

# SQuIDS: A Tool to Solve Time Evolution in finite dimensional (open) Quantum Systems

An Application to Neutrino Oscillations

arxiv:1412.3832

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#### Outline:

- 1. Introduction (Quantum Evolution with Density Matrices)
- 2. SQuIDS (Overview and Exercises)
  - 2.1 The Const class (Overview + Exercise)
  - 2.2 The SU\_vector class (Overview + Exercise)
  - 2.3 The SQuIDS class (Overview + Exercise)

#### Motivation

Task: Solve time evolution of finite dimensional quantum (sub-)systems:

- ► Flavor oscillations
- Quantum computation
- ► Systems with finitely many energy levels
- Spins

Time evolution of closed quantum system: Schrödinger equation

$$i\frac{\partial}{\partial t}|\psi\rangle = \hat{H}|\psi\rangle \qquad (\hbar = 1)$$
 (1)



# Density matrices instead of state vectors

Often: finite dimensional system S coupled to a complicated (but uninteresting) environment E

- → Get rid of Environment (keyword: partial trace)
- → Just consider degrees of freedom of interest

#### Consequence: Decoherence

- $\triangleright$  Subsystem cannot be described by pure state  $|\psi\rangle$
- Mixed state: Described by density matrix  $\varrho = \sum_i p_i |\psi_i\rangle \langle \psi_i|$



# Example: Neutrino oscillations in matter

$\mathcal{S}\simeq\mathbb{C}^3$	E
flavor degrees of freedom	all remaining d.o.f (momenta, spins,)
$ \psi angle = \sum_{lpha=\mathbf{e}}^{ au} \psi_{lpha}   u_{lpha} angle$	ightarrow infinite dimensional

- ▶ We are not at all interested in E
- Only the  $\nu$  flavor composition is interesting to us
- **But**: E significantly influences flavor d.o.f.
- ⇒ Need effective description!



# Time evolution of the density matrix

Master equation(s): (multiple density matrices possible)

$$\frac{\mathrm{d}\varrho_j}{\mathrm{d}t} = -i[\hat{H}_j(t),\varrho_j(t)] + \{\Gamma_j(t),\varrho_j(t)\} + F_j[\{\varrho_k\}_k,t]$$

- ▶ Why multiple  $\varrho_i$ ? E.g.: One per energy bin!
- $\hat{H} = \hat{H}_0 + \hat{H}_1(t)$ : Unitary evolution
- Γ: Decoherence
- ▶ F: Other non-linear effects (coupling between  $\varrho_i$ )



### Simple Example: $\nu$ Oscillations in Vacuum

#### Neutrino Experiment:

- ► Fixed baseline *L*
- $\triangleright$  N energy bins  $\{E_i\}_i$
- $\hat{H}^{j} = \hat{H}_0^{j} = E_j \cdot \mathbb{I} + \frac{1}{2E_i} \mathbb{M}^2$
- Γ ≡ 0
- $F \equiv 0$
- $\blacktriangleright \ \varrho_j(t) = \sum_{\alpha=e}^{\tau} \phi_{\alpha}^j(t) |\nu_{\alpha}\rangle\langle\nu_{\alpha}|$

$$egin{aligned} \mathbb{M} &= \sum_{j,k=1}^{3} (\mathbb{M}_0)_{jk} |
u_j
angle \langle
u_k| \ &= \sum_{lpha,eta=e}^{ au} (\mathbb{M}_1)_{lphaeta} |
u_lpha
angle \langle
u_eta| \ &= \sum_{lpha,eta=e}^{ au} (\mathbb{M}_1)_{lphaeta} |
u_lpha
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u_lpha
angle |
u_0 &= u_3 \ &= u_{
m PMNS}^{\dagger} \mathbb{M}_0 U_{
m PMNS} 
\end{aligned}$$



# Some Important Considerations

The master equation simplifies to

$$\frac{\mathrm{d}\varrho_j}{\mathrm{d}t} = -i[\hat{H}_0^j, \varrho_j(t)]$$

Further simplifications

- ightharpoonup Consider in mass basis:  $\hat{H}_0^j$  is diagonal
- ► Depends only on commutator!
  - $[A, \mathbb{I}] = 0 \Rightarrow [\hat{H}_0^j, \varrho_i(t)] = [\hat{H}_0^j \epsilon_i \mathbb{I}, \varrho_i(t)]$

$$ilde{H}^j = rac{1}{2 E_j} egin{pmatrix} 0 & 0 & 0 & 0 \ 0 & \Delta m_{21}^2 & 0 \ 0 & 0 & \Delta m_{31}^2 \end{pmatrix}$$



# Some Important Considerations (ctd.)

Can solve  $H_0$  evolution analytically!

