Hazma☆

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Abstract

Hazma does many wonderful things.

Keywords: Dark matter, indirect detection

1. Introduction

Hazma is great.

Notation/units. Unless otherwise noted, the units used throughout hazma are

• Masses: MeV

• Decay widths: MeV

• Cross sections: cm³/s

2. Installation and Package Structure

2.1. Installation

Before installing hazma, the user needs to have a few of well-established packages: cython, scipy, numpy, and scikit-image. Theses are easily installed by using PyPi. If the user has PyPi installed on their system, then these packages can be installed using

 $^{^{\}mbox{\tiny \triangle}}$ Code and tutorials available at github.com/LoganAMorrison/Hazma. $\it Email~addresses:$ a.m.coogan@uva.nl (Adam Coogan), loanmorr@ucsc.edu (Logan A. Morrison)

pip install cython, scipy, numpy, scikit-image, matplotlib

Hazma can be installed in the same way, using:

pip install hazma

This will download a tarball from the PyPi repository, compile all the c-code and install hazma on the system. Alternatively, the user can install hazma by downloading the package from https://github.com/LoganAMorrison/Hazma.git. Once downloaded, navigate to the package directory using the command line and run either

pip install .

or

python setup.py install

Note that since hazma makes extensive usage of the package cython, the user will need to have a c-compiler installed on their system. For more information, see the cython installation guide: https://cython.readthedocs.io/en/latest/src/quickstart/install.html

3. Overview: The theory Class

Should we have boxes for different functions and constructors specifying type signatures? Should that be left for the actual docs in the code? Seems like it, based on other publications in this journal.

Instances of the theory class represent dark matter models, and are the main objects used to perform analyses in Hazma. Two theory, each containing a DM and mediator particle, come built into hazma. The Lagrangians defining these theories and the domains of validity are described in the following subsection. theory possesses methods to compute quantities such as DM annihilation cross sections, gamma ray spectra and constraints, and CMB constraints. This section catalogues the definitions of these quantities used by hazma along with minimal examples of how to use the associated methods.

3.1. Built-in Theories

Each of the models that ship with hazma contain two BSM particles:

- A dark matter particle;
- A mediator M that interacts with the DM as well as Standard Model particles.

The Lagrangians can be expressed as

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{DM} + \mathcal{L}_{M} + \mathcal{L}_{Int(M)}, \tag{1}$$

which consists of the SM Lagrangian, the free Lagrangians for the mediator and DM, and the mediator's interactions with the DM and SM fields. Both the dark matter and the mediator are taken to be uncharged under the Standard Model gauge group. The Lagrangian is defined in terms of the microscopic degrees of freedom of the Standard Model (quarks, leptons and gauge bosons). However, at the energy scale of interest for self-annihilations of nonrelativistic MeV dark matter, quarks and gluons are not the corrent strongly-interacting degrees of freedom. Instead, the microscopic Lagrangian must be matched onto the effective Lagrangian for pions and other mesons using the techniques of chiral perturbation theory (chPT). The models currently implemented in hazma utilize leading-order chPT.

As with an effective field theory (EFT), chPT has a limited range of validity. Observables in chPT are computed in terms of an expansion in a small parameter, the meson momentum p divided by the mass scale $\Lambda_{\rm ChPT} \sim 4\pi f_\pi \sim 1~{\rm GeV}$, where f_π is the pion decay constant. As $p^2 \to \Lambda_{\rm ChPT}$, higher-order Feynman diagrams in the chPT expansion provide contributions to observables comparable to leading order ones. This suggests that leading-order chPT cannot be trusted for computing dark matter self-annihilation cross sections when $m_{\rm DM} \gtrsim 500~{\rm GeV}$. In fact, the convergence of the chPT expansion is disrupted at a lower mass scale by the lowest-lying hadronic resonances, the ρ ($m_{\rho}=770~{\rm GeV}$) and the $f_0(500)$ ($m_{f_0(500)}\sim 450~{\rm MeV}$). Figure 1 illustrates the resulting regions

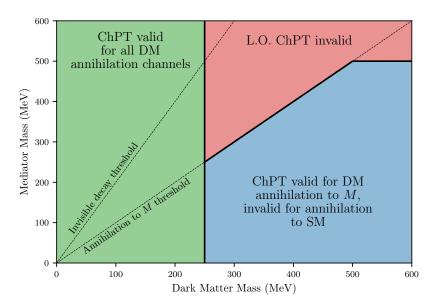


Figure 1: Values of the DM mass $m_{\rm DM}$ and mediator mass m_M for which the leading-order chiral perturbation theory calculations used to implement the models built into hazma can be trusted.

