

# Machine Learning Techniques to Enhance Event Reconstruction in Water Cherenkov Detectors

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NuFact 2022, Salt Lake City, Utah, 4<sup>th</sup> August 2022

# Water Cherenkov Neutrino Experiments

Current generation **Super-K** and **T2K** and next generation **Hyper-K** are world-leading neutrino experiments.

Broad & ambitious physics programmes covering many neutrino sources as well as proton decay measurements.

Water Cherenkov detector technology provides huge target mass with excellent particle ID and reconstruction capabilities.

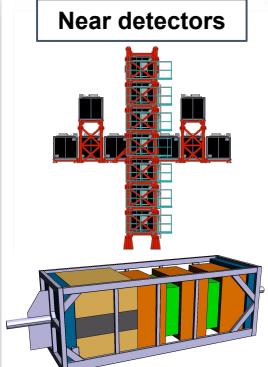
See also: L. Kormos (T2K, Mon 2:20pm  
M. Friend (J-PARC, Wed 8:30am)



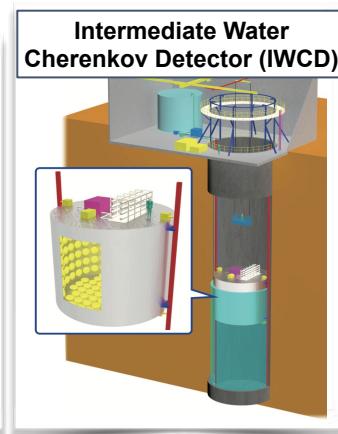
Water Cherenkov Test-beam Experiment (WCTE) at CERN



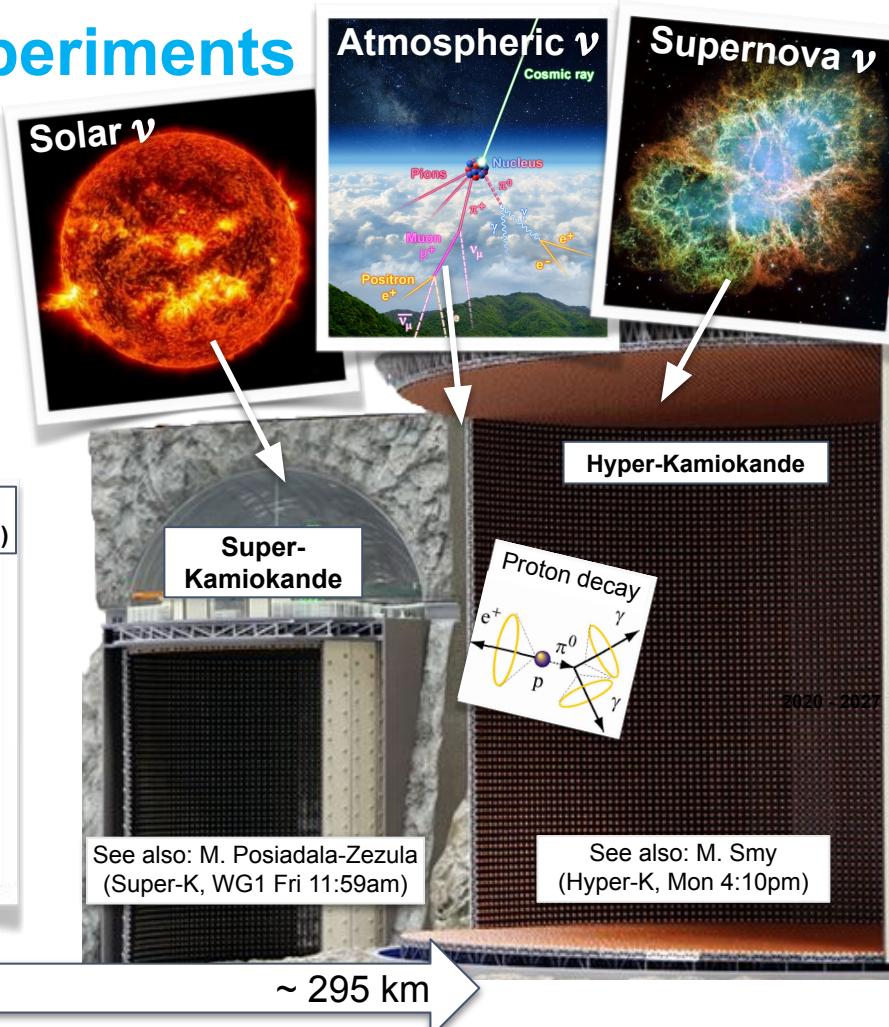
J-PARC  $\nu$  beam



Near detectors



Intermediate Water Cherenkov Detector (IWCD)



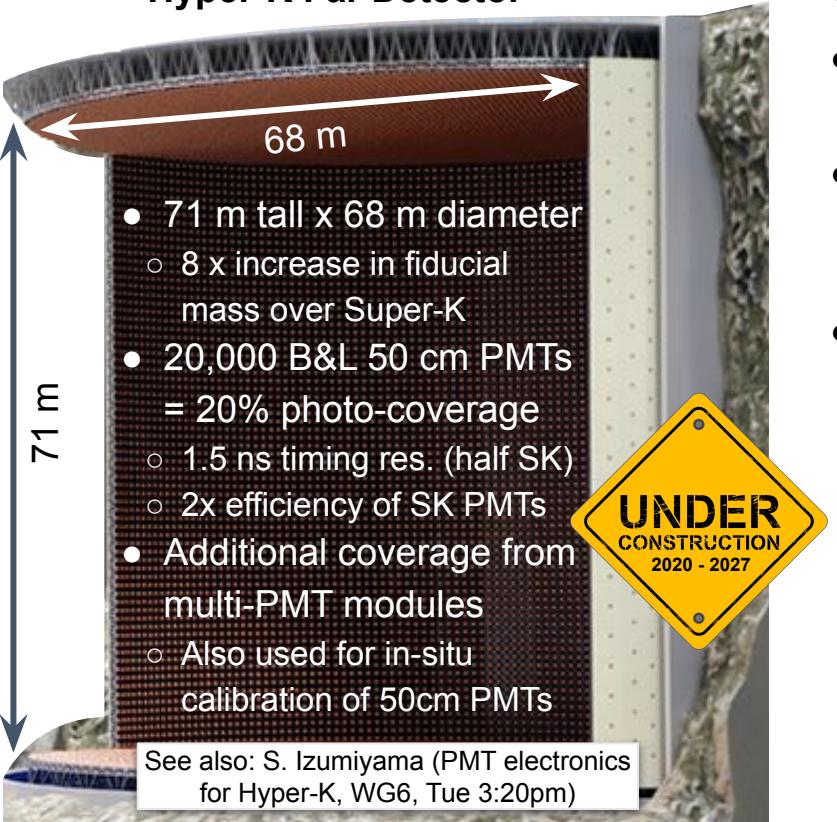
280 m

~ 1 km

~ 295 km

# Hyper-K's WC Detectors

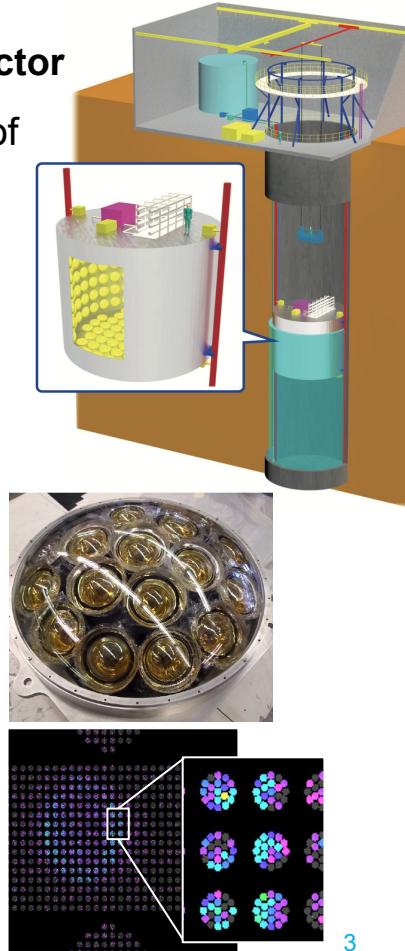
## Hyper-K Far Detector



## Intermediate Water Cherenkov Detector

- Measures  $\nu$  flux and cross-section of beam at ~1 km from source
- Moves vertically in ~50 m tall pit
  - spans off-axis angles of  $\nu$  beam for different  $\nu$  energy spectra
- 6 m tall x 8 m diameter tank with ~500 multi-PMT modules (mPMTs)
  - 8 cm PMTs:
    - Better position resolution
    - < 1 ns timing resolution
  - Additional directionality information
  - mPMTs will also be used for WCTE
  - Also in consideration for portion of far detector photo-coverage

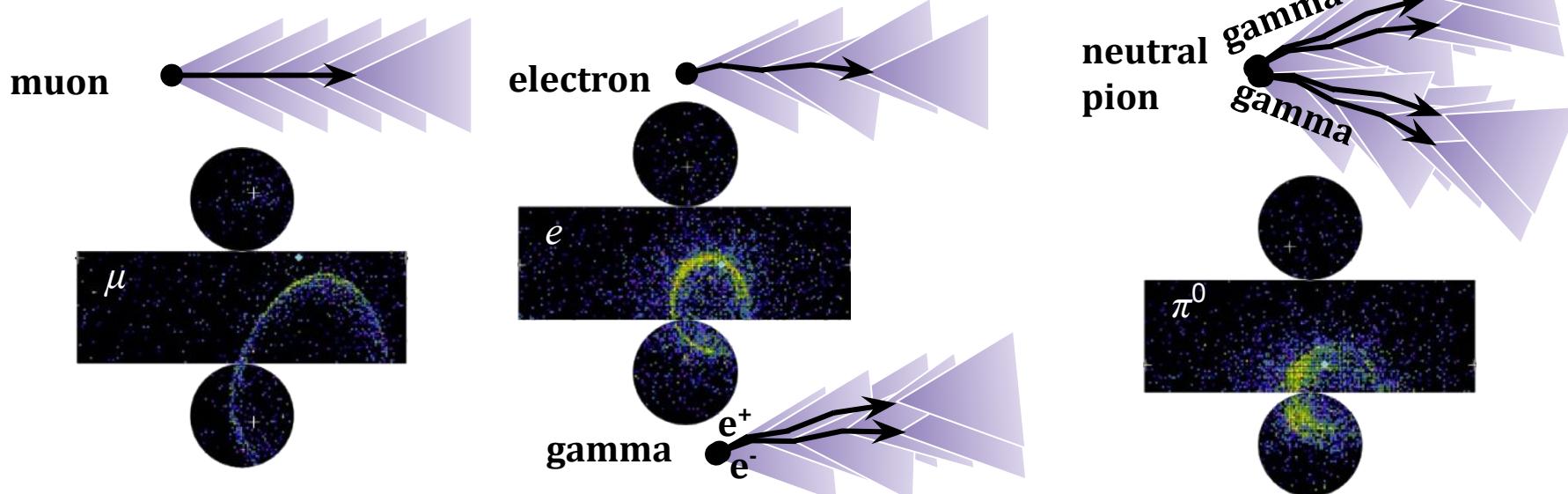
See also: R. Akutsu (mPMTs for IWCD & WCTE, WG6 Tue 3pm)



# Reconstruction in WC detectors

Classification: Particle type identification (PID)

- Different particles produce different types of rings



Regression: reconstructing particle's properties:

- Location and time of PMT hits allows triangulating position and direction
- Amount of charge observed at PMTs gives estimate of energy

# Machine learning reconstruction for WC

Limit of traditional maximum-likelihood reconstruction methods (fiTQun) is being reached

- Computation time is becoming a limiting factor
  - Larger far detector with more PMTs increases computation time
  - Smaller intermediate detector requires scaled down resolutions
  - Improving resolutions requires more complex algorithms with fewer approximations

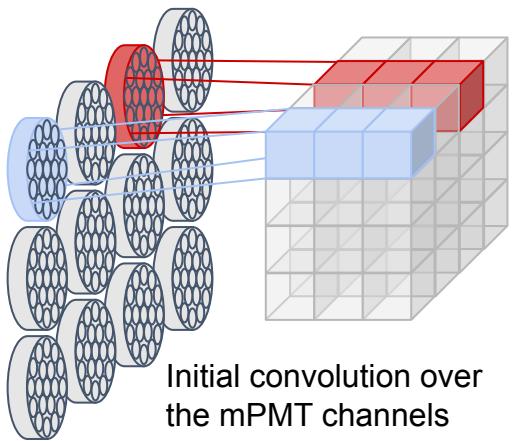
ML and deep neural networks have potential to push reconstruction further

- Very successful in areas of computer vision and image processing
- Potential to use all information without detector model approximations
- Very fast to run once neural networks have been trained
  - fiTQun on CPU: 1 event takes more than 1 minute
  - ML reconstruction on GPU: 100,000 events per minute
  - Opens opportunities for analyses with huge datasets not currently possible

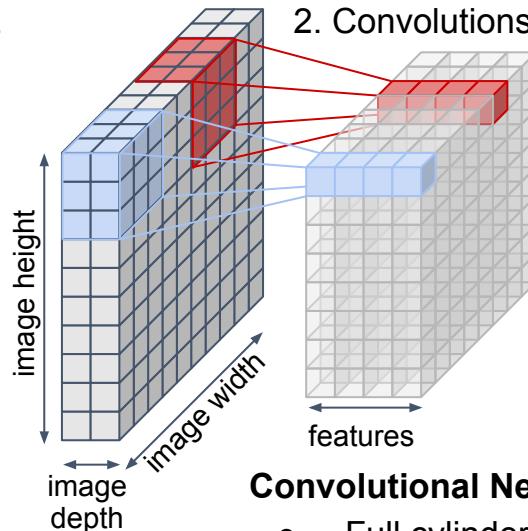
See also: A. Yankelevich (ML for solar  $\nu$  in SK, WG1+WG6 Thu 3:04pm)

# Deep network architectures for IWCD

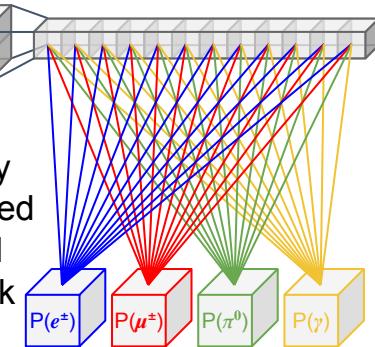
## 1. Convolution over mPMTs



## 2. Convolutions & down-samples



## 3. Fully connected neural network



## Convolutional Neural Network based on ResNet-18

- Full cylinder of mPMTs is unwrapped onto flat image
- One pixel per multi-PMT

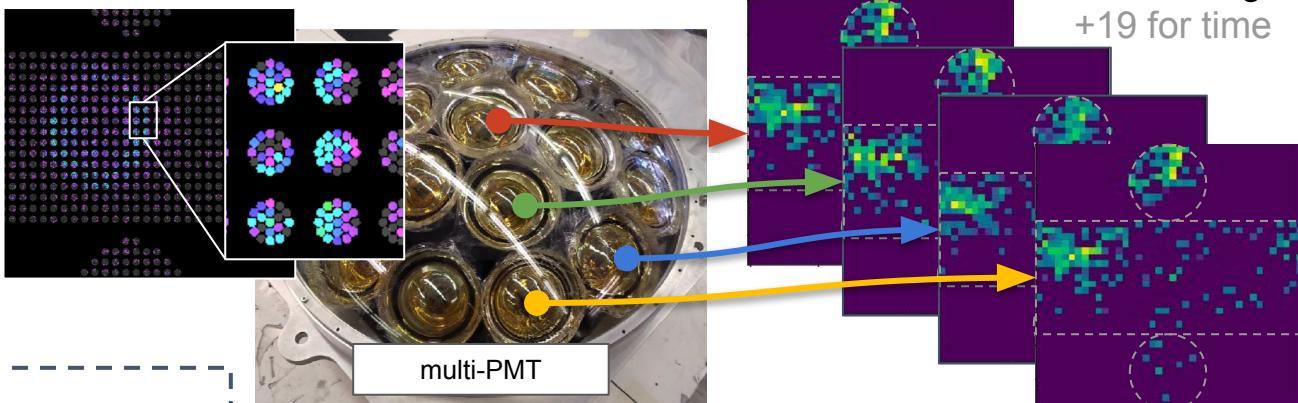
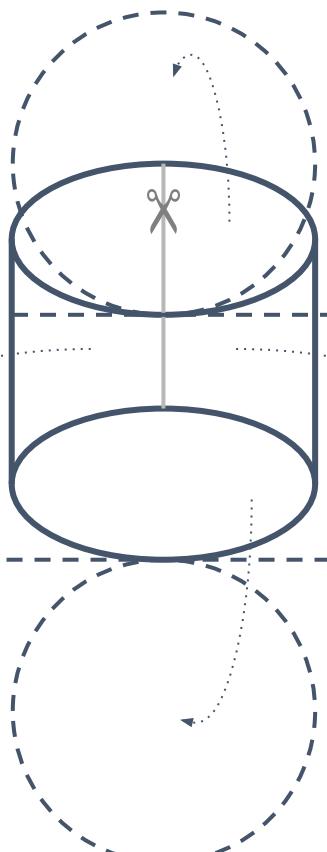
## Point Cloud Neural Network based on PointNet

