

# MIDAS

## (Meteor Identification and Detection Automation System)

CE901 - Msc Dissertation Project

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# 1

# Introduction

# Introduction to MIDAS

## What is MIDAS?

- MIDAS (Meteor Identification and Detection Automation System): A project to automate meteor detection in spectrogram images from BRAMS (Belgian RAdio Meteor Stations).

## Why MIDAS?

- Manual analysis of large spectrogram datasets is inefficient and time-consuming.
- Machine learning enables accurate and efficient detection of meteors.

## Models Evaluated in MIDAS:

- YOLO11m: High-performing object detection model.
- MeteorNet: Custom neural network model.
- Hybrid Model: MobileNet\_v3 + Random Forest classifier.
- Fusion Model: Combines YOLO11m and MeteorNet for enhanced accuracy.

## Key Deliverable:

- A web application enabling users to upload spectrogram images and detect meteors automatically.

## Impact:

- Provides valuable tools for meteor observation, advancing research in atmospheric science and machine learning.

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## Related works

# Background

## What is a Meteor?

- A meteor, or "shooting star," is the streak of light from a meteoroid entering Earth's atmosphere.
- The light is caused by intense heat from friction between the meteoroid and atmospheric gases.

## What Meteors Leave in Space?

- Meteoroids vaporize in the atmosphere, leaving behind ionized particle trails.
- These trails can reflect radio waves detectable by ground-based stations.

## Radio Signals and Meteors:

- Ionized trails reflect radio waves, enabling meteor detection via radio signals.
- Reflected radio waves produce characteristic signals for analysis.

## Importance of Meteor Detection

- Helps understand meteoroid behaviour, composition, and movement in Earth's atmosphere.
- Provides insights into atmospheric science and space debris interactions.

## Role of Spectrogram Images

- Represents frequency-time data of meteor echoes.
- Captures ionized trails created by meteors that reflect radio waves.
- Enables detailed analysis of meteors' size, shape, and movement using Doppler shifts.

# Previous works in meteor detection

## Optical Techniques:

- Long-exposure photography to study meteor showers (Hawkes & Jones, 1975).
- Provided insights into brightness, frequency, and spatial distribution.

## Radio-Based Methods:

- Forward-scattering techniques reflect radio waves off ionized meteor trails (McKinley, 1961).
- Enabled 24/7 meteor monitoring in all weather conditions.

## Current Machine Learning Approaches:

- Dilated CNNs outputs image different than the original image.
- Used sliding windows over labeled spectrograms. Developed a CNN model for pixel-level detection of meteor echoes (Stan Draulans, 2019).
- Explored simple CNN, transposed CNN, and dilated CNN for improved detection. Dilated CNN achieved the highest F1-score (0.870) due to effective noise handling (Jean Lobet, (2021)).

# Limitations of previous works

## Optical Techniques:

- Dependent on clear skies and darkness.
- Cannot detect faint meteors or those obscured by clouds.
- Time-consuming and labor-intensive.