where the leading-order chPT calculations can and cannot be trusted in the $(m_{\rm DM}, m_{\rm mediator})$ plane. An important consequence is that annihilation into kaons and heavier mesons can be ignored since that would require DM masses far above the chPT range of validity. This is important to keep in mind when using the models provided with hazma.¹

While the preceding discussion applies irrespective of the mediator and DM spin and quantum numbers, the rest of this subsection specializes to the two models that come with hazma: scalar_mediator, which contains a real scalar mediator S and vector_mediator, where the mediator is a vector V. In both cases the DM is taken to be a Dirac fermion, so that

$$\mathcal{L} = \bar{\chi}(i\partial \!\!\!/ - m_{\chi})\chi. \tag{2}$$

The rest of this subsection presents $\mathcal{L}_{\mathrm{Int}(S)}$ and $\mathcal{L}_{\mathrm{Int}(V)}$ at the level of quarks and gluons as well as the Lagrangians obtained by performing the chPT matching. Snippets are provided to demonstrate how to construct each **theory** and change its parameters. The interaction Lagrangians, matching procedure and a brief overview of the chiral Lagrangian are explained in detail in a companion paper [?].

3.1.1. scalar_mediator

The free Lagrangian for a real scalar is

$$\mathcal{L}_S = \frac{1}{2} (\partial_\mu S)(\partial^\mu S) - \frac{1}{2} m_S^2 S^2, \tag{3}$$

where m_S is the scalar's mass. The interactions with the light fundamental SM degrees of freedom

$$\mathcal{L}_{\text{Int}(S)} = -S \left(g_{S\chi} + g_{Sf} \sum_{f} \frac{y_f}{\sqrt{2}} \bar{f} f \right)$$
 (4)

$$+\frac{S}{\Lambda}\left(g_{SF}\frac{\alpha_{\rm EM}}{4\pi}F_{\mu\nu}F^{\mu\nu}+g_{SG}\frac{\alpha_s}{4\pi}G^a_{\mu\nu}G^{a\mu\nu}\right).$$
 (5)

¹Of course, user-defined models (see Section 5) need not use leading-order chPT.

The sum runs over fermions with mass below the GeV scale $(f = e, \mu, u, d, s)$. Note that the coupling g_{Sf} is outside the sum. The Yukawas are defined to be $y_f = \sqrt{2}m_f/v_h$, with the Higgs vacuum expectation value (vev) defined as $v_h = 246$ GeV. The parameter Λ is the mass scale at which S acquires (non-renormalizable) interactions with the photon and gluon.

After performing the matching onto the chiral Lagrangian and expanding to leading order in the pion fields, the resulting interaction Lagrangian is

$$\mathcal{L}_{\text{Int}(S)} = \frac{2g_{SG}}{9\Lambda} S \left[(\partial_{\mu} \pi^{0})(\partial^{\mu} \pi^{0}) + 2(\partial_{\mu} \pi^{+})(\partial^{\mu} \pi^{-}) \right]$$
 (6)

$$+\frac{4ieg_{SG}}{9\Lambda}SA^{\mu}\left[\pi^{-}(\partial_{\mu}\pi^{+})-\pi^{+}(\partial_{\mu}\pi^{-})\right]$$
 (7)

$$-\frac{B(m_u + m_d)}{6} \left(\frac{3g_{Sf}}{v_h} + \frac{2g_{SG}}{3\Lambda}\right) S\left[(\pi^0)^2 + 2\pi^+\pi^-\right]$$
 (8)

$$+\frac{B(m_u + m_d)g_{SG}}{81\Lambda} \left(\frac{2g_{SG}}{\Lambda} - \frac{9g_{Sf}}{v_h}\right) S^2 \left[(\pi^0)^2 + 2\pi^+\pi^-\right]$$
 (9)

$$+\frac{4e^2g_{SG}}{9\Lambda}S\pi^+\pi^-A_{\mu}A^{\mu}$$
 (10)

$$-g_{S\chi}S\bar{\chi}\chi - g_{Sf}S\sum_{\ell=e,\mu}\frac{y_{\ell}}{\sqrt{2}}\bar{\ell}\ell + \frac{\alpha_s}{4\pi\Lambda}g_{SG}SG^a_{\mu\nu}G^{a\mu\nu}.$$
 (11)

The parameters for the scalar model are attributes of the scalar_mediator class. Their names in hazma are

$$(m_Y, m_S, g_{SY}, g_{Sf}, g_{SG}, g_{SF}, \Lambda) \leftrightarrow (m_X, m_S, g_{SXX}, g_{Sff}, g_{SG}, g_{SFF}, lam).$$

The following snippet shows how to create an instance of scalar_mediator, change the value of a parameter, and print its new value:

Talk about Higgs portal and heavy fermion subclasses!