- Applies to point-cloud of PMT hits in 3D space
- Uses  $1 \times 1$  convolutions and learns transformations applied to points

PointNet MLP (convolution over point cloud features)

# Image-like data for CNN

19 for charge  
+19 for time

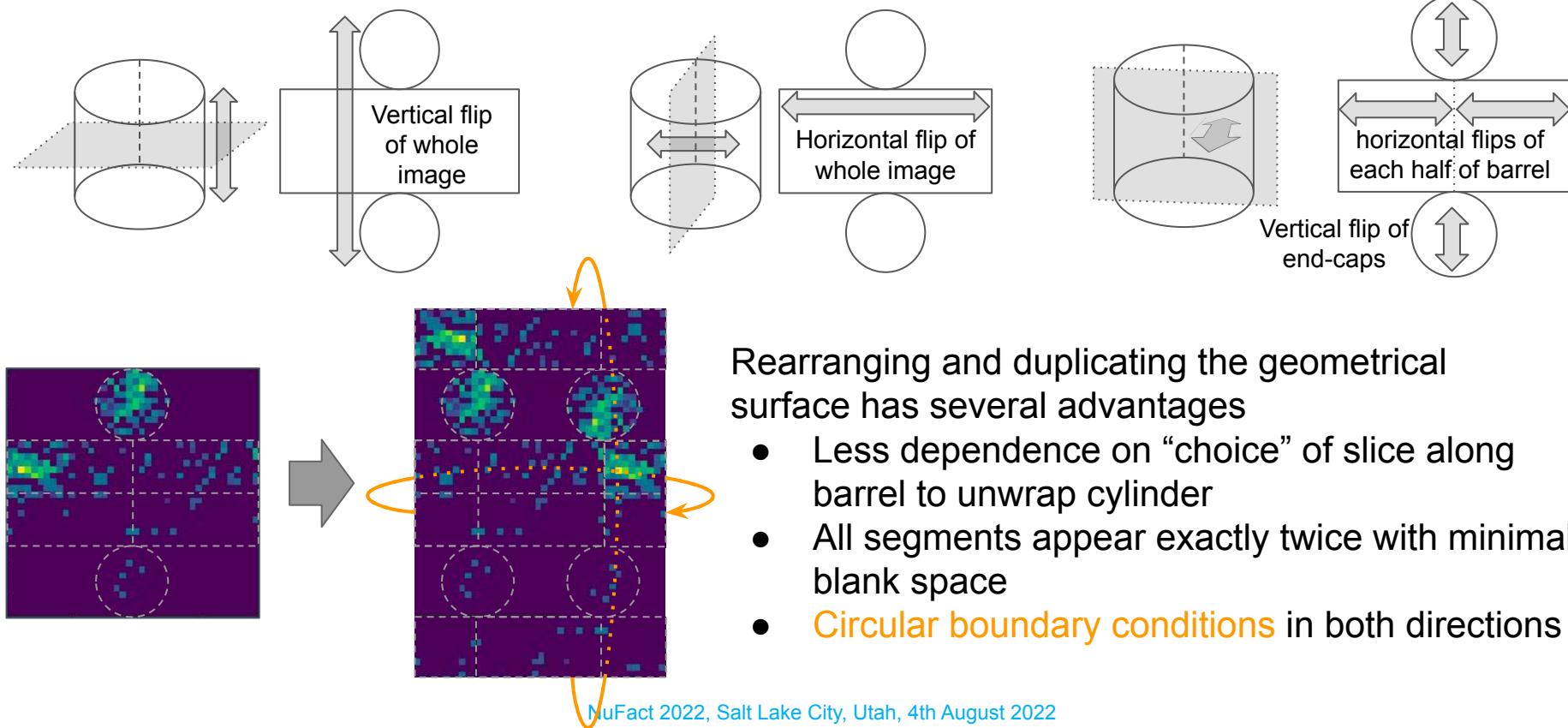


Full cylinder of mPMTs is unwrapped onto flat image

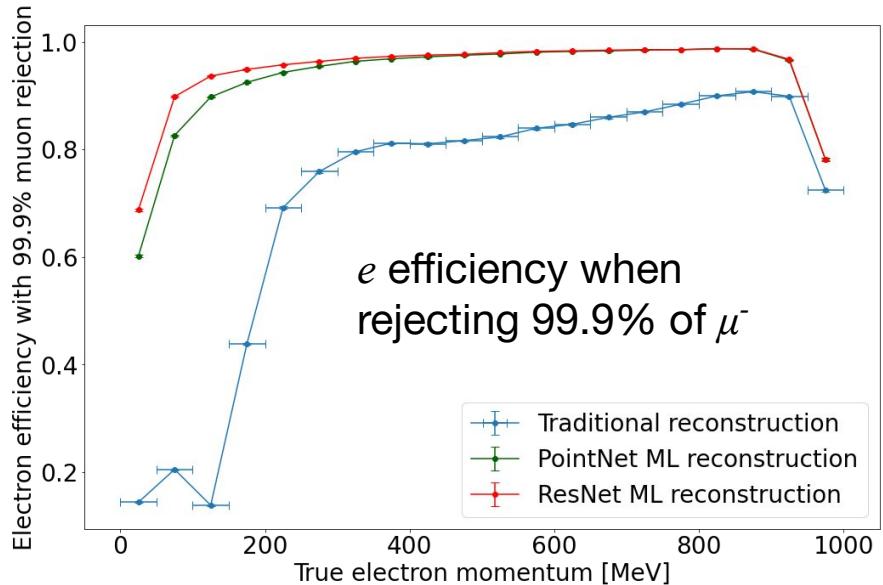
- One pixel per multi-PMT
- Charge (& time) of 19 PMTs per mPMT
- No special treatment at barrel / end-cap boundary
  - Alternative projections from cylinder to grid have also been explored

# Data Transformations and Augmentation

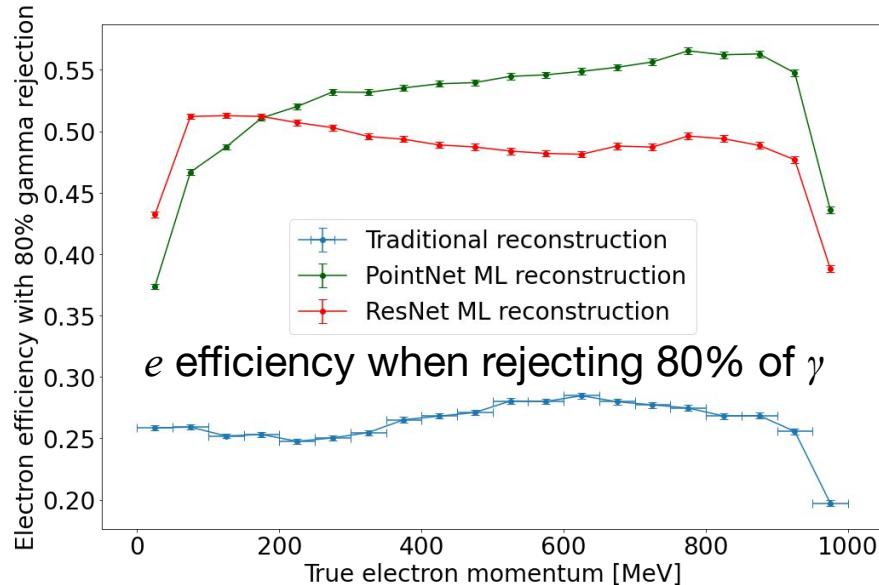
Applying random transformations using detector symmetry effectively increases dataset



# Classification for PID in IWCD



*e* efficiency when  
rejecting 99.9% of  $\mu^-$



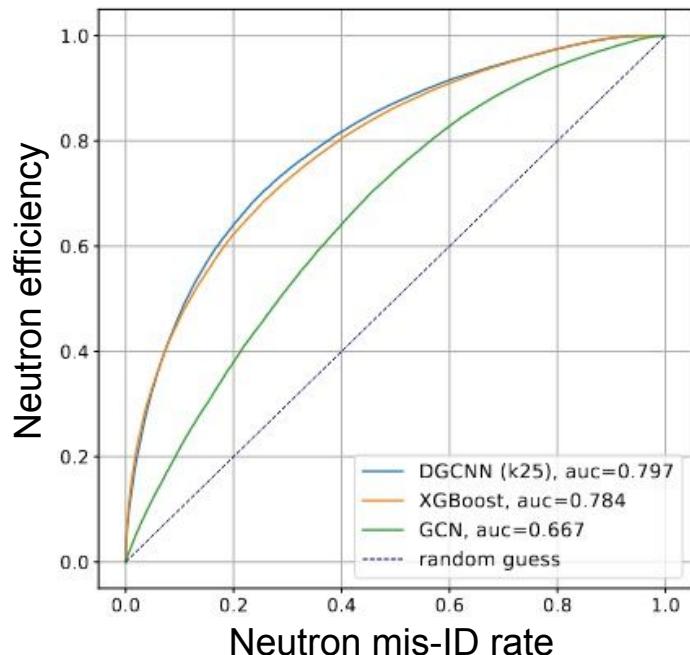
- $\nu_\mu$  beam produces mostly  $\mu$ , need rejection factor of 1000 for  $\nu_e$  measurement
- Improved performance across energy range
- ResNet performs slightly better than PointNet for  $e$  vs  $\mu$  classification

- $\gamma$  and  $e$  almost indistinguishable in water Cherenkov detectors
- Discrimination has not been possible before
- PointNet performs better than ResNet for  $e$  vs  $\gamma$  classification

# Classification for neutron captures

- At lower energies, images can be very sparse and CNNs tend to perform less well
- Alternative networks like graph networks may be more useful
  - Each PMT is a node on a graph
  - Time, charge, position are node features
  - Graph can be defined by nearest neighbors in Graph Convolution Network (GCN, arXiv:1609.02907)
  - Graph can be learned dynamically in Dynamic Graph Convolutional Neural Network (DGCNN, arXiv:1801.07829)
- Tested classifying neutron captures vs electron background
  - Signal: ~8 MeV gamma cascade
  - Background: beta decays of isotopes produced by cosmic muon spallation
  - Compared performance to baseline using BDT (XGBoost) with features including number of hits, hit isotropy, etc.

B. Jamieson, et al., arXiv:2206.12954

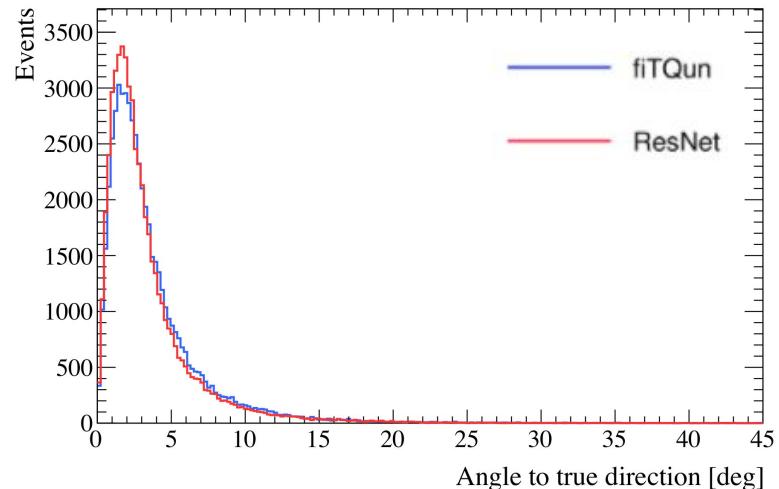
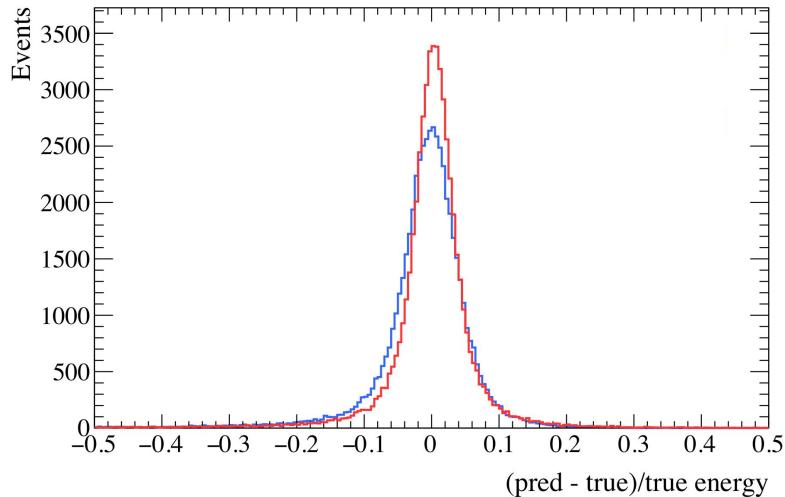
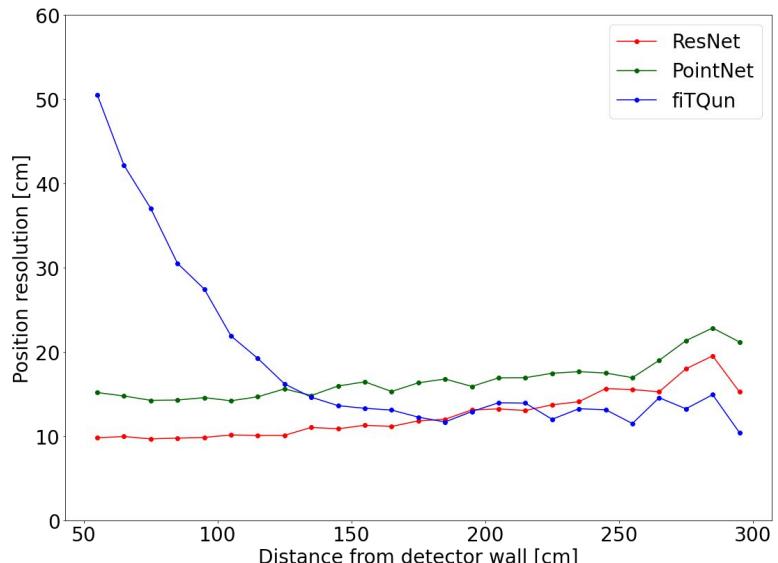


DGCNN outperforms BDT baseline, while GCN underperforms

# Position, direction, energy reconstruction

Using same IWCD data and networks as classification

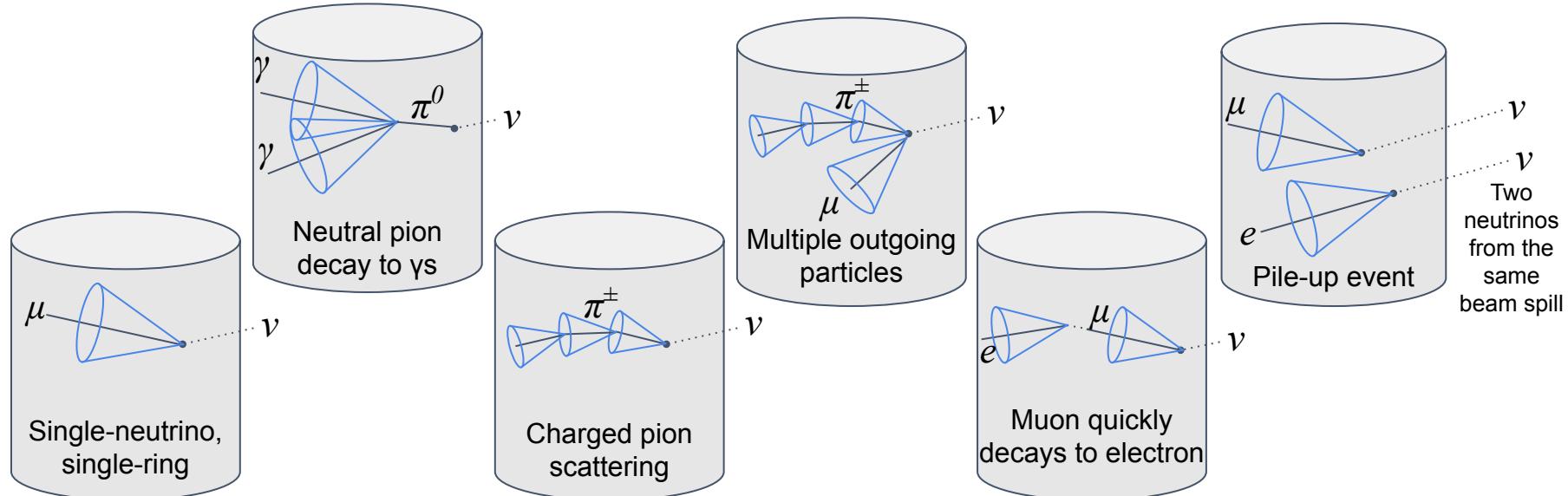
- Output reconstructed quantities instead of PID variables
- Improved performance over traditional reconstruction (fiTQun), particularly for particles close to detector wall