## Radio-Based Methods:

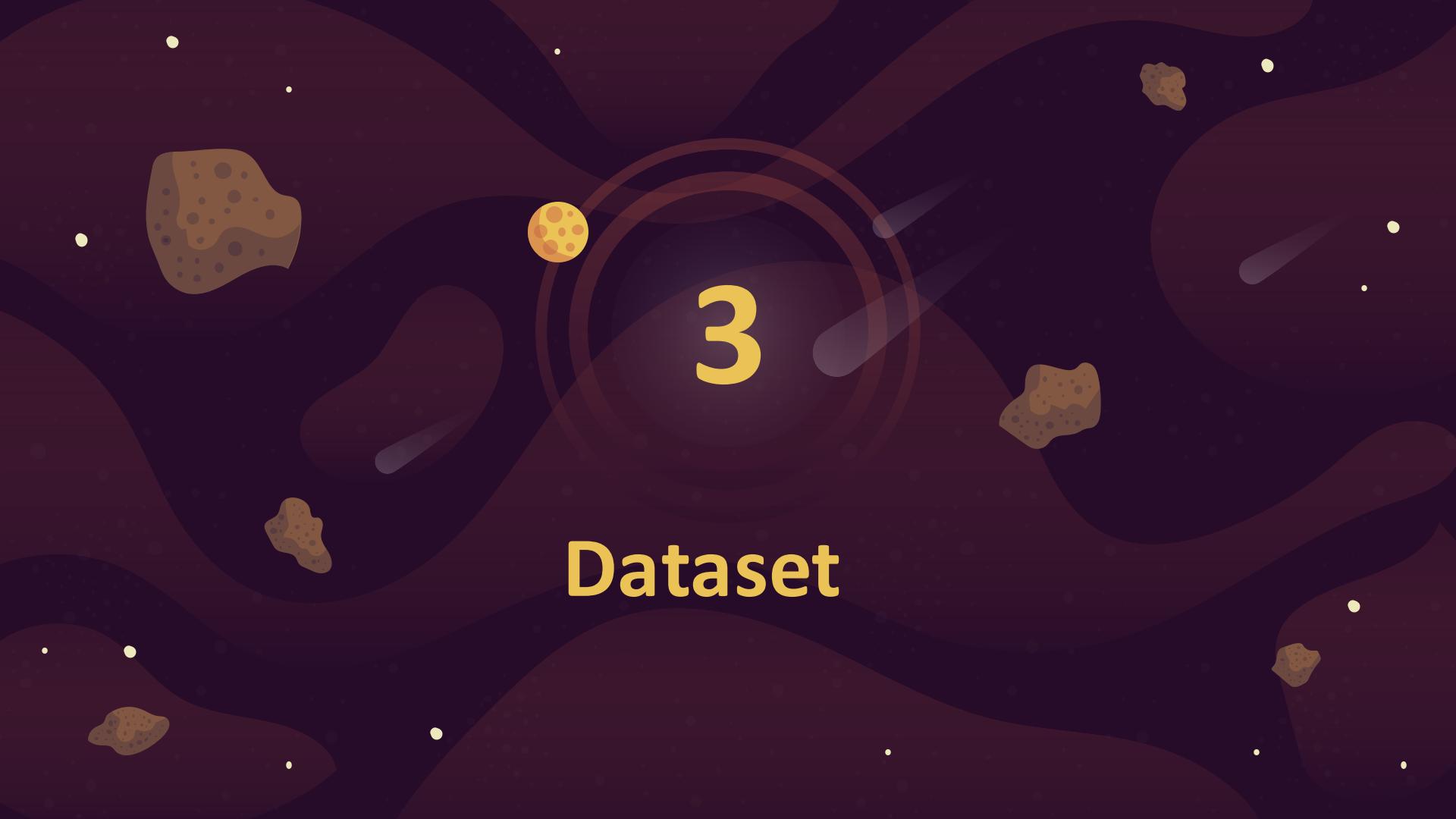
- Generate massive spectrogram data requiring advanced analysis tools.
- Manual interpretation is slow and prone to errors.

## Current Machine Learning Approaches:

- Challenges in detecting irregularly shaped meteor echoes.
- bounding boxes are inconsistent with the actual meteor shapes.
- Over-smoothing misses finer details of meteor echoes.
- Military aircrafts are incorrectly identified as meteor echoes.

# Objectives

- Download spectrogram images from BRAMS downloader.
- Label meteor objects within spectrogram images.
- Train Machine Learning Models(YOLO11, MeteorNet, and Hybrid Model).
- Evaluate performance using McNemar's test.
- Build MIDAS web application using trained models.
- Automate meteor detection from BRAMS spectrogram images.



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# Dataset

# Data preparation

## **Data Collection:**

- Downloaded 8,000 spectrograms using the BRAMS Data Downloader (09/09/2024 to 10/09/2024).
- Selected diverse locations for varied-quality and varied meteor observations.

## **Data Filtering:**

- Selected 1300 images for labeling.
- Diversity: Balanced mix of underdense and overdense echoes, normal and military aircrafts echoes, beacon signal, broadband signal echoes and background only.

