

CONDITIONS FOR PMC FORMATION IN 2002-2008 ESTIMATED FROM TIMED/SABER MEASUREMENTS

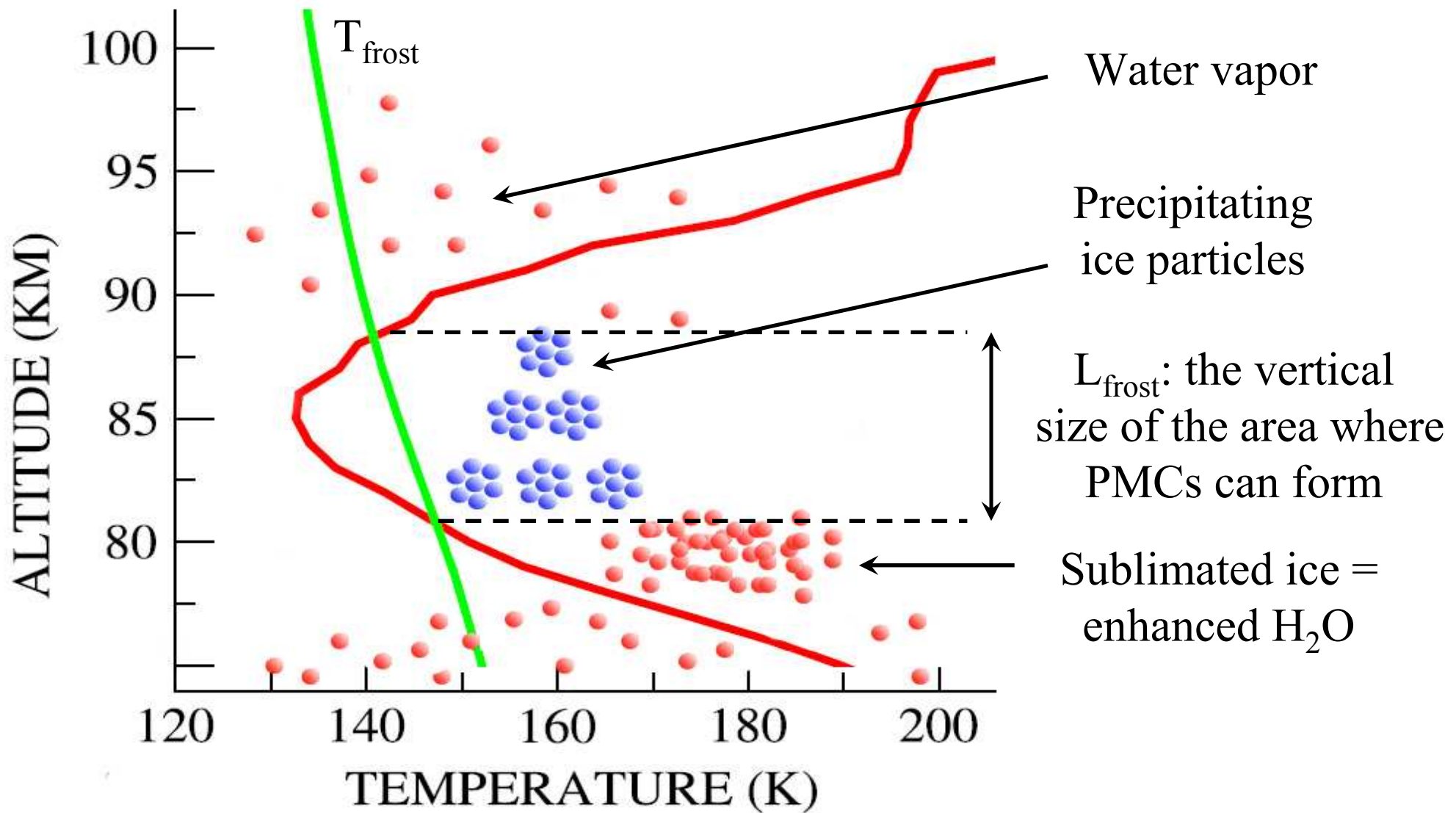