# Multi-ring and multi-vertex events

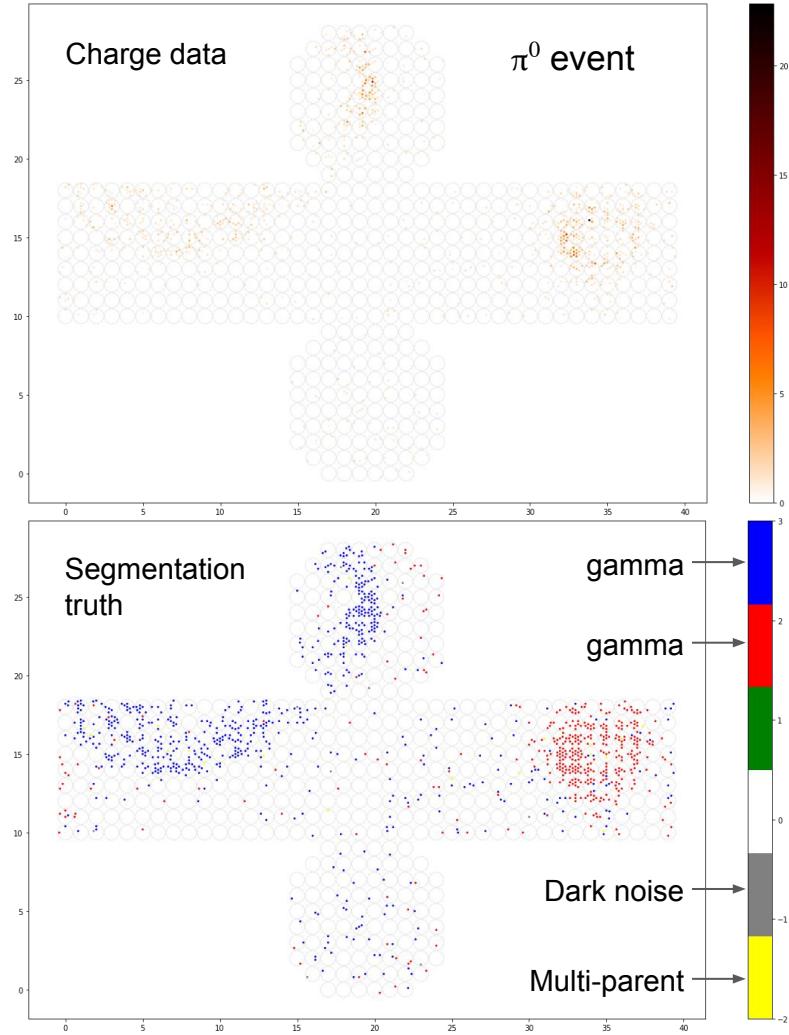
Need to develop ability to identify and reconstruct multi-ring and multi-vertex events

- Single-neutrino interactions can produce various multi-ring event topologies
- Pile-up of neutrino interactions is possible for IWCD due to proximity to beam source



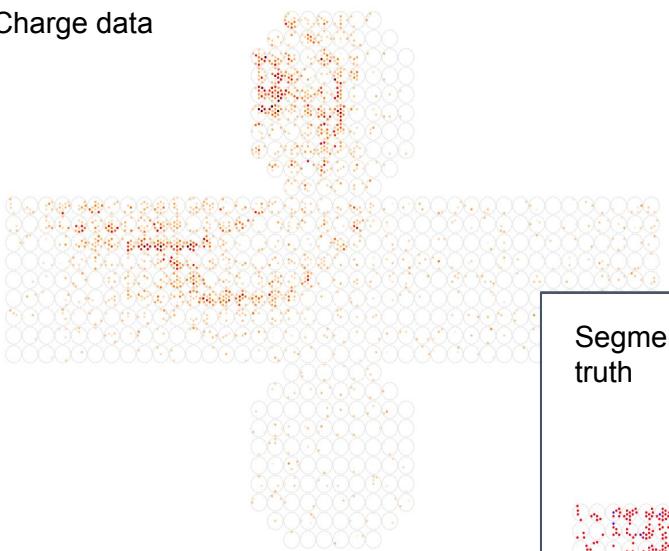
# Segmentation networks

- Classification networks can be extended to perform segmentation
  - Deconvolutions and upsampling reverse convolutions and downsampling
  - Provides output value for each pixel
  - Currently using U-Net and FRRN
- Starting development with  $\pi^0$  events
  - $\pi^0$  decay to produce two  $\gamma$  rings
  - Higher energy  $\pi^0$  have overlapping rings

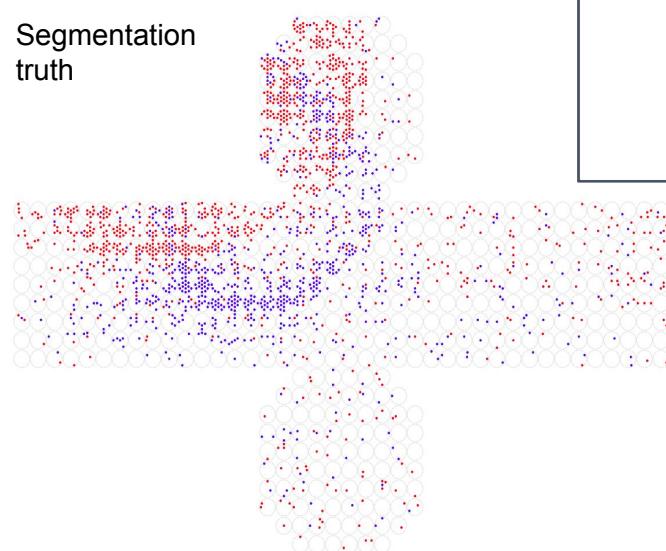


# Segmentation results

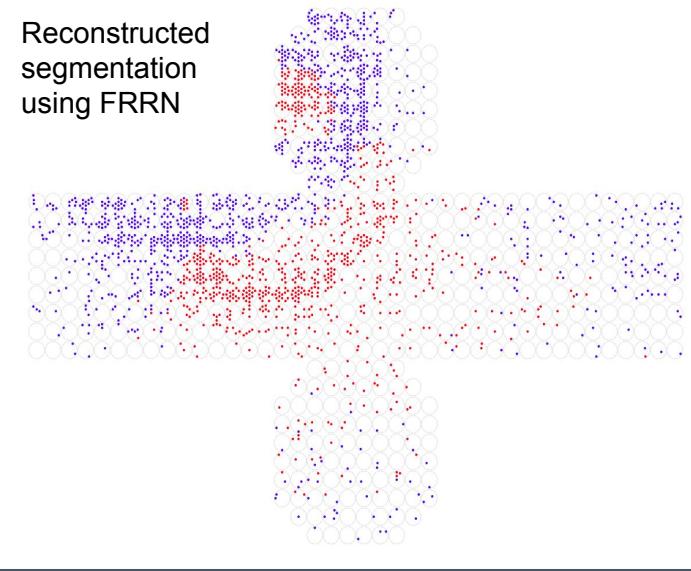
Charge data



Segmentation truth



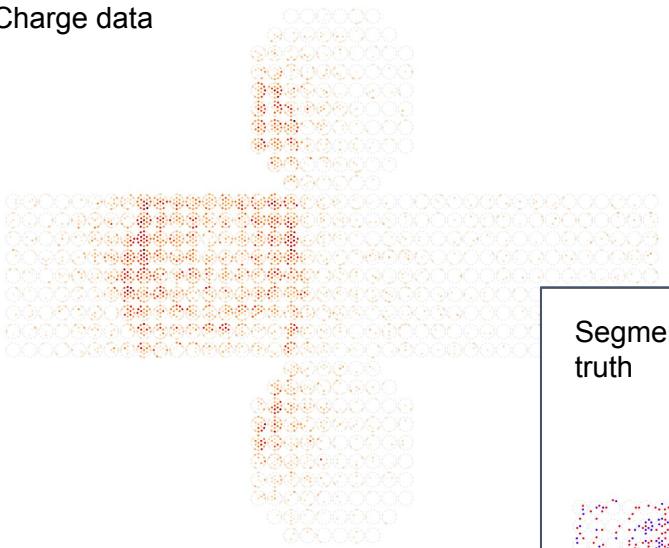
Reconstructed  
segmentation  
using FRRN



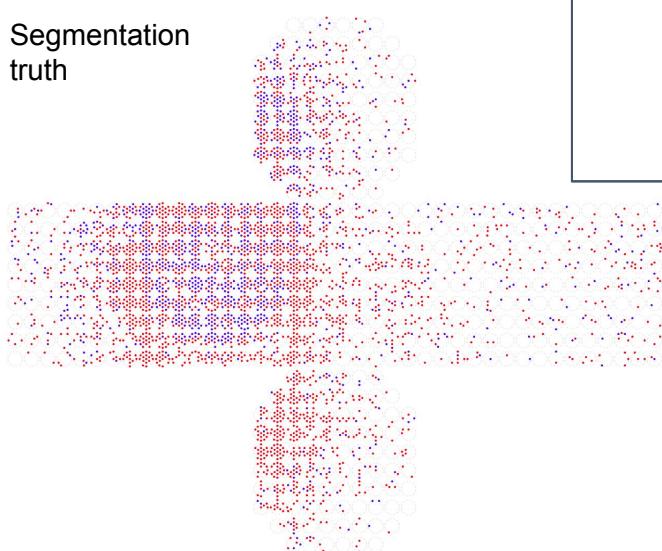
Works well with  
separated or partially  
overlapping rings

# Segmentation results

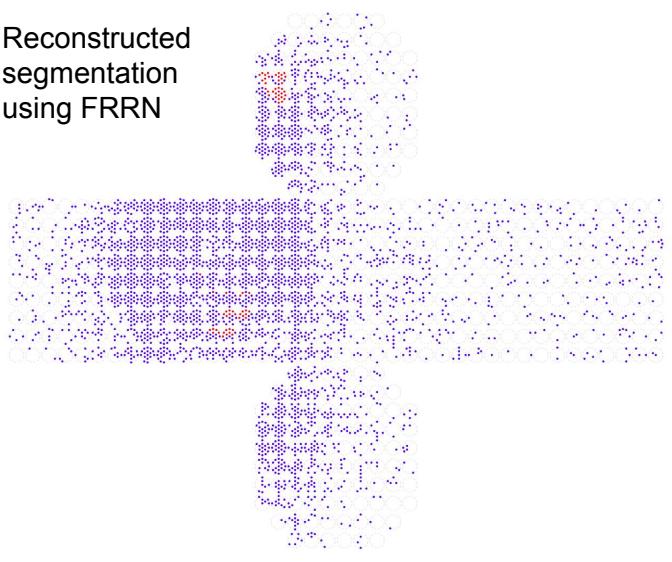
Charge data



Segmentation truth



Reconstructed  
segmentation  
using FRRN



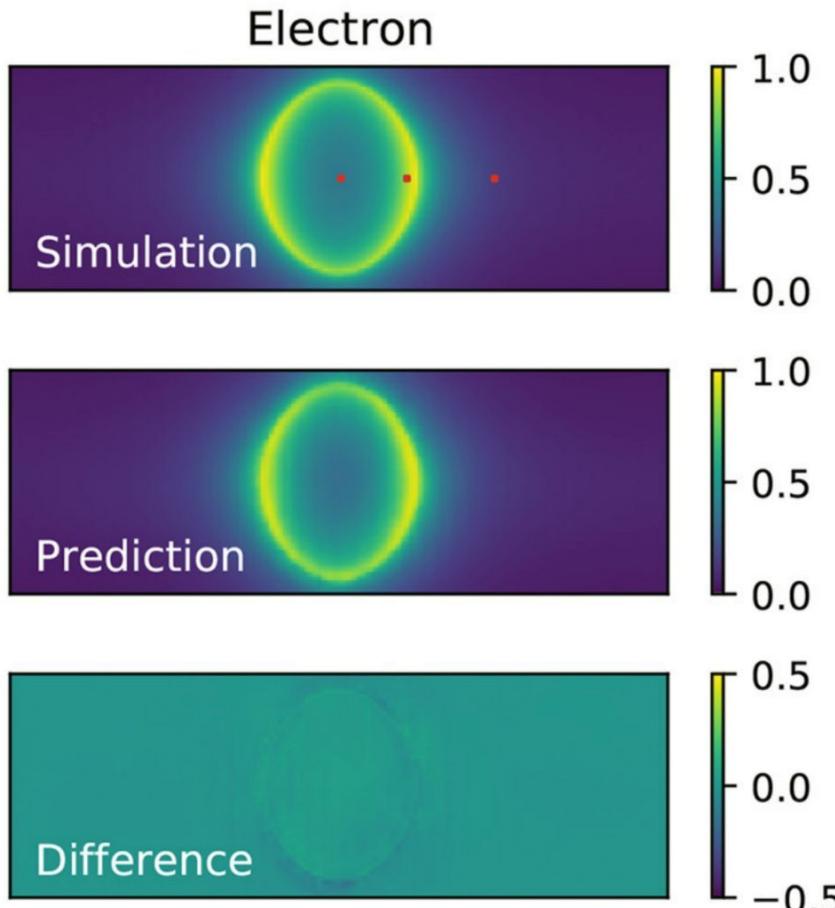
Poor reconstruction  
with some more  
overlapping rings

# Generative networks

Generative networks create synthetic data, with several possible applications

- Often used for faster or more accurate simulations or modifying data for different purposes
- Novel use as part of hybrid approach with maximum likelihood event reconstruction
  - Limitations of traditional reconstruction arise from computational complexity of likelihood function
  - Generative network can quickly produce Cherenkov rings used in likelihood calculation without physics model approximations
  - Predict parameters of Gaussian mixture model for charge & time likelihood functions at each PMT

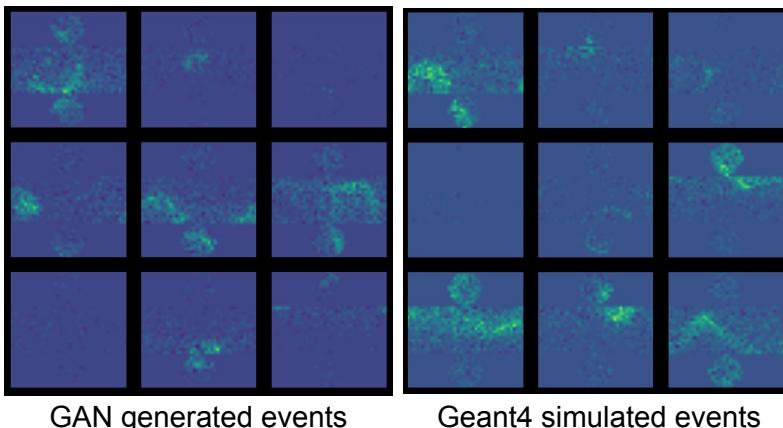
M. Jia, et al., arXiv:2202.01276



# Generative networks

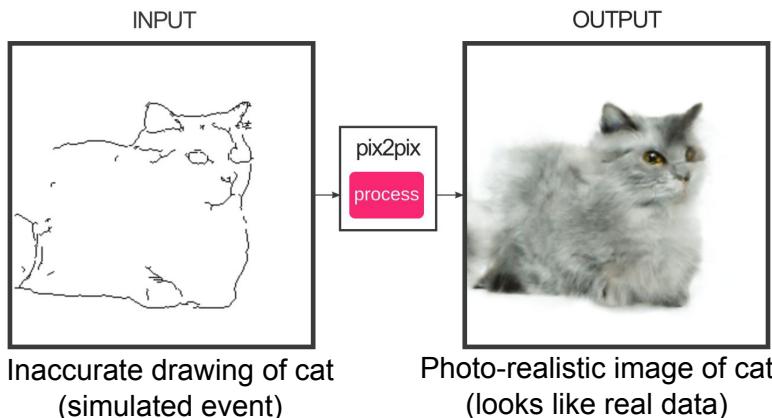
Generative networks create synthetic data, with several possible applications

