

# SQuIDS: A Tool to Solve Time Evolution in finite dimensional (open) Quantum Systems

An Application to Neutrino Oscillations

arxiv:1412.3832

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#### Outline:

- 1. Introduction (Quantum Evolution with Density Matrices)
- 2. SQuIDS (Overview and Exercises)
  - 2.1 The Const class (Overview + Exercise)
  - 2.2 The SU\_vector class (Overview + Exercise)
  - 2.3 The SQuIDS class (Overview + Exercise)

#### Motivation

Task: Solve time evolution of finite dimensional quantum (sub-)systems:

- Flavor oscillations
- Quantum computation
- Systems with finitely many energy levels
- Spins

Time evolution of closed quantum system: Schrödinger equation

$$i\frac{\mathrm{d}}{\mathrm{d}t}|\psi\rangle = \hat{H}|\psi\rangle \qquad (\hbar = 1)$$
 (1)



# Density matrices instead of state vectors

Often: finite dimensional system S coupled to a complicated (but uninteresting) environment E

- → Get rid of Environment (keyword: partial trace)
- → Just consider degrees of freedom of interest

#### Consequence: Decoherence

- ightharpoonup Subsystem cannot be described by pure state  $|\psi\rangle$
- Mixed state: Described by density matrix  $\varrho = \sum_i p_i |\psi_i\rangle \langle \psi_i|$



## Example: Neutrino oscillations in matter

$\mathcal{S}\simeq\mathbb{C}^3$	E
flavor degrees of freedom	all remaining d.o.f (momenta, spins,)
$ \psi angle = \sum_{lpha=\mathbf{e}}^{ au} \psi_{lpha}   u_{lpha} angle$	ightarrow infinite dimensional

- ▶ We are not at all interested in E
- Only the  $\nu$  flavor composition is interesting to us
- **But**: E significantly influences flavor d.o.f.
- ⇒ Need effective description!



# Time evolution of the density matrix

Master equation(s): (multiple density matrices possible)

$$\frac{\mathrm{d}\varrho_j}{\mathrm{d}t} = -i[\hat{H}_j(t), \varrho_j(t)]$$

- ▶ Why multiple  $\varrho_i$ ? E.g.: One per energy bin!
- $\hat{H} = \hat{H}_0 + \hat{H}_1(t)$ : Unitary evolution



# Time evolution of the density matrix

Master equation(s): (multiple density matrices possible)

$$\frac{\mathrm{d}\varrho_j}{\mathrm{d}t} = -i[\hat{H}_j(t), \varrho_j(t)] + \{\Gamma_j(t), \varrho_j(t)\}$$

- ▶ Why multiple  $\varrho_i$ ? E.g.: One per energy bin!
- $\hat{H} = \hat{H}_0 + \hat{H}_1(t)$ : Unitary evolution
- ► Γ: Decoherence, Attenuation



# Time evolution of the density matrix

Master equation(s): (multiple density matrices possible)

$$\frac{\mathrm{d}\varrho_j}{\mathrm{d}t} = -i[\hat{H}_j(t),\varrho_j(t)] + \{\Gamma_j(t),\varrho_j(t)\} + F_j[\{\varrho_k\}_k,t]$$

- ▶ Why multiple  $\varrho_i$ ? E.g.: One per energy bin!
- $\hat{H} = \hat{H}_0 + \hat{H}_1(t)$ : Unitary evolution
- Γ: Decoherence, Attenuation
- ▶ F: Other non-linear effects (coupling between  $\varrho_i$ )



## Simple Example: $\nu$ Oscillations in Vacuum

Neutrino Experiment:

- Fixed baseline L
- $\triangleright$  *N* energy bins  $\{E_j\}_j$
- $\blacktriangleright \hat{H}^j = \hat{H}_0^j = E_j \cdot \mathbb{I} + \frac{1}{2E_i} \mathbb{M}^2$
- Γ ≡ 0
- $F \equiv 0$
- $ightharpoonup arrho_j(t) = \sum_{lpha,eta=e}^{ au} \phi_{lphaeta}^j(t) |
  u_lpha
  angle \langle
  u_eta|$



