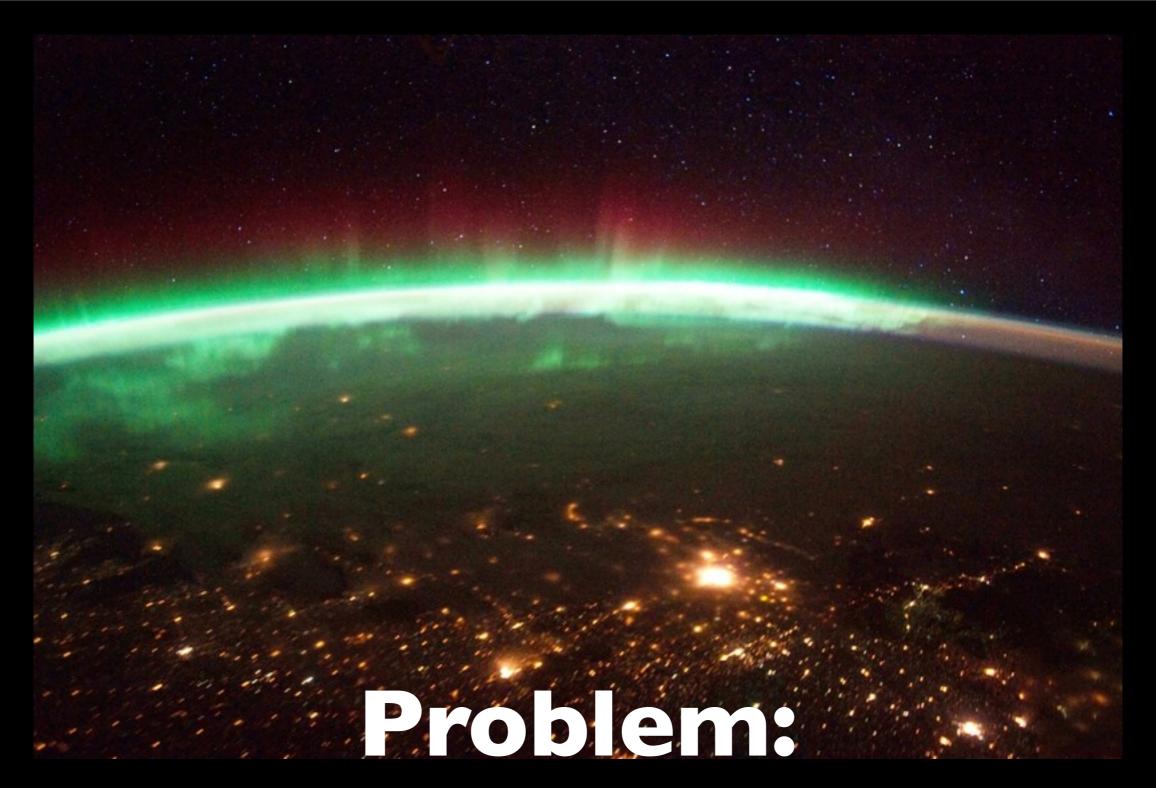


Ondrej Recnik - UWaterloo Computer Science (BSc June 2013)

Laura Culp - UWaterloo Computer Science (BMath/CS June 2013)

Fangda Li - UToronto Astronomy & Physics (BSc June 2013)

NASA SpaceApps Toronto



Use ISS image sequences as a low-cost method to determine camera orientation and extract aurora coverage on Earth

Ist Attempt: Brute Force Algorithm

Brute force algorithm

Known bright star coordinates (RA, DEC)

Star centroiding & hot pixel filtering algorithm

Choose 3 parameters: RA, DEC, rotation angle (neglect plate constant distortion*)

SNR Condition:
Mean (3x3 pixel box)
> SNR x Median
(9x9 pixel box)

