

Celestial Holography

1/13/25 - 1/16/25 @ Sanya

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4 × 1.5 hour lectures

- 1) Motivation + IR Triangle Primer
- 2) Asymptotic Symmetries & Soft Theorems
- 3) Celestial Amplitudes & the Holographic Dictionary
- 4) Holographic Symmetry Algebras & Future Directions

soft physics book: 1703.05448

celestial lectures: 2108.04801

survey up to '21: 2111.11392

short summary: 2310.04932

Lecture 1: Motivation + IR Triangle Primer

Goals: why flat holography?

What's different about flat spacetimes?

set up Penrose
diagram for $\mathbb{R}^{1,3}$

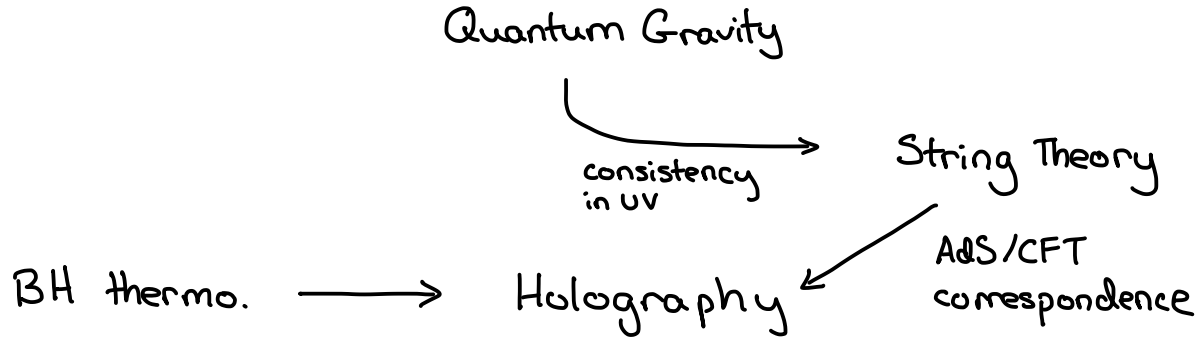
motivate IR triangle

How have people tackled these questions?

flat lim of AdS

Carrollian CFTs

Celestial CFT



Holographic Principle: A theory of quantum gravity can be encoded in a lower dim theory w/out gravity at the spacetime boundary.

Celestial Holography: want to apply the holo. princ. to $\Lambda=0$ spacetimes.

Two Approaches

Top Down: Find stringy construction.

Bottom Up: Match symmetries, identify consistency conditions.

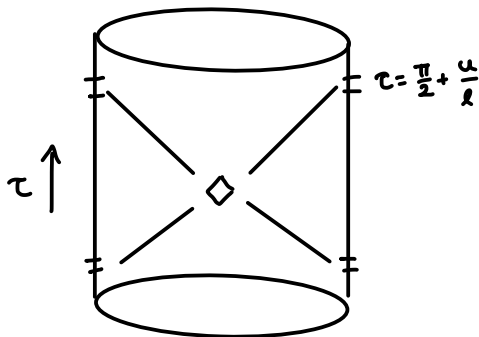
↑ we will follow this approach

— Camellian CFTs
— dim. red. to Celestial CFT

- causal structure of the boundary is different
- $\Lambda=0$ spacetimes have an enhanced asymptotic sym. group
 - IR Triangle

$$G_{\mu\nu} + \Lambda g_{\mu\nu} = 8\pi G T_{\mu\nu}$$

$$\Lambda < 0$$

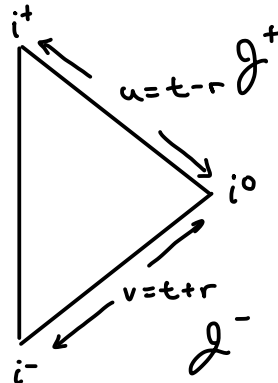


$$ds^2 = -\left(1 + \frac{r^2}{\ell^2}\right) dt^2 + \left(1 + \frac{r^2}{\ell^2}\right)^{-1} dr^2 + 2r^2 \gamma_{z\bar{z}} dz d\bar{z}$$