▶ Pass to interaction picture:

$$\begin{split} \tilde{\varrho}(t) &:= \exp(iH_0t)\varrho \exp(-iH_0t) \\ \Rightarrow \dot{\varrho} &= -i[H_0, \rho] + \exp(-iH_0t)\dot{\tilde{\varrho}} \exp(iH_0t) \end{split}$$

- Can subtract  $-i[H_0, \varrho]$  on both sides of master equation
- ► Must transform all terms to interaction picture (SQuIDS does that automatically and efficiently)

# Some Important Considerations (ctd.)

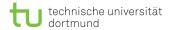
All matrices in our system are hermitian:  $A^{\dagger} = A$ 

- ▶ Hermitian  $n \times n$  matrices form  $N = n^2$  dimenional real vector space
- ► Convenient basis: SU(n) generators  $\sigma_i$  (e.g. n=2: Pauli matrices + identity)
- ▶ Decompose:  $\varrho = \sum_{i=0}^{n^2-1} \rho_i \cdot \sigma_i$
- ▶ Components  $\rho_i$  form  $n^2$  dimensional vector called SU\_vector in the following

# Summary

What did we learn so far (in general):

- 1. We passed to density matrix formulation (allows for mixed states)
- 2. Formulated master equation
- 3. Can subtract  $\epsilon_0 \cdot \mathbb{I}$  from  $\hat{H}$  (only energy diff. important)
- 4. Can solve  $\hat{H}_0$  exactly (interaction picture)  $\varrho \to e^{i\hat{H}_0t}\varrho e^{-i\hat{H}_0t}$
- 5. Can represent  $\varrho, H, \ldots$  as  $n^2$  dimensional, real vector (SU\_vector)!
  - → Efficient and preserves hermiticity automatically!



# **SQuIDS**



# Clone the Repo!

Git Repository includes all needed files (slides, code templates)

#### Instructions

cd to the location where you want to place the repo
git clone https://github.com/BObsen/sm\_to\_bsm\_neutrino.git
cd sm\_to\_bsm\_neutrino



#### Overview

SQuIDS mainly consists out of 3 interconnected classes:

- 1. squids::Const
  - ► Implements all sorts of constants of nature
  - ► Conversion between natural units and other unit systems
  - ▶ Stores system parameters (mixing angles, energy differences, ...)



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- 2. squids::SU\_vector
  - Represents hermitian matrices efficiently
  - ► Implements all sorts of operations on them (Unitary transformations, trace, commutator, . . . )



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- 2. squids::SU\_vector
  - Represents hermitian matrices efficiently
  - ► Implements all sorts of operations on them (Unitary transformations, trace, commutator, ...)
- 3. squids::SQuIDS
  - ► Abstract base class, uses squids::Const and squids::SU\_vector
  - ▶ Implements time evolution of the system of density matrices
  - ▶ Includes methods for taking expectation values etc



# SQuIDS - The Const Class



# Const - Construction of objects

Can only be default constructed:

```
squids::Const units;
```

- Constructs squids::Const object called units
- ► This object contains
  - ▶ Different physical constants  $(G_F, N_A, G, m_p, ...)$
  - ▶ Values of km, s, J, kg, etc. in natural units
  - Yet unspecified values for:
    - **b** basis change from  $B_0$  to  $B_1$  (e.g. mass and flavor basis)
    - energy differences which can be used for the hamiltonian  $\hat{H}_0$



#### Const - Unit conversion

Easily convert between SI and natural units:

```
The units are to be read as [unit we want] / [eV^{\alpha}], e.g.: km / [eV^{-1}]: [Value in eV^{-1}] = [Value in km] \cdot km / [eV^{-1}] [Value in km] = L / km
```



# Const - Setting / Getting Mixing Angles

Furthermore you can store system parameters:

```
\label{eq:squids::Const params;} $$\sup_{\cdot \in \mathcal{B}_{12} = 24^\circ$} $$params.SetMixingAngle(0, 1, 24 * params.degree); $$$ $$\setminus Sets \, \delta_{13} = 2^\circ$ $$params.SetPhase(0, 2, 2 * params.degree); $$$$$$ $$\setminus Sets \, \Delta E_{10} = 7eV$ $$params.SetEnergyDifference(1, 7 * params.eV); $$
```

- ► Substitute Set for Get: returns corresponding value
- ▶ Energy differences: Only convenience parameters simplifying definition of  $\hat{H}_0$



