

**Intensive photometry of southern Be variables. I. Winter objects (\*)**J. Cuypers <sup>(1)</sup>, L. A. Balona <sup>(2)</sup> and F. Marang <sup>(2)</sup><sup>(1)</sup> Koninklijke Sterrenwacht van België, Ringlaan 3, B-1180 Brussel, Belgium<sup>(2)</sup> South African Astronomical Observatory, P.O. Box 9, Observatory 7935, Cape, South Africa*Received May 29, accepted July 20, 1989*

**Summary.** — We present results of an intensive photometric campaign on some bright southern Be stars to search for periodic light variations. In order to obtain good phase coverage, observations were conducted from two sites with different longitude : ESO and SAAO. Most of the stars observed are indeed variable with periods close to one day (the expected rotational period for these stars). We present our results for winter objects.

**Key words :** Be stars — variable stars — photometry — stellar pulsations.

**1. Introduction.**

Periodic variability in Be stars has attracted great interest in recent years. The discovery of unexpected short-period variations in these stars has renewed the search for a viable mechanism of mass ejection. The current consensus of opinion is that non-radial oscillations, driven by an unknown mechanism, is the likely cause of the mass loss leading to the Be phenomenon (see Percy, 1987, for a recent review). This conclusion is mostly based on line-profile variations seen in some Be stars.

The small-scale “moving bumps” seen in the line profiles of some Be stars is attributed to high-order non-radial pulsations (NRP). But moving bumps are also seen in B stars without emission lines and even in  $\delta$  Scuti stars (Walker and Yang, 1987). It seems that NRP is not solely responsible for the mass ejection. It is also known that one of the distinguishing features of line-profile variations in Be stars is the presence of low-order line-profile variations which some attribute to low-order NRP (Penrod, 1987). But again, low-order NRP is present in non-Be stars as well ( $\beta$  Cep and  $\delta$  Sct). Many  $\beta$  Cep variables are rotating more rapidly than some Be stars. High rotation coupled with NRP does not seem to be a sufficient condition for the Be phenomenon. The information from photometric observations has not been adequately considered and it is possible that the conclusion that NRP is responsible for the mass loss could be premature.

A viable hypothesis should be able to explain the distribution of observed periods, the fraction of Be stars which are periodic, their photometric amplitudes and evolution of the shapes of their light curves. Such information can only be obtained by intensive photometric monitoring of a large sample of Be stars. Owing to the fact that periods are close to one day, it is very important to obtain observations from different longitudes. These observations should be obtained over a short time as the light curve can evolve quite rapidly. These demands can only be met by an international campaign.

In this paper we present results of such a campaign conducted from ESO and SAAO as well as some observations obtained at SAAO only. We observed 17 bright southern Be stars ; most of them turned out to be short-period variables. These stars were chosen from the *Bright Star* catalogue on the sole criterion that they should be well placed for observing. A list of candidate and comparison stars is given in table I. Results for variable program stars are given in tables IIa and IIb.

**2. Observations and reductions.**

The observations at SAAO were obtained with the Volks photometer attached to the 0.5-m reflector at Sutherland. These were made mostly through a single filter, Strömgren  $b$ , which was chosen as a compromise between the expected larger amplitude towards shorter wavelength and the

(\*) Observations collected at the European Southern Observatory and the South African Astronomical Observatory.  
Send offprint requests to : L.A. Balona.

decreased sensitivity to atmospheric extinction at longer wavelengths. Some four-colour observations were also made. For the brightest stars, a neutral density filter had to be used to prevent over-illumination of the photocathode. In these circumstances we calibrated the ND filter by observing the comparison stars with and without the filter. The rms deviation suggests that this calibration is good to within 3 millimag (the expected observational error of one measurement). Systematic errors due to the use of the neutral density filters are probably negligible owing to the narrow bandpass of Strömgren *b*.

In the reduction procedure, mean extinction coefficients were obtained for each night using the comparison stars. Corrections for transparency changes during the night were made by calculating the deviation for each comparison star. Stars in close proximity were treated as a group and mean transparency corrections were applied to the group, which included the Be stars. Comparison stars which were variable were detected at this stage and could be eliminated. We searched for micro-variability in the comparison stars by calculating their periodograms ; again such stars could be identified and eliminated. Finally, the mean magnitude of all comparison stars was adjusted to be in exact agreement with the mean *b*-magnitude for the same stars observed at ESO. This zero-point correction ensured that the Be stars observed from the two sites were on the same system.