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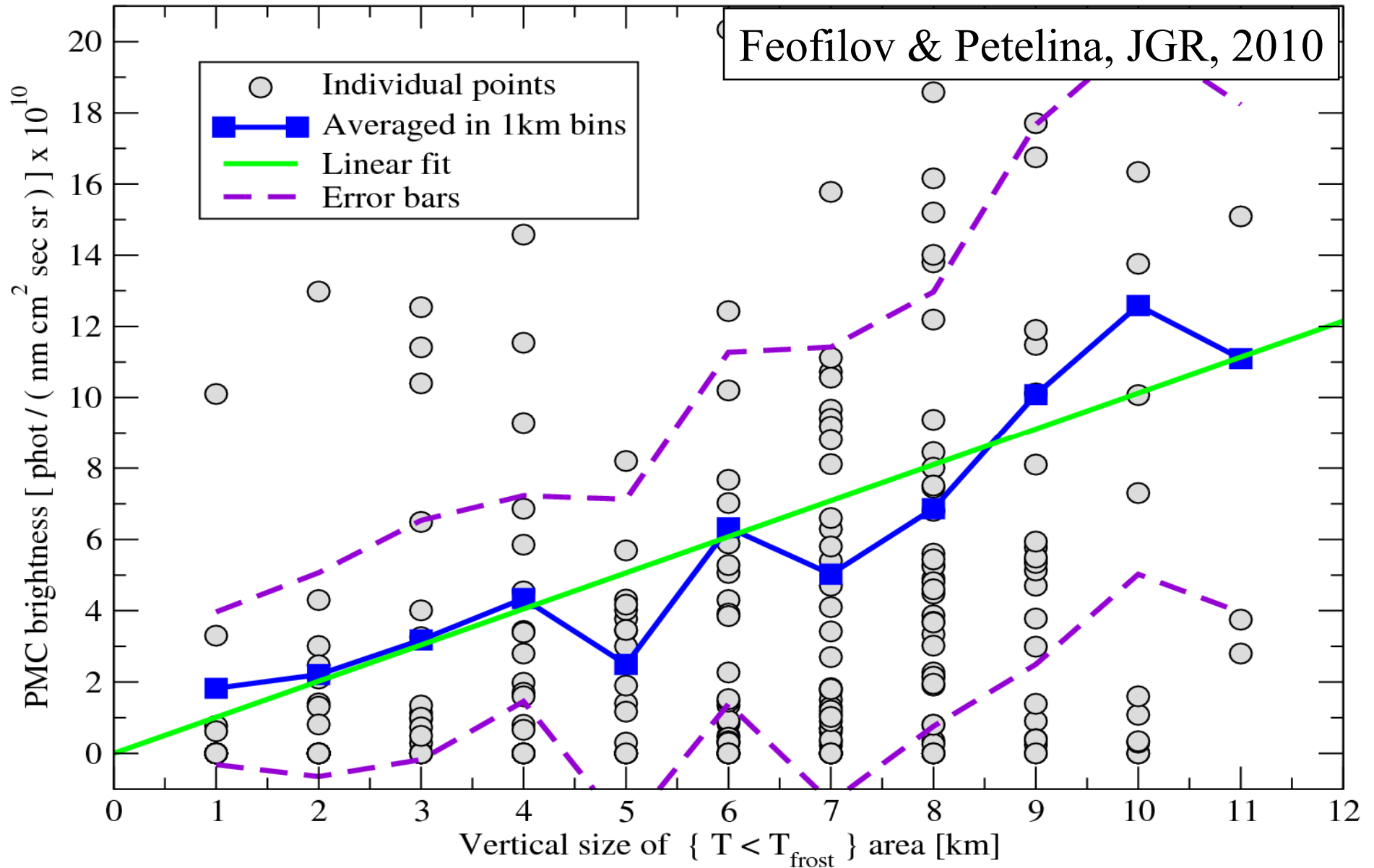
A simplified model of PMC formation