- Generative Adversarial Networks (GANs)
  - Train generative network and classifier together
  - Classifier is trained to distinguish generated data from training data
  - Generative network is trained to generate data the classifier cannot distinguish
- Can train using real data (e.g. calibration data, control samples, or general ‘unlabelled’ data)
  - Avoid biases / systematics from imperfect detector simulation models
- Potential use for noise reduction
  - Train generative network to produce de-noised events from noisy events
- Potential uses for detector calibration
  - Train network to modify simulated events to more closely match real data



GAN generated events

Geant4 simulated events



# Summary

Hyper-Kamiokande, the next-generation water Cherenkov neutrino detector has begun construction to start operation in 2027

- Both the far detector and IWCD will require new techniques to improve reconstruction, suppress backgrounds and reduce systematics

Machine learning can bypass the model approximations of old methods

- ResNet CNN and PointNet architectures already outperforming traditional methods
  - Improved reconstruction of particle position, direction and energy
  - Classification of particle types improves on existing selections and enables new analyses
- Additional benefit of huge increase in speed of reconstruction

Exploring other areas where machine learning can provide benefits

- Segmentation of multi-ring and multi-vertex events looks promising
- Generative networks allow hybrid ML/traditional reconstruction and novel approaches to handle detector calibration and modelling



**Hyper-Kamiokande**



**WatChMaL.org**

# Appendix

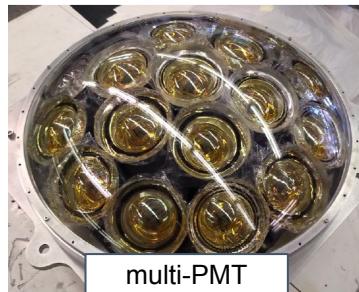
# Hyper-K Detector

8 x increase in fiducial mass over Super-K

- 71 m tall x 68 m diameter = 258 kt total mass  
188 kt fiducial mass

New photo-detector technology for increased sensitivity

- 20,000 B&L 50 cm PMTs = 20% photo-coverage
  - 1.5 ns timing resolution (half that of SK PMTs)
  - Double quantum efficiency of SK PMTs
- Additional photo-coverage from multi-PMT modules
  - 8 cm PMTs grouped in modules of 19 PMTs
  - Improved position, timing, direction resolution
  - Also used for in-situ calibration of 50cm PMTs

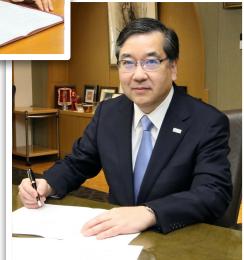


# The Hyper-K Experiment

February 2020: Budget approved by Japanese government

May 2020: Univ. of Tokyo President and KEK Director General signed MOU:

Univ. of Tokyo to construct & operate Hyper-K detector  
KEK to upgrade & operate J-PARC neutrino beam



# Hyper-K's WC detectors

## Hyper-K far detector

3rd generation of WC detectors at Kamioka

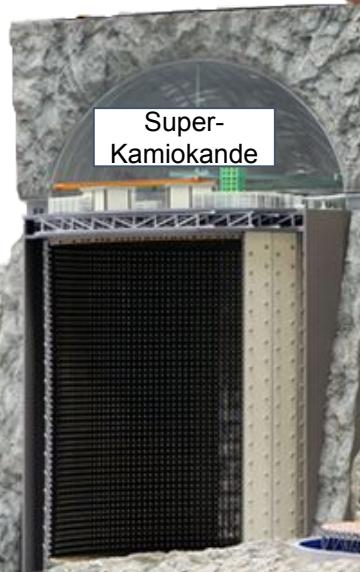
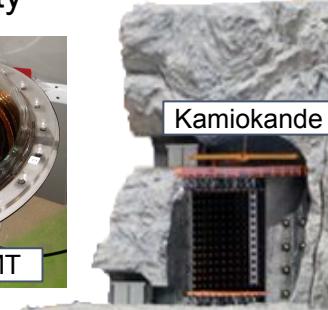
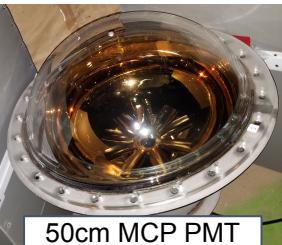
8 x increase in fiducial mass over Super-K

72 m tall x 68 m diameter = 258 kt total mass

188 kt fiducial mass

Baseline design: 40,000 B&L 50 cm PMTs  
= 40% photo-coverage

New photo-detector technology to  
provide increased sensitivity



# Hyper-K's WC detectors

## Intermediate detector (IWCD)

Located ~ 1 km from beam source

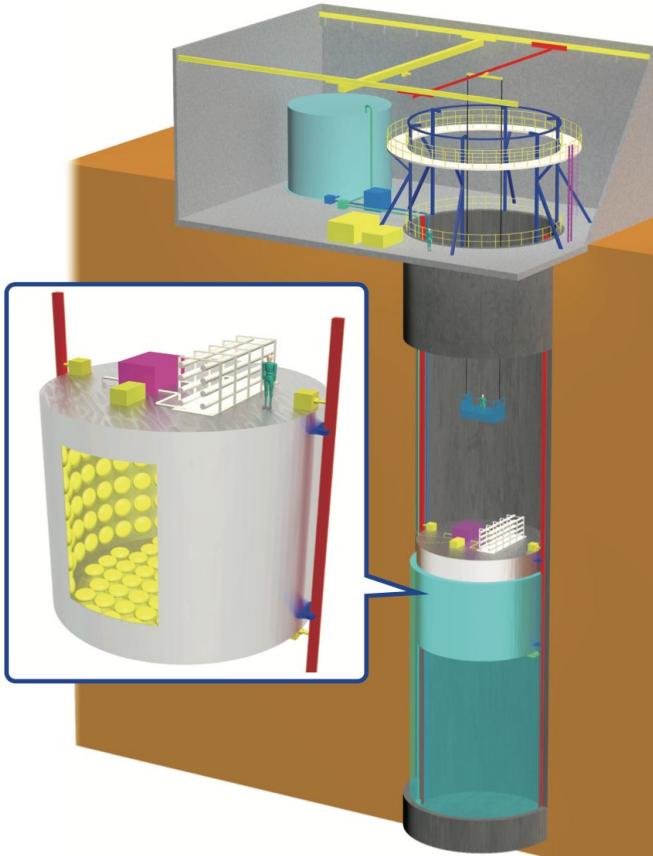
6 m tall x 8 m diameter inner detector

~ 500 multi-PMT modules

Measure combination of flux and cross-section to reduce systematics at far detector

High event rate, same detector technology and target nuclei as far detector

Moves vertically in ~50 m tall pit  
measuring different off-axis angles gives different  $\nu$  energy spectra



# Hyper-K's WC detectors

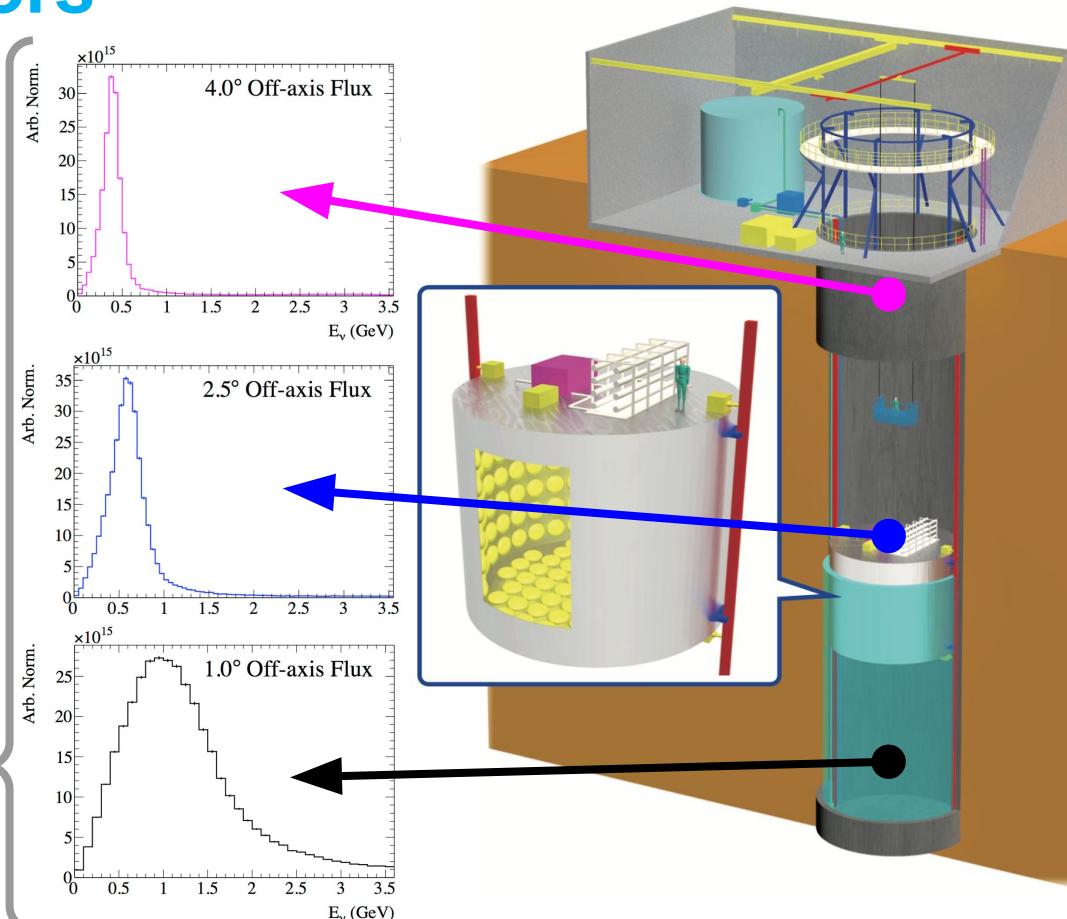
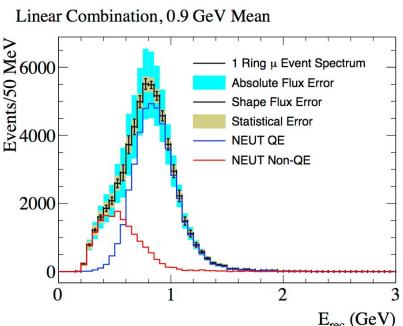
## Off-axis spanning detector

$\nu$  energy spectrum depends on angle off-axis to the neutrino beam

Far detector @  $2.5^\circ$  for peak at  $\sim 600$  MeV

Moving IWCD varies angle, allowing measurements at different energies

Linear combinations allows mimicking monochromatic beam or far-detector spectrum



# Hyper-K's WC detectors

## Multi-PMT modules

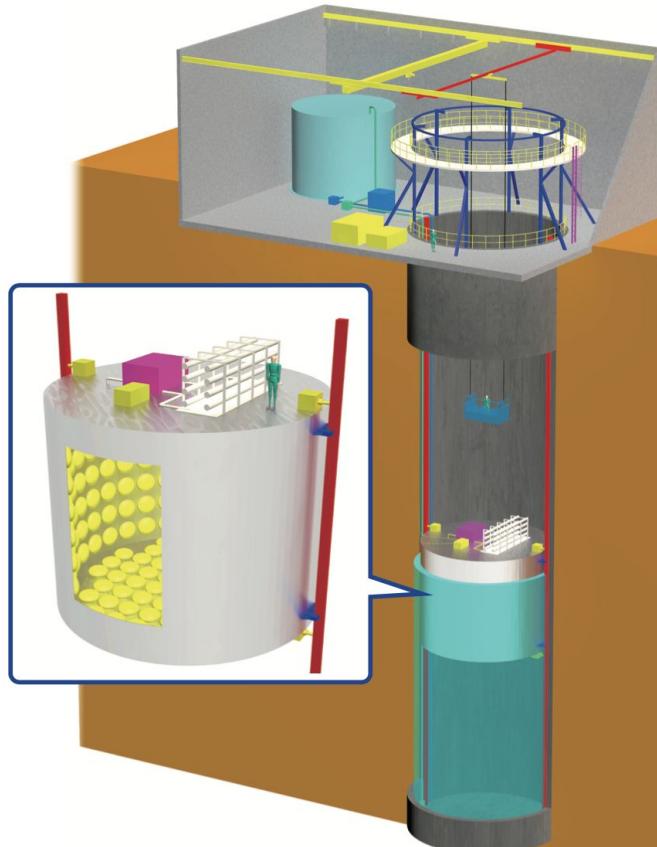
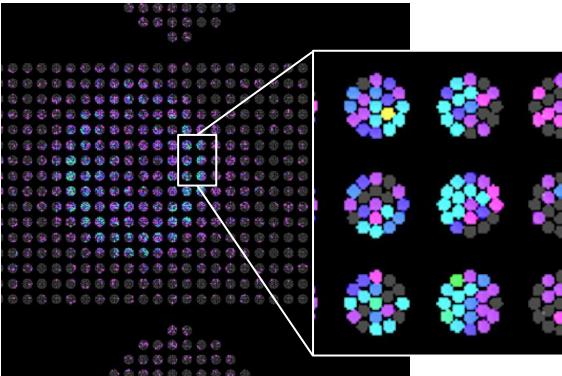
8 cm PMTs: Better position resolution

< 1 ns timing resolution

Additional directionality information

Need reconstruction to exploit additional information

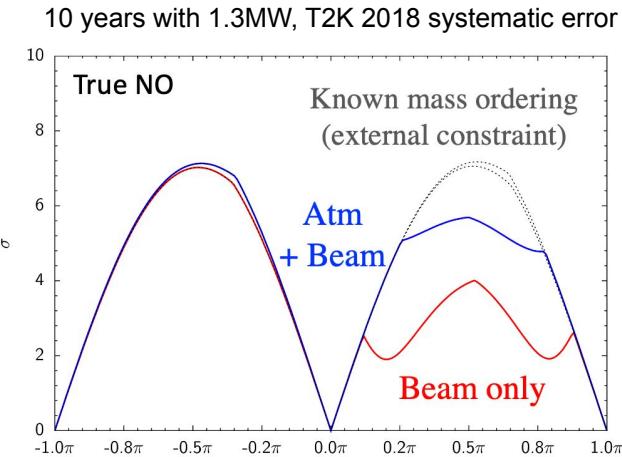
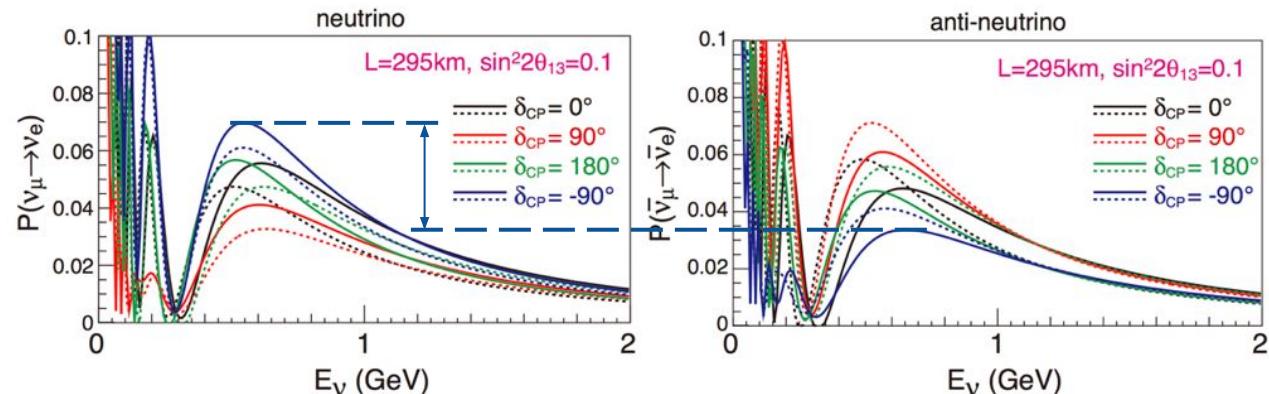
**Necessary for smaller detector size**