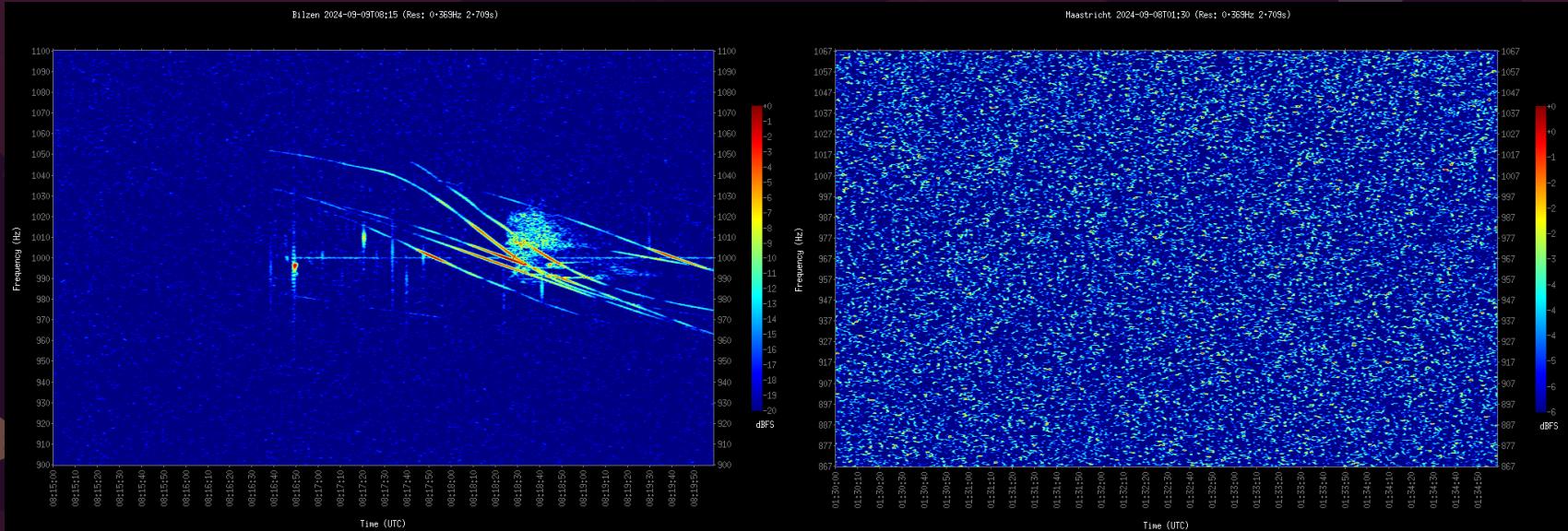
## **Data Annotation:**

- Annotated using LabelMe for bounding boxes around meteor echoes.
- Augmentation (horizontal/vertical flips) increased dataset to 3,400 samples.
- Focused on meteor echoes only, avoiding clutter from irrelevant objects.
- Format: Class ID, normalized center coordinates (x, y), width, height.

## **Data Splitting:**

- Training: 3,000 samples.
- Validation: 200 samples.
- Testing: 200 samples.
- Ensured proper learning, validation, and fair evaluation.

# Spectrogram images with and without objects



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# Methodology

# Model Development

## YOLO11m:

- SOTA object detection model.
- **Configuration:**
  - Pretrained weights: YOLO11m.pt
  - Optimizer: SGD, Learning Rate: 0.001
  - Augmentations: Auto-augment, horizontal/vertical flips
  - Training Epochs: 300 (early stopping: 10 epochs) Dataset: Trained on 3 sizes (3000, 2000, 1000 samples).
  - Input: 960×960 pixels,
  - Batch Size: 16
  - Confidence Threshold: 0.3, IoU: 0.6

## MeteorNet:

- Custom CNN for object detection.
- **Configuration:**
  - Input: 256×256 pixels, Batch Size: 16
  - Optimizer: Adam, Loss: Binary cross-entropy
  - Metrics: Precision, Recall, F1-Score, IoU
  - Data Preparation: Created binary masks for meteor regions from YOLO labels.
  - Training on subsets of 3000, 2000, and 1000 samples.
  - Features: High segmentation precision for meteor echoes.

# Model Development

## Hybrid Model (MobileNet\_V3 + Random Forest) :

- **Purpose:** Lightweight feature extraction with classification.
- **Configuration:**
  - MobileNet\_V3: Feature extractor, Random Forest: Classifier
  - Input: Resized to 256×256 pixels
  - Random Forest Estimators: 50, Class Weight: Balanced.
  - Feature Patches: 16×16 pixels using MobileNet\_V3.
  - Outcome: Lower accuracy than YOLO11m and MeteorNet.

## Fusion Detection Technique :

- **Overview:** Combines YOLO11m and MeteorNet.
- **Goal:** Improve meteor detection accuracy by leveraging strengths of both models.
- **Pipeline:**
  - **YOLO11m:** Detects potential meteor regions using bounding boxes.
  - **MeteorNet:** Segments meteor regions for more precise bounding boxes.
  - **Bounding Box Combiner:** Merges overlapping/adjacent boxes for consolidated detections.
- **Outcome:**
  - Clear visual representation of meteor objects in spectrograms.
  - Enhanced accuracy through complementary model integration.