Spherical projection to (x, y) pixel coordinates

(x, y) pixel coordinates of stellar centroids



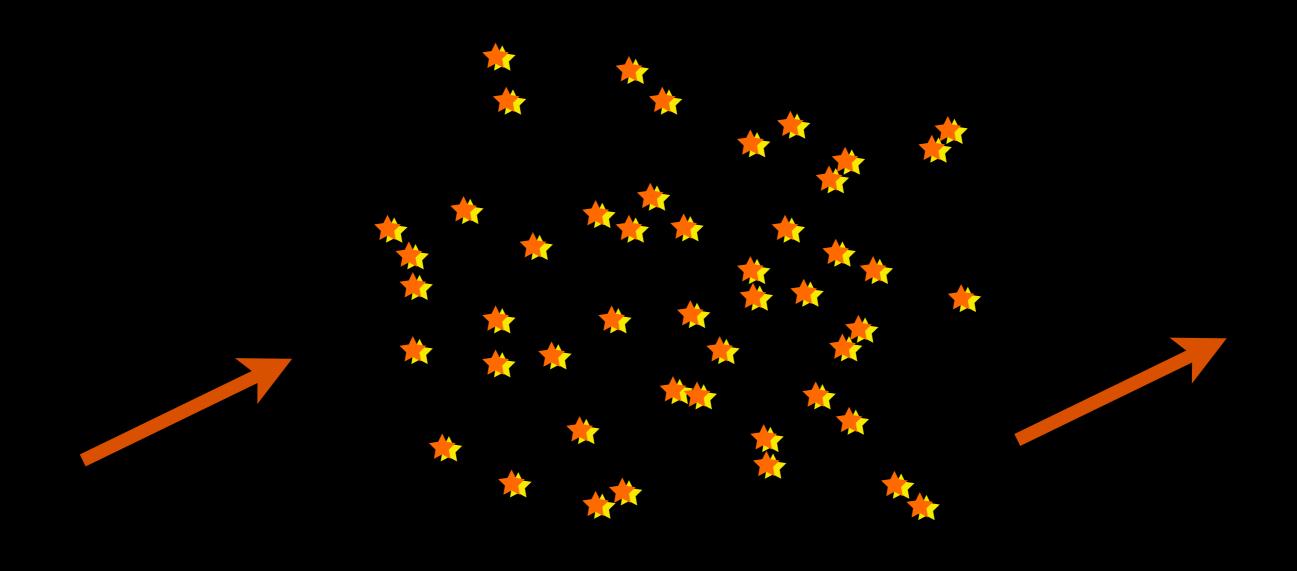
Sum distances to nearest neighbour

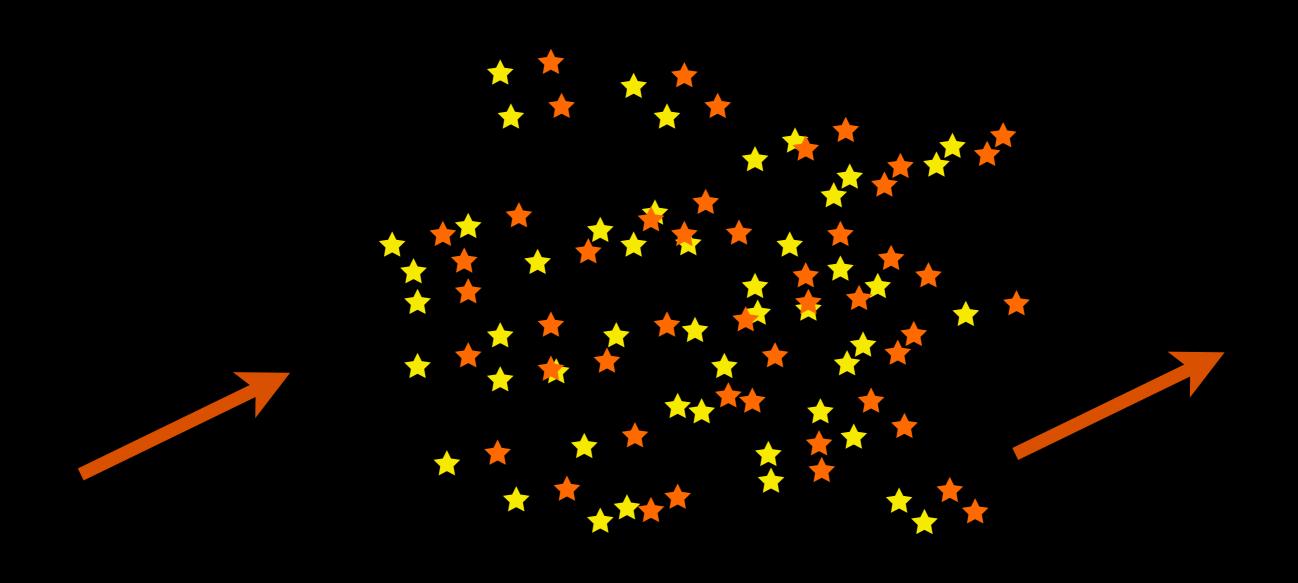




3D least squares optimization









> 10^12 image centres must be searched for least squares optimization on 50 stars in 70 degree FOV