*Also under investigation: Combining 50 cm PMTs + multi-PMT modules in far detector*

# Hyper-K's physics goals

## Long-baseline neutrino oscillations: CP violation

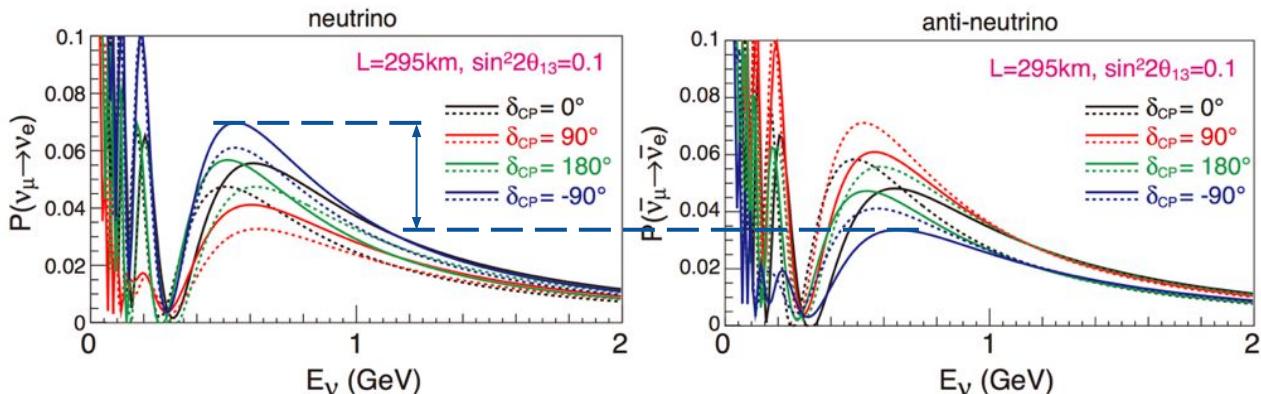


Combine beam and atmospheric neutrino observations for maximum sensitivity

- $\delta_{CP}$  precision comes mostly through difference in  $P(\nu_\mu \rightarrow \nu_e)$  vs  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$
- Effect of  $\delta_{CP}$  can be degenerate with normal vs inverted mass ordering
- Atmospheric  $\nu$ 's gain sensitivity to mass ordering by exploiting matter effect of Earth on oscillations

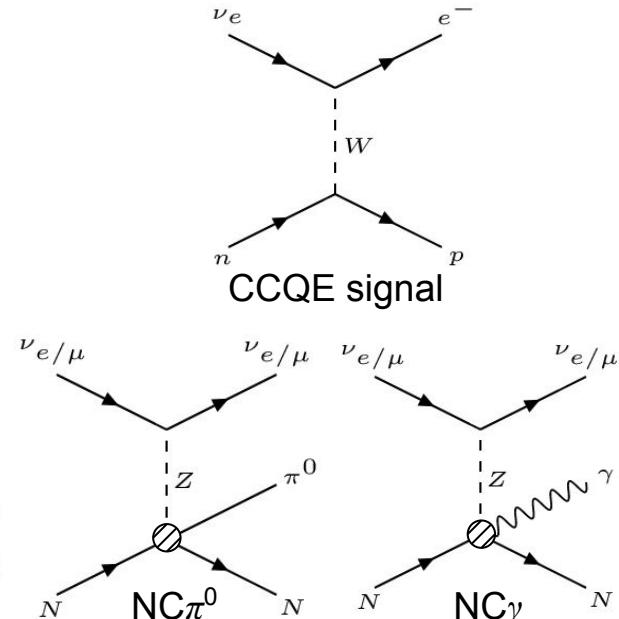
# Hyper-K's physics goals

## Long-baseline neutrino oscillations: CP violation



Oscillation maximum is at around 0.6 GeV

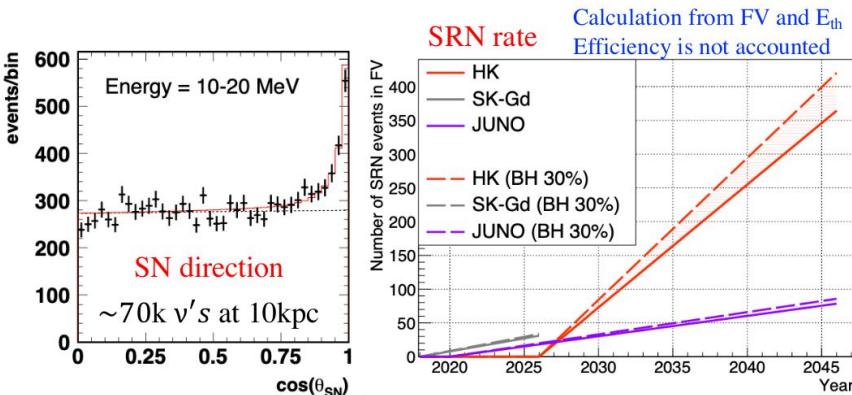
- Dominant signal  $\nu_e$  interaction is charged current quasielastic (CCQE)
- Potential background sources:
  - Neutral current interactions ( $\nu_e$  or  $\nu_\mu$ ) producing neutral pions or gammas
  - Muons from  $\nu_\mu$  misidentified as electrons from  $\nu_e$



# Hyper-K's physics goals

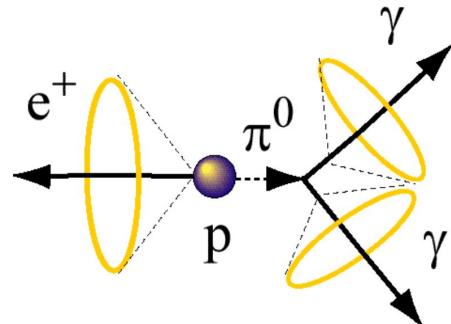
## Neutrino astrophysics

- Solar  $\nu$ 's: day/night asymmetry; hep  $\nu$ 's;  ${}^8\text{B}$   $\nu$  spectrum upturn
- Supernova  $\nu$ 's: 1000's  $\nu$  events for nearby supernova pointing, time & spectrum analysis; search for supernova relic  $\nu$ 's



## Proton decay

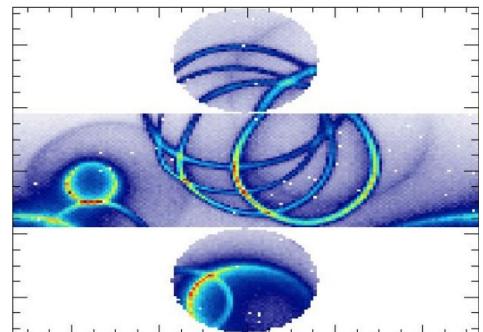
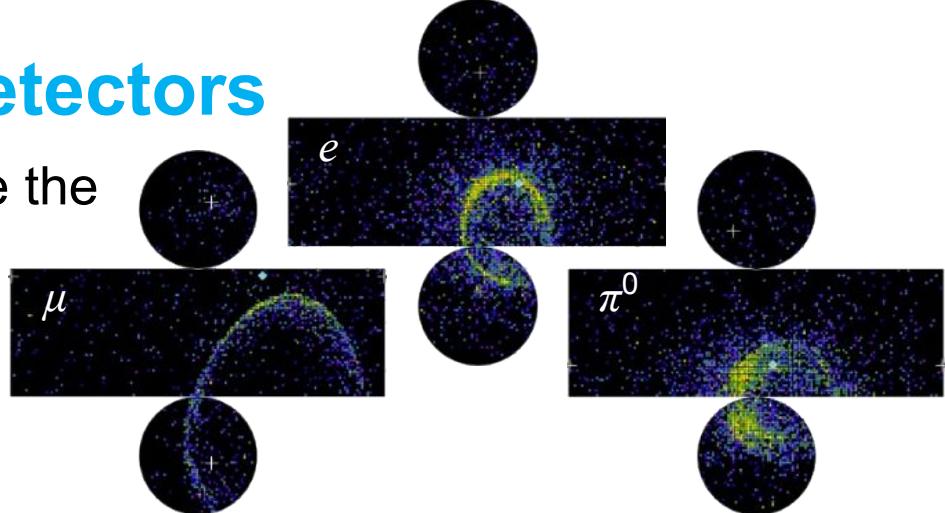
- Search to order of magnitude greater lifetime than current limit
- $10^{35}$  years for  $p \rightarrow e^+ + \pi^0$
- $3 \times 10^{34}$  years for  $p \rightarrow \bar{\nu} + K^+$



# Reconstruction for WC detectors

Take raw detector data and determine the physics that occurred

- Particle type identification
  - Separate signal events from background
- Particle momentum, direction, position
  - Kinematics essential to determine incoming neutrino energy for neutrino oscillation probability
- Separating & reconstructing multi-ring events
  - Events with multiple particles / rings contribute to both signal & background
  - Multiple neutrinos can interact around the same time in IWCD

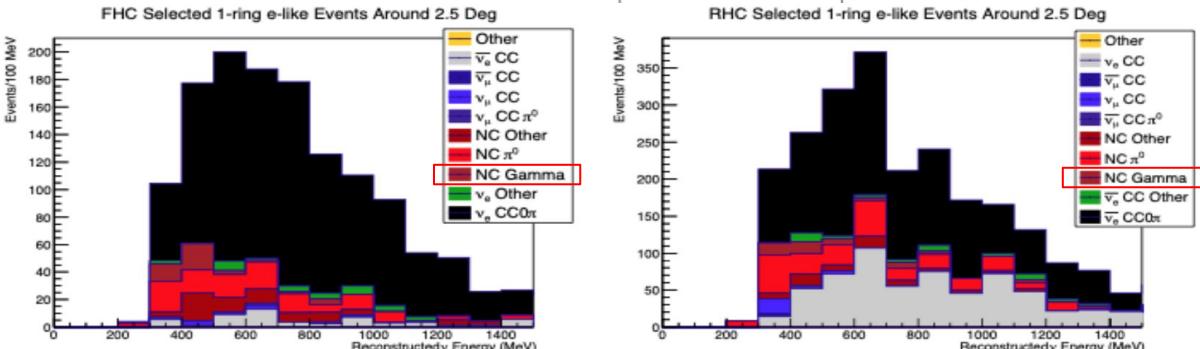


# Physics Motivations

New opportunities beyond simple reconstruction improvement

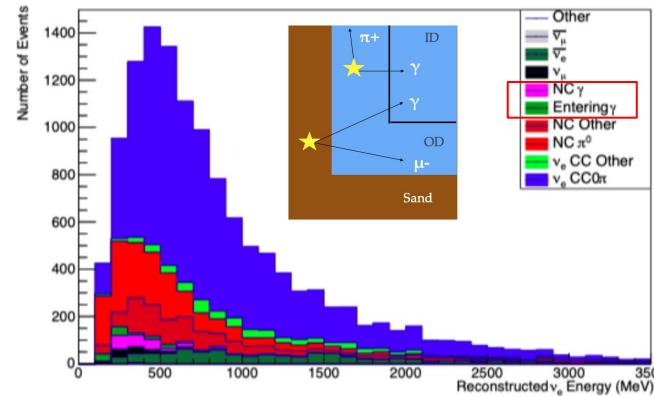
- NC  $\gamma$  discrimination and measurement

Systematic Type	Requirement	Motivation
$\sigma(\nu_e)/\sigma(\nu_\mu)$ to $\sigma(\bar{\nu}_e)/\sigma(\bar{\nu}_\mu)$ ratio	2.9-3.7%	CP violation search and precision $\delta_{CP}$ measurement.
$\sigma(\nu_e)/\sigma(\nu_\mu)$ , $\sigma(\bar{\nu}_e)/\sigma(\bar{\nu}_\mu)$	3-5%	$\theta_{23}$ octant and precision $\theta_{23}$ measurement.
Intrinsic $\nu_e$ , $\bar{\nu}_e$ and NC background normalizations	3-5%	CP violation search and precision $\delta_{CP}$ measurement.



- Bottom-up calibration: Enable multitude of detector parameter variations

2.7-4.0 degree off-axis range  
Selected 1-ring e-like events



- Potential neutron tagging application

# Traditional reconstruction method

fiTQun: Likelihood-based reconstruction for higher energies

- Originally developed for Super-K detector
  - Based on algorithm of MiniBooNE: <https://arxiv.org/abs/0902.2222>
- Uses full information of unhit PMTs + time & charge of hit PMTs:

$$L(\mathbf{x}) = \prod_j^{unhit} P_j(\text{unhit}|\mathbf{x}) \prod_i^{hit} P_i(\text{hit}|\mathbf{x}) f_q(q_i|\mathbf{x}) f_t(t_i|\mathbf{x})$$

Likelihood to maximise      Candidate event hypothesis      Probability of no hit at PMT      Probability of hit at PMT      Hit charge probability density      Hit time probability density

- Probabilities calculated based on direct + scattered + reflected light
- Likelihood ratios used to distinguish particle types and single-ring / multi-ring event topology hypotheses

# Machine learning reconstruction

**WatChMaL**: cross-collaboration group formed to explore ML for WC

Common challenges for ML with WC detectors

- Cylindrical geometry
- High-resolution, sparse data

Many physics goals

- Maximise precision of new detectors
- Reconstruct complex event topologies
- Discriminate electron and gamma rings
- Improving detector calibration & systematics