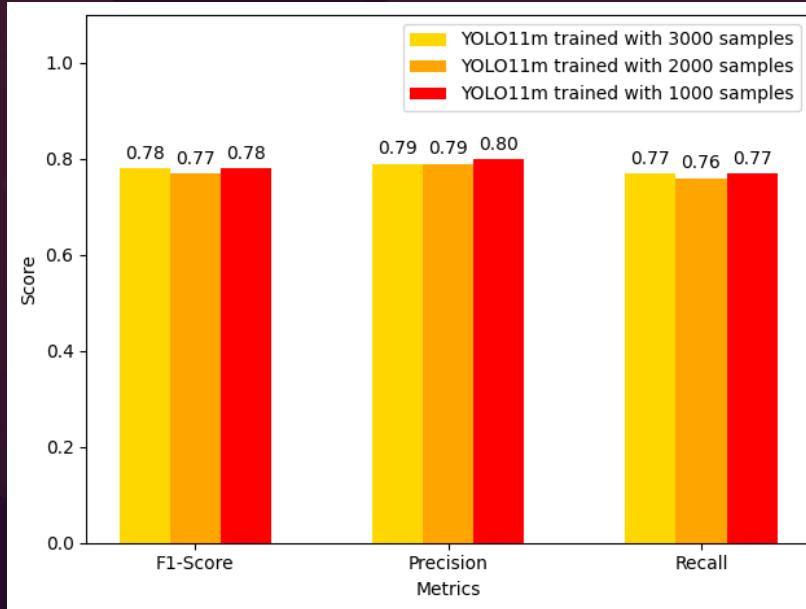


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# Evaluation

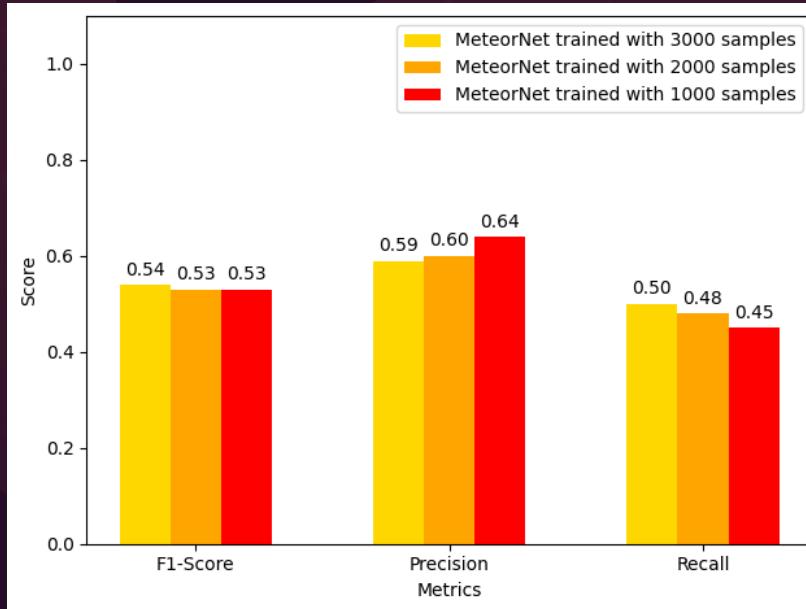
# Comparative Analysis of YOLO11m Performance

- YOLO11m was trained on 3 configurations: 3000, 2000, and 1000 samples.
- Training on 1000 samples achieved the best balance, while 3000 samples excelled in precision.



# Comparative Analysis of MeteorNet Performance

- MeteorNet was trained on 3 configurations: 3000, 2000, and 1000 samples.
- 3000 samples provide the best balance of metrics and is the optimal choice for MeteorNet in MIDAS.