## Simple Example: $\nu$ Oscillations in Vacuum

#### Neutrino Experiment:

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- $\hat{\mathcal{H}}^j = \hat{\mathcal{H}}_0^j = E_j \cdot \mathbb{I} + \frac{1}{2E_i} \mathbb{M}^2$
- $\Gamma \equiv 0$
- $F \equiv 0$
- $\triangleright \ \varrho_i(t) = \sum_{\alpha \beta = e}^{\tau} \phi_{\alpha\beta}^j(t) |\nu_{\alpha}\rangle\langle\nu_{\beta}|$

$$egin{aligned} \mathbb{M} &= \sum_{j,k=1}^{3} (\mathbb{M}_0)_{jk} |
u_j
angle \langle
u_k| \ &= \sum_{lpha,eta=e}^{ au} (\mathbb{M}_1)_{lphaeta} |
u_lpha
angle \langle
u_eta| \ &= \sum_{lpha,eta=e}^{ au} (\mathbb{M}_1)_{lphaeta} |
u_lpha
angle \langle
u_eta| \ &= \left(\begin{matrix} m_1 & 0 & 0 \\ 0 & m_2 & 0 \\ 0 & 0 & m_3 \end{matrix}
ight) \ &= U_{\mathrm{PMNS}}^{\dagger} \mathbb{M}_0 U_{\mathrm{PMNS}} \end{aligned}$$

$$\mathbb{M}_1 = U_{\mathrm{PMNS}}^\dagger \mathbb{M}_0 U_{\mathrm{PMNS}}$$

### Some Important Considerations

The master equation simplifies to

$$\frac{\mathrm{d}\varrho_j}{\mathrm{d}t} = -i[\hat{H}_0^j,\varrho_j(t)]$$

Further simplifications

- ► Consider in mass basis:  $\hat{H}_0^j$  is diagonal
- Depends only on commutator!

$$[A, \mathbb{I}] = 0 \quad \Rightarrow \quad [\hat{H}_0^j, \varrho_j(t)] = [\hat{H}_0^j - \epsilon_j \mathbb{I}, \varrho_j(t)]$$

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$$\begin{split} [A,\mathbb{I}] &= 0 \quad \Rightarrow \quad [\hat{H}_0^j,\varrho_j(t)] = [\hat{H}_0^j - \epsilon_j \mathbb{I},\varrho_j(t)] \\ \epsilon_j &:= \langle \nu_1 | \hat{H}_0^j | \nu_1 \rangle = E_j + \frac{m_1^2}{2E_i} \quad \text{(Ground state energy)} \end{split}$$



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u_1 | \hat{H}_0^j | 
u_1 
angle = E_j + rac{m_1^2}{2E_j} & ext{(Ground state energy)} \ &\Rightarrow \tilde{H}^j = rac{1}{2E_j} egin{pmatrix} 0 & 0 & 0 \ 0 & \Delta m_{21}^2 & 0 \ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} \end{aligned}$$



# Some Important Considerations (ctd.)

Can solve  $H_0$  evolution analytically!

► Switch to interaction picture:

$$\begin{split} \tilde{\varrho}(t) &:= \exp(iH_0t)\varrho \exp(-iH_0t) \\ \Rightarrow \dot{\varrho} &= -i[H_0, \rho] + \exp(-iH_0t)\dot{\tilde{\varrho}} \exp(iH_0t) \end{split}$$

- ► Can subtract  $-i[H_0, \varrho]$  on both sides of master equation
- Must transform all terms to interaction picture (SQuIDS does that automatically and efficiently)

# Some Important Considerations (ctd.)

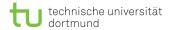
All matrices in our system are hermitian:  $A^{\dagger} = A$ 

- ▶ Hermitian  $n \times n$  matrices form  $N = n^2$  dimenional real vector space
- ► Convenient basis: SU(n) generators  $\sigma_i$  (e.g. n=2: Pauli matrices + identity)
- ▶ Decompose:  $\varrho = \sum_{i=0}^{n^2-1} \rho_i \cdot \sigma_i$
- ▶ Components  $\rho_i$  form  $n^2$  dimensional vector called SU\_vector in the following

# Summary

What did we learn so far (in general):

- 1. We passed to density matrix formulation (allows for mixed states)
- 2. Formulated master equation
- 3. Can subtract  $\epsilon_0 \cdot \mathbb{I}$  from  $\hat{H}$  (only energy diff. important)
- 4. Can solve  $\hat{H}_0$  exactly (interaction picture)  $\varrho \to e^{i\hat{H}_0t}\varrho e^{-i\hat{H}_0t}$
- 5. Can represent  $\varrho, H, \ldots$  as  $n^2$  dimensional, real vector (SU\_vector)!
  - → Efficient and preserves hermiticity automatically!



