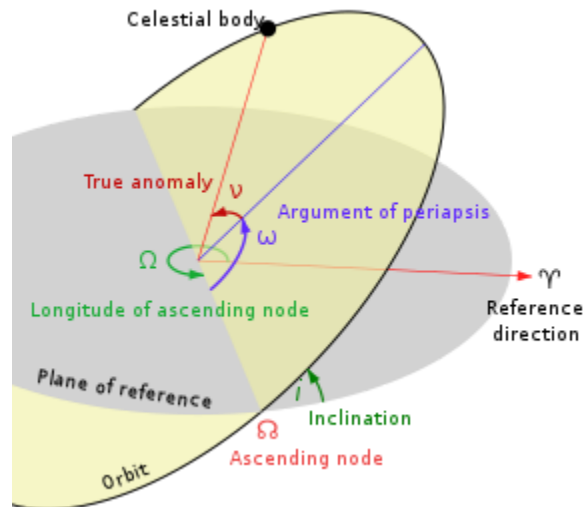


- 1.(80points) The SPE Approximation: The purpose of this question is to illustrate:
- How a non-Keplerian orbit can be represented using time-variable orbital elements.
 - How the secularly precessing ellipse is a better approximation to the non-Keplerian orbit than a Keplerian ellipse.

a.(5 points) Define a metric for the difference between any two orbits (recall—"orbits" can refer to an ephemeris, or a time-series of position & velocity).



The metric I am most acclimated with to define the differences between orbits is the orbital elements.

- **Semimajor Axis - a** : determines the size of the orbit.
- **Eccentricity - e** : determines the shape of the orbit
- **True Anomaly - ν** : determines location on the orbit
- **Argument of Periapsis - ω** : determines orientation of the ellipse (orbit plane)
- **Inclination - i** : also orients the orbit plane
- **Longitude of the Ascending Node - Ω** : also orients the orbit plane

b. Difference between IPE and TE

This is at the final time 3 days later.

rvDifference =

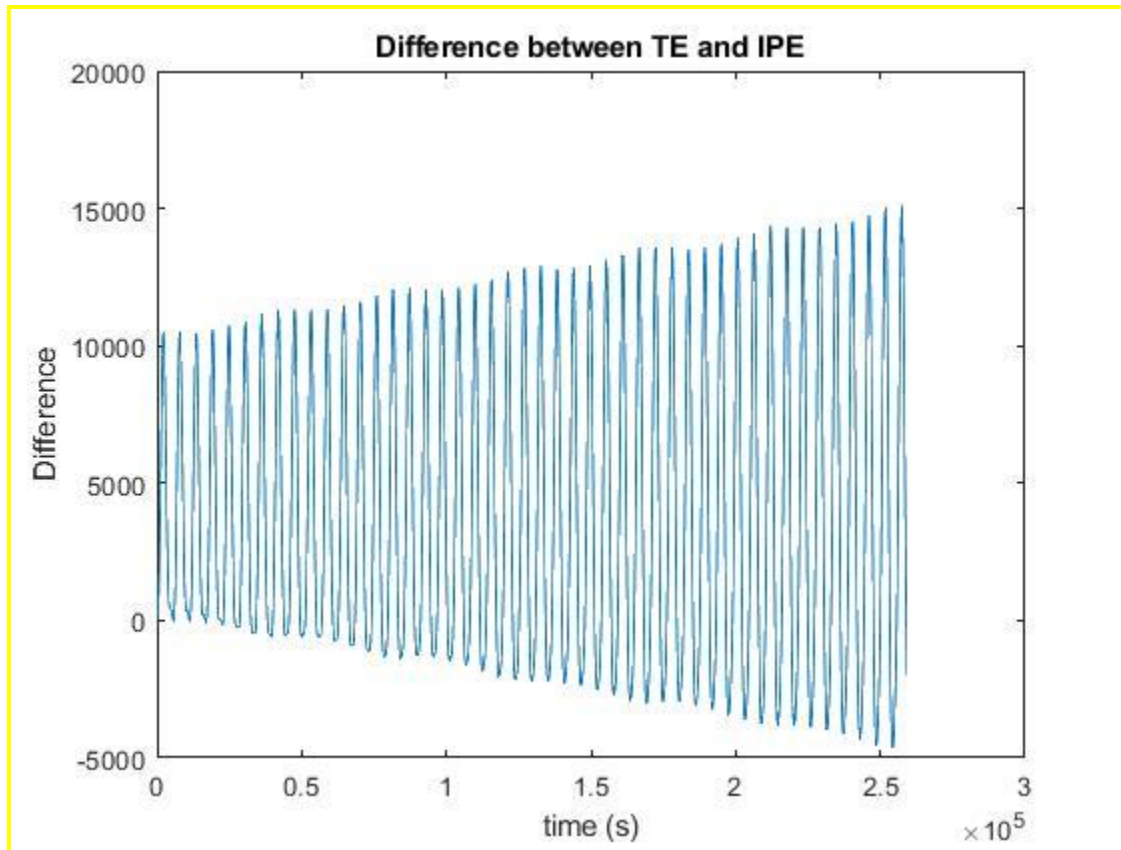
1.0e+06 *
 1.8097 0.9899 -1.0549 -0.0011 -0.0005 -0.0023