$$= -\left(1 + \frac{r^2}{\ell^2}\right) du^2 - 2du dr + 2r^2 \gamma_{z\bar{z}} dz d\bar{z}$$

flat limit $\left\{ \begin{array}{l} \longrightarrow -du^2 - 2du dr + 2r^2 \gamma_{z\bar{z}} dz d\bar{z} \\ \text{large } r \longrightarrow \sim r^2 \left(-\frac{1}{\ell^2} du^2 + 2\gamma_{z\bar{z}} dz d\bar{z} \right) \end{array} \right.$

$$\Lambda = 0$$

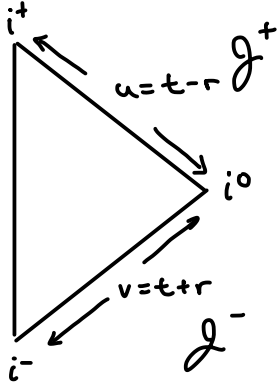


$$z = e^{i\phi} \tan \frac{\theta}{2}, \quad \gamma_{z\bar{z}} = \frac{2}{(1+z\bar{z})^2}$$

$$t = u + \ell \arctan \frac{r}{\ell}$$

$$\Lambda \rightarrow 0, \ell \rightarrow \infty \text{ where } \ell^2 = \frac{3}{|\Lambda|}$$

$$c \sim \frac{1}{\ell} \text{ Comollian limit}$$



starting w/ null coordinates

$$u = t - r, \quad v = t + r$$

we introduce rescaled coords.

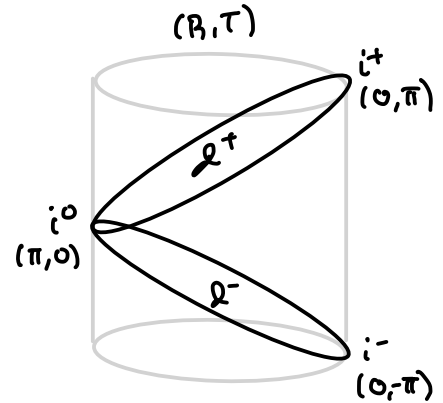
$$u = \tan U, \quad v = \tan V$$

w/ $U, V \in (-\frac{\pi}{2}, \frac{\pi}{2})$. Then the metric in terms of

$$T = U + V, \quad R = V - U$$

is conformal to a patch of $S^3 \times \mathbb{R}$

- massive particles begin at i^- and end at i^+
- massless particles begin at l^- and end at l^+
- spacelike geodesics end at i^0



Asymptotically flat spacetimes

- general solns to $G_{\mu\nu} = 8\pi G T_{\mu\nu}$
- have the same conf. boundary as $\mathbb{R}^{1,3}$ (ignoring this)
- have a larger asymptotic symmetry group

$$\begin{aligned}
 ds^2 = & \underbrace{-du^2 - 2dudr + 2r^2 \gamma_{z\bar{z}}}_{\text{flat metric}} + \underbrace{\frac{2m_B}{r} du^2}_{\text{Bondi mass}} + \underbrace{\left(D_z^2 C_{z\bar{z}} + \frac{1}{r} \left[\frac{4}{3} N_z - \frac{1}{4} D_z (C_{z\bar{z}} C^{\bar{z}z}) \right] \right)}_{\substack{\text{cov. deriv. on } S^2 \\ \text{angular mom. asp.}}} dudz + \text{c.c.} \\
 & + r C_{z\bar{z}} dz^2 + \text{c.c.} + \dots \\
 & \quad \uparrow \text{radiative dof} \quad \uparrow \text{subleading in } r \quad \quad \partial_u m_B, \partial_u N_z \text{ constrained by eom}
 \end{aligned}$$

$$\text{Asymptotic Symmetries} = \frac{\text{Allowed Symmetries}}{\text{Trivial Symmetries}}$$