# **SQuIDS**

# Clone the Repo!

Git Repository includes all needed files (slides, code templates)

#### Instructions

cd to the location where you want to place the repo
git clone https://github.com/BObsen/sm\_to\_bsm\_neutrino.git
cd sm\_to\_bsm\_neutrino



#### Overview

SQuIDS mainly consists out of 3 interconnected classes:

- 1. squids::Const
  - ► Implements all sorts of constants of nature
  - ► Conversion between natural units and other unit systems
  - ▶ Stores system parameters (mixing angles, energy differences, ...)



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- 2. squids::SU\_vector
  - Represents hermitian matrices efficiently
  - ► Implements all sorts of operations on them (Unitary transformations, trace, commutator, . . . )



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- 2. squids::SU\_vector
  - Represents hermitian matrices efficiently
  - ► Implements all sorts of operations on them (Unitary transformations, trace, commutator, ...)
- 3. squids::SQuIDS
  - ► Abstract base class, uses squids::Const and squids::SU\_vector
  - ▶ Implements time evolution of the system of density matrices
  - ▶ Includes methods for taking expectation values etc



# SQuIDS - The Const Class



# Const - Construction of objects

Can only be default constructed:

```
squids::Const units;
```

- Constructs squids::Const object called units
- ► This object contains
  - ▶ Different physical constants  $(G_F, N_A, G, m_p, ...)$
  - ► Values of km, s, J, kg, etc. in natural units
  - Yet unspecified values for:
    - $\blacktriangleright$  basis change from  $B_0$  to  $B_1$  (e.g. mass and flavor basis)
    - energy differences which can be used for the hamiltonian  $\hat{H}_0$



#### Const - Unit conversion

Easily convert between SI and natural units:

```
squids::Const units; double L = 10 * units.cm; \\ 506773 eV^{-1} double T = 1 * units.year; \\ 4.79116e+22 eV^{-1} double GF = units.GF * (units.GeV * units.GeV); \\ 1.16638e-05 \text{ GeV}^{-2}
```

```
The units are to be read as [unit we want] / [eV^{\alpha}], e.g.: km / [eV^{-1}]: [Value in eV^{-1}] = [Value in km] \cdot km / [eV^{-1}] [Value in km] = L / km
```



# Const - Setting / Getting Mixing Angles

Furthermore you can store system parameters:

- ► Substitute Set for Get: returns corresponding value
- ▶ Energy differences: Only convenience parameters simplifying definition of  $\hat{H}_0$



# SQuIDS - The Const Class: Exercise

#### Const class exercise

- 1. Declare a default constructed const class object
- 2. Answer the following questions:
  - 2.1 How many  $eV^{-1}$  correspond to 300 km
  - 2.2 How many radians correspond to 25°
  - 2.3 If you are 24 years old, how many  $eV^{-1}$  are you old?
- 3. Set the mixing parameters for three neutrino generations to:
  - $\theta_{12} = 33.48^{\circ}$
  - $\theta_{13} = 8.55^{\circ}$
  - $\theta_{23} = 42.3^{\circ}$
- 4. Set the energy differences to:
  - $\Delta m_{21}^2 = 7.5 \cdot 10^{-5} \,\mathrm{eV}^2$
  - $\Delta m_{31}^{21} = 2.45 \cdot 10^{-3} \, \text{eV}^2$



# SQuIDS - The $SU_vector\ Class$

From SM to BSM - 2023

#### SU Vector - Contructors

#### Construct a SU Vector:

► Default: No entries, no size

```
squids::SU_vector rho;
```

▶ Defined Dimension: entries are zero, size is  $dim(\mathcal{H}) > 0$ 

And more: construct from array, gsl\_matrix\_complex, std::vector

. . .



### SU Vector - Special Matrices

Construct a SU Vector to Standard Form:

• Identity: Corresponding matrix is the identity  $(Id)_{ij} = \delta_{ij}$ 

```
squids::SU_vector Id
= squids::SU_vector::Identity(dim);
```

▶ Projector on state  $k_0$ : only the  $(k_0, k_0)$  element is 1, i.e.

$$(\mathbb{P}_{k_0})_{ij} = \delta_{ik_0}\delta_{jk_0}$$

```
squids::SU_vector P_k0
= squids::SU_vector::Projector(dim, k0);
```

And more: See documentation (also in the repo)



#### SU Vector - Functions

Manipulate a SU Vector:

▶ Rotate from one basis to another  $(B_0 \leftrightarrow B_1)$ : (using angles specfied in params)

```
rho.RotateToB1(params); \\B0 to B1
rho.RotateToB0(params); \\B1 to B0
```

► Return dim of Hilbert space, Size of vector:

```
rho.Dim(); \\returns dim(H)
rho.Size(); \\returns number of elements
```

Furthermore: You can add, subtract multiply (also with scalars), assign, access elements via rho[] etc.