1

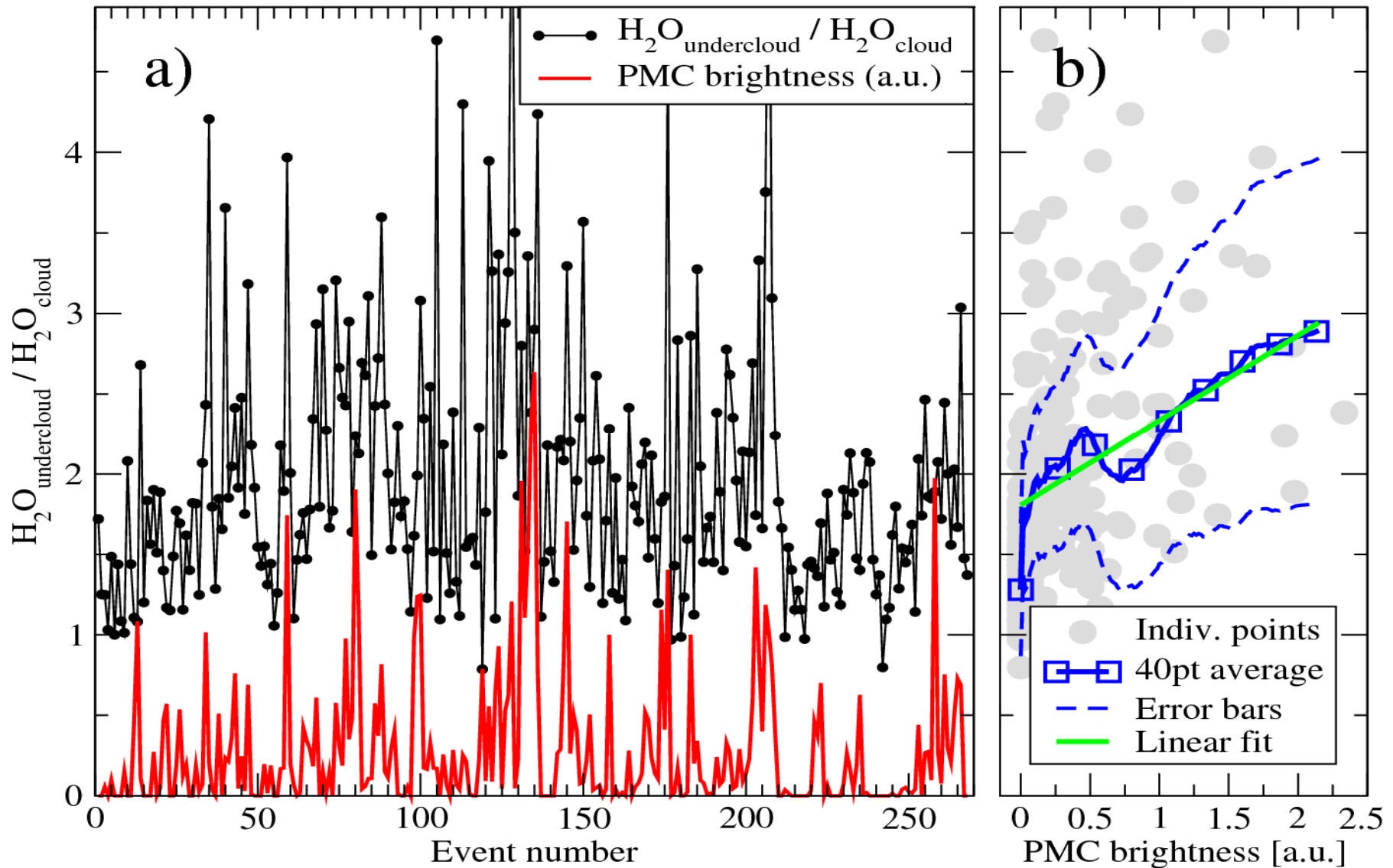


OSIRIS PMC brightness vs integrated L_{frost}

2



*PMC brightness correlated with H_2O VMR
in and below the cloud (Feofilov & Petelina, 2010)*