Bondi, van der Burg,
Metzner, Sachs '62

look at diffeos ξ which preserve gauge
s.t. $\mathcal{L}_\xi g_{\mu\nu} \sim$ same falloffs as above

the # of ξ which act non-triv. on rad. data is $\infty!$

$$\begin{aligned} \xi = & f \partial_u - \frac{1}{r} (D^z f \partial_z + D^{\bar{z}} f \partial_{\bar{z}}) + D^z D_z f \partial_r \\ & + (1 + \frac{u}{2r}) \gamma^z \partial_z - \frac{u}{2r} D^{\bar{z}} D_z \gamma^{\bar{z}} \partial_{\bar{z}} \\ & - \frac{1}{2} (u+r) D_z \gamma^z \partial_r + \frac{u}{2} D_z \gamma^z \partial_u + \text{c.c.} + \dots \end{aligned}$$

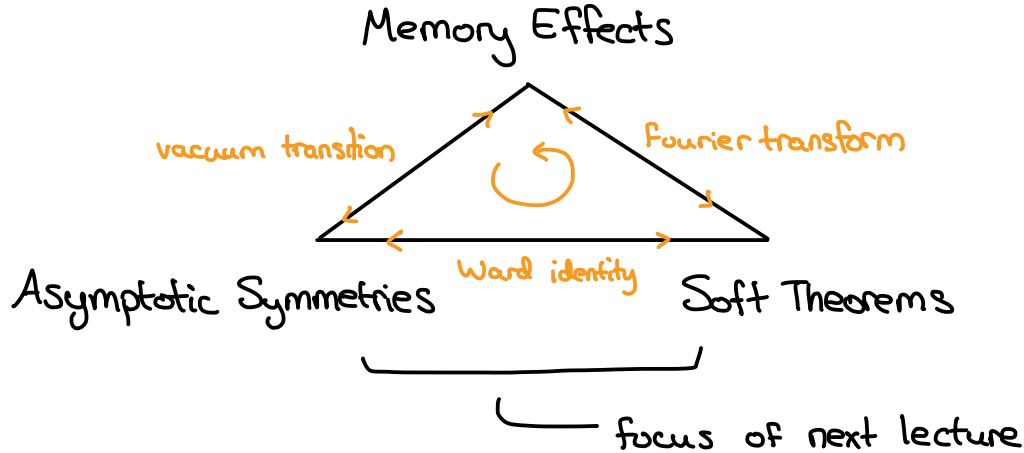
$f(z, \bar{z})$ supertranslations $\infty!$

$\gamma(z)$ superrotations

vs. Poincare $f = c_1 + c_2 \frac{z+\bar{z}}{1+z\bar{z}} + c_3 \frac{i(\bar{z}-z)}{1+z\bar{z}} + c_4 \frac{1-z\bar{z}}{1+z\bar{z}}$