# SQuIDS - The SU\_vector Class: Exercise

#### SU vector exercise

- 1. Declare an empty SU vector corresponding to a 3D Hilbert space
- 2. Initialize an array of projectors for the three mass eigenstates  $(B_0)$
- 3. Rotate them to the flavor basis  $(B_1)$
- 4. Initialize a SU vector corresponding to the matrix  $(B_0)$

$$\Delta \mathbb{M}^2 := egin{pmatrix} 0 & 0 & 0 \ 0 & \Delta m_{21}^2 & 0 \ 0 & 0 & m_{31}^2 \end{pmatrix}$$
 (2)



# SQuIDS - The SQuIDS Class



# SQuIDS - General Overview pt. I

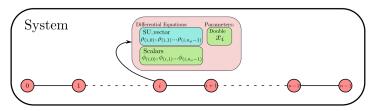
squids::SQuIDS is a so called base class:

- ► You derive your own class from it
- ▶ It provides fundamental functionality through member functions
  - ► Integration of the master equation
  - ▶ Efficient memory management and storage of density matrices
- ► Already includes several members storing basic data of your system
- Virtual member functions:
  - ➤ Some you have to define for the class to work (HO, constructors, init, ...)
  - ► Some you can ignore if your system doesn't need them



# SQuIDS - General Overview pt. II

Distribution of density matrices across the nodes [arxiv:1412.3832]:



- $\triangleright$  n nodes  $x_i$  (energy bins, angles, whatever feature of your system)
- At each node:
  - nrho density matrices
  - ns scalars (not important for us)
- → Specify initial data, HO, HI, GammaRho, etc.
- $\Rightarrow$  SQuIDS evolves the whole system (+ can take expectation values of observables)



#### SQuIDS - Constructors

Constructors / Initializors: (Set bkg. params and allocate memory)

```
SQuDIS(); \\ default
SQuDIS(uint nx, uint dim, unit nrho, uint nscalar,
double ti = 0);
void ini(uint nx, uint dim, unit nrho, uint nscalar,
double ti = 0);
```

- ▶ nx: Number of x nodes x\_i
- dim: Dimensions of Hilbert space
- nrho / nscalar: # density matrices / scalars per node
- ti: initial time, defaults to zero

Call them in your own constructor with the system parameters you need!

#### SQuIDS - Member variables

SQuIDS is set up to include the following member variables:

- ▶ std::vector<double> x: x range
- ▶ uint nsun: Dim. of Hilbertspace
- Const params: squids::Const object containing system parameters
- And many more!

You can (and should) access these from within your own member functions!



```
int Set_xrange(double xini, double xend, std::string
scale); \\ Sets x = {xini, ..., xend} with lin or log
scale
```



```
int Set_xrange(double xini, double xend, std::string
scale); \\ Sets x = {xini, ..., xend} with lin or log
scale
double Get_x(uint i) const; \\ Returns x[i]
```



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double Get_x(uint i) const; \\ Returns x[i]
const * Const Get_params() const; \\ Returns Const member
```



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```



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int Evolve(double dt); \\ Evolves System by dt
double GetExpectationValue(SU_vector op, uint irho, uint
ix) const; \\ Calculates exp. val. of op at node ix with
density matrix irho at current time
```



Some very useful predefined member functions are:

```
int Set_xrange(double xini, double xend, std::string
scale); \\ Sets x = {xini, ..., xend} with lin or log
scale
double Get_x(uint i) const; \\ Returns x[i]
const * Const Get_params() const; \\ Returns Const member
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double GetExpectationValue(SU_vector op, uint irho, uint
ix) const; \\ Calculates exp. val. of op at node ix with
density matrix irho at current time
```

Of course there are more but these are most important for us!