3.1.2. vector_mediator

For the vector mediator the free part of the Lagrangian is

$$\mathcal{L}_V = -\frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \frac{1}{2} m_V^2 V_\mu V^\mu, \tag{12}$$

where m_V is the mass of the vector. The interactions considered are

$$\mathcal{L}_{\text{Int}(V)} = V_{\mu} \left(g_{V\chi} \bar{\chi} \gamma^{\mu} \chi + \sum_{f} g_{Vf} \bar{f} \gamma^{\mu} f \right) - \frac{\epsilon}{2} V^{\mu\nu} F_{\mu\nu}. \tag{13}$$

The sum again runs over the light fermions $(f = e, \mu, u, d, s)$, and V may have different couplings to each of these. The last term is a kinetic mixing between the photon and V, which can be eliminated by transforming the photon $A_{\mu} \to A_{\mu} - \epsilon V_{\mu}$. Upon this field redefinition V acquires an ϵ -suppressed interaction with the SM fermions, which is captured by changing the fermion couplings

$$g_{Vf} \to g_{Vf} - \epsilon e Q_f,$$
 (14)

where Q_f is the electric charge of the fermion f and e > 0 is the electron's charge.

Matching onto the chiral Lagrangian and isolating the terms contributing at leading order to the quantities computed in hazma gives

$$\mathcal{L}_{\text{Int}(V)} = -i(g_{Vu} - g_{Vd})V^{\mu} \left(\pi^{+}\partial_{\mu}\pi^{-} - \pi^{-}\partial_{\mu}\pi^{+}\right)$$
(15)

$$+ (g_{Vu} - g_{Vd})^2 V_{\mu} V^{\mu} \pi^+ \pi^- \tag{16}$$

$$+2e(Q_u - Q_d)(g_{Vu} - g_{Vd})A_{\mu}V^{\mu}\pi^{+}\pi^{-}$$
(17)

$$+\frac{1}{8\pi^2 f_{\pi}} \epsilon^{\mu\nu\rho\sigma} (\partial_{\mu} \pi^0) \tag{18}$$

$$\times \left\{ e(2g_{Vu} + g_{Vd}) \left[(\partial_{\nu} A_{\rho}) V_{\sigma} + (\partial_{\nu} V_{\rho}) A_{\sigma} \right] \right. \tag{19}$$

$$+3(g_{V_{IJ}}^{2}-g_{V_{IJ}}^{2})(\partial_{\nu}V_{\rho})V_{\sigma}\}.$$
 (20)

The base vector_mediator class contains the parameters²

```
\begin{split} (m_\chi, m_V, g_{V\chi}, g_{Vu}, g_{Vd}, g_{Vs}, g_{Ve}, g_{V\mu}) \\ & \leftrightarrow (\texttt{mx}, \texttt{mv}, \texttt{gvxx}, \texttt{gvuu}, \texttt{gvdd}, \texttt{gvss}, \texttt{gvee}, \texttt{gvmumu}). \end{split}
```

The following snippet shows how to instantiate and alter the parameter values for vector_mediator:

For convenience, a subclass called KineticMixing is also provided to handle the important case where V couples to the SM purely through the kinetic mixing term ($\epsilon \neq 0$, $g_{Vf} = 0$ for all f). The parameters for this subclass are

$$(m_{\gamma}, m_{V}, g_{V\gamma}, \epsilon) \leftrightarrow (\text{mx}, \text{mv}, \text{gvxx}, \text{eps}).$$

While the underlying parameters gvuu, ..., gvmumu can be accessed by instances of KineticMixing, they cannot be set directly since they are fully determined by eps:

²Note that gvss is not currently used since vector_mediator is based on leading-order chPT.

```
>>> from hazma.vector_mediator import KineticMixing
>>> km = KineticMixing(mx=150., mv=1e3, gvxx=1., eps=0.1)
>>> km.eps
0.1
>>> km.gvuu
-0.020187846690459792 # = -0.1 * 2/3 * sqrt(4 pi / 137)
>>> km.gvuu = 0.1
AttributeError: Cannot set gvuu
```