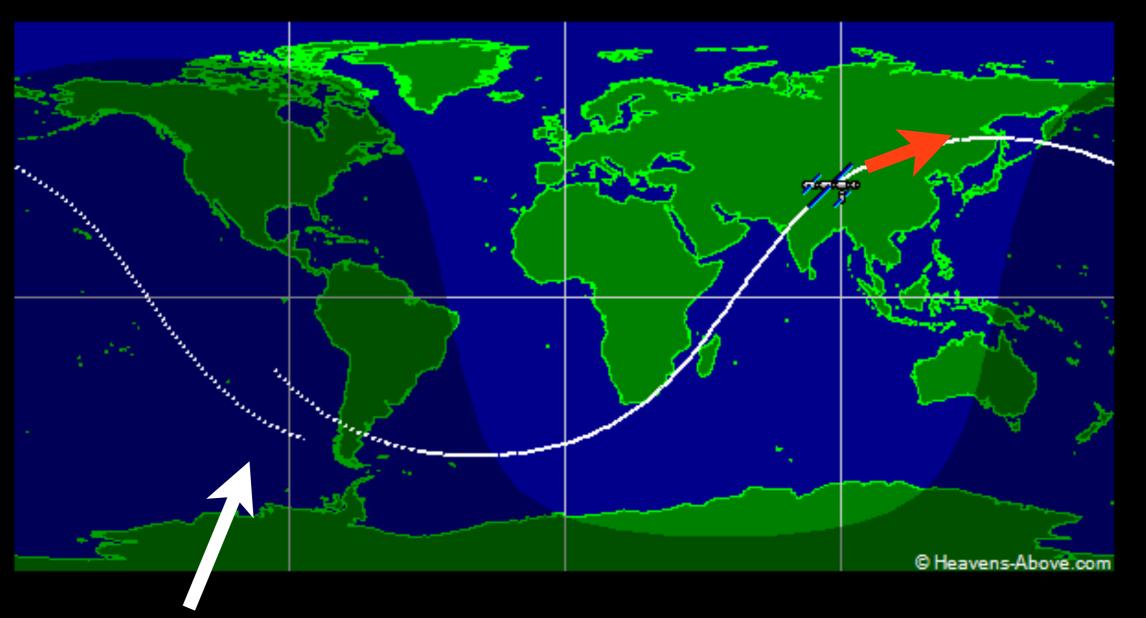


Requires a more refined method...

