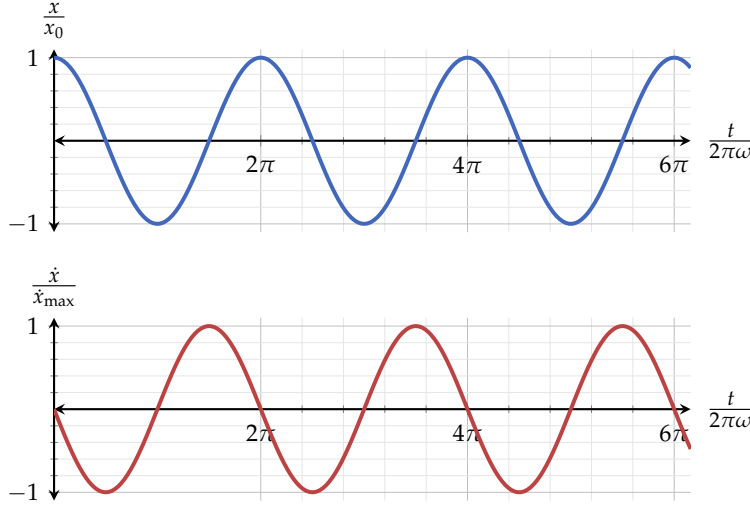


Figure 1.2: Simple harmonic oscillator, top (blue): position  $x$  vs. time  $t$ . The axes units are such that a full period of the oscillation takes  $\Delta t = 2\pi$ , and that the minimum and maximum values of the position are  $\pm 1$ , respectively. Bottom (red): velocity vs. time on the same time axes, and a velocity axis which is scaled such that  $v_{\max} = 1$ .

It would be useful to understand how does the spring-mass system evolve given a specific combination of position and velocity. This can be done by plotting the **phase space** of the system (Figure 1.3): on the horizontal axis we specify the position  $x$  of the mass, and on the vertical axis the velocity  $\dot{x}$ .

In the case of a simple harmonic oscillator, the evolution of the system is shown on the phase space plot as ellipses - and when we use normalized coordinates ( $\tilde{x} = \frac{x}{x_0}$  and  $\tilde{\dot{x}} = \frac{\dot{x}}{v_{\max}}$ ), the ellipses turn into perfect circles - just like we see in Figure 1.3. These circles

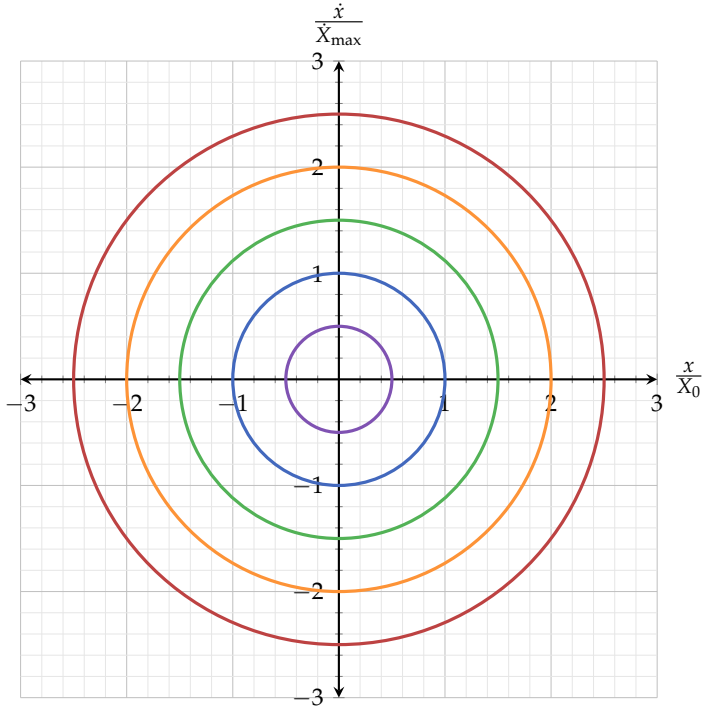


Figure 1.3: Phase space plot of simple harmonic oscillators with different momenta (either their masses are different, or their initial distance are different). The axes are scaled such that their units are the initial position  $X_0$  and maximum velocity  $\dot{X}_0$ , respectively, of the second oscillator (trajectory drawn in blue).

or ellipses are paths of constant energy: this can be seen when we “translate” the plot to have the potential energy and kinetic energy as horizontal and vertical axes, respectively.