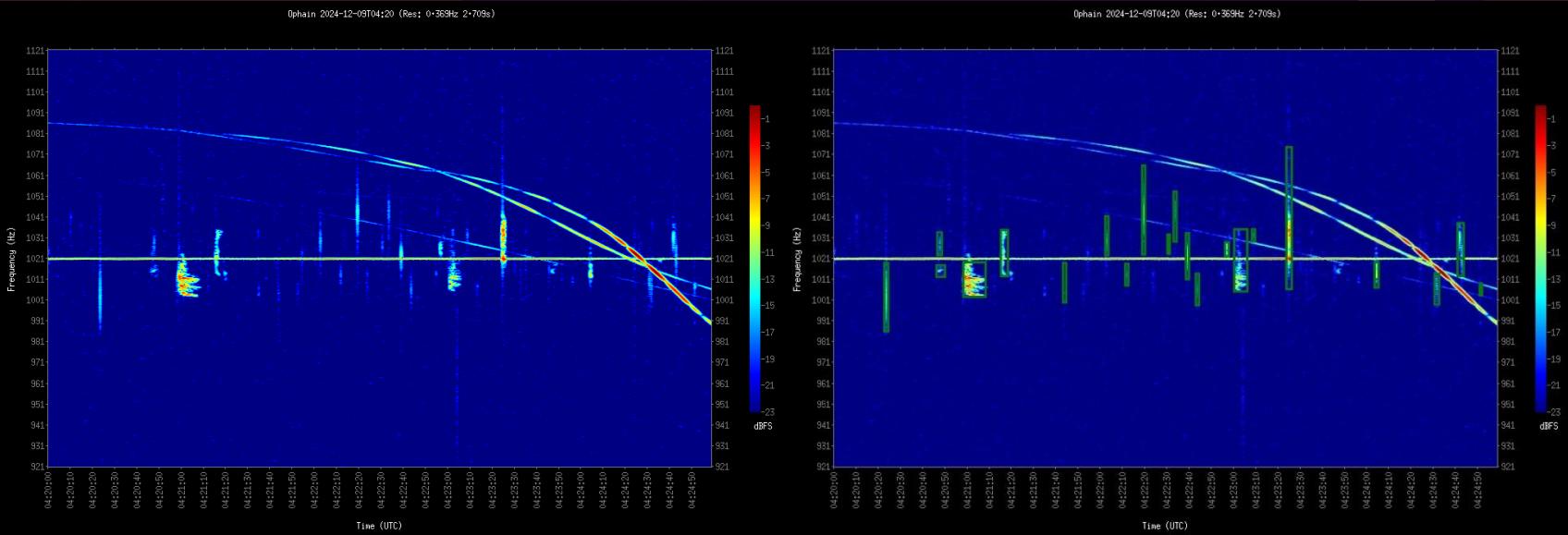
# Analysis of Hybrid Model Performance

- Trained on 3000 samples, not tested on smaller datasets.
- Less precise compared to YOLO11m (well-centered) and MeteorNet (detailed boundaries). Often misplaced or covering extra areas.
- The hybrid model had a weaker balance of true positive detection and false positive avoidance.
- YOLO11m and MeteorNet outperformed the hybrid model in accuracy and reliability.

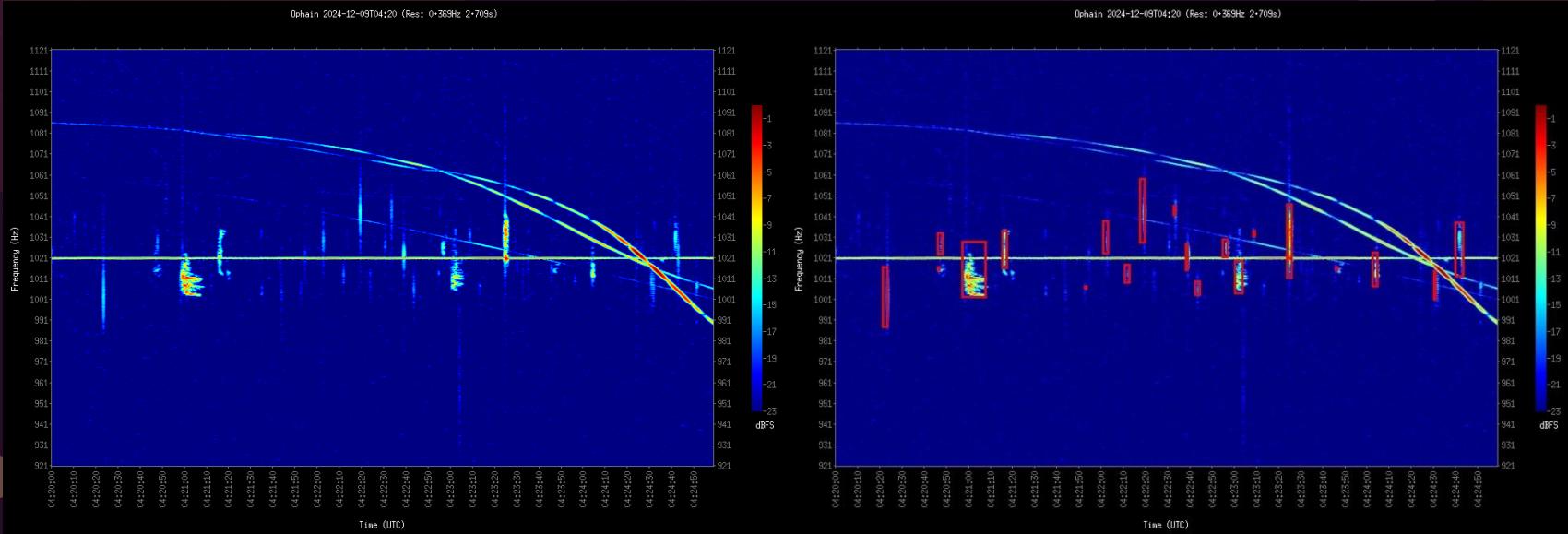
# Analysis of Fusion Technique Performance

- Combines YOLO11m and MeteorNet outputs.
- Improved meteor detection accuracy by leveraging strengths of both models.
- Detected meteors missed by YOLO11m or MeteorNet alone.
- Merged overlapping bounding boxes for better placement.
- Resolved inconsistencies from individual model predictions.
- Removed false positives from MeteorNet if not confirmed by YOLO11m.
- Resulted in cleaner, more reliable predictions.