10! $\gamma = a + bz + c\bar{z}^2$

IR Triangle

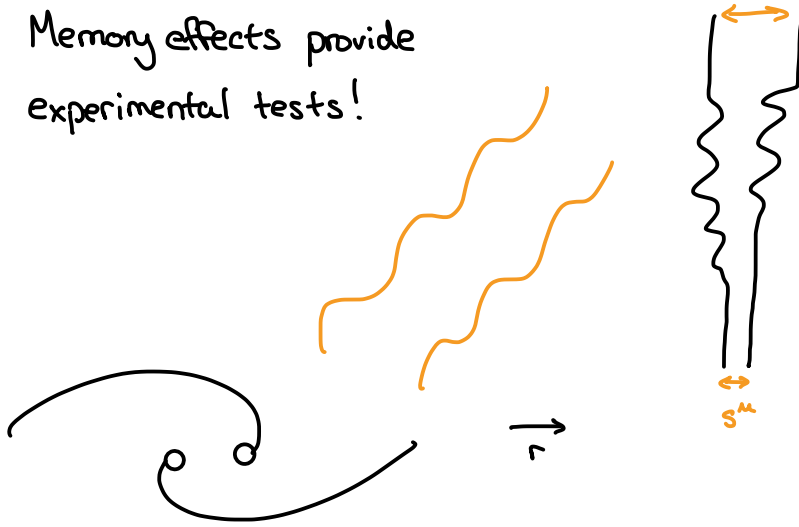


- leading EDM
- leading grav.
- subleading grav.

...

pattern \rightarrow predictions

Memory effects provide experimental tests!



$$\partial_c^2 S^\mu = R^\mu_{\lambda\rho\nu} t^\lambda t^\rho S^\nu$$

$$\downarrow$$

$$\partial_u^2 S^{\bar{z}} = \frac{\partial^2 \bar{z}}{2r} \partial_u^2 C_{zz} S^z$$

$$\downarrow$$

$$\Delta S^{\bar{z}} = \frac{\partial^2 \bar{z}}{2r} \Delta C_{zz} S^z$$

geod. dev.

$$t^\lambda \partial_\lambda \sim \partial_u, \tau \sim u$$

$$R_{\tau zu \tau} \sim -\frac{1}{2} r \partial_u^2 C_{zz}$$

Ström. lec. ex. 13

\exists non-triv. tail behavior of grav. waveform

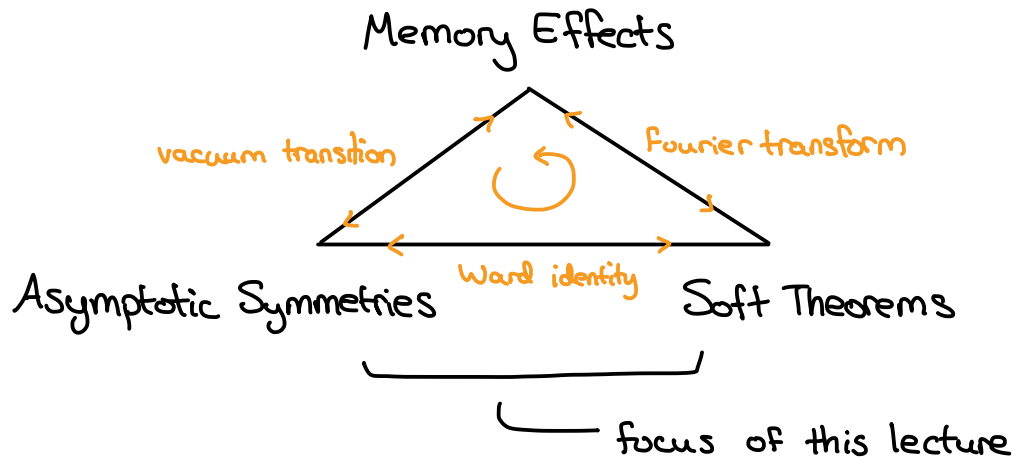
\rightarrow meas. w/ asymp. detectors mem. effect

$\rightarrow \square \Theta(u) \xrightarrow{\text{F.T.}} \frac{1}{\omega} \sim$ soft pole

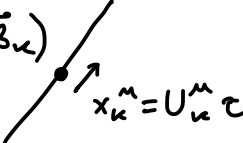
$\rightarrow \Delta C_{zz} = -2 D_z^2 \Delta C$ vac. trans.