Recall that for close mechanical systems, the total energy  $E$  is given by the sum of the sum of the *kinetic energy*  $K$  and the *potential energy*  $U$ :

$$E = K + U. \quad (1.7)$$

The kinetic energy is a function of the velocity  $v = \dot{x}$ :

$$K = \frac{1}{2}\dot{x}^2 = \frac{p^2}{2m}, \quad (1.8)$$

and the potential energy is a function of the position  $x$  via the force  $F$ :

$$F = -\frac{dU}{dx}. \quad (1.9)$$

In the case of an harmonic oscillator  $F = -kx$ , and therefore

$$U = \int F dx = -k \int x dx = -\frac{1}{2}kx^2 + c. \quad (1.10)$$

Since we can add to the potential any constant with no change to the force derived from it (remember that the derivative of a constant is zero), we can simply set the potential at  $x = 0$  to  $U(x = 0) = 0$ , meaning that the integral constant is  $c = 0$ . Altogether, we get the following system energy:

$$E = \frac{p^2}{2m} - \frac{1}{2}kx^2, \quad (1.11)$$

Essentially, this is equal to applying the following transformations to the phase space plot:

$$x \rightarrow -\frac{1}{2}kx^2, \quad (1.12)$$

$$\dot{x} \rightarrow \frac{p^2}{2m} = \frac{1}{2m}m^2\dot{x}^2 = \frac{1}{2}m\dot{x}^2. \quad (1.13)$$

Since the transformation “stretches” both axes by the same power (up to the constants  $m$  and  $k$ ), the shapes remain the same as in the original plot.

The total energy of the system can then be extracted from the radius of each circle path:

$$E = \sqrt{K^2 + U^2} = \sqrt{\frac{p^2}{4m^2} + \frac{k^2x^4}{4}} = \frac{1}{2} \left( \frac{p^2}{m^2} + k^2x^4 \right). \quad (1.14)$$

When  $K = 0$  (i.e.  $p = 0$ ) the entire energy is stored as potential energy:

$$E = \frac{1}{2}\sqrt{k^2x^4} = \frac{1}{2}kx^2. \quad (1.15)$$

And when  $U = 0$  (i.e.  $x = 0$ ) the entire energy is stored as kinetic energy:

$$E = \frac{1}{2}\sqrt{\frac{p^2}{m^2}} = \frac{p}{2m} = \frac{1}{2}m\dot{x}^2. \quad (1.16)$$

### 1.5.2 Damped Harmonic Oscillator

We can make the harmonic oscillator model a bit more realistic if we add a damping force proportional to the velocity of the mass. This corresponds e.g. to introducing friction into the model. The form of the damping force is  $F_{\text{damp}}(t) = -c\dot{x}(t) = -c\dot{x}$ , for some real damping coefficient  $c \geq 0$  (in the case where  $c = 0$  we get back the simple harmonic oscillator model). Thus, the overall Newton 2nd law equation of the system has the following form:

$$m\ddot{x} = -c\dot{x} - kx. \quad (1.17)$$

The term  $-c\dot{x}$  always acts in the opposite direction to the velocity due to the minus sign. Re-arranging Equation 1.17 we get the more “canonical” form

$$m\ddot{x} + c\dot{x} + kx = 0, \quad (1.18)$$

We then use the solution  $x(t) = e^{\lambda t}$ , substituting it into Equation 1.18:

$$F_{\text{total}}(t) = \lambda^2 m e^{\lambda t} + c e^{\lambda t} + k e^{\lambda t} \quad (1.19)$$

$$= e^{\lambda t} (m\lambda^2 + c\lambda + k) \quad (1.20)$$

$$= 0. \quad (1.21)$$

Since Equation 1.21 is true for all  $t$  (and in any case  $e^{\lambda t} \neq 0$  for all  $t$ ), for the equation to be true it must be that the quadratic equation in  $\lambda$  equals zero. Using the quadratic formula we get that

$$\lambda_{1,2} = \frac{-c \pm \sqrt{c^2 - 4km}}{2m}. \quad (1.22)$$

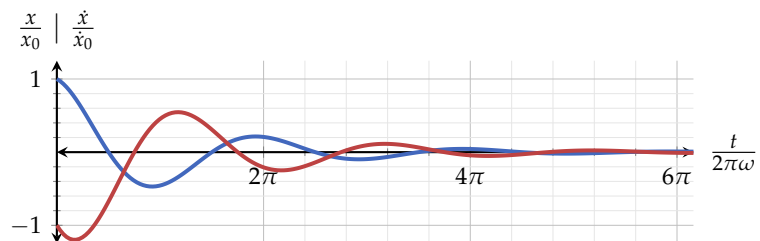
For the solutions to  $\lambda$  to be real numbers,  $c^2 \geq 4km$  - otherwise the term in the square root is negative. When  $c^2 > 4km$  we call the system **overdamped**, and the case where  $c^2 = 4km$  is the **critical damping** value. MORE TEXT.