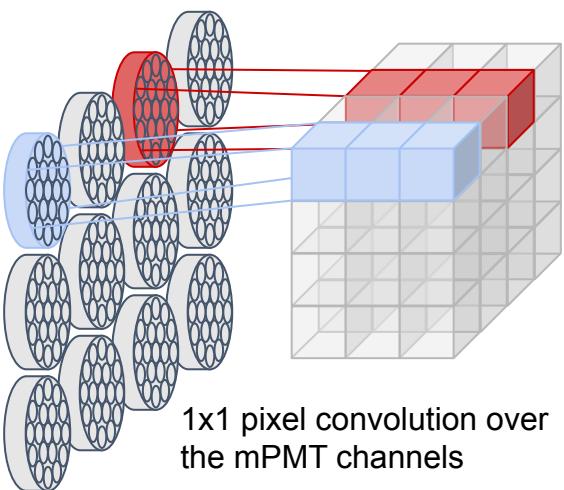


WatChMaL.org

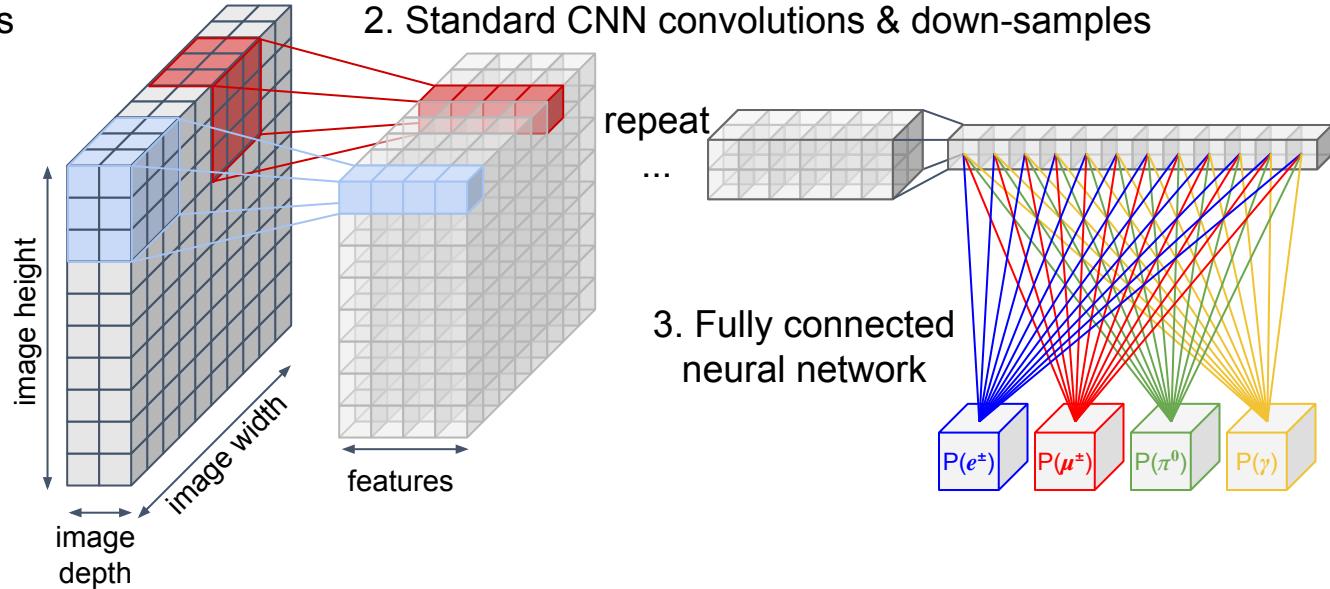


# CNN Architecture

1. Convolution over mPMTs



2. Standard CNN convolutions & down-samples

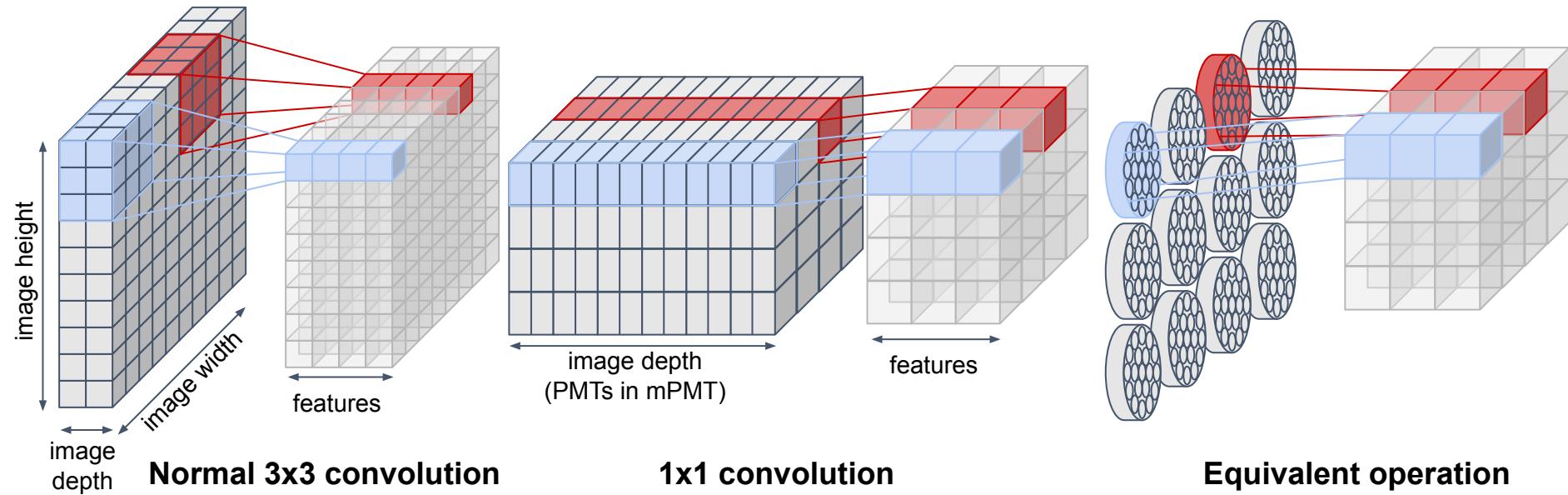


Network based on ResNet-18 CNN architecture [arXiv:1512.03385]

- Replaced initial 7x7 pixel convolution with 1x1 convolutions over all channels
  - Equivalent to convolution over the 19 PMTs within each mPMT

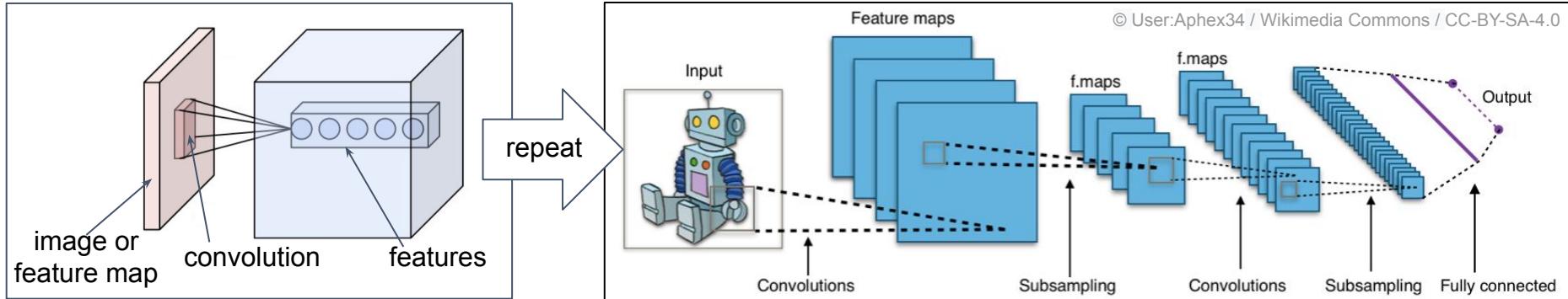
# CNN architecture

Treating each PMT inside mPMT as a channel, starting with  $1 \times 1$  convolution  
→ equivalent to doing a ‘convolution’ over each mPMT



# CNN architecture

Convolutional neural networks hugely successful in image processing

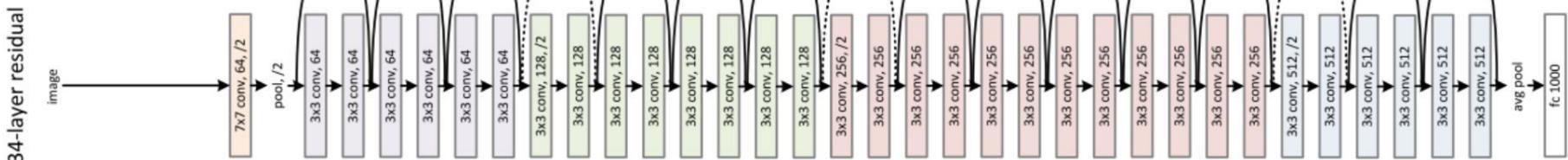
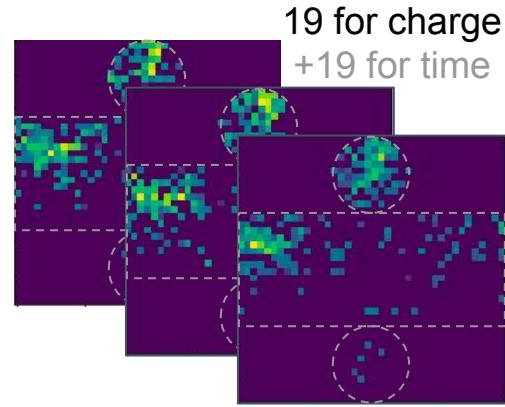


- Start with image with pixel values ('features'): T and Q at each PMT
- Scan many small (e.g. 3x3) convolution kernels across image
  - Increases number of features
- Downsample image (e.g. 2x2 max-pooling)
  - Decreases number of pixels
- End with 1-D array of features, feed into traditional fully-connected neural network
- Learn convolution and final network weights through 'back-propagation' of loss

# CNN architecture

Full cylinder of mPMTs is unwrapped onto 40x40 image

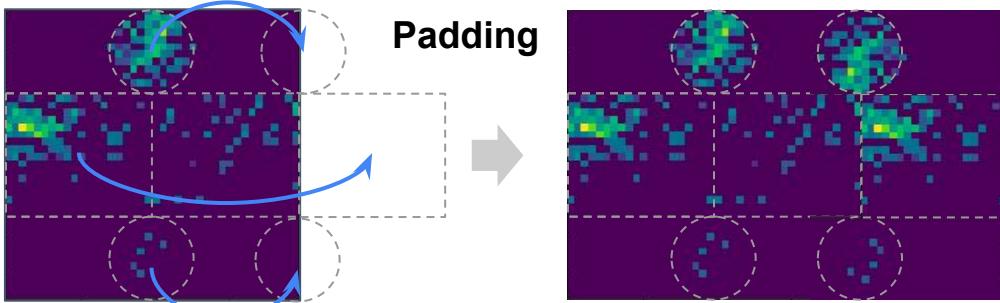
- 38 channels: charge & time of 19 PMTs per mPMT
- No special treatment for geometrical effects at boundary between barrel and end-caps
- Data augmented by reflecting / rotating around tank axis



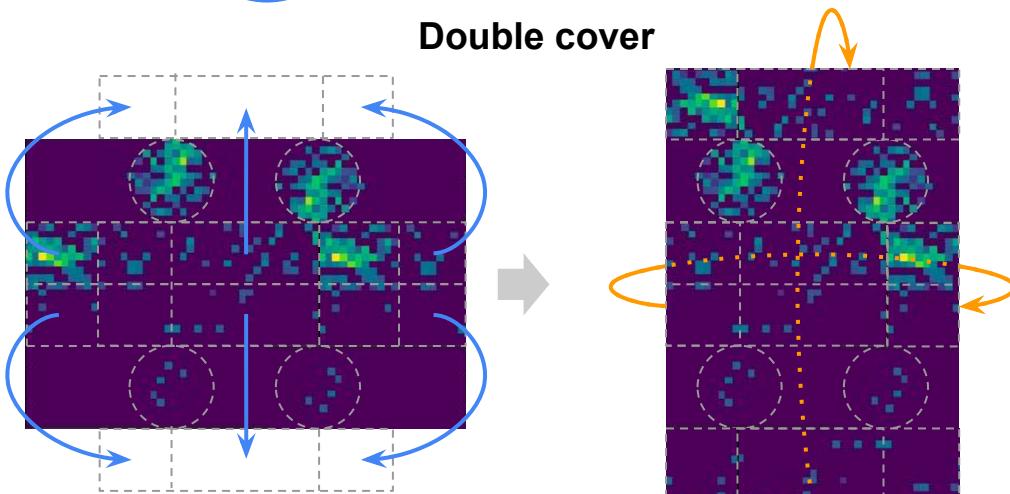
Mostly using ResNet-18 architecture [arXiv:1512.03385]

- Initial 1x1 convolution added to act on the 19 PMTs of each mPMT
- Also explored deeper networks with small improvement

# 'Double cover' images



**Double cover**



'Padding' the image improves accuracy for some events

- Original image 'slices' along barrel at arbitrary position
- Some events have rings that span this slice
- Repeat part of the image after rotating tank to help CNN learn events where ring is sliced

Rearranging and duplicating in a more complex pattern has additional advantages

- All segments appear exactly twice
- **Circular boundary conditions** in both directions
- Minimal blank space

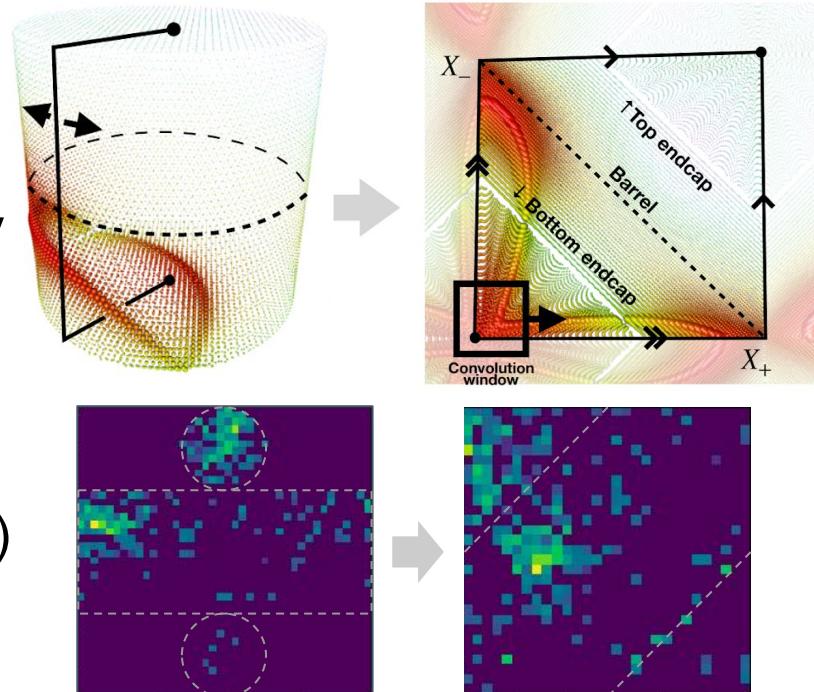
# Topological map to square

Alternative map onto square with boundary conditions preserving topology of cylinder

- Cut open along barrel to centre of end caps (solid line)
- Deform onto square, keeping density of PMTs constant
- Place mPMTs onto nearest pixel
- Use boundary conditions identifying edges of square (indicated by arrows)
  - Pad image with copy of pixels from the corresponding edge

$$X_{\pm} = W(\rho, z) \frac{\pi \pm \phi}{2\pi}$$
$$W(\rho, z) = \sqrt{\frac{\rho^2 + 2Rz + RH}{R^2 + RH}}$$

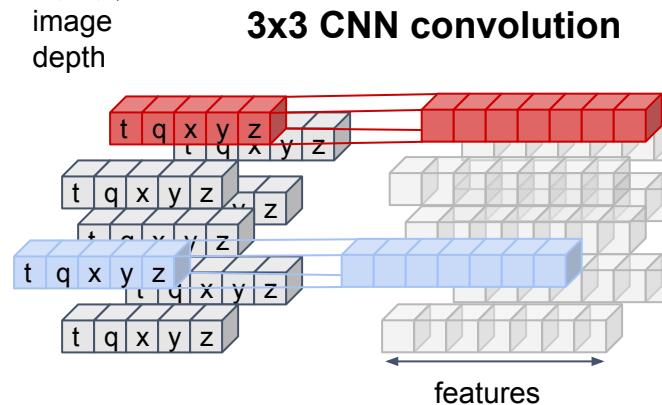
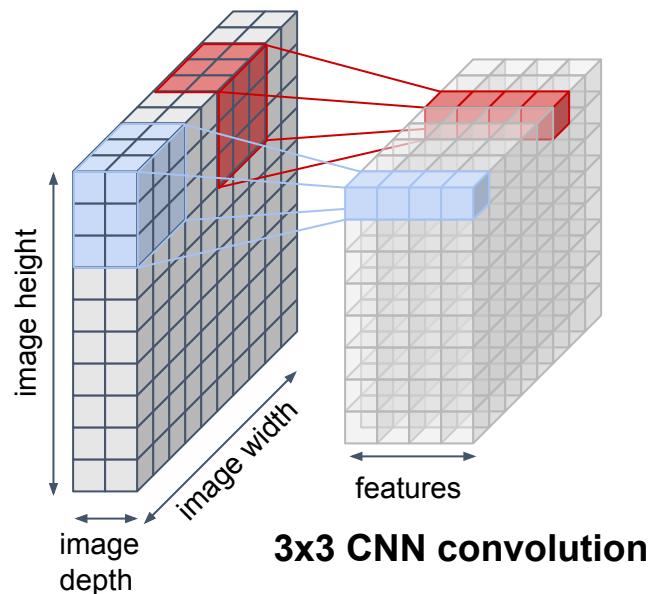
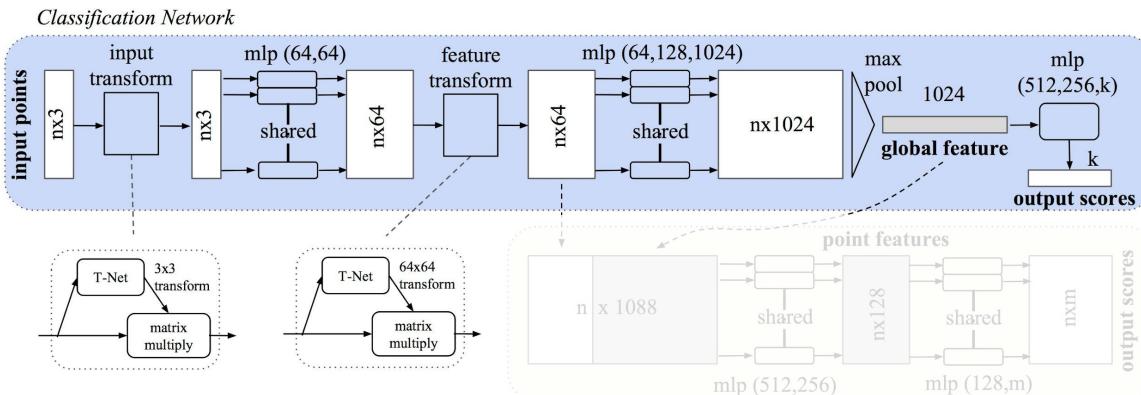
Solve differential eq. for constant Jacobian  
 $dX_+ dX_- = \left| \frac{\partial(X_+, X_-)}{\partial(\rho, \phi)} \right| d\rho d\phi$