oeDifference =

1.0e+03 *
 8.9975 0.0000 -0.0000 0.0002 0.0000 0.0001

C. Difference between oscillating elements of IPE and TE with time-scales

Assumed the timescale set as the same found in the data text file, 3 minute intervals for 3 days or 259200 seconds.

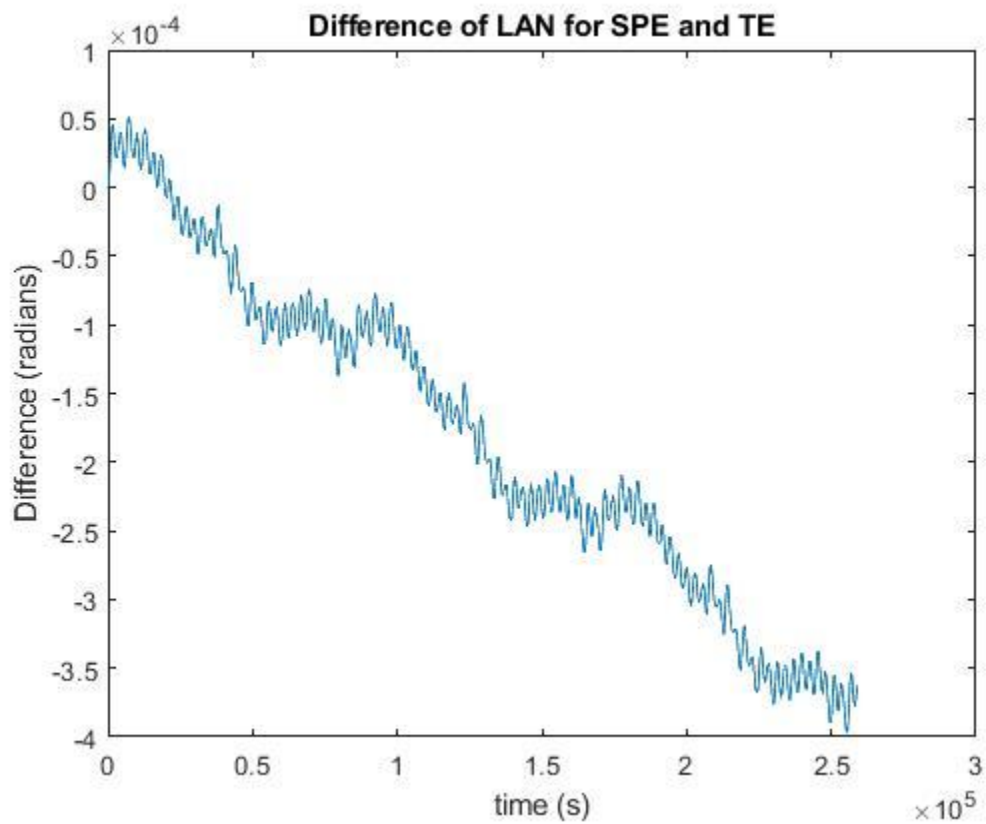
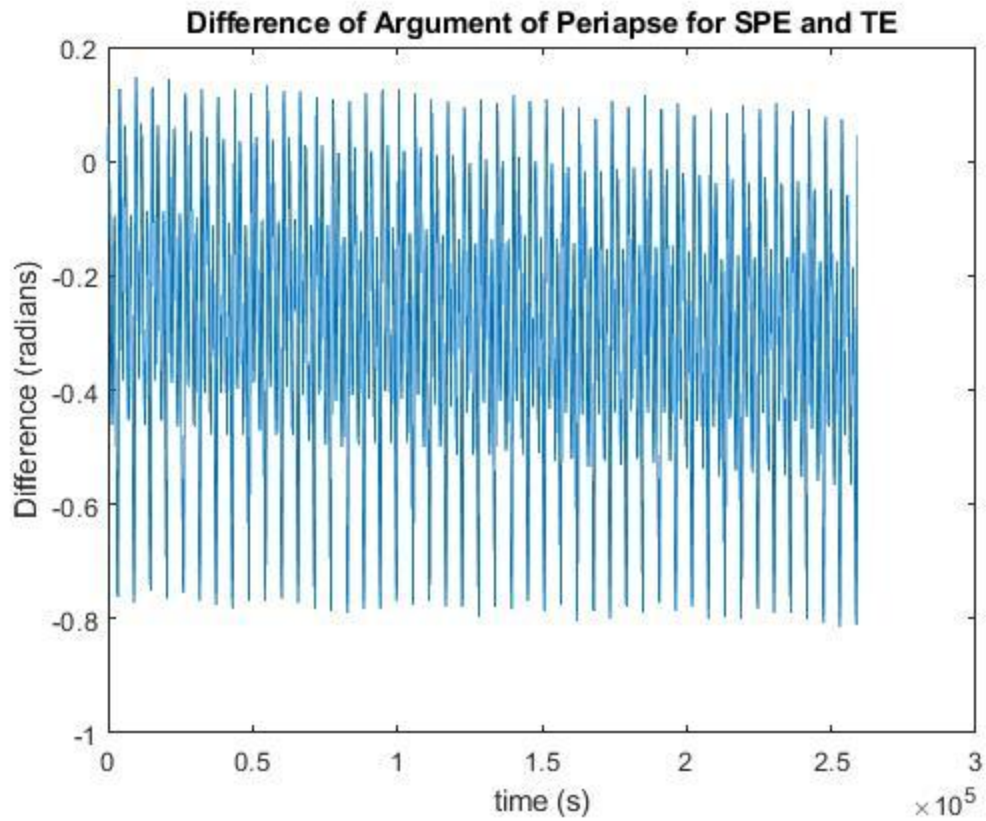