Lecture 2: Asymptotic Symmetries & Soft Theorems

Goal: Demonstrate Asymptotic sym. \Leftrightarrow soft thm. for $U(1)$ example



Let us start by considering the electromagnetic field for a set of moving point charges with charge Q_k and 4-velocity U_k^μ

$$U_k^\mu = \gamma_k(1, \vec{\beta}_k)$$


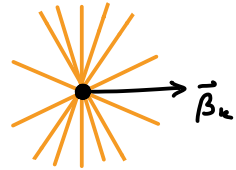
$$x_k^\mu = U_k^\mu \tau$$

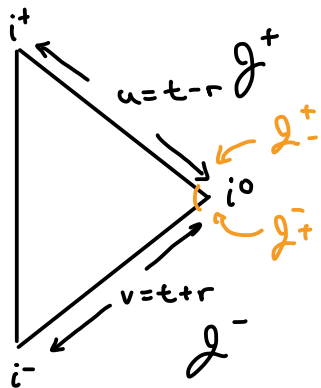
$$j_\mu^M(x) = \sum_{k=1}^n Q_k \int d\tau U_{k,\mu} \delta^{(4)}(x^\nu - U_k^\nu \tau)$$

The Liénard-Wiechert solution to $\nabla^\mu F_{\mu\nu} = e^2 j_\nu^M$ has

$$F_{rt}(t, \vec{x}) = \frac{e^2}{4\pi} \sum_{k=1}^n \frac{Q_k \gamma_k (r - t \hat{x} \cdot \vec{\beta}_k)}{|\gamma_k^2 (t - r \hat{x} \cdot \vec{\beta}_k)^2 - t^2 + r^2|^{3/2}}$$

\vec{E} field lines at $t=0$





Now let's examine how things behave near the conf. bndy.

$$\begin{aligned} \text{recall: } r \rightarrow \infty, u \text{ fixed} &\rightarrow J^+ \\ r \rightarrow \infty, v \text{ fixed} &\rightarrow J^- \end{aligned}$$

also $F_{r\epsilon} = F_{ru} = F_{rv}$ due to anti-sym

$$F_{ru}|_{J^+} = \frac{e^2}{4\pi r^2} \sum_{k=1}^n \frac{Q_k}{\gamma_k^2 (1 - \hat{x} \cdot \vec{\beta}_k)^2}$$

$$F_{rv}|_{J^-} = \frac{e^2}{4\pi r^2} \sum_{k=1}^n \frac{Q_k}{\gamma_k^2 (1 + \hat{x} \cdot \vec{\beta}_k)^2}$$

$$\left. \begin{aligned} F_{ru}|_{J^+} \\ F_{rv}|_{J^-} \end{aligned} \right\} \lim_{r \rightarrow \infty} r^2 F_{ru}(\hat{x})|_{J^+} = \lim_{r \rightarrow \infty} r^2 F_{rv}(-\hat{x})|_{J^-}$$

antipodal matching!

↑ this will play an important role

What is the ASG for $U(1)$ gauge theory?

$$S = -\frac{1}{4e^2} \int d^4x \sqrt{-g} F_{\mu\nu} F^{\mu\nu} + S_M$$

$$\frac{\delta}{\delta A_\mu} \Rightarrow \nabla^\mu F_{\mu\nu} = e^2 j_\nu^M \quad \text{where } F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad \text{and } \nabla^\mu j_\mu^M = 0$$

Now there is also a gauge sym $\delta_\epsilon A_\mu = \partial_\mu \epsilon(u, r, z, \bar{z})$ we should gauge fix

$$\nabla^\mu A_\mu = 0 \quad \text{still allows } \epsilon \text{ s.t. } \square \epsilon = 0$$

↑ residual gauge dof.

Consider the asymptotic expansion

$$\mathcal{O}(u, r, z, \bar{z}) = \sum_n r^{-n} \mathcal{O}^{(n)}(u, z, \bar{z})$$

Then solving $\square \mathcal{E} = 0$ order-by-order gives

$$(\square \mathcal{E})^{(n)} = 2(n-2) \partial_u \mathcal{E}^{(n-1)} + [\square^2 + (n-2)(n-3)] \mathcal{E}^{(n-2)}$$