#### SQuIDS - Evolution functions

Last but not least: The time independent Hamiltonian HO

```
SU_vector HO(double x, uint irho) const;
```

- ► Returns  $\hat{H}_0$  as SU\_vector object
- Assumes that  $\hat{H}_0$  diagonal (i.e. given in mass basis B0 for  $\nu$  oscillations)
- Cannot modify but can read member variables (const)
- ▶ irho: HO for density matrix at node x<sub>i</sub>



# SQuIDS - The SQuIDS Class: Exercise



## SQuIDS application: Neutrino oscillations in vacuum

General set up: Vacuum Oscillations Experiment with . . .

- ▶ Fixed baseline L ( $\hat{=}$  time variable)
- ▶ n logarithmic energy bins:  $E \in [10 \,\mathrm{MeV}, 10 \,\mathrm{GeV}]$
- ► All neutrinos are produced as electron neutrinos

Use: x nodes as energy nodes, one density matrix per node, no scalar functions, only H0 non-zero



# SQuIDS application: Neutrino oscillations in vacuum

- 1. Declare vacuum class publically derived from squids::SQuIDS class in vac/src/vacuum.hpp with methods
  - Default constructor
  - ► Initializing constructor
  - ► HO
  - ► GetProbabilities
  - Destructor if needed
- Define the corresponding member functions in vac/src/vacuum.cpp
  - ▶ nbins *x*-nodes corresponding to number of *E* bins (log scaled)
  - ▶ One density function per node, no scalar functions
  - nflavor neutrino flavors
  - ▶ Initial condition  $\rho_i(t=0) = \mathbb{P}_e$  for all  $i \in \{1, ..., nbins\}$
  - ► Work in mass basis!!!

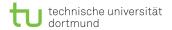
#### **SQuIDS**

# SQuIDS application: Neutrino oscillations in vacuum

- 3. Initialize an object from your class in vac/src/main.cpp (nbins = 1000, nflavor = 3,  $E \in [10 \,\mathrm{MeV}, 10 \,\mathrm{GeV}]$ )
- 4. Evolve the system for  $L=300~\mathrm{km}$
- 5. Save the oscillation probabilities  $P_{e\alpha}(E_i, L)$  to file(s)  $\alpha = e, \mu, \tau$
- 6. Plot them against the energy

#### Outlook

- ▶ So far we only scratched the surface of SQuIDS' abilities
- ▶ Setting  $\hat{H}_1(t, \varrho)$ ,  $\Gamma(t, \varrho)$ ,  $F(t, \varrho) \neq 0$  opens up possibilities to include decoherence, interactions, etc.
- ► These extra terms are solved numerically (using GSL)
- Examples for neutrinos:
  - ► Wave packet decoherence
  - ► Neutrino propagation through the sun / earth
  - ► Collective neutrino oscillations (early universe, supernovae, ...)
  - ► Active sterile oscillations by using more than three generations



# **BACK UP**



#### Installation

What do we need for this tutorial?

- ► A unix-like (sub-)system
  - ► Linux
  - ► Mac (+ Xcode developer tools!)
  - On Windows: WSL
- ► A C++ compiler
- ► Make, wget, Git

Use scripts install\_gsl.sh and install\_SQuIDS.sh from the repo!



# Installation (GSL)

```
cd $HOME
mkdir -p smToBsmLibs/gs1
wget ftp://ftp.gnu.org/gnu/gsl/gsl-latest.tar.gz
tar -zxvf gsl-latest.tar.gz
rm gsl-latest.tar.gz
cd $(find gsl-* | head -n 1)
./configure --prefix=$HOME/smToBsmLibs/gsl
make
make check
make install
LD_LIBRARY_PATH=$HOME/smToBsmLibs/gsl/lib:$LD_LIBRARY_PATH
export LD_LIBRARY_PATH
cd $HOME
rm -rf $(find gsl-* | head -n 1)
```



# Installation (SQuIDS)

```
cd $HOME
mkdir -p smToBsmLibs/SQuIDS
git clone https://github.com/jsalvado/SQuIDS.git
cd $(find SQuIDS* | head -n 1)
./configure --with-gsl-incdir=$HOME/smToBsmLibs/gsl/include \
--with-gsl-libdir=$HOME/smToBsmLibs/gsl/lib \
--prefix=$HOME/smToBsmLibs/SQuIDS
make
make test
make install
LD_LIBRARY_PATH=$HOME/smToBsmLibs/SQuIDS/lib:$LD_LIBRARY_PATH # linux only
export LD_LIBRARY_PATH # linux only
cd $HOME
rm -rf $(find SQuIDS* | head -n 1)
```