2nd Attempt: Incorporation of time domain information

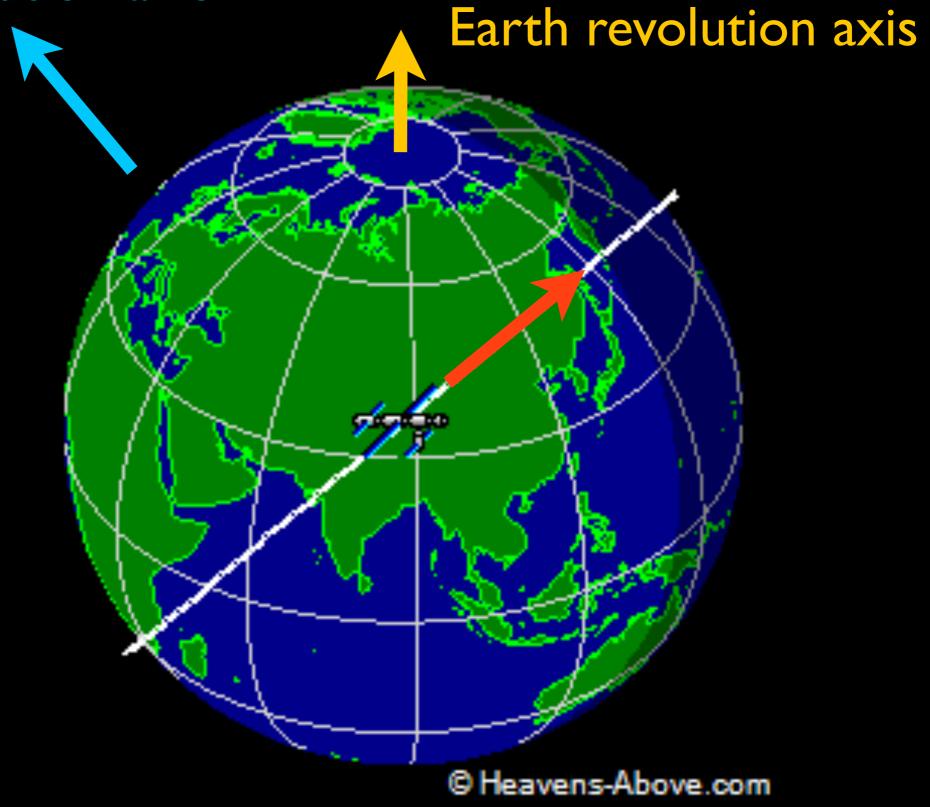
Spherical geometry of Star Trails

ISS Orbit

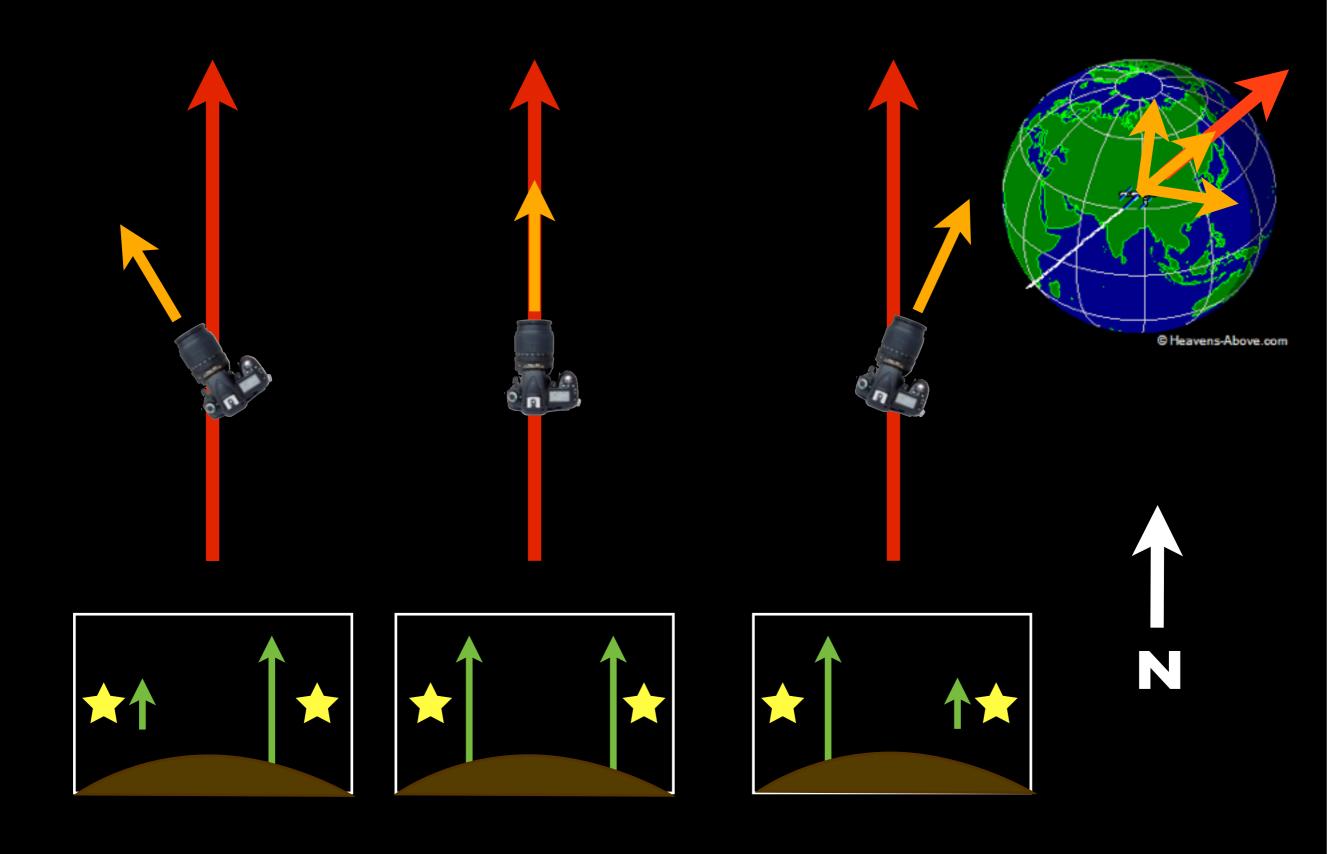