The SABER Instrument Aboard the TIMED Satellite

4

TIMED: Thermosphere, Ionosphere, Mesosphere
Energetics & Dynamics

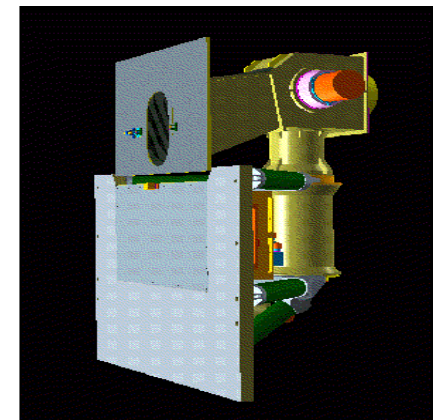
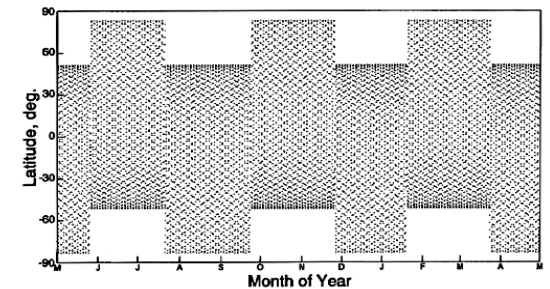
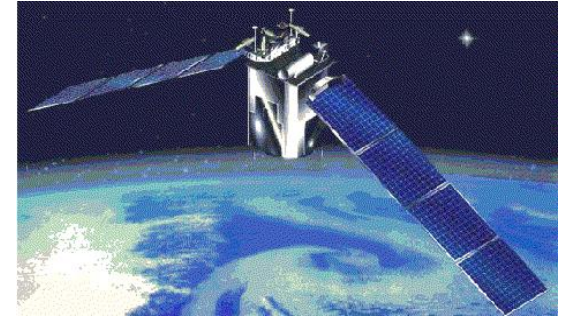
74.1° inclined 625 km orbit;

Latitudinal coverage: 83°S–52°N / 53°S–82°N

Data available since 25 January 2002

SABER: Sounding of the Atmosphere Using
Broadband Emission Radiometry

- Limb scanning infrared radiometer
(~10–100 km, ~2 km footprint)
- 10 broadband channels (1.27–17 μm)
- Products: **kinetic temperature**, **pressure**,
CO₂, O₃, **H₂O**, NO, O₂, OH, O, H