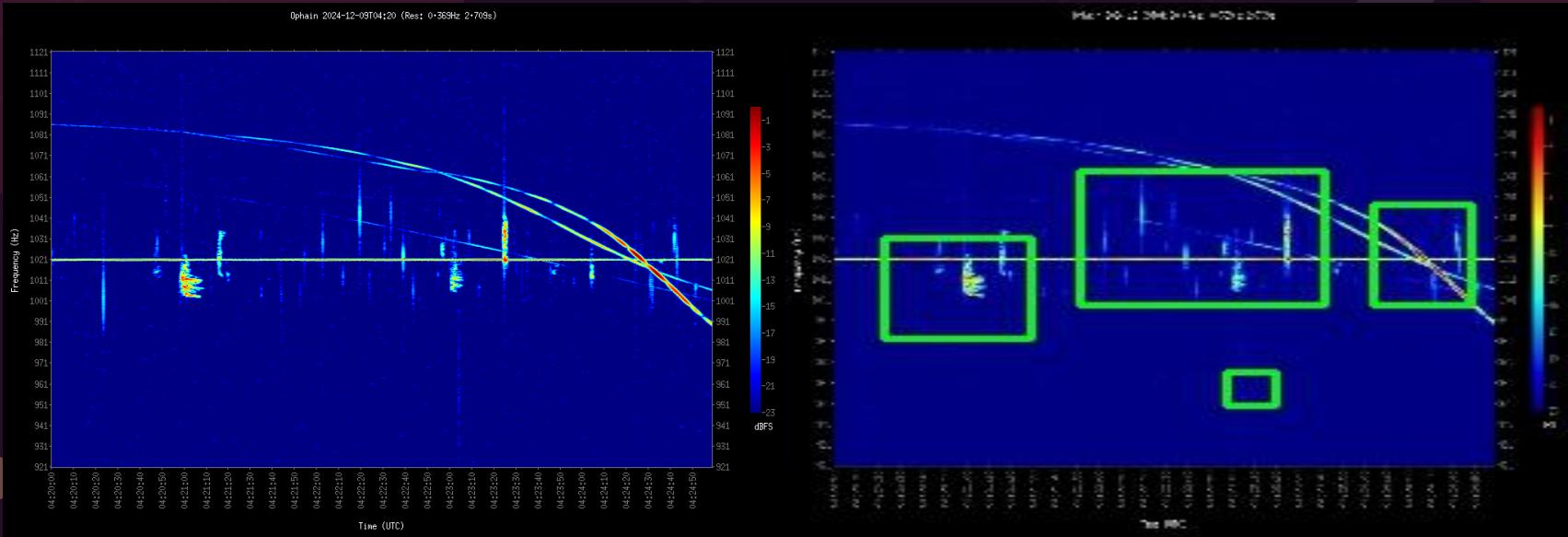
# Prediction power of Yolo11m



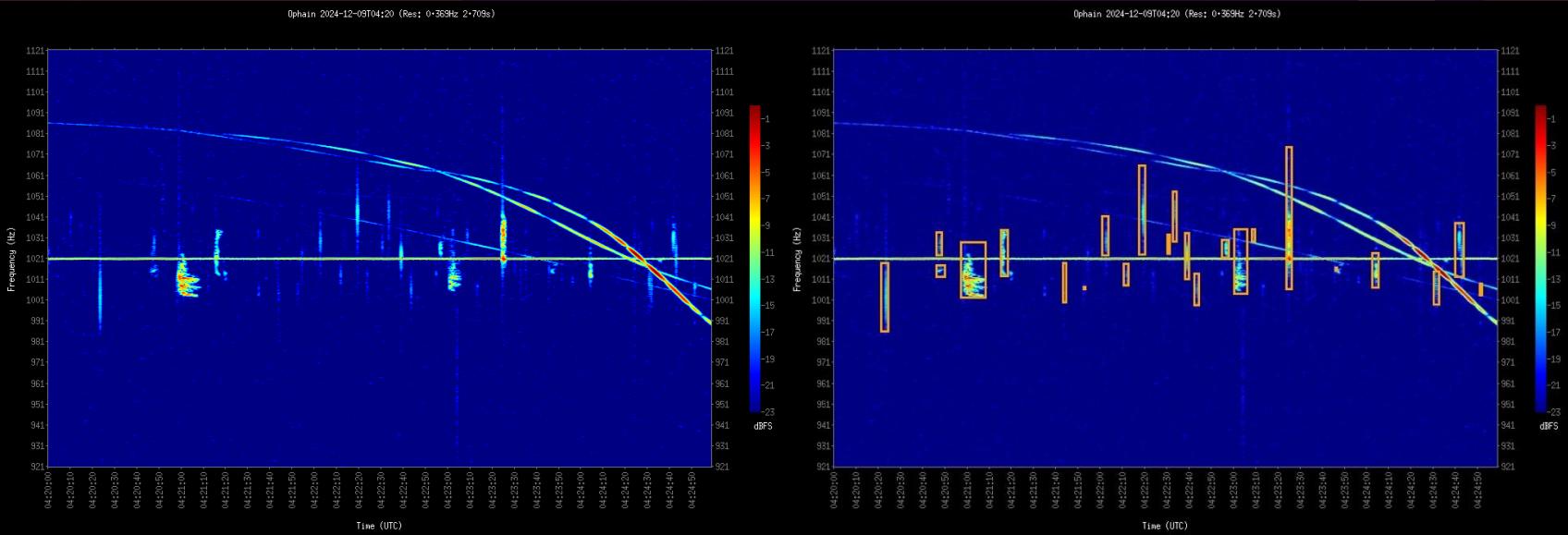
# Prediction power of MeteorNet



# Prediction power of Hybrid Model



# Prediction power of Fusion Technique



# McNemar's Test Results and Analysis

## Purpose of McNemar's Test:

- Compare the performance of YOLO11m, MeteorNet, and Fusion models.
- Evaluate strengths and weaknesses statistically.
- Excluded Hybrid model due to poor performance.

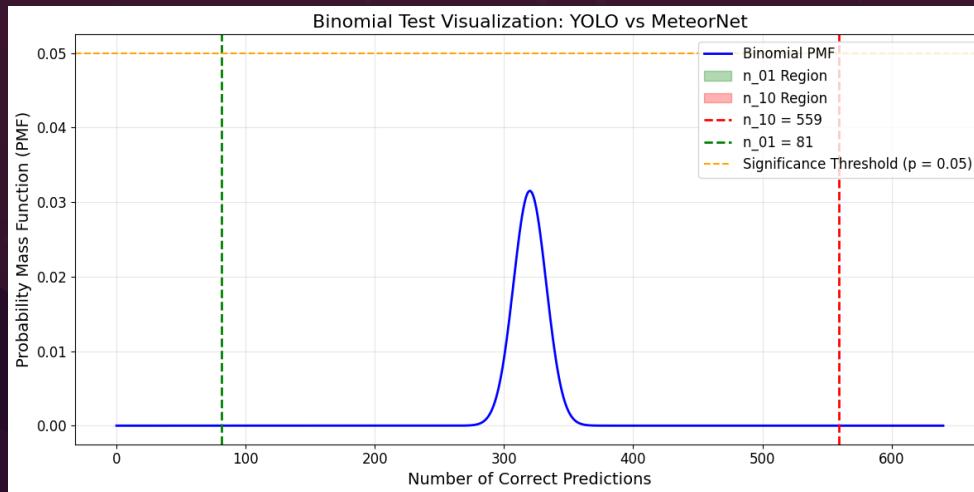
## Implications:

- Fusion model is the best-performing approach for meteor detection.