Precession due to rotation of Earth during ~90 min ISS period

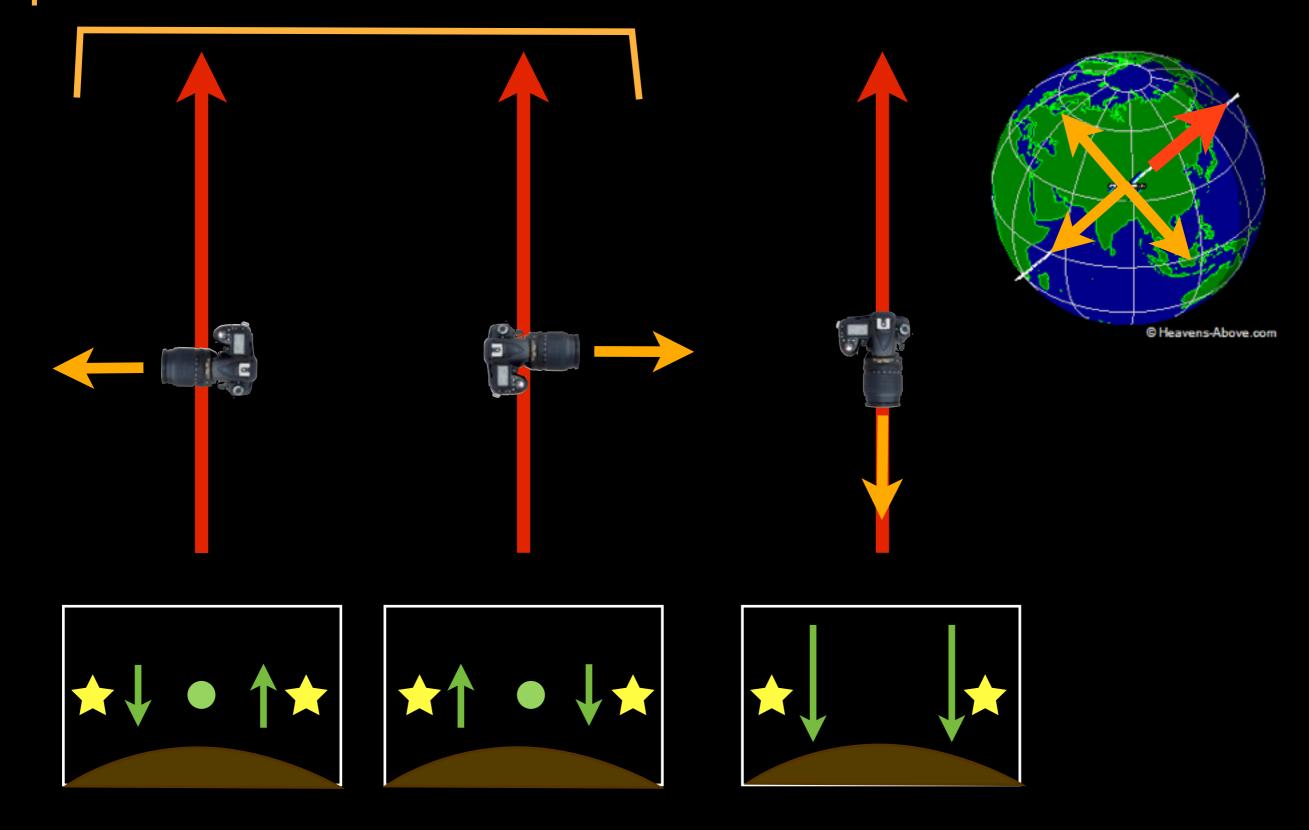
ISS revolution axis



Direction of ISS velocity given by d/dt(nadir)