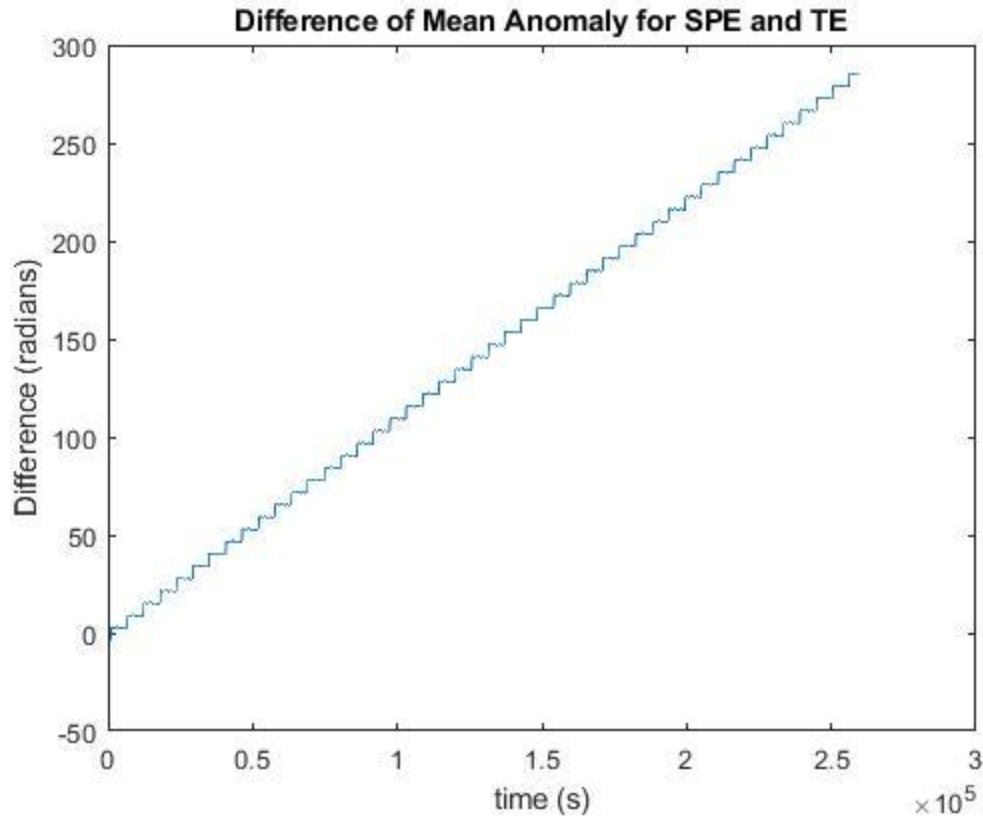
D. Difference between SPE and TE after 3 day propagation (final position and velocity in terms of orbital elements)

