Physics 12: Second-hand Data Test 2017

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When source of waves moves, the waves it emits change frequency relative to a stationary observer. This applies to both transverse and longitudinal (sound) waves. As a car moves away from you the frequency of the sound you hear is lower than the frequency it is emitting. A similar effect using radar waves is used by police to measure the speed of cars.

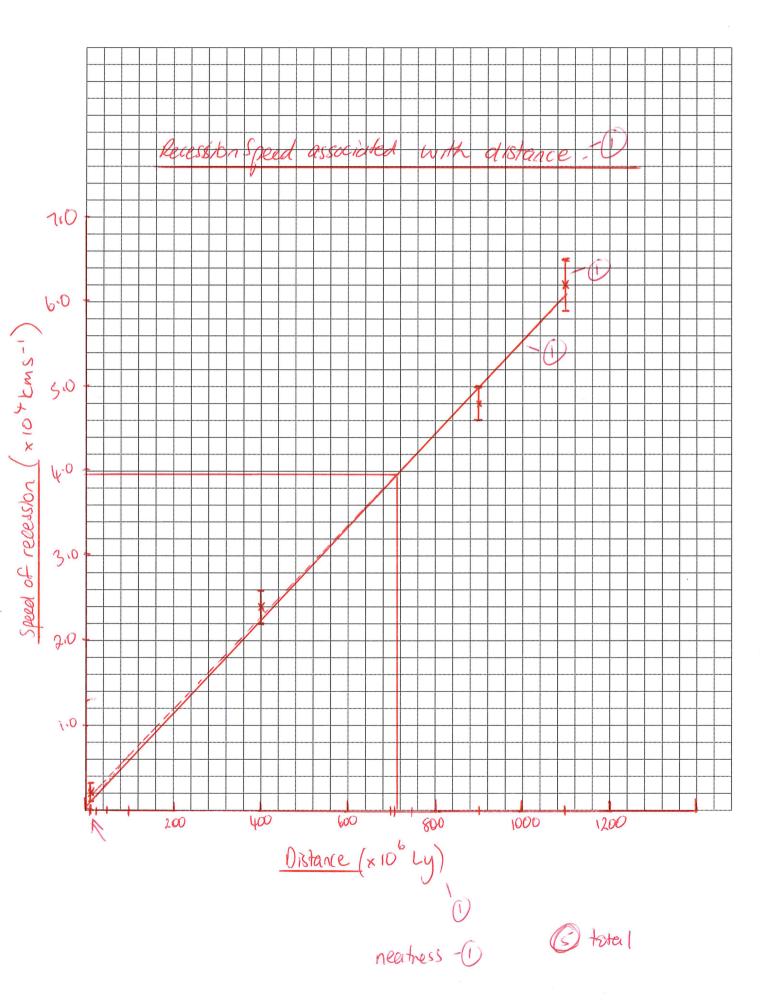
Thus if a source of electromagnetic waves such as a star is moving away from an observer on Earth then the frequencies of the lines in the star's emitted electromagnetic spectrum are shifted to lower values. This is known as red shift.

In 1920, Edwin Hubble measured the red shifts of several galaxies and discovered that most galaxies are moving away from the Earth, suggesting that the Universe is expanding. Hubble also found that the further away a galaxy is, the larger its red shift; that is, the faster it is moving.

The following data together with the associated errors were recorded by Hubble at Mount Wilson in California in the 1940s using an optical telescope.

Object name	Speed of recession (× 10 ⁴ km s ⁻¹)	Distance (× 10 ⁶ light years)
Virgo	0.2 ± 0.1	10.2
Corona Borealis	2.4 ± 0.2	400
Hydra	6.2 ± 0.3	1100
Kip	4.8 ± 0.2	900

(a) Graph these data on the graph paper below, including error bars. Plot recession speed (y-axis) against distance (x-axis) and draw a line of best fit. (5 marks)



(b) Use the graph to predict the recession speed of a galaxy that is 710 × 10⁶ light years from Earth. (2 marks

from Earth. 6-gph evidence - (1)

\$ 4.0 × 10 4kms-1 - (1)

(c) Hubble's Law can be stated as

$$v_{galaxy} = (H_o)(distance)$$

where the term H_o is called Hubble's constant.

Use your graph to calculate a value for H_o. Take care with the units. (4 marks)

(d) The shift in wavelength $\Delta\lambda$ due to recession of a spectral line of wavelength λ is given by the formula

$$v_{galaxy} = \left(\frac{\Delta \lambda}{\lambda}\right) c$$

where c is the speed of light, 3×10^8 m s⁻¹.

A line in the spectrum of ionised calcium has wavelength 393.3 nm when measured in the laboratory. When similar light from the elliptical galaxy NGC 4889 is measured its wavelength is 401.8 nm.

Determine the recession speed of this galaxy.