The overall position vs. time relationship of the underdamped system is given by

$$x(t) = x_0 e^{-\frac{c}{2m}t} \cos(\omega t), \quad (1.23)$$

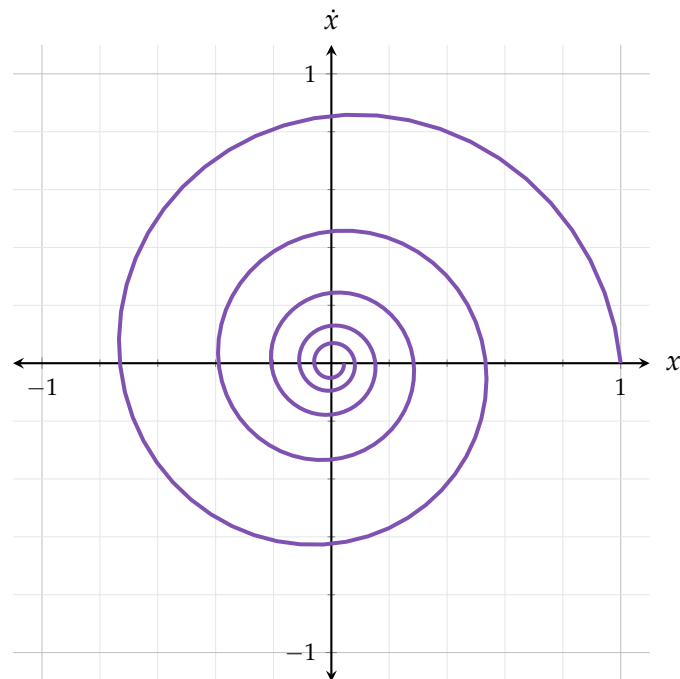
where as before  $\omega = \sqrt{\frac{k}{m}}$  (see Figure 1.4).

Figure 1.4: Underdamped harmonic oscillation. . .



The phase space plot of the damped system shows. . .

Figure 1.5: Phase space plot of an underdamped harmonic oscillator. . .



### 1.5.3 Simulating Harmonic Oscillators in Python

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# Simulating Orbital Mechanics

## 2.1 Preface

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## 2.2 Relevant Physical and Mathematical Background

### 2.2.1 Classic Gravitational Force

Already in the 17th century, *Isaac Newton* formulated the gravitational force existing between any two objects with masses greater than zero. The strength of the force is given by the equation

$$F = G \frac{m_1 m_2}{r^2}, \quad (2.1)$$

where  $m_1$  and  $m_2$  are the respective masses of the two objects,  $r$  is the distance between them, and  $G$  is a the *universal gravitational constant*,

$$G = (6.6743 \pm 0.0015) \times 10^{-11} \left[ \text{N m}^2 \text{ kg}^{-2} \right] \quad (2.2)$$

The direction of the force is the line connecting the centers of mass of the two objects. Due to Newton's third law, the forces acting on the two objects are equal and opposite: the force applied by  $m_1$  on  $m_2$ ,  $F_{1 \rightarrow 2}$ , is pointing **from**  $m_2$  **onto**  $m_1$ , and the force applied by  $m_2$  on  $m_1$ ,  $F_{2 \rightarrow 1}$  is pointing **from**  $m_1$  **onto**  $m_2$  - and is exactly opposite to  $F_{1 \rightarrow 2}$ , i.e. in vector notation

$$\vec{F}_{1 \rightarrow 2} = -\vec{F}_{2 \rightarrow 1}. \quad (2.3)$$

If the two objects have positions  $\vec{r}_1$  and  $\vec{r}_2$ , the vector pointing from object 1 to object 2 is

