Cosmology Practice

For scalar inflation

Input Potential

In[2]:=
$$V[\phi_{-}] := \frac{1}{2} m^2 \phi^2$$

$$ln[3]:= \rho[\phi_{-}] := \frac{1}{2} (D[\phi[t], t])^{2} + V[\phi]$$

In[4]:= H :=
$$\frac{1}{\sqrt{3} \star M} \sqrt{\rho [\phi]}$$

In[5]:=
$$\rho[\phi]$$

Out[5]=
$$\frac{m^2 \phi^2}{2} + \frac{1}{2} \phi' [t]^2$$

$$In[6]:= D[H, t]$$

Out[6]=
$$\frac{\phi'[t] \phi''[t]}{2 \sqrt{3} M \sqrt{\frac{m^2 \phi^2}{2} + \frac{1}{2} \phi'[t]^2}}$$

Inflation constants (Not approximation)

Slow Role Parameters

Basic

$$\ln[10] = A[\phi_{-}] = \frac{M^2 * D[V[\phi], \phi]^2}{2 * V[\phi]^2}$$

Out[10]=
$$\frac{2 \text{ M}^2}{\phi^2}$$

$$In[11]:= B[\phi_{-}] = M^2 \star \frac{D[D[V[\phi], \phi], \phi]}{V[\phi]}$$

Out[11]=
$$\frac{2 \text{ M}^2}{\phi^2}$$

$$\ln[12] = \mathbf{N}_{SR} = \mathbf{Assuming} \left[\mathbf{M} > 0 & & \phi_{\mathbf{I}} > 0 & & \phi_{\mathbf{E}} > 0, \text{ Integrate} \left[\frac{1}{\sqrt{2 * \mathbf{A}[\phi]} * \mathbf{M}}, \{\phi, \phi_{\mathbf{E}}, \phi_{\mathbf{I}}\} \right] \right] / \cdot \phi_{\mathbf{E}} \rightarrow 0$$

Out[12]=
$$\frac{\phi_i^2}{4 \text{ M}^2}$$

$$\log_{13} = \phi_{SR} := \text{Simplify} \Big[\text{Reduce}[N == N_{SR}, \phi_{I}], \text{ Assumptions} \rightarrow M > 0 \text{ \&\& } \phi_{i} > 0 \Big]$$

$$ln[14]:= RU = \{ToRules[\phi_{SR}]\}$$

$$\text{Out[14]} = \left\{ \left\{ \phi_\mathtt{i} \ \to -\, 2\ \mathtt{M}\ \sqrt{\mathtt{N}}\ \right\}\text{, } \left\{ 2\ \mathtt{M}\ \sqrt{\mathtt{N}}\ \to \phi_\mathtt{i} \right\} \right\}$$

In[15]:=
$$A[\phi_I]$$
 /. RU

Out[15]=
$$\left\{ \frac{1}{2 \text{ N}}, \frac{2 \text{ M}^2}{\phi_{\text{ii}}^2} \right\}$$

In[16]:=
$$\epsilon_v$$
 = Part[%, 1]

Out[16]=
$$\frac{1}{2 \text{ N}}$$

$$ln[17]:= B[\phi_I] /. RU$$

Out[17]=
$$\left\{ \frac{1}{2 \text{ N}}, \frac{2 \text{ M}^2}{\phi_{i}^2} \right\}$$

In[18]:=
$$\eta_{v}$$
 = Part[%, 1]

Out[18]=
$$\frac{1}{2 \text{ N}}$$

Tilt of Scalar Spectrum

Data: 0.9666+- 0.0062

$$ln[19] = n[N] = 1 - 2 \epsilon_v - \eta_v$$

Out[19]=
$$1 - \frac{3}{2 \text{ N}}$$

Out[20]= 0.0084

$$ln[21] = Abs[n[50] - 0.9666]$$

Out[21]= 0.0034

Tensor to Scalar Ratio

Data: r<0.168 (Planck TT + Low P)

-Euler-Mascheroni Constant

$$ln[22]:= \gamma_E := 0.5772$$

$$ln[23] = C_E = 4 (Log[2] + \gamma_E) - 5$$

Out[23]= 0.0813887

-Tensor to Scalar Ratio

$$\ln[24] = r[N_{-}] = 16 * \epsilon_{v} * \frac{(1 - (C_{E} + 1) \epsilon_{v})^{2}}{(1 - (3 * C_{E} + 1) \epsilon_{v} + C_{E} * \eta_{v})^{2}}$$

$$\text{Out[24]=} \ \ \frac{8 \, \left(1 - \frac{0.540694}{N}\right)^2}{\left(1 - \frac{0.581389}{N}\right)^2 \, N}$$

In[25]:= **r[60]**

Out[25]= 0.133516

ln[26] = r[50]

Out[26]= 0.160264

Plot

 $\label{eq:local_local_local_local_local_local} $$ \ln[27]:=$ P1 = ParametricPlot[\{n[N], r[N]\}, \{N, 50, 60\}, PlotRange \rightarrow \{\{.92, 1\}, \{0, 0.2\}\}] $$ $$ \left(\frac{1}{2} + \frac{1}{2} + \frac{1}{2$ 0.15 Out[27]= 0.10 0.05 0.92 0.93 0.94 0.95 0.96 0.97 0.98 0.99 1.00