(3 marks)

$$V = \frac{4068 - 393.3}{393.3} C = \frac{8.5}{393.3} C$$
(cancel $\times 10^{-9}$)
$$V = 0.0216 C$$

$$V = 6.48 \times 10^{6} \text{ ms}^{-1}$$

(e) Edwin Hubble could estimate the age of the Universe from his data by calculating the time for which one of the galaxies has been receding. Determine Hubble's value for the age of the Universe by using the data for Corona Borealis. (4 marks)

Additional working space

Corona Borealis V=2.4 ×10 4 kms-1
S = 400 × 106 Ly
Stolal = Vc x time = (3 x 108) x (365,25 x 24 x 60 x 60) (4
$S_{total} = (3 \times 10^8) \times (31557600) \times (4003 \times 10^6)$ $S_{total} = (9.467 \times 10^{15} \text{ m}) \times (400 \times 10^6)$ $S_{total} = 317869 \times 10^{15} \text{ m}$
S. = (9.467 × 10 15 m) × (400 8 × 10 6)
S = 3.7869 × 10tm
$V = \frac{S}{L} \qquad L = \frac{S + b + al}{V + cR}$
time =?
S=3.78x1024m
V=2.4×104 kms-1 = 2.4×10 ms-1
0 32 4 10 10 10 10 10 10
t = 15 = 3.7869 × 1024
$t = \sqrt{c_R} = 3.7869 \times 10^{24} \text{m}$ $2.4 \times 10^7 \text{ms}^{-1}$
~ '4 × 10 ms
t = 1.58 × 10 17 seconds
U - 1.30 ×10 Seconds

t = 5.00 × 10 9 years

Physics Stage 3: Practical Exam 3B

Question 22 (18 marks)

When source of waves moves, the waves it emits change frequency relative to a stationary observer. This applies to both transverse and longitudinal (sound) waves. As a car moves away from you the frequency of the sound you hear is lower than the frequency it is emitting. A similar effect using radar waves is used by police to measure the speed of cars.

Thus if a source of electromagnetic waves such as a star is moving away from an observer on Earth then the frequencies of the lines in the star's emitted electromagnetic spectrum are shifted to lower values. This is known as red shift.

In 1920, Edwin Hubble measured the red shifts of several galaxies and discovered that most galaxies are moving away from the Earth, suggesting that the Universe is expanding. Hubble also found that the further away a galaxy is, the larger its red shift; that is, the faster it is moving.

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(a) Graph this data on the graph paper below, including error bars. Plot recession speed (*y*-axis) against distance (*x*-axis) and draw a line of best fit. (5 marks)

Description	Marks
Graph axes labelled correctly	2
Points plotted correctly	1
Error bars include to size	1
Line of best drawn thru 0,0	1
	Total 5

Use the graph to predict the recession speed of a galaxy that is 710×10^6 light years (b) from Earth.

Description	Marks
Evidence they have used graph to predict the recession speed	1
Value 4 × 10 ⁷ m s ⁻¹	1
	Total 2

(c) Hubble's Law can be stated as

$$v_{galaxy} = (H_o)(distance)$$

where the term H_o is called Hubble's constant.

Use your graph to calculate a value for H_o. Take care with the units.

5.54×10-2.

(4 marks)

Description	Marks
Gradient of graph= 4 × 10 ⁷ m s ⁻¹	2
$H_0 = 5.6 \times 10^{-2} \text{ m s}^{-1} \text{ Ly}^{-1}$	2
(allow range of answers 4.5 to 6.5 due to line of best fit)	
	Total 4

The shift in wavelength $\Delta\lambda$ due to recession of a spectral line of wavelength λ is given (d) by the formula

$$v_{galaxy} = \left(\frac{\Delta \lambda}{\lambda}\right) c$$

where c is the speed of light, 3 × 10⁸ m s⁻¹.

A line in the spectrum of ionised calcium has wavelength 393.3 nm when measured in the laboratory. When similar light from the elliptical galaxy NGC 4889 is measured its wavelength is 401.8 nm.

Determine the recession speed of this galaxy.

(3 marks)

Description	Marks
recession speed = $\frac{(4.108 \times 10^{-9}) - (3.93 \times 10^{-9})}{3.93 \times 10^{-9}} \times 3 \times 10^{8}$	2
$= 6.48 \times 10^6 \text{ m s}^{-1}$	1
	Total 3

(e) Edwin Hubble could estimate the age of the Universe from his data by calculating the time for which one of the galaxies has been receding. Determine Hubble's value for the age of the Universe by using the data for Corona Borealis. (4 marks)

Description	Marks
1 light year = $3 \times 10^8 \times 365 \times 60 \times 60 \times 24 = 1.58 \times 10^{15}$ m	1
time = $\frac{(400 \times 10^6)(1.58 \times 10^{15})}{2.2 \times 10^3 \times 10^4}$	2
$2.2\times10^3\times10^4$	
= 1.37 × 10 ⁸ years	1
	Total 4