# SQuIDS - The Const Class: Exercise

#### Const class exercise

- 1. Declare a default constructed const class object
- 2. Answer the following questions:
  - 2.1 How many  $eV^{-1}$  correspond to 300 km
  - 2.2 How many radians correspond to 25°
  - 2.3 If you are 24 years old, how many  $eV^{-1}$  are you old?
- 3. Set the mixing parameters for three neutrino generations to:
  - $\theta_{12} = 33.48^{\circ}$
  - $\theta_{13} = 8.55^{\circ}$
  - $\theta_{23} = 42.3^{\circ}$
- 4. Set the energy differences to:
  - $\Delta m_{21}^2 = 7.5 \cdot 10^{-5} \, \text{eV}^2$
  - $\Delta m_{31}^{21} = 2.45 \cdot 10^{-3} \, \text{eV}^2$



# SQuIDS - The $SU_vector\ Class$

From SM to BSM - 2023

#### SU Vector - Contructors

#### Construct a SU Vector:

► Default: No entries, no size

```
squids::SU_vector rho;
```

▶ Defined Dimension: entries are zero, size is  $dim(\mathcal{H}) > 0$ 

And more: construct from array, gsl\_matrix\_complex, std::vector



# SU Vector - Special Matrices

Construct a SU Vector to Standard Form:

▶ Identity: Corresponding matrix is the identity  $(Id)_{ij} = \delta_{ij}$ 

```
squids::SU_vector Id
= squids::SU_vector::Identity(dim);
```

Projector on state  $k_0$ : only the  $(k_0, k_0)$  element is 1, i.e.

$$(\mathbb{P}_{k_0})_{ij} = \delta_{ik_0}\delta_{jk_0}$$

```
squids::SU_vector P_k0
= squids::SU_vector::Projector(dim, k0);
```

And more: See documentation (also in the repo)



#### SU Vector - Functions

Manipulate a SU Vector:

▶ Rotate from one basis to another  $(B_0 \leftrightarrow B_1)$ : (using angles specfied in params)

```
rho.RotateToB1(params); \\B0 to B1
rho.RotateToB0(params); \\B1 to B0
```

▶ Return dim of Hilbert space, Size of vector:

```
rho.Dim(); \\returns dim(H)
rho.Size(); \\returns number of elements
```

Furthermore: You can add, subtract multiply (also with scalars), assign, access elements via rho[] etc.



# SQuIDS - The SU\_vector Class: Exercise

#### SU vector exercise

- 1. Declare an empty SU vector corresponding to a 3D Hilbert space
- 2. Initialize an array of projectors for the three mass eigenstates  $(B_0)$
- 3. Rotate them to the flavor basis  $(B_1)$
- 4. Initialize a SU vector corresponding to the matrix  $(B_0)$

$$\Delta \mathbb{M}^2 := egin{pmatrix} 0 & 0 & 0 \ 0 & \Delta m_{21}^2 & 0 \ 0 & 0 & m_{31}^2 \end{pmatrix}$$
 (2)



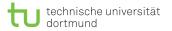
# SQuIDS - The SQuIDS Class



# SQuIDS - General Overview pt. I

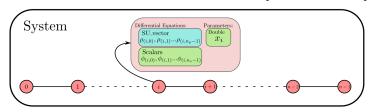
squids::SQuIDS is a so called base class:

- ► You derive your own class from it
- ▶ It provides fundamental functionality through member functions
  - ► Integration of the master equation
  - ▶ Efficient memory management and storage of density matrices
- ► Already includes several members storing basic data of your system
- Virtual member functions:
  - ➤ Some you have to define for the class to work (HO, constructors, init, ...)
  - ▶ Some you can ignore if your system doesn't need them



# SQuIDS - General Overview pt. II

Distribution of density matrices across the nodes [arxiv:1412.3832]:



- $\triangleright$  n nodes  $x_i$  (energy bins, angles, whatever feature of your system)
- At each node:
  - nrho density matrices
  - ▶ ns scalars (not important for us)
- → Specify initial data, HO, HI, GammaRho, etc.
- $\Rightarrow$  SQuIDS evolves the whole system (+ can take expectation values of observables)



### SQuIDS - Constructors

Constructors / Initializors: (Set bkg. params and allocate memory)

```
SQuDIS(); \\ default
SQuDIS(uint nx, uint dim, unit nrho, uint nscalar,
double ti = 0);
void ini(uint nx, uint dim, unit nrho, uint nscalar,
double ti = 0);
```

- nx: Number of x nodes x\_i
- dim: Dimensions of Hilbert space
- nrho / nscalar: # density matrices / scalars per node
- ti: initial time, defaults to zero

Call them in your own constructor with the system parameters you need!



# SQuIDS - Member variables

SQuIDS is set up to include the following member variables:

- ▶ std::vector<double> x: x range
- ▶ uint nsun: Dim. of Hilbertspace
- Const params: squids::Const object containing system parameters
- And many more!