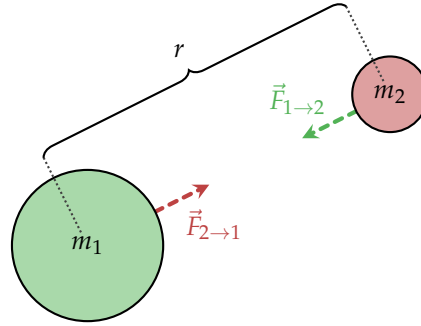
$$\vec{r}_{1 \rightarrow 2} = \vec{r}_2 - \vec{r}_1, \quad (2.4)$$

with the vector pointing from object 2 to object 1 having the exact opposite components, i.e.  $\vec{r}_{2 \rightarrow 1} = -\vec{r}_{1 \rightarrow 2}$ . The norms of  $\vec{r}_{1 \rightarrow 2}$

and  $\vec{r}_{2 \rightarrow 1}$  are simply  $r$  (the distance between the objects), and their directions are the unit vectors in the direction of  $\vec{r}_{1 \rightarrow 2}$  and  $\vec{r}_{2 \rightarrow 1}$ , respectively:

$$\begin{aligned}\hat{r}_{1 \rightarrow 2} &= \frac{\vec{r}_{1 \rightarrow 2}}{\|\vec{r}\|_{1 \rightarrow 2}} = \frac{\vec{r}_{1 \rightarrow 2}}{r}, \\ \hat{r}_{2 \rightarrow 1} &= \frac{\vec{r}_{2 \rightarrow 1}}{\|\vec{r}\|_{2 \rightarrow 1}} = \frac{\vec{r}_{2 \rightarrow 1}}{r} = -\hat{r}_{1 \rightarrow 2}.\end{aligned}\quad (2.5)$$

Figure 2.1: Gravitational force between two objects with masses  $m_1$  and  $m_2$ . Each object applies an attractive force on the other object, with norm  $F = G \frac{m_1 m_2}{r^2}$  (where  $r$  is the distance between the objects) and in the direction pointing from each object to the other object.



In total, the vector notation of the gravitational force applied by the objects on each other are

$$\begin{aligned}\vec{F}_{1 \rightarrow 2} &= G m_1 m_2 \frac{\hat{r}_{1 \rightarrow 2}}{r^2}, \\ \vec{F}_{2 \rightarrow 1} &= G m_1 m_2 \frac{\hat{r}_{2 \rightarrow 1}}{r^2} = -\vec{F}_{1 \rightarrow 2}.\end{aligned}\quad (2.6)$$

#### Note 2.1 Another gravity force vector notation

In some textbooks, Equation 2.6 are written without the unit vectors  $\hat{r}_{1 \rightarrow 2}$  and  $\hat{r}_{2 \rightarrow 1}$ , instead using the distance vectors and dividing by  $r^3$ , i.e.

$$\begin{aligned}\vec{F}_{1 \rightarrow 2} &= G m_1 m_2 \frac{\vec{r}_{1 \rightarrow 2}}{r^3}, \\ \vec{F}_{2 \rightarrow 1} &= G m_2 m_1 \frac{\vec{r}_{2 \rightarrow 1}}{r^3}.\end{aligned}$$

The result is of course the same as in Equation 2.6, since for any non zero vector  $\vec{v}$ ,

$$\frac{\vec{v}}{\|\vec{v}\|^3} = \frac{1}{\|\vec{v}\|^2} \frac{\vec{v}}{\|\vec{v}\|} = \frac{1}{\|\vec{v}\|^2} \hat{v}.$$

Let us look at an example of calculating the gravitational forces between two objects.

#### Example 2.1 Calculating a gravitational force

Let us calculate the gravitational forces between two objects

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$A$  and  $B$ , using the following parameters:

$$\vec{r}_A = \begin{bmatrix} 1 \\ -2 \\ 0 \end{bmatrix}, m_A = 1,$$

$$\vec{r}_B = \begin{bmatrix} 2 \\ 5 \\ -2 \end{bmatrix}, m_B = 2.$$

For the sake of simplicity, we use  $G = 1$  and don't consider units with this example.

The vector pointing from  $A$  to  $B$  is

$$\vec{r}_{A \rightarrow B} = \vec{B} - \vec{A} = \begin{bmatrix} 2 \\ 5 \\ -2 \end{bmatrix} - \begin{bmatrix} 1 \\ -2 \\ 0 \end{bmatrix} = \begin{bmatrix} 1 \\ 7 \\ -2 \end{bmatrix},$$

and the vector pointing from  $B$  to  $A$  is

$$\vec{r}_{B \rightarrow A} = -\vec{r}_{A \rightarrow B} = \begin{bmatrix} -1 \\ -7 \\ 2 \end{bmatrix}.$$

The distance  $r$  between the objects is the norm of either of the above vectors, so we'll use  $\vec{r}_{A \rightarrow B}$ :

$$r = \|\vec{r}\|_{A \rightarrow B} = \sqrt{1^2 + 7^2 + 2^2} = \sqrt{19} \approx 7.3485.$$