# PointNet architecture

PointNet designed to work on ‘point clouds’ rather than images of pixels

- Each hit PMT is a ‘point’ with time, charge & position, not fixed to grid
- Convolution-like operations act on each point’s charge, time and position
- Learn global transformations applied to all points
- Single pooling layer from all points to 1D array
- Can apply to any detector geometry

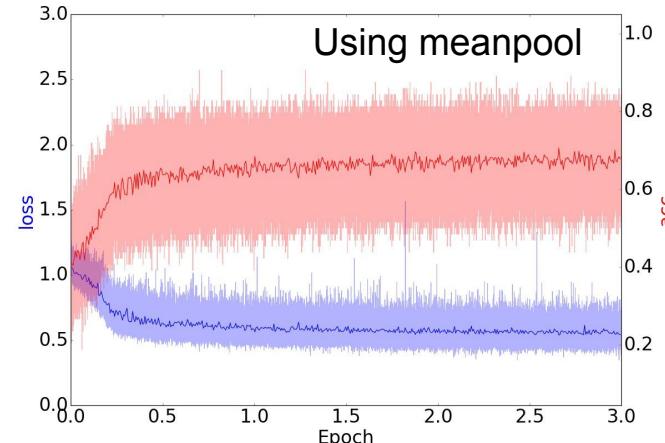
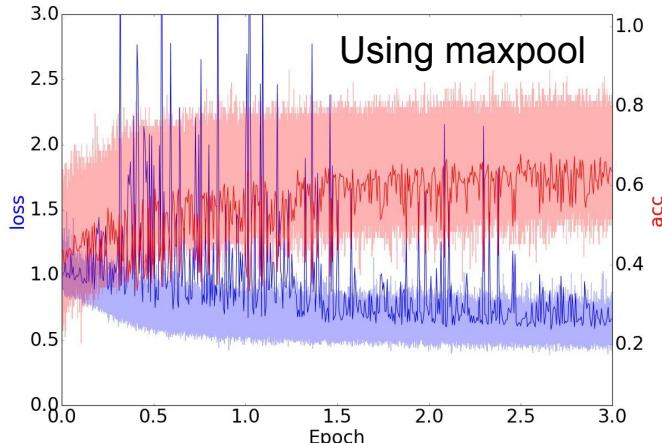


**PointNet MLP (1x1 convolution on point cloud)**

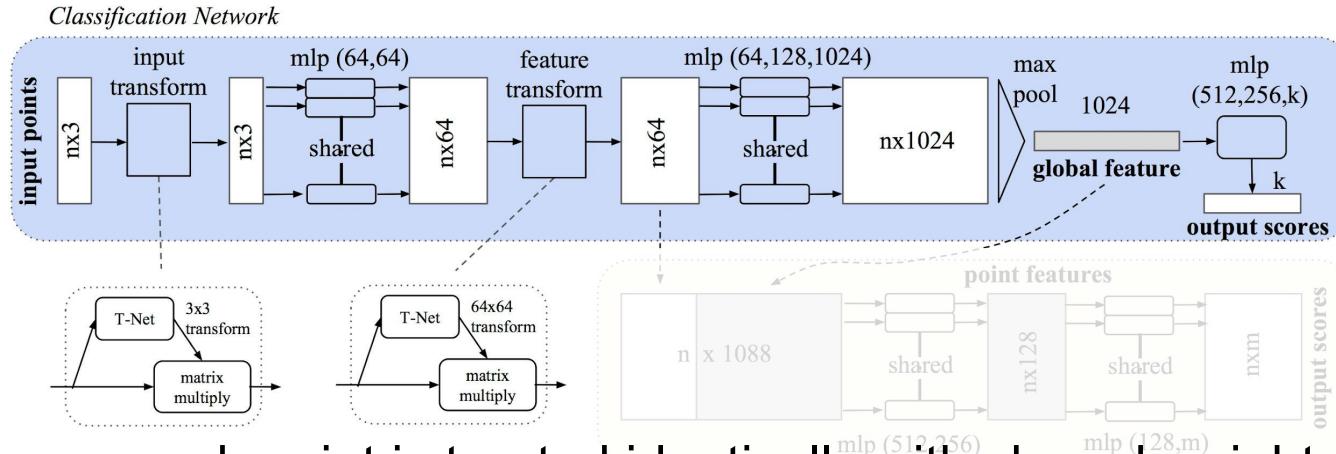
# PointNet architecture

Some changes to standard PointNet give improvements

- Severe overfitting until max. features reduced from 1024 to 256
  - Possibly due to limited batch size with larger network
  - Data augmentation could also help
- We find that mean pool works better than standard max pool here
  - PointNet usually picks key points to learn features, but aggregating information from all points seems better for our tasks



# PointNet architecture



In MLP layers, each point is treated identically with shared weights

- Similar to each pixel treated the identically in a CNN
- But without downsampling, information does not transfer between points

Instead 'T-Nets', resembling PointNet, learn transformations of the points

- Linear transformation is learnt to e.g. rotate all input vectors
- Feature transform allows global information to affect individual points

Single downsampling layer at the end of the network collapses all points

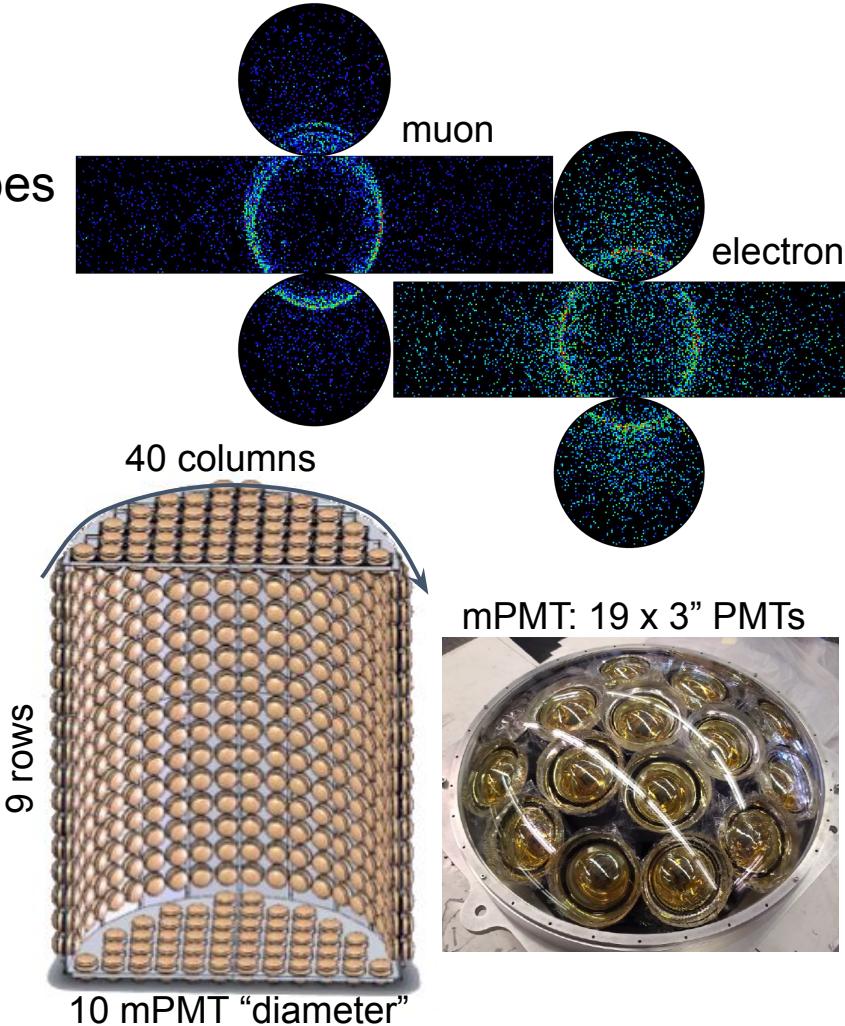
# Particle type classification

Initial studies to classify  $\mu / \pi^0 / e / \gamma$  particle types

- $\mu$  vs  $e$  is classified extremely well by traditional methods (>99% accuracy)
- $e$  vs  $\pi^0$  works reasonably well, but could be improved
- $e$  vs  $\gamma$  has not been used successfully with traditional methods

Simulated 3M of each type in IWCD detector

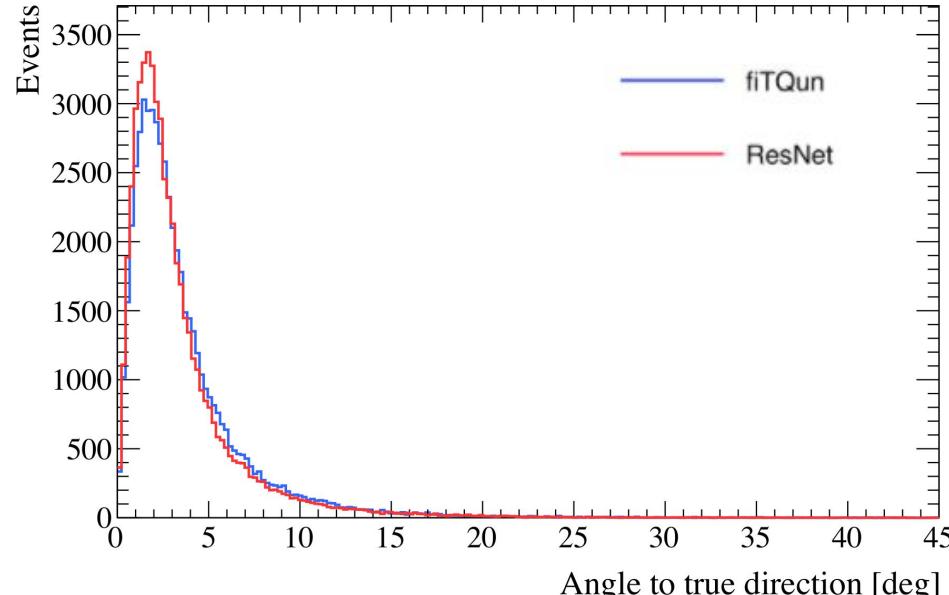
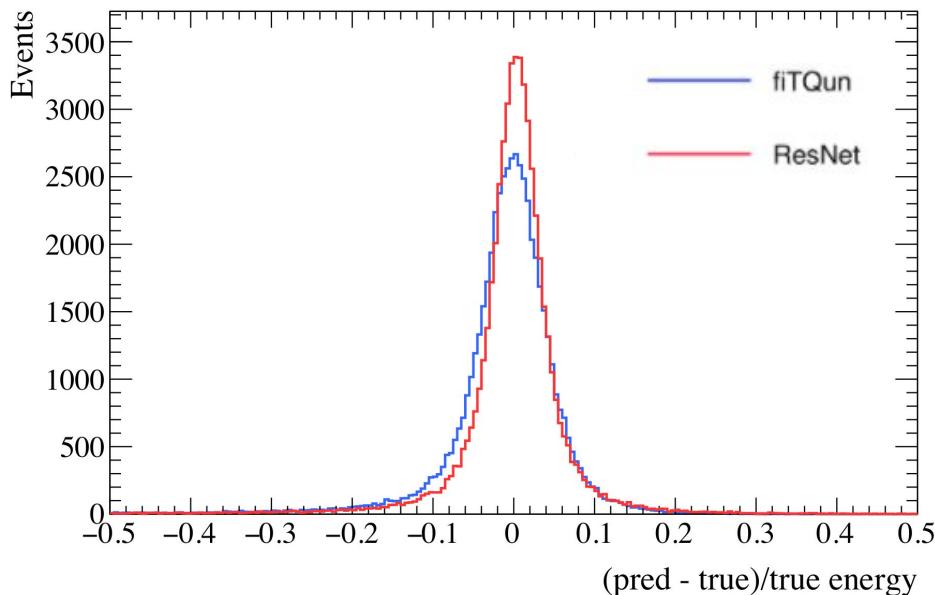
- 0 - 1 GeV energy above threshold
- Uniform positions, isotropic directions
- Split full dataset into 50% : 10% : 40% for training : validation : testing



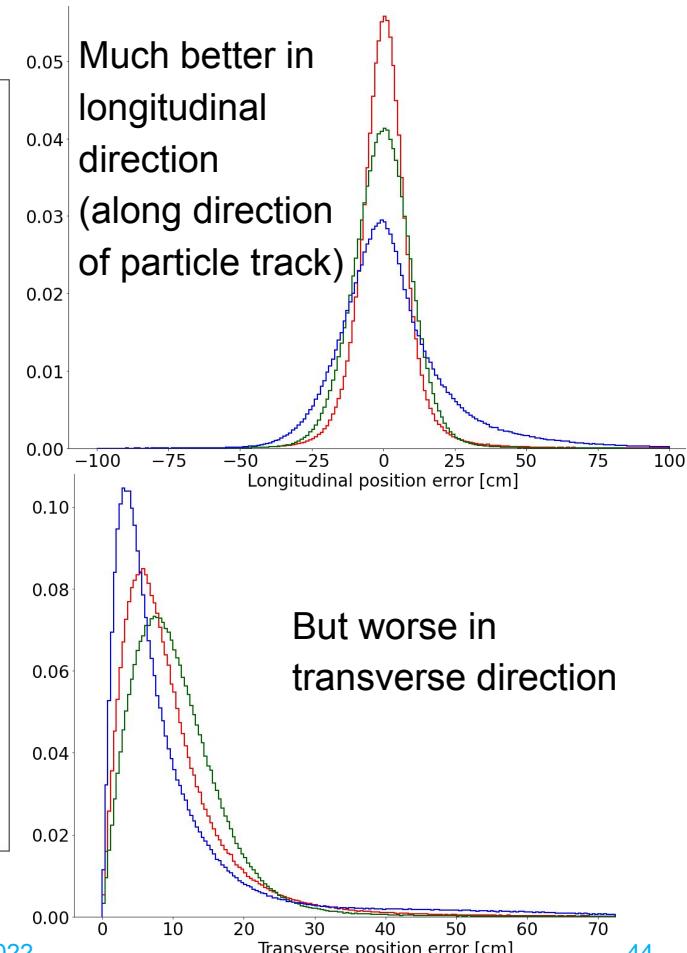
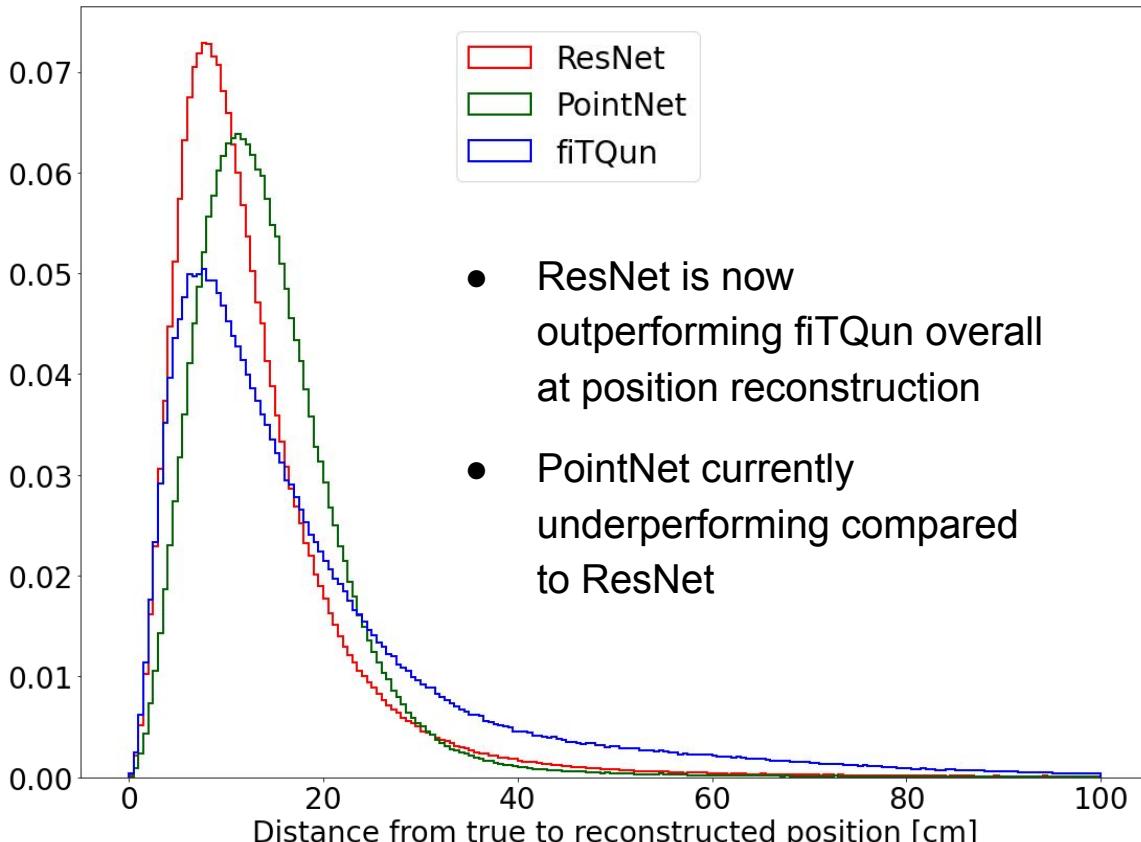
# Position, direction, energy reconstruction