- Statistical Significance: All comparisons have  $p\text{-value} < 0.05$ .
- Fusion improves detection coverage, bounding box accuracy, and reduces false positives.

# McNemar's Test Results and Analysis

## YOLO11m vs MeteorNet:

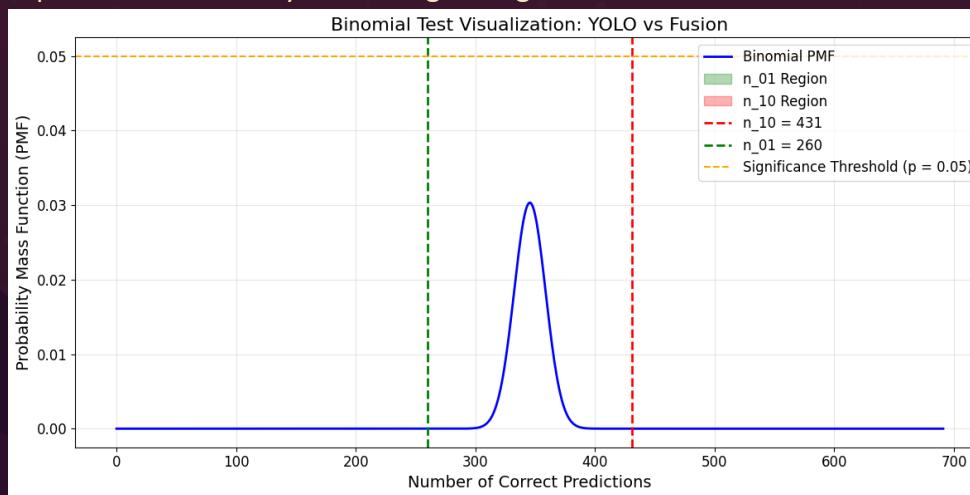
- Correct YOLO11m Predictions Missed by MeteorNet: 559 ( $n_{10}$ )
- Correct MeteorNet Predictions Missed by YOLO11m: 81 ( $n_{01}$ )
- Z-Score: 18.8946
- p-value is < 0.05
- Conclusion: YOLO11m outperforms MeteorNet significantly due to its design for object detection.