SPEdifference =

0 0 0 -0.2006 -0.0074 287.1639

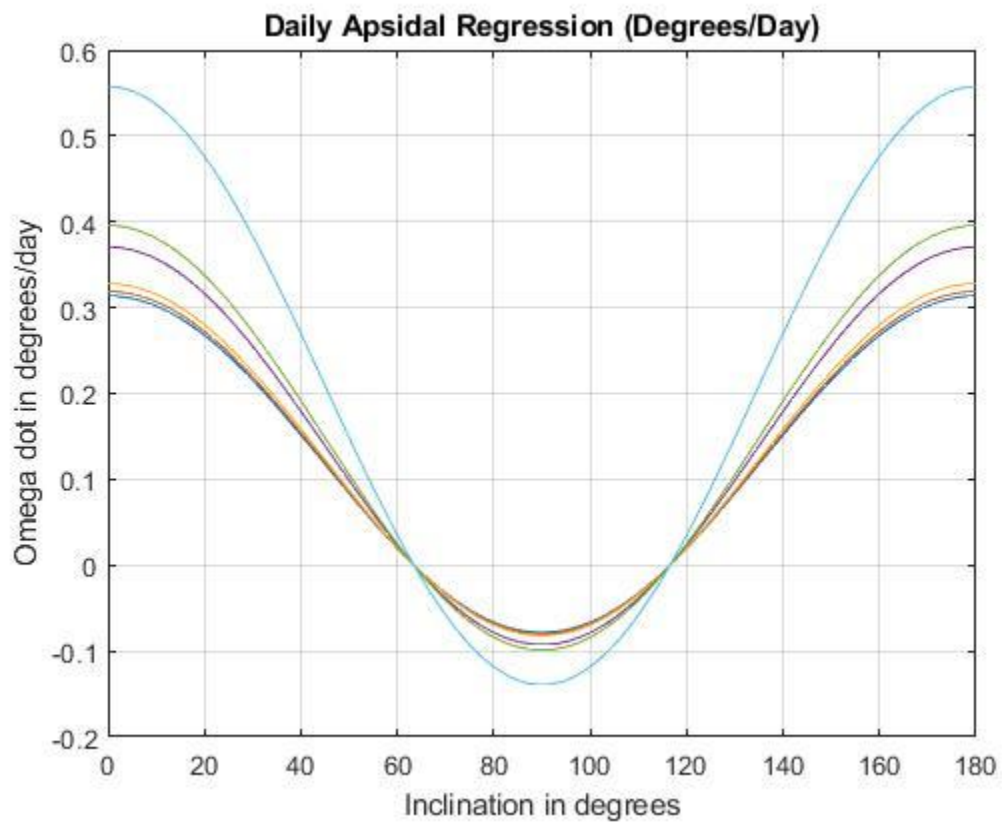
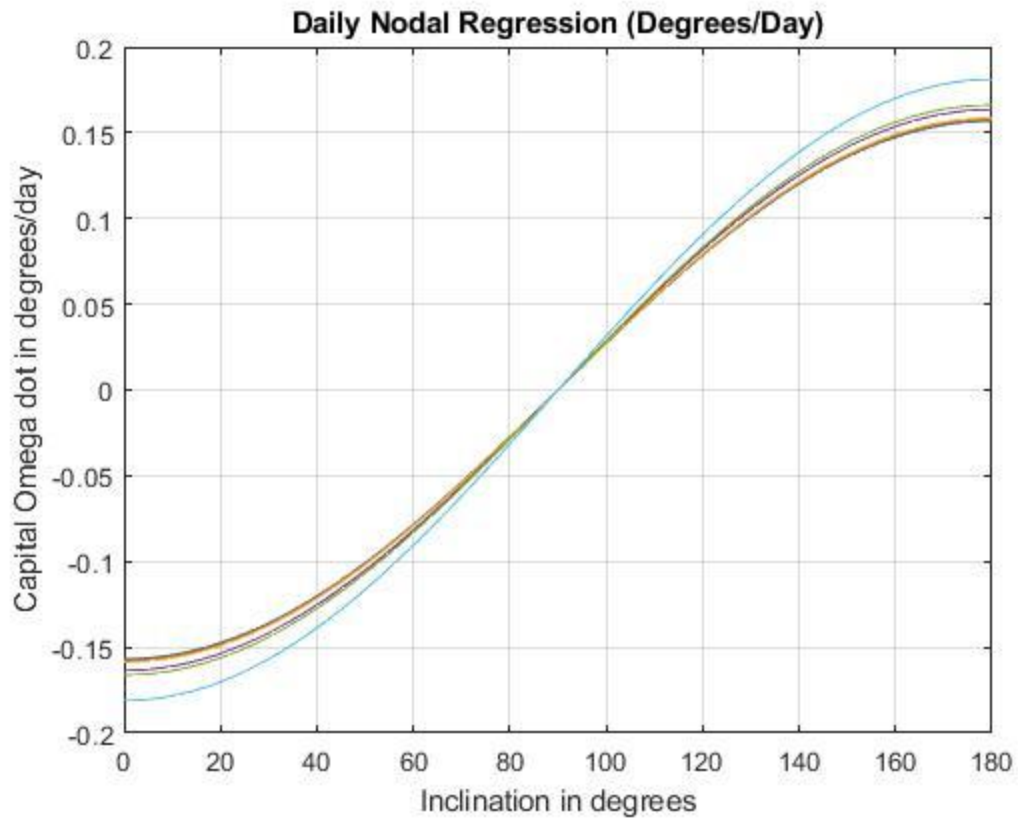
e. Plot the difference of the osculating elements





F. The difference found in parts b and c was massive compared to parts d and e. Although part c was calculated as a norm of position and velocity vectors, it has a difference magnitude $1e4$ to $1e8$ greater than the difference associated with Argument of Periapse, Longitude of Ascending Node, and Argument of Periapse. This was to be expected as parts c and d take into account the oblateness of the Earth and the J_2 scalar in order to produce a far more accurate propagation.