$\uparrow \mathcal{E}^{(1)}(u, z, \bar{z})$ free data

Can use this to set $A_u^{(1)} = 0$. Then

$$A_u \sim \mathcal{O}\left(\frac{1}{r^2}\right) \quad A_r \sim \mathcal{O}\left(\frac{1}{r^2}\right) \quad A_A \sim \mathcal{O}(1)$$

This residual gauge fixing still allows for a non-zero $\epsilon^{(0)}(z, \bar{z})$.

$$Q_\epsilon = \frac{1}{e^2} \int_{i_0} \epsilon(z, \bar{z}) \star F$$

generates the non-trivial $\delta_\epsilon A_A = \mathcal{L}_A \epsilon$ respecting our b.c.'s

\Rightarrow ASG \ni large $U(1)$ gauge trans.

Meanwhile bc of antipodal matching of $F_{ru}|_{\mathcal{I}^+}$ and $F_{ru}|_{\mathcal{I}^-}$:

$$Q_\epsilon^+ = \frac{1}{e^2} \int_{\mathcal{I}^+} \epsilon \star F = \frac{1}{e^2} \int_{\mathcal{I}^-} \epsilon \star F = Q_\epsilon^- \quad \text{if } \epsilon(z, \bar{z})|_{\mathcal{I}^+} = \epsilon(z, \bar{z})|_{\mathcal{I}^-}$$

$\underbrace{\hspace{10em}}$ our Ward id.
 $\underbrace{\hspace{10em}}$
 $\underbrace{\hspace{10em}}$ antipodal pts.

Can we see $\langle \text{out} | Q_E^+ S - S Q_E^+ | \text{in} \rangle = 0$ in S-matrix elements?

leading r-behavior of u-component of eom:

$$\partial_u F_{ru}^{(2)} + D^z F_{uz}^{(0)} + D^{\bar{z}} F_{u\bar{z}}^{(0)} + e^2 j_u^{(2)} = 0 \quad (\text{assume } m=0 \text{ charges})$$

$$\downarrow$$

$$Q_E^+ = -\frac{1}{e^2} \underbrace{\int_{\mathcal{I}^+} du d^2z \left(\int_z \varepsilon F_{u\bar{z}}^{(0)} + \int_{\bar{z}} \varepsilon F_{uz}^{(0)} \right)}_{Q_S \text{ (photon field)}} + \underbrace{\int_{\mathcal{I}^+} du d^2z \varepsilon \gamma_{z\bar{z}} j_u^{(2)}}_{Q_H \text{ (meas. change)}}$$

ditto for Q_E^- .

Now $\langle \text{out} | Q_S^+$ will change the state to one w/ additional photon
 $\langle \text{out} | Q_H^+$ no change in particle # soft thm!

$$A_\mu(x) = e \sum_{\alpha=\pm} \int \frac{d^3q}{(2\pi)^3} \frac{1}{2\omega_q} [\varepsilon_\mu^{\alpha*}(\vec{q}) a_\alpha(\vec{q}) e^{iq \cdot x} + \varepsilon_\mu^\alpha(\vec{q}) a_\alpha(\vec{q})^\dagger e^{-iq \cdot x}]$$

for our gauge $F_{uz}^{(0)} = \int_u A_z^{(0)} - \int_z A_u^{(0)}$ where $A_z^{(0)} = \lim_{r \rightarrow \infty} \int_z X^\mu A_\mu(x)$

now at large r fixed u : $e^{iq \cdot x} = e^{-i\omega_q u - i\omega_q r(1-\cos\theta)} \rightarrow e^{-i\omega_q u} \times \frac{1}{\omega_q r} \frac{\delta(\theta)}{\sin\theta}$