# McNemar's Test Results and Analysis

## YOLO11m vs Fusion:

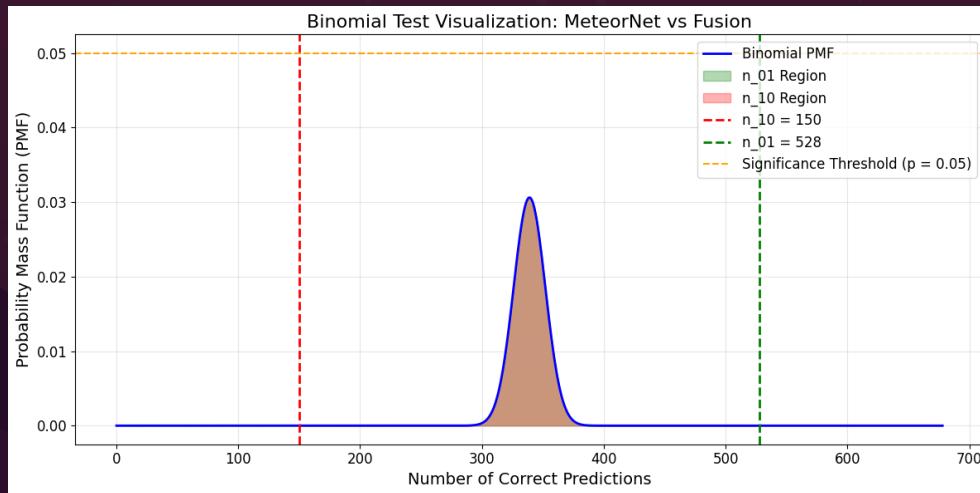
- Correct YOLO11m Predictions Missed by Fusion: 431 ( $n_{10}$ )
- Correct Fusion Predictions Missed by YOLO11m: 260 ( $n_{01}$ )
- Z-Score: 6.5051
- p-value is  $< 0.05$
- Conclusion: Fusion surpasses YOLO11m by combining strengths of YOLO11m and MeteorNet.



# McNemar's Test Results and Analysis

## MeteorNet vs Fusion:

- Correct MeteorNet Predictions Missed by Fusion: 150 (n\_10)
- Correct Fusion Predictions Missed by MeteorNet: 528 (n\_01)
- Z-Score: -14.5170
- p-value is < 0.05
- Conclusion: Fusion resolves MeteorNet's limitations with YOLO11m's precision.



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# Using MIDAS

# Operating MIDAS

MIDAS simplifies meteor detection with an intuitive GUI and integrated models, but its limitations in deployment and performance highlight areas for future improvement.

## Steps to Operate MIDAS:

1. **Model Selection:** Choose a detection model from the sidebar dropdown.
2. **Image Upload:** Drag and drop or browse to upload spectrogram images.
3. **Run Detection:** Click the Detect button to process images.
4. **View Results:** Results appear with annotated bounding boxes in the main display area.

## Color-Coded Outputs:

- YOLO11m: Green bounding boxes.
- MeteorNet: Red bounding boxes.
- Hybrid: Light Green bounding boxes.
- Fusion: Orange bounding boxes.

## Limitations:

- Requires manual installation of dependencies.
- Runs only on local machines.



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# Future Scope

# Future Scope

- Cloud Hosting: Deploy MIDAS on a cloud platform for remote accessibility. Eliminates the need for local setup and manual dependency installation.
- Integration with Meteor Networks: Connect MIDAS with systems like BRAMS for real-time meteor detection and classification.
- Dataset Expansion: Increase training data.
- Increasing the input size for MeteorNet should definitely improve its detection power.
- Implement automated hyperparameter tuning for all models to enhance performance.
- Wider Applications: Test MIDAS in different meteor observatories using spectrogram data. Explore deployment in real time meteor observation systems.

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# Conclusion

# Conclusion

- MIDAS achieved its goal of improving meteor detection in spectrogram images.
- The Fusion model emerged as the best-performing approach.
- YOLO11m: Reliable for precise detection but missed faint meteors.
- MeteorNet: Detected faint meteors effectively but missed few meteors.
- Hybrid Model: Underperformed due to misplaced bounding boxes and poor handling of complex patterns.

# Learnings

- Labelling techniques.
- Importance of balancing precision and recall.
- Combining models can lead to superior performance.
- Techniques like dropout, regularization, and learning rate adjustments enhanced model robustness.
- Hyperparameter tuning (e.g., learning rate, dropout) enhanced model performance.
- Evaluating the model performance using McNemars test.
- Accurate labelling improves model training and detection reliability.

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# Acknowledgements

# Acknowledgements

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# Thank You!