You can (and should) access these from within your own member functions!



# SQuIDS - Handy functions

Some very useful predefined member functions are:

```
int Set_xrange(double xini, double xend, std::string
scale); \\ Sets x = {xini, ..., xend} with lin or log
scale
```



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double Get_x(uint i) const; \\ Returns x[i]
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ix) const; \\ Calculates exp. val. of op at node ix with
density matrix irho at current time
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```

Of course there are more but these are most important for us!



### SQuIDS - Evolution functions

Last but not least: The time independent Hamiltonian HO

```
SU_vector HO(double x, uint irho) const;
```

- ► Returns  $\hat{H}_0$  as SU\_vector object
- Assumes that  $\hat{H}_0$  diagonal (i.e. given in mass basis B0 for  $\nu$  oscillations)
- ► Cannot modify but can read member variables (const)
- ▶ irho: HO for density matrix at node x<sub>i</sub>



# SQuIDS - The SQuIDS Class: Exercise



## SQuIDS application: Neutrino oscillations in vacuum

General set up: Vacuum Oscillations Experiment with . . .

- ▶ Fixed baseline L ( $\hat{=}$  time variable)
- ▶ n logarithmic energy bins:  $E \in [10 \,\mathrm{MeV}, 10 \,\mathrm{GeV}]$
- ► All neutrinos are produced as electron neutrinos

Use: x nodes as energy nodes, one density matrix per node, no scalar functions, only H0 non-zero



## SQuIDS application: Neutrino oscillations in vacuum

- 1. Declare vacuum class publically derived from squids::SQuIDS class in vac/src/vacuum.hpp with methods
  - Default constructor
  - ► Initializing constructor
  - ► HO
  - ► GetProbabilities
  - Destructor if needed
- Define the corresponding member functions in vac/src/vacuum.cpp
  - ▶ nbins *x*-nodes corresponding to number of *E* bins (log scaled)
  - ▶ One density function per node, no scalar functions
  - nflavor neutrino flavors
  - ▶ Initial condition  $\rho_i(t=0) = \mathbb{P}_e$  for all  $i \in \{1, ..., nbins\}$
  - ► Work in mass basis!!!

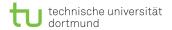
#### **SQuIDS**

## SQuIDS application: Neutrino oscillations in vacuum

- 3. Initialize an object from your class in vac/src/main.cpp (nbins = 1000, nflavor = 3,  $E \in [10 \,\mathrm{MeV}, 10 \,\mathrm{GeV}]$ )
- 4. Evolve the system for  $L=300~\mathrm{km}$
- 5. Save the oscillation probabilities  $P_{e\alpha}(E_i, L)$  to file(s)  $\alpha = e, \mu, \tau$
- 6. Plot them against the energy

#### Outlook

- ► So far we only scratched the surface of SQuIDS' abilities
- ▶ Including  $\hat{H}_1(t)$ ,  $\Gamma$ ,  $F \neq 0$  opens up possibilities to solve a vast range of systems



# **BACK UP**



#### Installation

What do we need for this tutorial?

- ► A unix-like (sub-)system
  - ► Linux
  - ► Mac (+ Xcode developer tools!)
  - On Windows: WSL
- ► A C++ compiler
- ► Make, wget, Git

Use scripts install\_gsl.sh and install\_SQuIDS.sh from the repo!



## Installation (GSL)

```
cd $HOME
mkdir -p smToBsmLibs/gsl
wget ftp://ftp.gnu.org/gnu/gsl/gsl-latest.tar.gz
tar -zxvf gsl-latest.tar.gz
rm gsl-latest.tar.gz
cd $(find gsl-* | head -n 1)
./configure --prefix=$HOME/smToBsmLibs/gsl
make
make check
make install
LD_LIBRARY_PATH=$HOME/smToBsmLibs/gsl/lib:$LD_LIBRARY_PATH
export LD_LIBRARY_PATH
cd $HOME
rm -rf $(find gsl-* | head -n 1)
```



## Installation (SQuIDS)

```
cd $HOME
mkdir -p smToBsmLibs/SQuIDS
git clone https://github.com/jsalvado/SQuIDS.git
cd $(find SQuIDS* | head -n 1)
./configure --with-gsl-incdir=$HOME/smToBsmLibs/gsl/include \
--with-gsl-libdir=$HOME/smToBsmLibs/gsl/lib \
--prefix=$HOME/smToBsmLibs/SQuIDS
make
make test
make install
LD_LIBRARY_PATH=$HOME/smToBsmLibs/SQuIDS/lib:$LD_LIBRARY_PATH # linux only
export LD_LIBRARY_PATH # linux only
cd $HOME
rm -rf $(find SQuIDS* | head -n 1)
```