The observations at ESO were made between 13 June and 6 July 1987. They were obtained with the six-channel *uvbyH $\beta$*  photometer on the Danish 0.5-m telescope. This instrument uses a grating and slots to define the passbands ; in consequence some systematic differences could be expected between the *b* magnitudes of the two observatories. In practice, these differences were very small (a few millimagnitudes) and posed no problem when combining the data sets. Because a new telescope control system with autocentering and automated observations had just been installed and still had to be checked during the observing run, the quality and quantity of the observations was not as high as could be expected at the start of the run. Later on, most of the problems were solved and nowadays the new, automated system has proved to be very reliable.

For the brightest stars, the same procedure was used as at SAAO. The reduction procedure was similar except that a polynomial model for transparency changes and zero-point drift was calculated for each night. Furthermore, a general least-squares solution which included all observations for all comparison stars provided best estimates for their mean magnitudes. Variable stars among them were easily identified by the large standard deviations. They were not included in obtaining the final magnitudes for the comparison stars. A small zero-point difference between these results and the standard *uvby* system is still possible, but this is unimportant for the study of their light variations. The rms scatter for the comparison stars is in all cases less than 5 millimag.

### 3. Period finding.

We employed two methods of searching for periodicities. One is the standard Fourier periodogram analysis technique for unequally spaced data ; the other is a modification of Stellingwerf's (1978) phase-dispersion minimization (PDM) technique. The first method is suited to the case where the underlying variation is approximately sinusoidal, but it can give misleading results when the variations are highly non-sinusoidal or when a multiple-wave light curve is present. It is known that many Be stars have a double-wave light curve which will not be detected by this method. Clearly it will be impossible to discriminate between a single-wave light curve and a double-wave light curve with nearly equal amplitudes.

The PDM technique will identify the period for all kinds of light curves, but since the general noise level is rather high in this method (Swingler, 1989), one has to be careful of spurious periodicities. Our final choice was based on a careful examination of the phase diagram resulting from the periods produced by the Fourier and PDM techniques. Our criterion was by nature rather subjective, but we attempted to choose the period which resulted in the curve with the least scatter. Except for one or two cases, there is little doubt that our choice is the correct one.

Besides the complication discussed above, a serious difficulty arises when the irregular light variations which are sometimes present in Be stars (attributed to effects of the circumstellar material) has a time scale comparable to the underlying periodic variation. In these cases it is not possible to obtain the correct period with any certainty. As some long-term activity is nearly always present, we normally removed the long-term trend by fitting a low-order polynomial to the data. This of course has the effect of reducing the scatter at the expense of some uncertainty in the shape and amplitude of the light curve. The derived period will in general be unaffected by this procedure.

It is difficult to derive reliable standard errors for the periods, but an estimate can be obtained from the half-width of the Fourier periodogram peaks. Typically, we observed for about four weeks in each season which leads to a standard deviation of about  $0.02 \text{ d}^{-1}$  in frequency which can be taken as a reliable guide to the error in the quoted frequencies. Sometimes we have combined many seasons in order to obtain a more accurate period. This is always possible to do, but invariably there will be many possible choices of frequency (each separated by approximately one cycle per year) which will produce as good a fit. In this case the real error is not reduced by combining the data.

### 4. The results.

#### *v Cen*

This star is a single-line spectroscopic binary (Palmer, 1906). Wilson (1914) determined the orbital elements and obtained a circular orbit with  $P = 2.62516 \text{ d}$ . Hendry and Bahng (1981) found H $\alpha$  to have double emission

surrounding an absorption core. This observation indicates that  $\nu$  Cen could be classified as a Be star, but it is probable that the emission originates as a result of binary interaction rather than an intrinsic property of the star itself. Spectroscopic observations suggest that this star may be a  $\beta$  Cep variable with a period of about 0.17 d and an amplitude of 5 km s $^{-1}$  (Rajamohan, 1977 ; Kubiak and Seggewiss, 1982 ; Ashoka *et al.*, 1985). Waelkens and Rufener (1983) do not find any evidence for short period light variability, except for a variation with the same period as the binary period.

