

GalSim Library Quick Reference

1. Overview

BARNEY TODO: Tidy this whole thing up, make it look a lot less ugly, maybe use an entirely different document class.

The GalSim Library provides a number of Python classes and methods for simulating astronomical images. The fundamental work flow will normally be something like:

- Construct a representation of your desired astronomical object as a single GalSim `GSOBJect` instance or in combination using the special `Add` and `Convolve` compound-type `GSOBJects` — see Section 2.
- *Optional:* Apply transformations such as shear or magnification using the methods of the resulting `GSOBJect` instance — see Section 3.
- Draw the object into a GalSim `Image` object representing a postage stamp image of your astronomical object. This can be done using the `draw()` or `drawShoot()` methods carried by all `GSOBJects` for rendering images (`drawShoot` uses photon shooting) — see Section 3.
- *Optional:* Add noise to the `Image` using one of the GalSim random deviate classes — see Section 4.
- *Optional:* Add the postage stamp `Image` to a subsection of a larger `Image` instance, or to a larger structure containing multiple `Image` instances each derived from `GSOBJects` as described above — see Section 5.
- Save the `Image(s)` to file in FITS (Flexible Image Transport System) format — see Section 5.

There are many examples of this workflow in the directory `GalSim/examples/`, showing most of the GalSim library in action, in the scripts named `demo1.py` – `demo8.py`.

We now provide a brief, reference description of the GalSim classes and methods which can be used in this workflow. Where possible this has been hyperlinked to the online GalSim documentation generated by *doxygen* where a more detailed description can generally be found.

2. The GSOBJects

There are currently 12 types of GSOBJect. The first ten listed are ‘simple’ or ‘atomic’ GSOBJects that can be initialized by providing values for their required or optional parameters; the last two are ‘compound’ classes used to represent combinations of GSOBJects. They are summarized in the following hyperlinked list, in the order in which the classes appear in `GalSim/galsim/base.py`:

- Gaussian — *a 2D Gaussian light profile.*
- Moffat — *a Moffat profile, used to approximate PSFs.*
- AtmosphericPSF — *currently an image-based implementation of a Kolmogorov PSF (see below), but expected to evolve to use an image of a stochastically modelled atmospheric PSF in the near future.*
- Airy — *an Airy PSF for ideal diffraction through a circular aperture, supports central obscuration.*
- Kolmogorov — *the Kolmogorov PSF for long-exposure images through a turbulent atmosphere.*
- OpticalPSF — *a simple model for non-ideal (aberrated) propagation through circular or square apertures with obscuration.*
- Pixel — *used for integrating light onto square or rectangular pixels.*
- Sérsic — *the Sérsic family of galaxy light profiles.*
- Exponential — *the Exponential disc, a Sérsic with index $n = 1$.*
- DeVaucouleurs — *commonly used to model galaxy bulge profiles, a Sérsic with index $n = 4$.*
- RealGalaxy — *models galaxies using real data, including a correction for the original PSF. Requires the download of external data for full functionality.*
- Add — *a compound object used for summing multiple GSOBJects.*
- Convolve — *a compound object used for convolving multiple GSOBJects.*

For more information and initialization details for each GSOBJect, the Python docstring for each class is available by typing

```
>>> print galsim.<GSOBJect_name>.__doc__
```

within the Python interpreter. Alternatively follow the hyperlinks on the class names above to view the *doxygen* documentation based on the Python docstrings.

3. Important GSOBJECT methods

A number of methods are shared by all the GSOBJECTs of Section 2, and are also to be found in `GalSim/galsim/base.py` within the definition of the GSOBJECT base class. In what follows, we assume that a GSOBJECT labelled `obj` has been instantiated using one of the calls described in the documentation linked above. For example,

```
>>> obj = galsim.Sersic(n=3.5, half_light_radius=1.743).
```

Some of the most important and commonly-used methods for such an instance are:

- `obj.copy()` — *return a copy of the GSOBJECT.*
- `obj.getFlux()` — *get the flux of the GSOBJECT.*
- `obj.scaleFlux(flux_ratio)` — *multiply the flux of the GSOBJECT by flux_ratio.*
- `obj.setFlux(flux)` — *set the flux of the GSOBJECT to flux.*
- `obj.applyDilation(scale)` — *apply a dilation of the linear size of the GSOBJECT by a factor scale.*
- `obj.applyMagnification(scale)` — *dilate linear size by scale and GSOBJECT flux by scale², conserving surface brightness.*
- `obj.applyShear(*args, **kwargs)` — *apply a shear to the GSOBJECT, handling a number of different input conventions.*
- `obj.applyRotation(theta)` — *apply a rotation of theta (positive direction anti-clockwise) to the GSOBJECT, where theta is a galsim.Angle instance (see Section 6).*
- `obj.applyShift(dx, dy)` — *apply a (dx, dy) shift to this object.*
- `obj.draw(...)` — *draw an image of the GSOBJECT using Discrete Fourier Transforms and interpolation to perform the image rendering.*
- `obj.drawShoot(...)` — *draw an image of the GSOBJECT by shooting a finite number of photons to perform the image rendering. The resulting image therefore contains stochastic noise, but the rendering is otherwise very close to exact.*

Once again, for more information regarding each GSOBJECT method, the Python docstring is available

```
>>> print obj.<method_name>.__doc__
```

within the Python interpreter. Alternatively follow the hyperlinks on the class names above to view the *doxygen* documentation based on the Python docstrings. You will see that many of the GSOBJECT instances also have their own specialized methods, often for retrieving parameter values. Examples are `obj.getSigma()` for the Gaussian or `obj.getHalfLightRadius()` for many of the GSOBJECTs.

4. Random deviate classes and methods

A short summary of the 8 random deviates currently implemented in GalSim, with a short description of their distributions:

- `UniformDeviate` — *uniform distribution in the interval $[0, 1)$.*
- `GaussianDeviate` — *Gaussian distribution with mean and standard deviation σ .*
- `BinomialDeviate` — *Binomial distribution for N trials each of probability p .*
- `PoissonDeviate` — *Poisson distribution with a single mean rate.*
- `CCDNoise` — *Distribution following a basic CCD noise model, depending on gain and read_noise .*
- `WeibullDeviate` — *Weibull distribution family (includes Rayleigh and Exponential) with shape parameters a and b .*
- `GammaDeviate` — *Gamma distribution for parameters α and β .*
- `Chi2Deviate` — *χ^2 distribution for degrees of freedom parameter n .*

Unfortunately the random deviate classes are not yet fully integrated within the *doxygen* documentation, due to their being C++ with compiled Python wrappers. This means that the class names above are not hyperlinked.

However, the full docstrings are available in `galsim/random.py`, so please refer there for more information, or type

```
>>> print galsim.<RandomDeviate_name>.__doc__
```

within the Python interpreter.

We now illustrate the most commonly-used methods of the random deviates. If we assume that some random deviate instance has been instantiated as `dev`, for example

```
>>> dev = galsim.GaussianDeviate(sigma=3.9, mean=50.),
```

Some of the most important and commonly-used methods for such an instance are:

- `dev.applyTo(image)` — *adds a deviate distributed according to the distribution represented by `dev` to each element in in a supplied Image instance `image` (see Section 5).*
- `dev()` — *this direct call method returns a new random number drawn from the distribution represented by `dev`.*

5. Image classes and methods

- `ImageS`
- `ImageI`
- `ImageF`
- `ImageD`

- `img.read()`
- `img.addNoise()`
- `img.write()`

6. Miscellaneous classes and methods

- `Angle`
- `Ellipse`