Optical and ISS revolution axes coincide



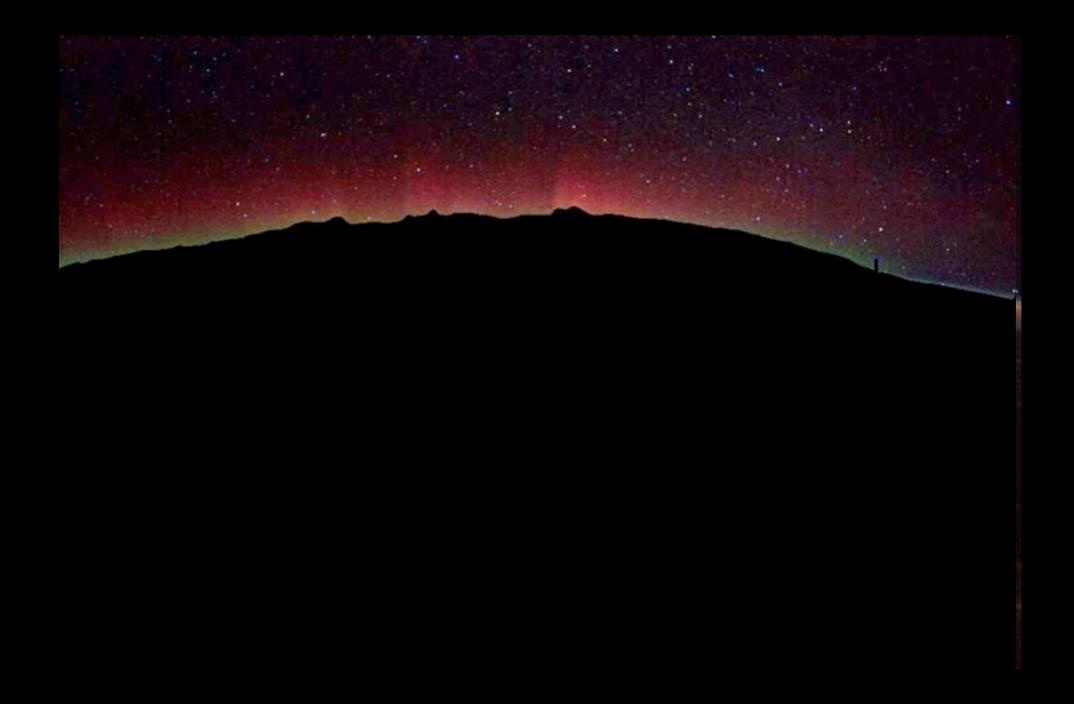
Star Trail Computational Method

Orientation via Star Trails



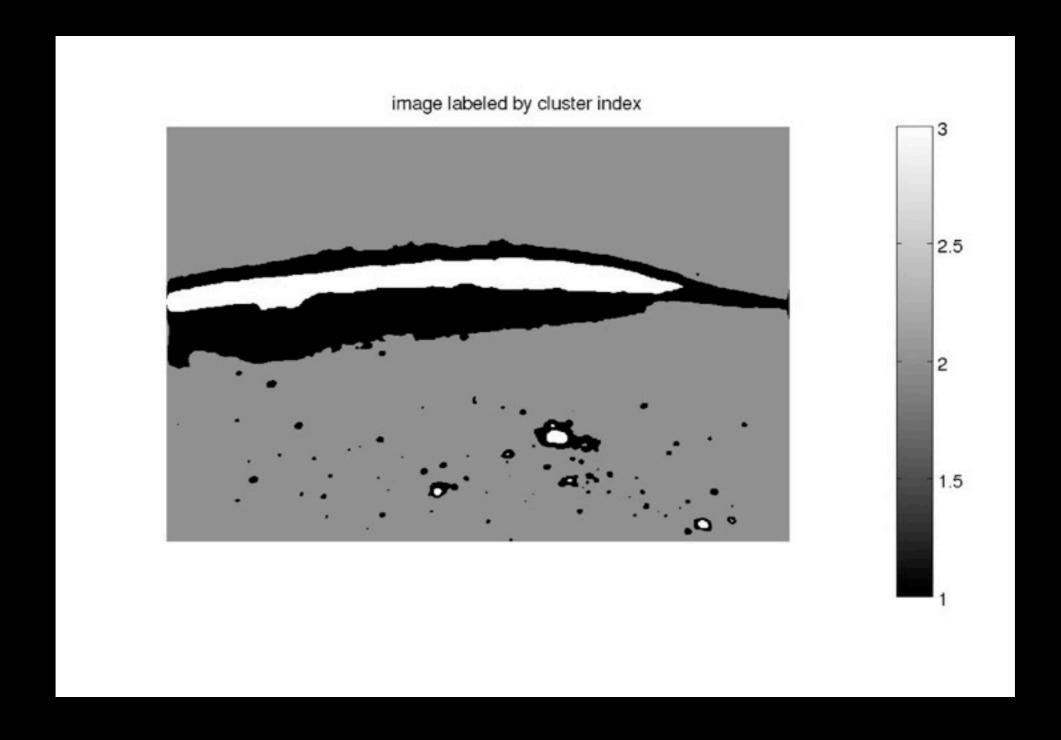
Remove Earth from image via K-means image segmentation

Orientation via Star Trails



Remove Earth from image via K-means image segmentation

Orientation via Star Trails



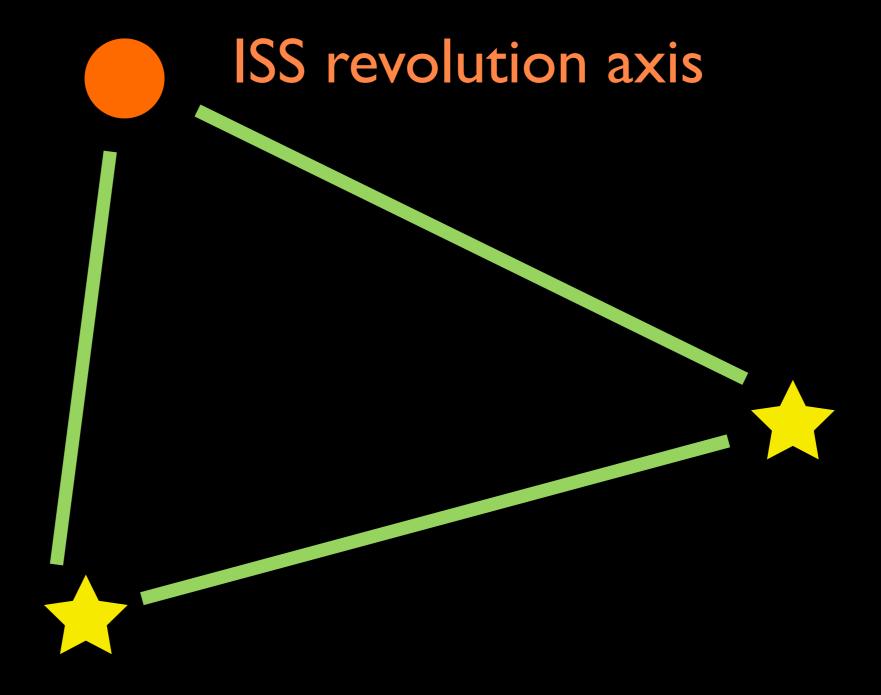
Remove Earth from image via K-means image segmentation