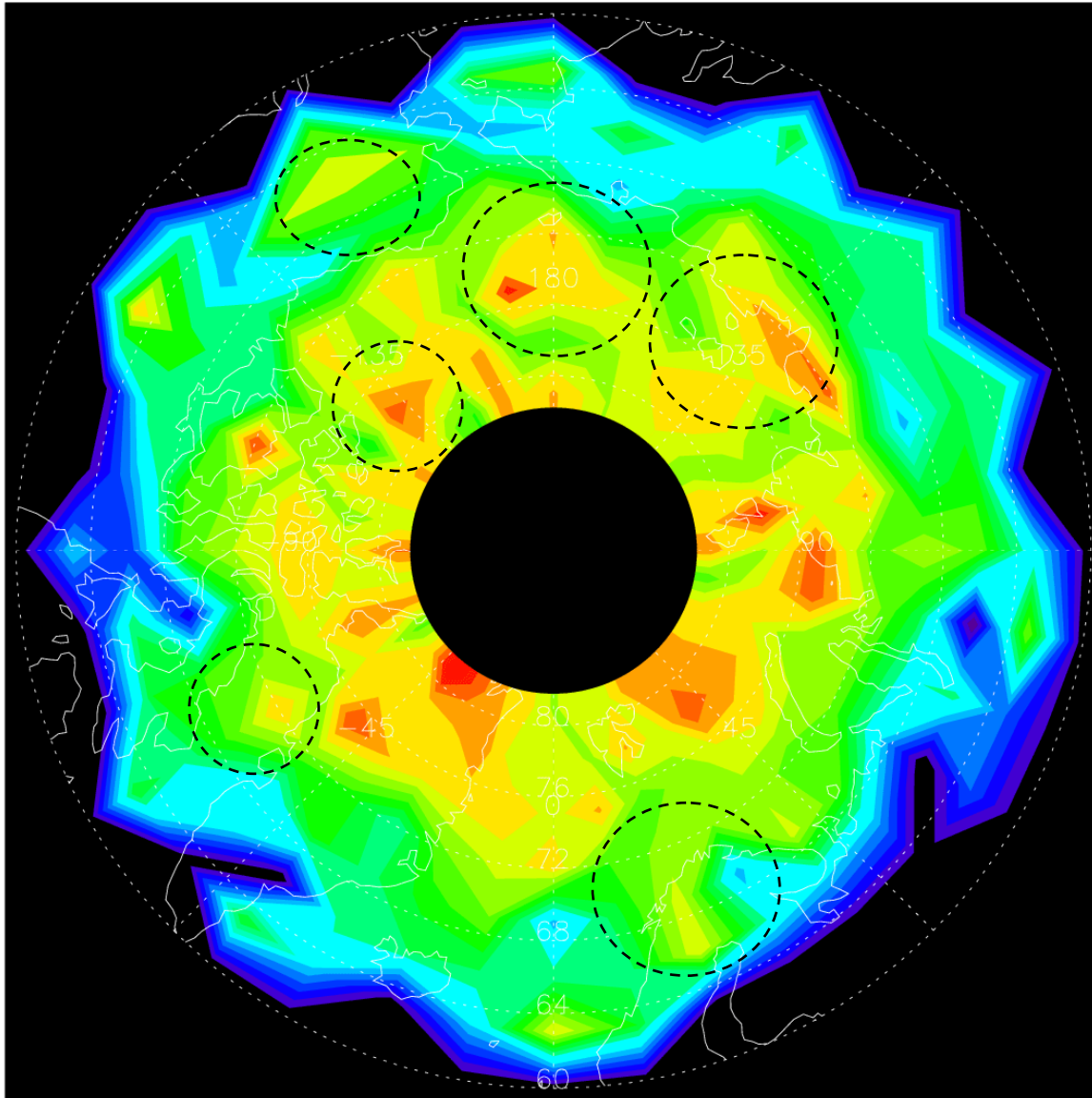
General idea and approach

- A simplified model of PMC particles lifecycle is shown on Fig. 1. It does not take into account meridional transport effects that are important in some cases. However, as Fig. 2 and 3 show (see also [Feofilov and Petelina, JGR, 2010]), there is an obvious correlation between the *local* atmospheric parameters and PMC brightness.
- The latter appears to be proportional to L_{frost} , the vertical size of the area where $T < T_{\text{frost}}$, and to the ratio of H_2O VMR below and in the cloud that are related to the ice sublimation and freeze drying effects, respectively.
- In this work we continued studying the polar summer mesosphere using SABER/TIMED instrument (Fig. 4) and plotted the maps of L_{frost} and $(\text{H}_2\text{O}_{\text{undercloud}} / \text{H}_2\text{O}_{\text{cloud}})$ ratio for polar summer seasons of 2007-2008 (see the examples on Fig. 7-9).

- Fig. 7 shows the L_{frost} distribution that gives the estimate for the PMC brightness as follows from Fig. 2.
- Fig. 8 demonstrates the ratio of $\text{H}_2\text{O}_{\text{undercloud}}/\text{H}_2\text{O}_{\text{cloud}}$ for the same conditions. Both figures demonstrate the similar latitudinal behavior: the values shown tend to increase to the pole. The fine structures seen on these plots are similar. This proves the concept used in [Feofilov and Petelina, 2010].
- Fig. 9 shows the PMC albedo measured by CIPS instrument onboard AIM satellite for the same days. The meridional behavior is consistent with that seen in Fig. 7 and 8 while the fine structure does not always repeat that of those figures. This is linked with high temporal variability of the mesospheric area within the day.

