GalSim Library Quick Reference

1. Overview

BARNEY TODO: Tidy this whole thing up, make it look a lot less ugly, maybe use an entirely different document class.

The GalSim Library provides a number of Python classes and methods for simulating astronomical images. The fundamental work flow will normally be something like:

- Construct a representation of your desired astronomical object as a single GalSim GSObject instance or in combination using the special Add and Convolve compound-type GSObjects see Section 2.
- *Optional*: Apply transformations such as shear or magnification using the methods of the resulting GSObject instance see Section 3.
- Draw the object into a GalSim Image object representing a postage stamp image of your astronomical object. This can be done using the draw() or drawShoot() methods carried by all GSObjects for rendering images (drawShoot uses photon shooting) see Section 3.
- Optional: Add noise to the Image using one of the GalSim random deviate classes see Section 4.
- Optional: Add the postage stamp Image to a subsection of a larger Image instance, or to a larger structure containing multiple Image instances each derived from GSObjects as described above see Section 5.
- Save the Image(s) to file in FITS (Flexible Image Transport System) format see Section 5.

There are many examples of this workflow in the directory GalSim/examples/, in the example scripts BasicDemo.py, MultiObjectDemo.py and RealDemo.py. We now provide a brief, reference description of the GalSim classes and methods which can be used in this workflow.

2. The GSObjects

There are currently 12 types of GSObject. The first ten listed are 'simple' or 'atomic' GSObjects that can be initialized by providing values for their required or optional parameters; the last two are 'compound' classes used to represent combinations of GSObjects. They are summarized in the following hyperlinked list, in the order in which the classes appear in GalSim/galsim/base.py:

- Gaussian a 2D Gaussian light profile.
- Moffat a Moffat profile, used to approximate PSFs.
- AtmosphericPSF currently an image-based implementation of a Kolmogorov PSF (see below),
 but expected to evolve to use an image of a stochastically modelled atmospheric PSF in the near future.
- Airy an Airy PSF for ideal diffraction through a circular aperture, supports central obscuration.
- Kolmogorov the Kolmogorov PSF for long-exposure images through a turbulent atmosphere.
- OpticalPSF a simple model for non-ideal (aberrated) propagation through circular or square apertures with obscuration.
- Pixel used for integrating light onto square or rectangular pixels.
- Sersic the Sérsic family of galaxy light profiles.
- \circ Exponential the Exponential disc, a Sérsic with index n=1.
- o DeVaucouleurs commonly used to model galaxy bulge profiles, a Sérsic with index n=4.
- RealGalaxy models galaxies using real data, including correction for the original PSF.
- Add a compound object used for summing multiple GSObjects.
- o Convolve a compound object used for convolving multiple GSObjects.

For more info and initialization details for each GSObject, type

```
>>> print galsim.<GSObject_name>.__doc__
```

within the Python interpreter, or follow the hyperlinks on the class names above to view the *doxygen* documentation based on the Python docstrings.

3. Important GSObject methods

A number of methods are shared by all the GSObjects of Section 2, and are also to be found in GalSim/galsim/base.py within the definition of the GSObject base class. In what follows, we assume that a GSObject labelled obj has been instantiated using one of the calls described in the documentation linked above. For example,

```
>>> obj = galsim.Sersic(n=3.5, half_light_radius=1.743).
```

Some of the most important and commonly-used methods for such an instance are:

- o obj.copy() returns a copy of the GSObject.
- obj.getFlux() gets the flux of the GSObject.
- o obj.scaleFlux(flux_ratio) multiplies the flux of the GSObject by flux_ratio.
- o obj.setFlux(flux) sets the flux of the GSObject to flux.
- obj.applyDilation(scale) applies a dilation of the linear size of the GSObject by a factor scale.
- \circ obj.applyMagnification(scale) dilate linear size by scale and GSObject flux by $scale^2$, conserving surface brightness.
- applyShear(*args, **kwargs) apply a shear to the GSObject, handling a number of different input conventions.
- o obj.applyRotation(theta) apply a rotation of theta (positive direction anti-clockwise) to the GSObject, where theta is a galsim. Angle instance (see Section 6).
- o applyShift (dx, dy) apply a (dx, dy) shift to this object.
- o obj.draw(...) draws an image of the GSObject using Discrete Fourier Transforms and interpolation to perform the image rendering.
- obj.drawShoot(...) draws an image of the GSObject by shooting a finite number of
 photons to perform the image rendering. The resulting image therefore contains stochastic noise, but
 the rendering is otherwise very close to exact.

4. Random deviate classes and methods

5. Image classes and methods

6. Miscellaneous classes and methods