Similar ResNet and PointNet architectures as used for classification

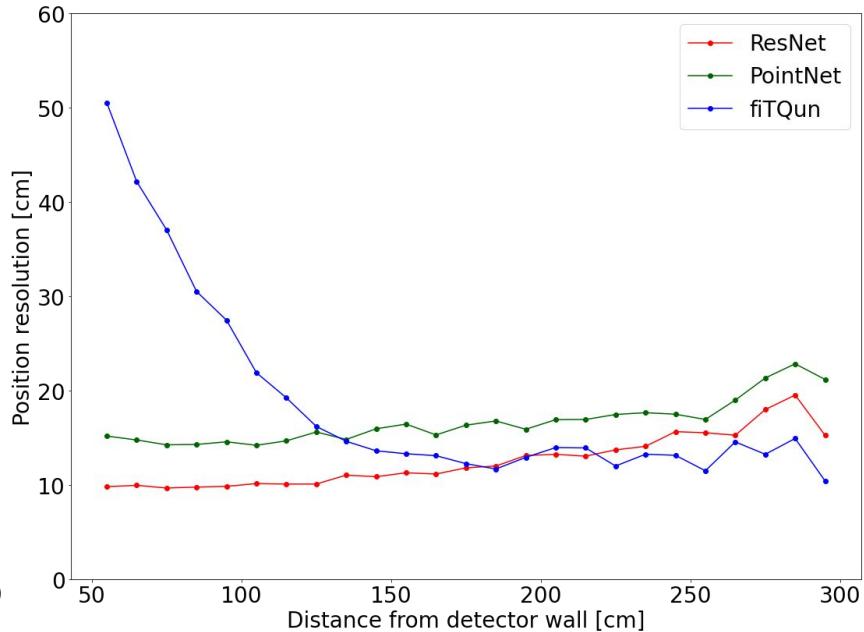
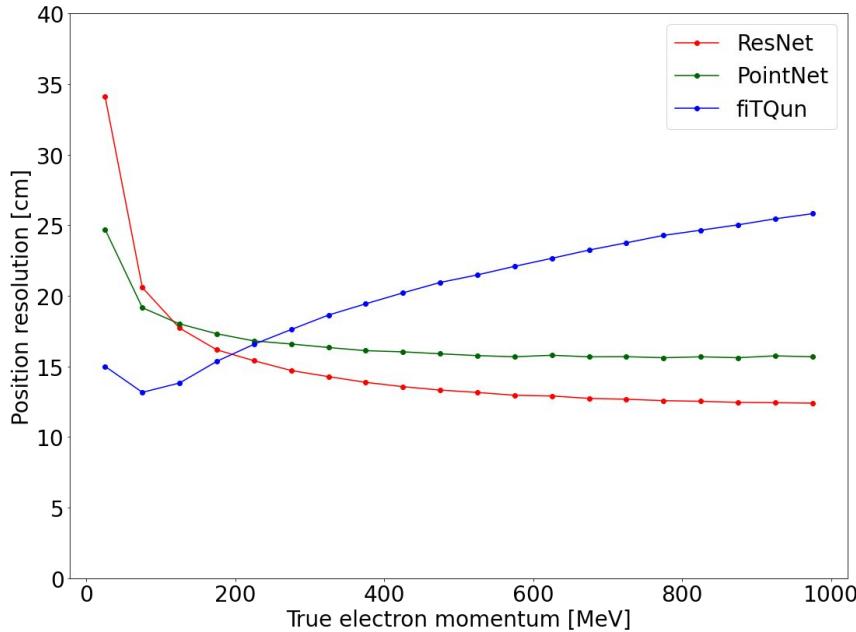
- Output reconstructed quantities instead of classification variables
- Use Huber loss to minimise true-reconstructed residuals
- ResNet is outperforming fiTQun at energy and direction reconstruction



# Position, direction, energy reconstruction

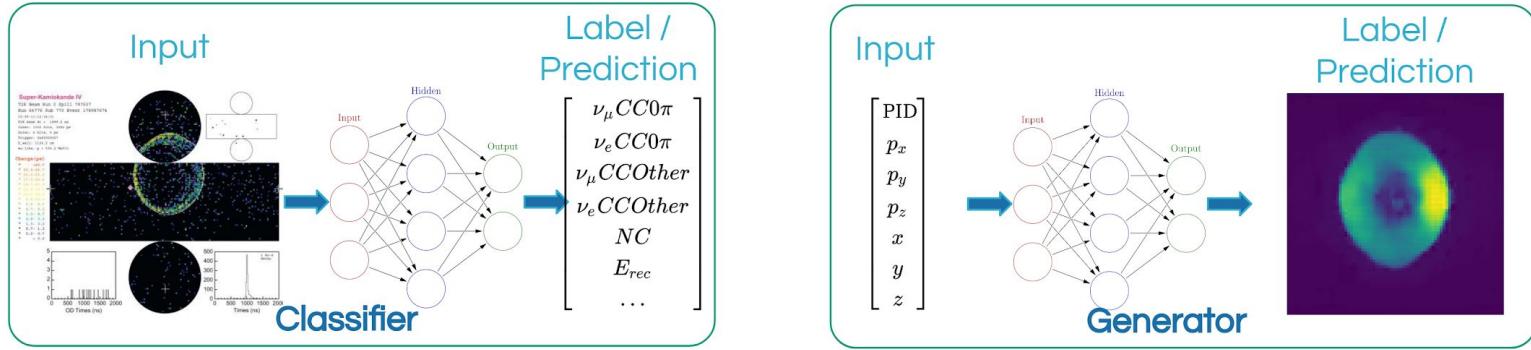


# Position, direction, energy reconstruction



- Improvement in reconstruction with ML mainly in events close to detector wall
  - Approximations in likelihood calculation break down when close to PMTs
  - Could allow expansion of detector fiducial volume to allow increased statistics
- ML reconstruction could be improved at lower energy
  - potentially struggles to learn reconstruction of sparse events

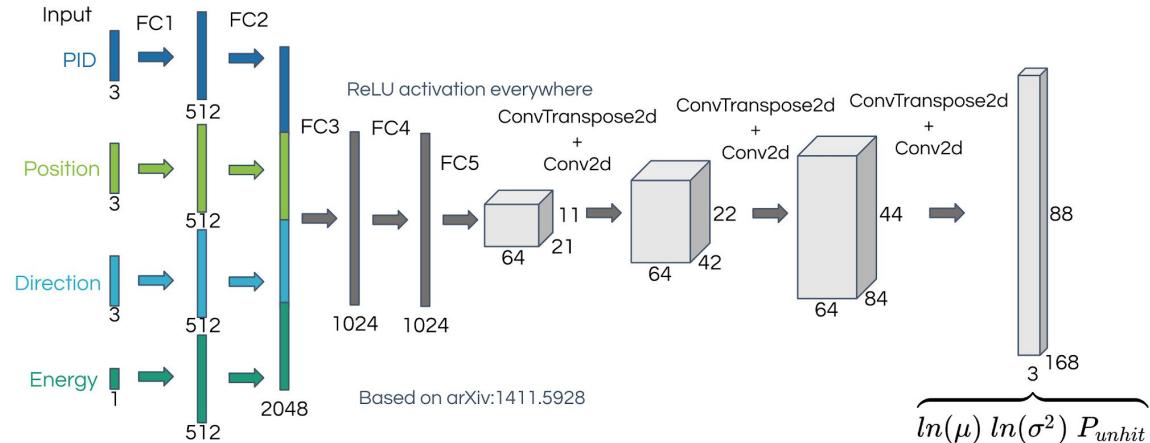
# Cherenkov ring generator



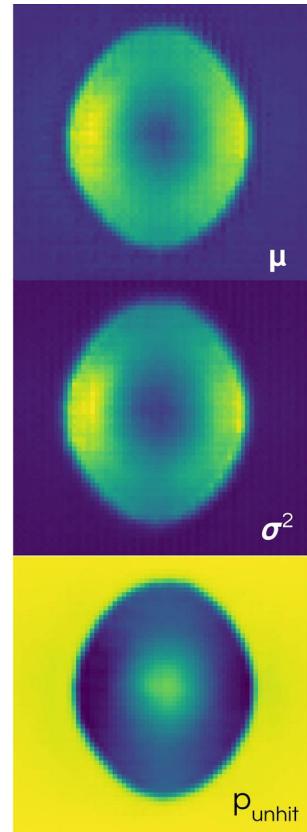
Investigating hybrid method using generative network

- Generative network can predict PMT hit charge and time
- Use to replace likelihoods in traditional reconstruction
- Combine learning ability of CNN with physics domain knowledge of traditional reconstruction
- Simple replacement for existing reconstruction in full analysis chain

# Cherenkov ring generator



$$\text{Loss} = -\ln(\mathcal{L}) = -\sum_{unhit} \ln(P_{unhit}) - \sum_{hit} \ln(1 - P_{unhit}) - \sum_{hit} \frac{1}{2} \left[ \ln(2\pi\sigma^2) + \frac{(q_{obs} - \mu)^2}{\sigma^2} \right]$$



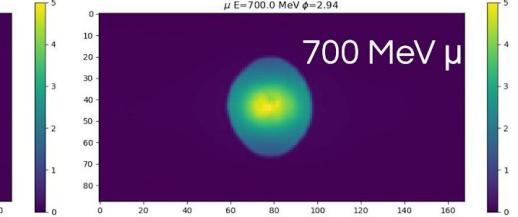
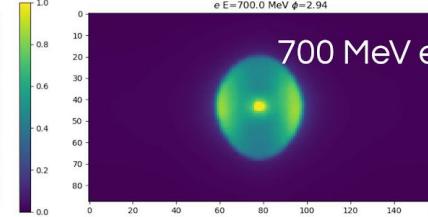
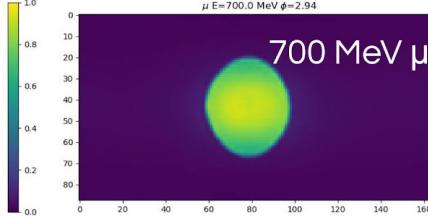
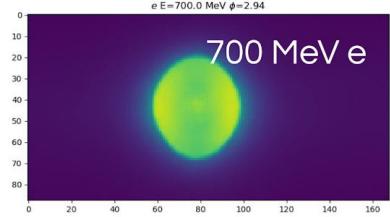
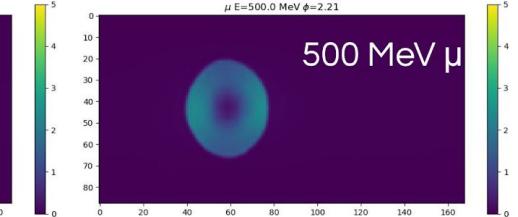
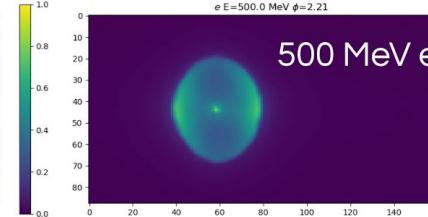
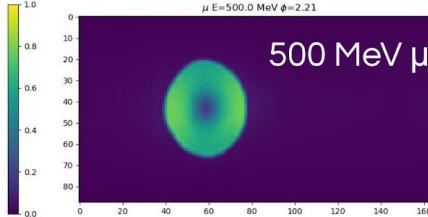
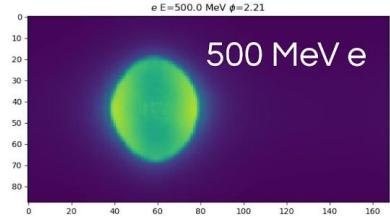
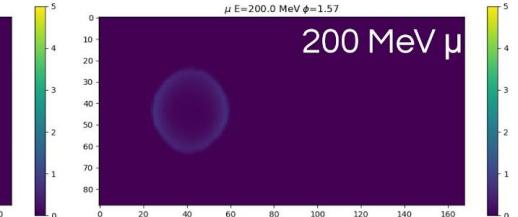
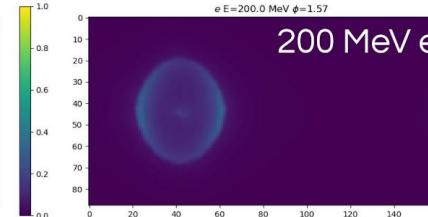
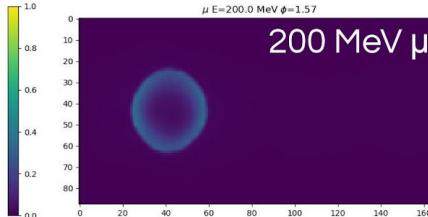
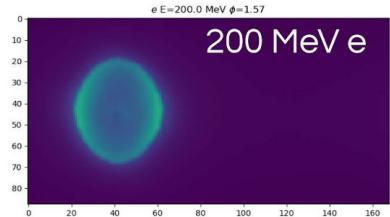
Network outputs likelihoods for hits observed at PMT

- Probability of PMT being hit
- Gaussian pdf ( $\mu, \sigma$ ) for charge

# Cherenkov ring generator

Hit probability  
x  
Mean charge

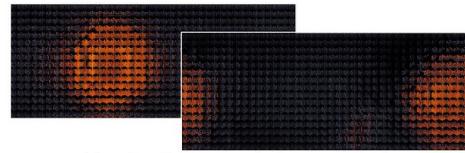
Hit probability



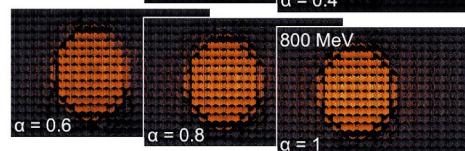
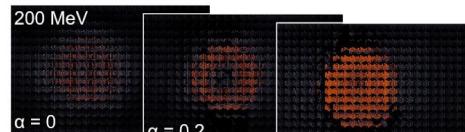
# Generative networks

Also considering using generative networks for improved detector systematics

- Train generative network to reproduce real data: removed dependence on MC
- Train GAN to take simulated event and make it look like real data
  - Reduce detector systematics by ‘fixing’ mismodelled detector simulation
- Initial work on VAE showed some promise, but struggled with noise and sharp details
- Now we are investigating GANs

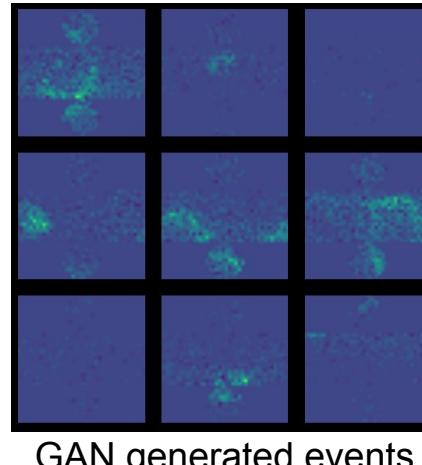


Randomly generated new events

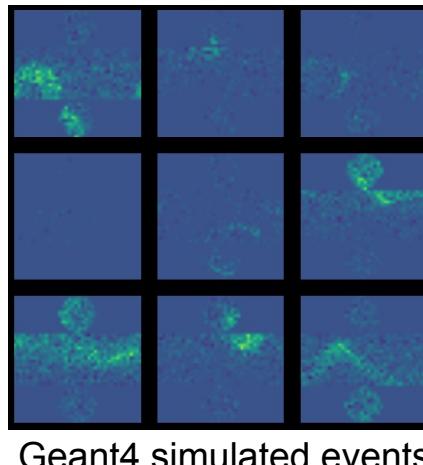


Interpolate between 200 MeV and 800 MeV events

arXiv:  
1911.02369



GAN generated events



Geant4 simulated events