Vertical size of $T < T_{\text{frost}}$ area, L_{frost} , 2007, 183-196

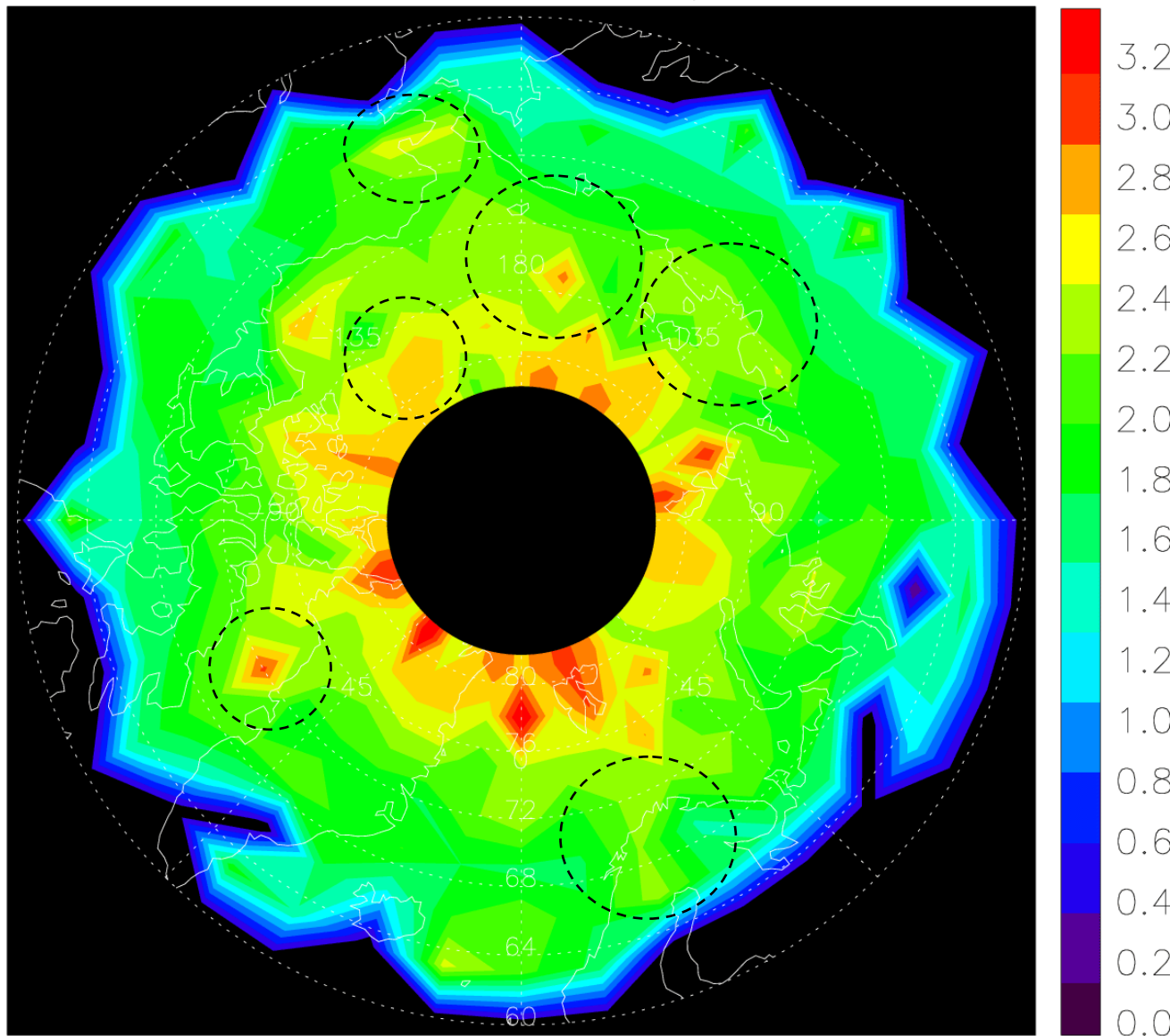
7



- T_{frost} estimated in accordance with Murphy and Koop, 2005.
- SABER V1.07 data (P,T)
- H_2O profiles retrieved in accordance with Feofilov et al, 2009.
- $3^\circ \text{ lat} \times 10^\circ \text{ lon}$ bins
- Two-week averages of L_{frost} values is shown.
- Larger L_{frost} values “predict” brighter clouds.
- Latitudinal behavior agrees with that of CIPS PMC brightness (see Fig. 9).