3.2. Cross section, decay widths and branching fractions

theory. $list_annihilation_final_states()$. Lists the final states into which the DM can potentially annihilate. These are based on the terms appearing in $\mathcal{L}_{Int(M)}$. Note that the actual branching fractions into a given final state may be zero if the relevant couplings are zero or the center of mass energy for the annihilation is smaller than the sum of the final state particles' masses.

```
>>> ScalarMediator.list_annihilation_final_states()
['mu mu', 'e e', 'g g', 'pi0 pi0', 'pi pi', 's s']
>>> VectorMediator.list_annihilation_final_states()
['mu mu', 'e e', 'pi pi', 'pi0 g', 'v v']
```

This is not quite right. For example, it excludes the final states $e^+e^-\gamma$, $S\pi^0\pi^0$, $\pi^+\pi^-\gamma\gamma$ for the scalar and $V\pi^0$ for the vector. While that's fine since those final states are so suppressed, the wording here needs to be precise.

theory.annihilation_cross_sections(cme). For an instance th of theory and a center of mass energy cme, this returns a dict whose keys are the elements of th.list_annihilation_final_states() and 'total' and whose values are the corresponding cross sections for annihilating into those states. In the theories included with hazma, the propagators used for the mediators take into account their widths.

In the following example, the DM is lighter than the mediator, but the center of mass energy is large enough to permit annihilations into the mediator:

theory.annihilation_branching_fractions(cme). For an instance the of theory and a center of mass energy cme, this returns a dict whose keys are the elements of th.list_annihilation_final_states() and whose values are the corresponding branching fractions for annihilating into those states.

```
>>> km = KineticMixing(mx=170., mv=1e3, gvxx=1., eps=0.2)
>>> bfs = km.annihilation_branching_fractions(400.)
>>> print(bfs)
{'pi0 g': 1.5585058782195258e-05,
   'e e': 0.4855695299567697,
   'mu mu': 0.4698109970123726,
   'v v': 0.0,
   'pi pi': 0.04460388797207543}
```

theory.partial_widths(). Returns a dict whose keys are states into which the mediator can decay and whose values are the corresponding decay widths.

- 3.3. Gamma ray and positron spectra
- 3.4. Gamma ray limits

Explain procedure

3.5. Cosmic Microwave Background limits

Explain procedure

3.6. Gamma-ray and Positron Spectra

In hazma, their are two main types of spectra: Final state radiation (FSR) and decay spectra. Processes where there is a photon or electron in the immediate final state are classified as FSR. For example, the process of two dark matter particles annihilating through a scalar mediator into two charged fermions or mesons and a photon, i.e. $\overline{\chi}\chi \to S^* \to \overline{f}f\gamma$ or $\overline{\chi}\chi \to S^* \to \pi^+\pi^-\gamma$. Decay spectra arise when one of the final state particles is unstable and is able to decay radiatively, like the muon, charged pion or neutral pion. If this is the case, the spectrum for the state A decaying or annihilating into the state XB with X being some state and B then decaying radiatively to the state $C\gamma$, can be approximated using the narrow width approximation as

$$\frac{dN(A \to XB \to XC\gamma)}{dE} \sim BF(A \to XB) \frac{dN(B \to X\gamma)}{dE}$$
 (21)

An example of such a process is two dark matter particles annihilating into two muons or two pions: $\overline{\chi}\chi \to S^* \to \mu^+\mu^-$ or $\overline{\chi}\chi \to S^* \to \pi^+\pi^-$. For both cases, the final state particles are unstable and will decay. For the decay spectra, it is easiest to compute the gamma-ray or positron spectra in the rest frame of the decaying particle. Then, to compute the spectra in another frame, the spectrum must be boosted.

3.6.1. Boost Integrals

4. How Hazma Works

5. Advanced Usage

• Make a list or graphic showing what the user has to implement.

6. Conclusion

7. Bibliography

Here are two sample references: [1, 2].

References

- [1] R. Feynman, F. Vernon Jr., Annals of Physics 24 (1963) 118–173. doi:10. 1016/0003-4916(63)90068-X.
- [2] P. Dirac, Physica 19 (1953) 888–896. doi:10.1016/S0031-8914(53) 80099-6.