The direction vectors are therefore

$$\hat{r}_{A \rightarrow B} = \frac{1}{7.3485} \begin{bmatrix} 1 \\ 7 \\ -2 \end{bmatrix} = \begin{bmatrix} 0.1361 \\ 0.9526 \\ -0.2722 \end{bmatrix},$$

$$\hat{r}_{B \rightarrow A} = -\hat{r}_{A \rightarrow B} = \begin{bmatrix} -0.1361 \\ -0.9526 \\ 0.2722 \end{bmatrix}.$$

The gravity force which  $A$  applies onto  $B$  is then

$$\vec{F}_{A \rightarrow B} = \overbrace{G \frac{m_1 m_2}{r^2}}^{=2 \times 1} \hat{r}_{A \rightarrow B} = \frac{2}{54} \begin{bmatrix} 0.1361 \\ 0.9526 \\ -0.2722 \end{bmatrix} = \begin{bmatrix} 0.0050 \\ 0.0353 \\ -0.101 \end{bmatrix}.$$

and similarly,

$$\vec{F}_{B \rightarrow A} = -\vec{F}_{A \rightarrow B} = \begin{bmatrix} -0.0050 \\ -0.0353 \\ 0.101 \end{bmatrix}.$$



In the case where we only consider two objects, and choose our frame of reference such that one of the objects is stationary - an analytical solution to the spatial trajectory taken by the second object is known and well studied. It is called a **Keplerian orbit**, and it always takes the form of a conic section.

### 2.2.2 Conic Sections

A conic section (sometimes simply just called “a conic”) is a 2-dimensional shape resulting from the intersection of a plane and a cone (see Figure 2.2). Depending on the angle  $\alpha$  by which the plane intersects the cone relative to the cone’s side, the resulting shape can be one of 3 general types (here  $\theta$  is the cone’s angle):

1. If  $\alpha > \theta$  the intersection is an **ellipse**. If in addition  $\alpha = 90^\circ$  the ellipse becomes a **circle**.
2. If  $\alpha = \theta$  the intersection is a **parabola**.
3. If  $\alpha < \theta$  the intersection is a **hyperbola**.

Figure 2.2: An intersection of a cone and a plane. Both the cone and plane are infinite - the cone extends infinitely “down”, but also has a second “inverted” part on the top, also extending to infinity. In the case here shown, the intersection is an ellipse. Image reproduced with modifications from !SOURCE!

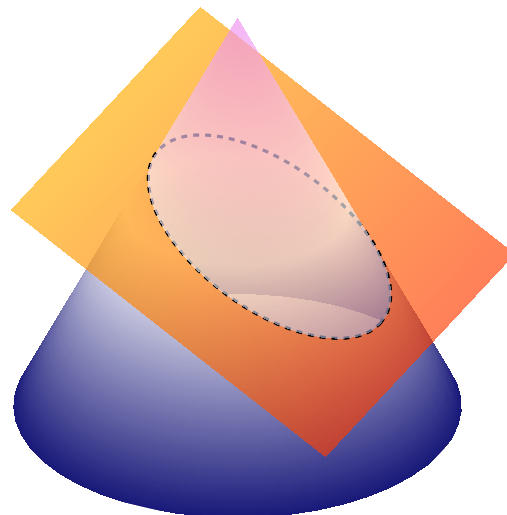
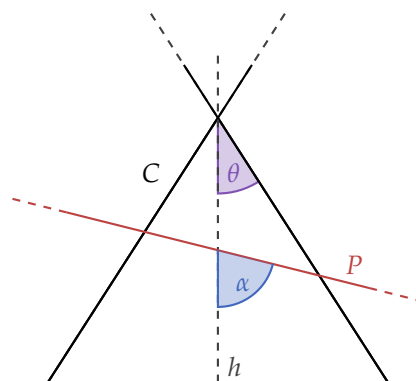


Figure 2.3: Side view of an infinite cone  $C$  and an infinite plane  $P$  intersecting it. The angle between  $P$  and the cone’s height line  $h$  is  $\alpha$ , and the angle between the cone’s surface and  $h$  is  $\theta$ . In this figure  $0 < \theta < \alpha < 90^\circ$ , and thus the shape formed by the intersection of  $C$  and  $P$  is an ellipse.



2.3 *Forward Euler Method*

2.4 *Backward Euler Method*

2.5 *Verlet Integration*

2.6 *Runge-Kutta Method*