Problem 2.



CODE APPENDIX:

%% Applied Orbital Mechanics HW#3

%Mo & to specified by initial conditions

%constants

mu = 3.986004415*10¹⁴;

ae = 6378136.3;

we = 7.292115*10⁻⁵;

g=9.81;

j2=1.082*10⁻³;

%

%% PART B

%t = 0.0 sec

rv0=[0.447927332846917398E+07 0.245170029746151669E+07 -0.461804936123361625E+07 -0.455055498

oe0 = hw6rv2oe(rv0,mu);

options = odeset('RelTol', 1e-14, 'AbsTol', 1e-14);

[t,rvf_prop]=ode45(@(t,r) body2(t,r,mu),[0:180:259200], rv0, options);

rvf_propagated = [rvf_prop(end,1), rvf_prop(end,2), rvf_prop(end,3), rvf_prop(end,4), rvf_prop(end,5), rvf_prop

oef_prop = hw6rv2oe(rvf_propagated,mu);

%t = 259200.0 sec

rvf=[0.187794820127492165E+07 0.835255100187982782E+06 0.653630524325667322E+07 0.645165445

oef = hw6rv2oe(rvf,mu);

%[a e i w bigOmega v(nu)]

rvDifference = rvf_propagated - rvf;

oeDifference = oef_prop - oef;

%% PART C

file = load('hw2-p01.txt');

timescale = (file(:,1));

rvTE = (file(:,2:7));

```

for b = 1:size(rvTE,1)
    range_prop = norm(rvf_prop(b,1:3));
    range_TE = norm(rvTE(b,1:3));
    rvSPediff_plot(b) = range_TE - range_prop;
end

figure(1)
plot(timescale,rvSPediff_plot)
xlabel('time (s)')
ylabel('Difference')
title('Difference between TE and IPE')
%The graph diverges, IPE value becomes less and less accurate as the orbit
%is propagated.

%% PART D
a0=oe0(1);
e0=oe0(2);
i0=oe0(3);
w0=oe0(4);
bigW0=oe0(5);

E0=2*atan(sqrt((1-oe0(2))/(1+oe0(2))))*tan(oe0(6)));
M0=E0-e0*sin(E0);
nu0= oe0(6);

oe0_M=oe0;
oe0_M(6) = M0;
n = sqrt(mu/a0^3);

% T = 2*pi*sqrt(oe0(1)^3/mu);
% Tquarter = T/4;
%
% E0=2*atan(sqrt((1-oe0(2))/(1+oe0(2))))*tan(oe0(6)));
% M0=sqrt(mu/(oe0(1)^3))*T-2*pi+E0-oe0(2)*sin(E0);

bigWbar = -1.5*n*(ae/a0)^2*j2/sqrt(1-e0^2)*cos(i0);
wbar = -0.75*n*(ae/a0)^2*j2/(1-e0^2)^2*(1-5*(cos(i0))^2);
Mbar = n*(1-(0.75*(ae/a0)^2*j2/(1-e0^2)^(3/2)*(1-3*(cos(i0))^2)));

bigWspe = bigW0 + bigWbar*(timescale);

```

```

wspe = w0 + wbar*(timescale);
Mspe = M0 + Mbar*(timescale);

spe=zeros(1441,6);
spe(:,1)=a0;
spe(:,2)=e0;
spe(:,3)=i0;
spe(:,4)=wspe;
spe(:,5)=bigWspe;
spe(:,6)=Mspe;

% Ef_M=zeros(1441,1)
% for iter=1:1441
%   Ef_M(i)=hw7kepler(Mspe(iter),e0);
% end
%
% nu_M = 2*atan(sqrt((1+e0)/(1-e0))*tan(Ef_M/2));
% speNu=spe
oeTE=zeros(1441,6);
for i = 1:1441
    oeTE(i,:)= hw6rv2oe(rvTE(i,:), mu);
end
Ee=2*atan(sqrt((1-e0)/(1+e0))*tan(oeTE(:,6)/2);
Me=Ee-e0*sin(Ee);
oeTE_M(:,6)=Me;

speFinal = [a0 e0 i0 wspe(end) bigWspe(end) Mspe(end)];
SPEdifference = speFinal - oe0_M'

diffwe=spe(:,4)-oeTE(:,4);
diffWe=spe(:,5)-oeTE(:,5);
diffMe=spe(:,6)-oeTE(:,6);
diffMe=diffMe+2*pi*(diffMe<-.5);
% speInitial=zeros(1440,6);
% speInitial(:,1)=a0;
% speInitial(:,2)=e0;
% speInitial(:,3)=i0;
% speInitial(:,4)=wspe;
% speInitial(:,5)=bigWspe;
% speInitial(:,6)=Mspe;

