

```
# Return rounded mass value
return np.round(total_mass, 3)
```

```
In [67]: # These are the 'testing' locations of the galaxies data
mw_path = '../SampleData/MW_000.txt'
m31_path = '../SampleData/M31_000.txt'
m33_path = '../SampleData/M33_000.txt'
```

```
In [65]: # List of Galaxy names
names = 'MilkyWay', 'M31', 'M33'

# List of total mass components for the Milky Way: Halo, Disk, and Bulge
mw_masses = ( GalaxyMass(mw_path, ptype) for ptype in [1, 2, 3] )
m31_masses = ( GalaxyMass(m31_path, ptype) for ptype in [1, 2, 3] )
m33_masses = ( GalaxyMass(m33_path, ptype) for ptype in [1, 2] )

# Column labels! According to assignment
columns = ['Galaxy Name', 'Halo Mass (10^12 Msun)', 'Disk Mass (10^12 Msun)',
           'Bulge Mass (10^12 Msun)', 'Total (10^12 Msun)', 'f_bar']

# Pass data to a Pandas DataFrame, create initial table
mass_breakdown = pd.DataFrame([['MilkyWay', *mw_masses, 0, 0],
                                ['M31', *m31_masses, 0, 0],
                                ['M33', *m33_masses, 0, 0] ],
                               columns=columns)

mass_breakdown
```

Out[65]:

	Galaxy Name	Halo Mass (10 ¹² Msun)	Disk Mass (10 ¹² Msun)	Bulge Mass (10 ¹² Msun)	Total (10 ¹² Msun)	f_bar
0	MilkyWay	1.975	0.075	0.010	0	0.0
1	M31	1.921	0.120	0.019	0	0.0
2	M33	0.187	0.009	0.000	0	NaN

```
In [66]: # Index component mass columns (1,2,3), and sum rowwise to get total mass per galaxy (10^12 Msun)
total_masses = mass_breakdown[columns[1:4]].sum(axis=1)

# Reassign values for total mass column
mass_breakdown['Total (10^12 Msun)'] = total_masses
mass_breakdown
```

Out[66]:

	Galaxy Name	Halo Mass (10 ¹² Msun)	Disk Mass (10 ¹² Msun)	Bulge Mass (10 ¹² Msun)	Total (10 ¹² Msun)	f_bar
0	MilkyWay	1.975	0.075	0.010	2.060	0.0
1	M31	1.921	0.120	0.019	2.060	0.0
2	M33	0.187	0.009	0.000	0.196	NaN

```

In [73]: # Baryon mass per galaxy
baryon_masses = mass_breakdown[columns[2:4]].sum(axis=1)

# baryon fraction per galaxy
f_bar = baryon_masses / total_masses

# Local group baryon mass, and DM mass
local_baryon_mass = baryon_masses.sum()
local_DM_mass = mass_breakdown['Halo Mass (10^12 Msun)'].sum()

# Total local group masss (10^12 Msun)
local_mass = total_masses.sum()

# local group baryon fraction
local_fbar = local_baryon_mass / local_mass

# Reassign values for baryon fraction mass
mass_breakdown['f_bar'] = f_bar

print(f'Local Group Baryon Mass = \t{local_fbar}')
mass_breakdown

```

Local Group Baryon Mass = 0.053985171455050975

Out[73]:

	Galaxy Name	Halo Mass (10^12 Msun)	Disk Mass (10^12 Msun)	Bulge Mass (10^12 Msun)	Total (10^12 Msun)	f_bar
0	MilkyWay	1.975	0.075	0.010	2.060	0.041262
1	M31	1.921	0.120	0.019	2.060	0.067476
2	M33	0.187	0.009	0.000	0.196	0.045918

```

In [77]: print(f"MW to M31 Halo Mass Ratio = \t{mass_breakdown['Halo Mass (10^12 Msun)'][0]/mass_breakdown['Halo Mass (10^12 Msun)'][1]}")
print(f'MW to M31 Baryon Mass Ratio = \t{baryon_masses[0]/baryon_masses[1]}')

```

MW to M31 Halo Mass Ratio = 1.028110359187923
 MW to M31 Baryon Mass Ratio = 0.6115107913669064

Questions (some of the questions sort of overlap)

1. How does the total mass of the MW and M31 compare in this simulation? What galaxy component dominates this total mass?

Their total mass are equal! However M31's baryon mass is slightly higher than the MW, and the opposite for the halo mass. This baryon mass is mostly concentrated on the bulge; M31's bulge is ~x2 heavier than MW's bulge.

1. How does the stellar mass of the MW and M31 compare? Which galaxy do you expect to be more luminous?

Andromeda is expected to be brighter due to the slightly higher baryon mass ratio.

1. How does the total dark matter mass of MW and M31 compare in this simulation (ratio)? Is this surprising, given their difference in stellar mass?

The MW to M31 Halo mass ratio is ~1.03. While the baryon ratio is ~0.61. This means that the MW's dark matter mass makes up the lack of baryon. However, it is surprising that the ratios are so different. They essentially have the same dark matter mass; so they should probably also have the same baryon mass.

1. What is the ratio of stellar mass to total mass for each galaxy (f_{bar})? In the Universe, $\Omega_b/\Omega_m \sim 16\%$. of all mass is locked up in baryons (gas & stars) vs dark matter. How does this ratio compare to the baryon fraction you computer for each galaxy? Given than the total gas mass in the disk of these galaxies is negligible compared to the stellar mass, any ideas for why the universal baryon fraction might differ from that in these galaxies?

The baryon fractions are on the table above. Our f_{bar} values are ~1/2 and 1/3 of the universal baryon fraction. Throughout the galaxy's life, the baryon mass fraction changes due to the mainly gravitational interaction between baryon and dark matter; baryon particles interact in many more ways than dark matter. This interaction can take in form of AGN feedback & outflows.