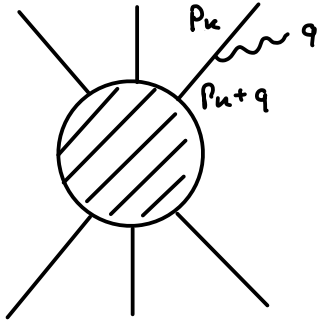
$$\Rightarrow A_z^{(0)} = \frac{-i}{8\pi^2} \frac{\sqrt{2}e}{1+z\bar{z}} \int_0^\infty d\omega_q [a_+(\omega_q \hat{x}) e^{-i\omega_q u} - a_-(\omega_q \hat{x})^\dagger e^{i\omega_q u}]$$

$$\Rightarrow \int du F_{uz}^{(0)} = \frac{-1}{8\pi} \frac{\sqrt{2}e}{1+z\bar{z}} \lim_{\omega \rightarrow 0^+} [\omega a_+(\omega \hat{x}) + \omega a_-(\omega \hat{x})^\dagger]$$

so $\langle \text{out} | Q_S^\dagger S - S Q_S^- | \text{in} \rangle = - \langle \text{out} | Q_H^\dagger S - S Q_H^- | \text{in} \rangle$ from our Ward id

↪ but we know this from soft thms!

Weinberg tells us these insertions have a universal form!



$$\langle \text{out} | a_+(\vec{q}) S | \text{in} \rangle = e \sum_{\text{out-in}} \frac{Q_k p_k \cdot \epsilon^+}{p_k \cdot q} \langle \text{out} | S | \text{in} \rangle + \mathcal{O}(\omega_q^0)$$

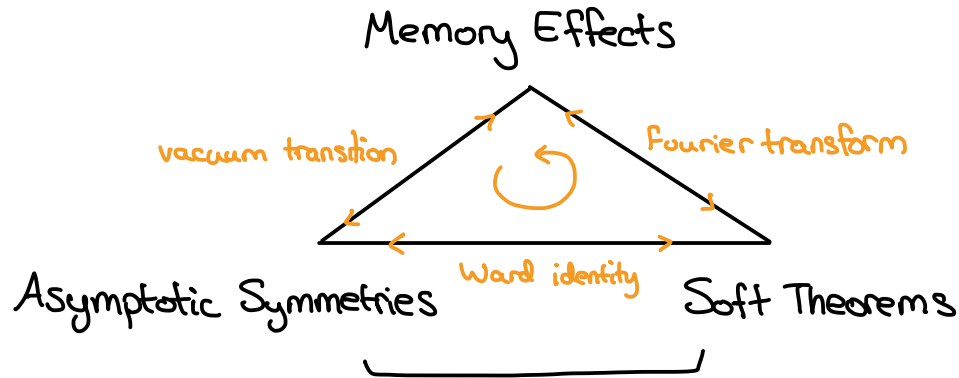
Plugging in the soft theorem

$$\langle \text{out} | \int d\bar{u} F_{u\bar{z}}^{(0)} S | \text{in} \rangle = -\frac{e^2}{4\pi} \sum_{\text{out-in}} \frac{Q_k}{z - \bar{z}_k} \langle \text{out} | S | \text{in} \rangle$$

we indeed have $\langle \text{out} | Q_\epsilon^+ S - S Q_\epsilon^- | \text{in} \rangle = 0$

Ward id \iff soft thm!

Looking back at the IR triangle



- leading EDM large $U(1)$
- leading grav. Supertranslations
- NEW! \Rightarrow • subleading grav. Superrotations
- ...

pattern \rightarrow predictions

But look! $j^+ := Q_S^+(\varepsilon = \frac{1}{z-w}) = -4\pi \int du F_{uz}$ obeys

$$\langle j(z) \mathcal{O}_1(z_1, \bar{z}_1) \dots \mathcal{O}_n(z_n, \bar{z}_n) \rangle = \sum_k \frac{Q_u}{z - z_k} \langle \mathcal{O}_1(z_1, \bar{z}_1) \dots \mathcal{O}_n(z_n, \bar{z}_n) \rangle$$

ASG \Rightarrow 2D Kac-Moody sym. of S-matrix?

Next time: reorganize scattering to
make these symmetries manifest!

↑ CCFT!