Star Trails

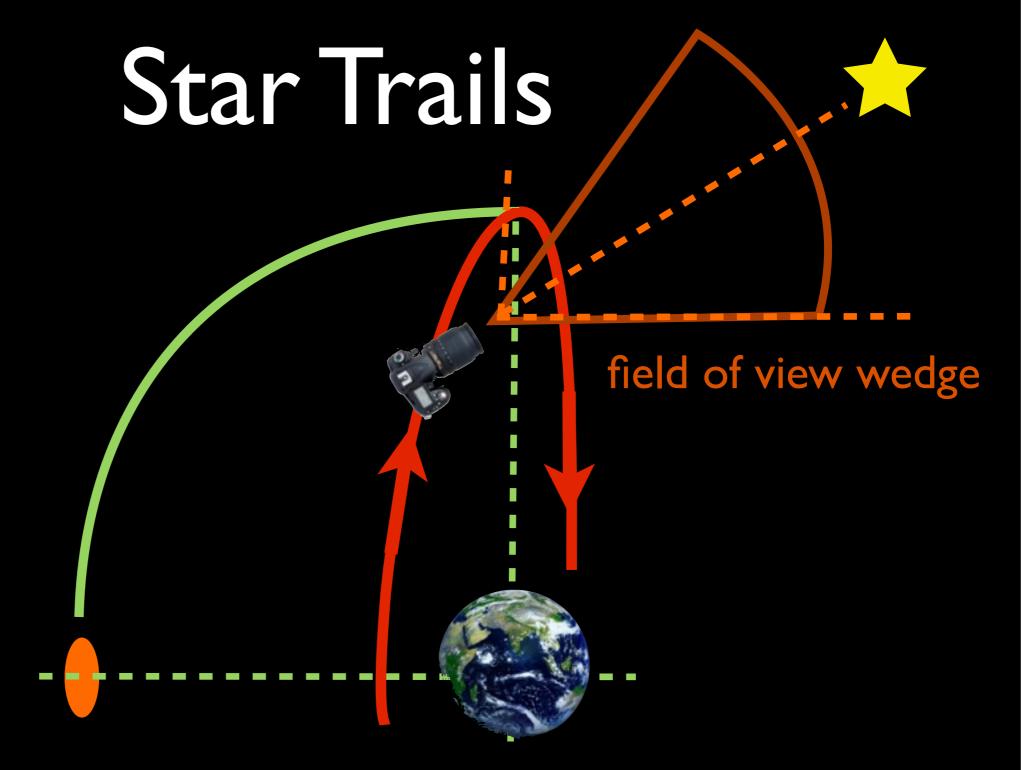


Angular distance to pole is a function of pixel velocity

Star Trails



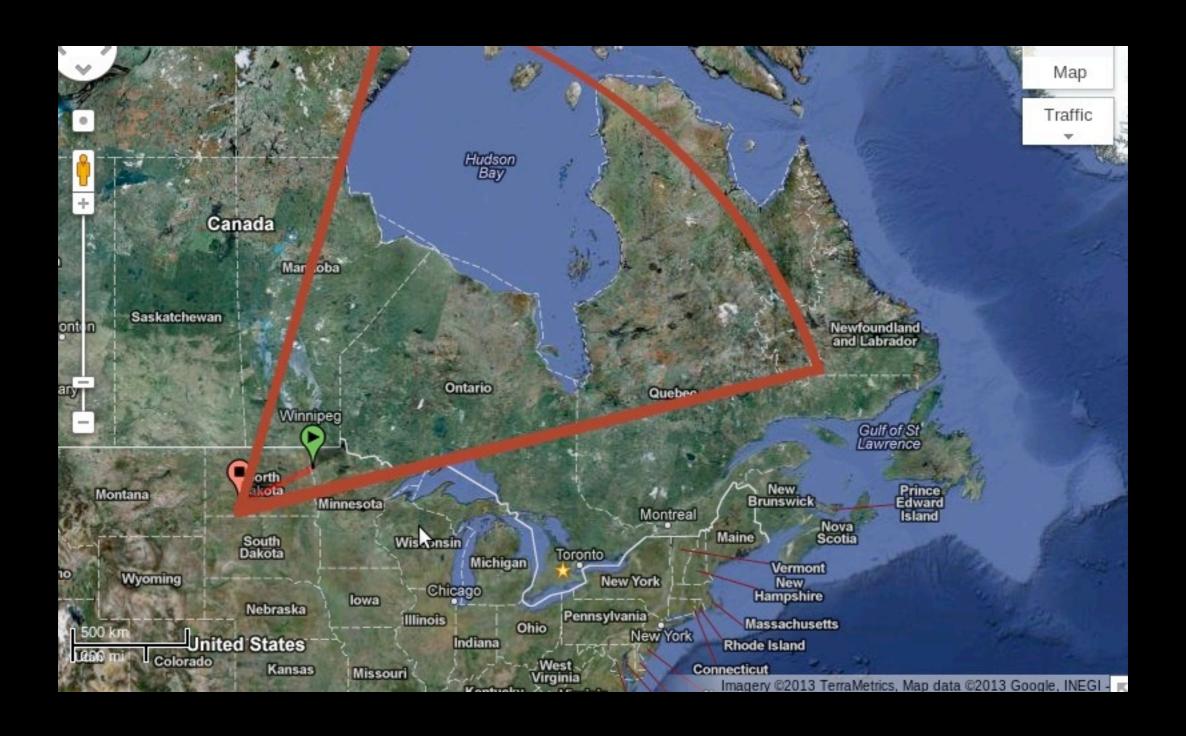
Angular distance to pole is a function of pixel velocity