$H_2O_{undercloud} / H_2O_{cloud}$ ratio, 2007, 183-196

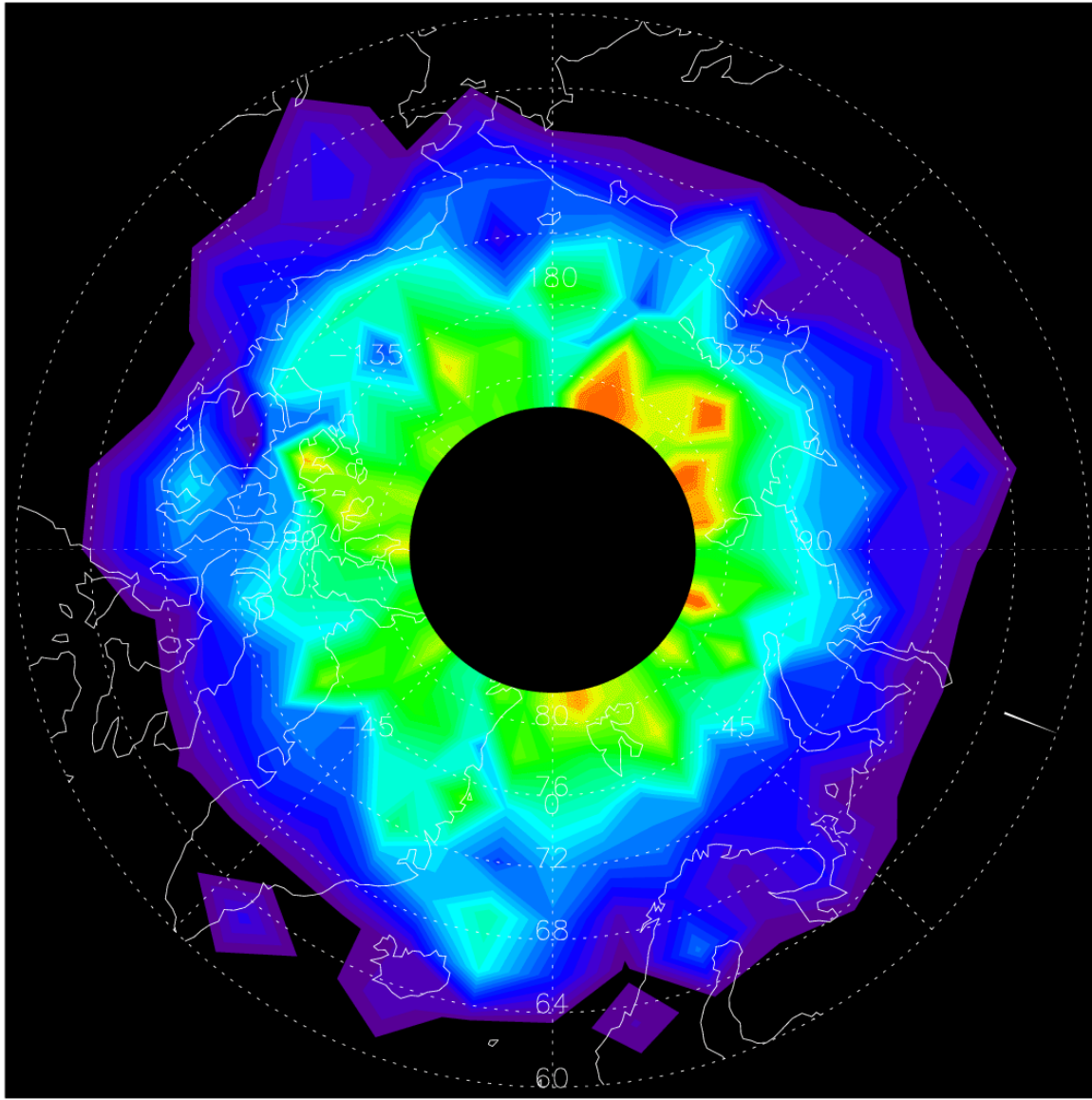
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- Parameters are the same as on Fig. 7.
- Ratio of H_2O VMR below and in the cloud is shown.
- Only the cases with $T < T_{frost}$ = “allowed” PMCs are considered.
- Typical ratio for the non-PMC cases is 2.
- Note the similarities in the structures seen on this figure and Fig. 7 marked by dashed circles.

PMC albedo measured by CIPS, 2007, 183-196

9



- CIPS 03.20 level 4 data.
- 3° lat \times 10° lon bins.
- Daily overlaps with SABER selected.
- Two-week average of daily “snapshots” presented.
- Latitudinal behavior agrees with that seen on Fig. 7 and 8 though the structures do not always match.