```



```

% spe=[oe0,speInitial];

%%% PROBLEM E

figure(2)
plot(timescale,diffwe)
title('Difference of Argument of Periapse for SPE and TE')
xlabel('time (s)')
ylabel('Difference (radians)')

figure(3)
plot(timescale,diffWe)
title('Difference of LAN for SPE and TE')
xlabel('time (s)')
ylabel('Difference (radians)')

figure(4)
plot(timescale,diffMe)
title('Difference of Mean Anomaly for SPE and TE')
xlabel('time (s)')
ylabel('Difference (radians)')
% rv_spe(1,:) = rv0;
%
% w_spe = zeros(1441, 1);
% M_spe = zeros(1441, 1);
% W_spe = zeros(1441, 1);
% E_spe = zeros(1441, 1);
% nu_spe = zeros(1441, 1);
% for k = 2:size(timescale,1)
%   w_spe(k) = w0 + wbar*timescale(k);
%   M_spe(k) = M0 + Mbar*timescale(k);
%   W_spe(k) = bigW0 + bigWbar*timescale(k);
%   E_spe(k) = keplerEqn(M_spe(k),e0);
%   nu_spe(k) = 2*atan(sqrt((1+e0)/(1-e0))*tan(E_spe(k)/2));
%   oe_spe = [a0 e0 i0 w_spe(k) W_spe(k) nu_spe(k)];
%   rv_spe(k,:) = hw6oe2rv(oe_spe,mu);
% end
%
% for k = 1:size(rvTE,1)
%   range_spe = norm(rv_spe(k,1:3));

```

```

% range_TE = norm(rvTE(k,1:3));
% rvSPEdiff_plot(b) = range_TE - range_spe;
% end
%
% figure(2)
% plot(timescale,rvSPEdiff_plot)
% xlabel('time (s)')
% ylabel('Difference')
% title('Difference between TE and SPE')

%% PROBLEM 2
muMars = 4.282837e13;
aMars = 3396190;
J2Mars = 1.932;
am2 = aMars + 100000;
I2 = linspace(0,pi,100)';
e2 = [0 .09 .148 .283 .331 .5];

nM = sqrt(muMars/(am2^3));

O2 = zeros(100,6);
w2 = zeros(100,6);

for p = 1:6
    O2(:,p) = (-3/2)*nM*((aMars/am2)^2)*J2Mars/sqrt(1-(e2(p)^2))*cos(I2);
    w2(:,p) = (-3/4)*nM*((aMars/am2)^2)*J2Mars/((1-(e2(p)^2))^2)*(1-5*(cos(I2).^2));
end

I2plot = I2*180/pi;
Omega2plot = O2*180/pi;
w2plot = w2*180/pi;

figure(5)
plot(I2plot, Omega2plot)
grid on

```

```
title('Daily Nodal Regression (Degrees/Day)')  
xlabel('Inclination in degrees')  
ylabel('Capital Omega dot in degrees/day')
```

```
figure(6)  
plot(I2plot, w2plot)  
grid on  
title('Daily Apsidal Regression (Degrees/Day)')  
xlabel('Inclination in degrees')  
ylabel('Omega dot in degrees/day')
```