ISS revolution axis

Angular distance to pole is a function of pixel velocity

Field of View Wedge

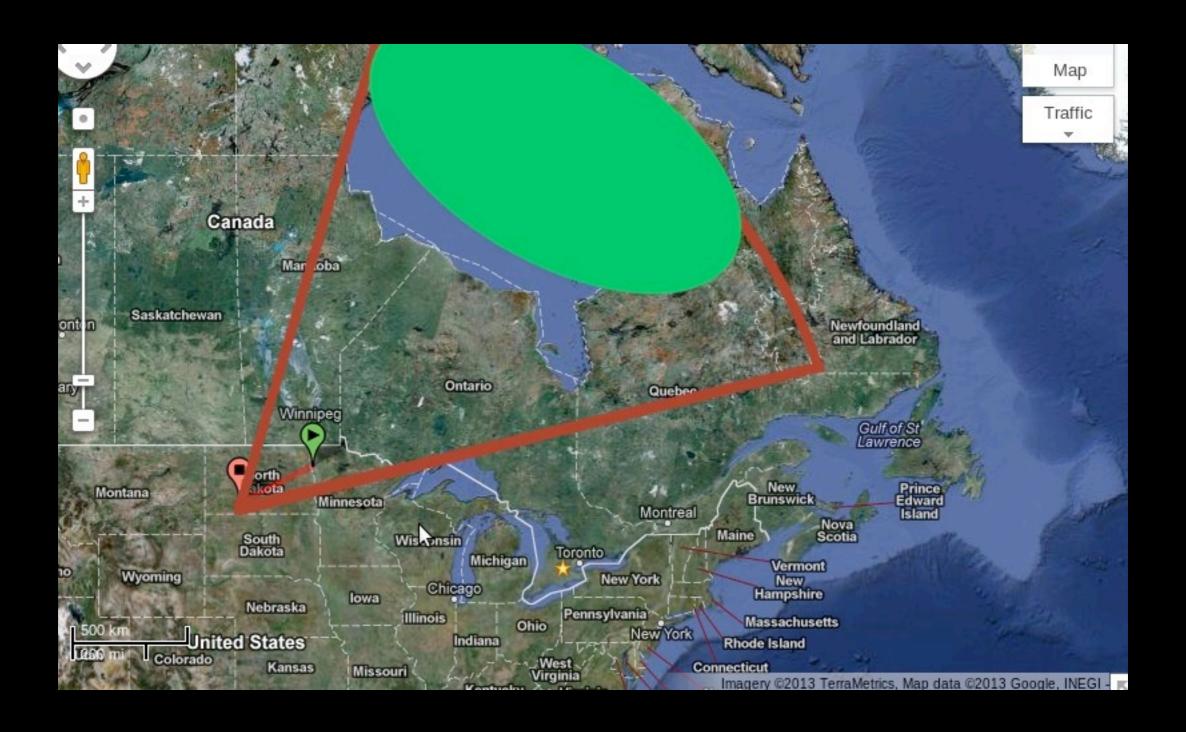


Aurora Extraction



Isolate aurora via K-means image segmentation

Aurora in FoV Wedge



Star Trail Method Summary

Direction of ISS velocity given by d/dt(nadir)

Rotation of camera given by lengths of star trails

Pitch of camera given by altitude and direct image measurement

Efficient camera orientation method; can extrapolate the location of the aurora