Our photometric observations (SAAO only) extend over two seasons, 1987 and 1988. The 1988 observations are particularly numerous and clearly show a variation with the same period as the orbital period (Fig. 1). Minimum light corresponds to maximum positive radial velocity. As discussed by Waelkens and Rufener (1983), the resulting light curve is most probably a reflection effect. A careful investigation did not reveal any signs of a variation at any other frequency with an amplitude exceeding 2 millimags.

#### $\mu$ Cen

This star is known as a "pole-on" Be star ( $v$  in  $i = 155$  km s $^{-1}$ ). Ghosh *et al.* (1987) have summarized the history of H $\alpha$  emission ; they reported an outburst in March/April 1987, about three months before the first set of our observations. A full discussion of an earlier outburst (February 1987) is given in Baade *et al.* (1988). Thimm (private communication) reported from spectra taken on May 29, 30 and on June 1, 1987 that H $\alpha$  changed from pure absorption to emission in 3 days. This was only a few days before the start of our observations. All indications are that this Be Star was in a very active phase during this time.

$\mu$  Cen was one of the first Be stars in which periodic line-profile variability was found (Baade, 1984). Baade observed three or four nearly equally-spaced moving bumps. There were also low-order changes in the line profile in which the symmetry of the whole line varied, but with a longer time scale. The period of the low-order variations was estimated to be 0.505 d, five times the period of the moving bumps. Baade interpreted these results in terms of NRP with  $\ell = m = 2$  and  $\ell = m = 10$ . In our opinion the poor time coverage places considerable doubt on the accuracy of this period which is practically equal to half a day, since the aliasing problem must be quite considerable. More recently, Baade (1987a) obtained 9 nights of line profile observations in which this period is apparently confirmed.

Since two kinds of line-profile variations are seen in this star, Baade (1987b) claims that it is multiperiodic and hence proof that the NRP interpretation is the only feasible one. Harmanec (1987a) has suggested that the short period structure in Be stars can be represented by integer ratios, in which case we are dealing not with incompatible periods but harmonics of a longer period (which he identifies as the period of rotation). Baade (1987b) has criticized this suggestion. Since the period ratio in  $\mu$  Cen is 5 : 1 according

to Baade, this star is perhaps not a convincing case of multi-periodicity.

We observed  $\mu$  Cen for 29 nights in June/July 1987, 25 nights in March/April 1988 and 7 nights in June 1988. The latter two data sets are SAAO observations only. The 1987 data show that the star was photometrically very active at this time. There are two major excursions in brightness, each with a light range of 0.07 mag, separated by five nights.

A Fourier periodogram analysis shows no obvious trace of coherent periodicity : the three largest peaks being at frequencies of 0.05, 0.95 and 0.16 d $^{-1}$ . These first frequencies appear to be artifacts caused by the large brightness excursions. The PDM periodogram shows essentially the same structure except for a large peak at 0.48 d $^{-1}$ . This does not seem to be associated with the random brightness fluctuation and is the only coherent period we can extract from our 1987 data. The resulting light curve is a double-wave with a period of 2.10 d (Fig. 2).

Since our results from 1987 did not lead to a convincing period determination, we re-observed the star from SAAO in 1988. This time the star was less active, but surprisingly the periodograms look very similar to those of 1987. In particular the frequency at 0.48 d $^{-1}$  was very strongly present. This seems to indicate that the 2.10 d period is physically real. As a final check we combined the data from both seasons which allowed us to refine the period to 2.1017 d. A plot of the light curve is shown in figure 2. The double-wave nature is very clear, but the large scatter indicates that there is an additional source of variation. A large increase in amplitude followed by a swift decrease can explain some of the apparent scatter in 1987. The peak-to-peak amplitude decreased from 0.08 mag in 1987 to 0.04 mag in 1988, though the large amplitude in 1987 could be mostly a result of a non-periodic brightness excursion during a few nights. But this cannot be invoked as the cause of the large scatter in 1988 ; this phenomenon is very common in the light curves of Be stars and can be termed "flickering".

In the light of these results, we cannot confirm Baade's 0.5 d period (frequency 2.0 d $^{-1}$ ) in  $\mu$  Cen. We can possibly reconcile our results with this period if we note that the single-wave frequency (0.96 d $^{-1}$ ) is the one-day alias of Baade's period. We certainly did not find any indication of a 0.5 d period in our data. As a test, we analyzed Baade's (1984) radial velocity observations for the HeI line and constructed the phase diagram using a period of 2.1017 d. The resulting velocity curve has considerably greater scatter than the one constructed with 0.505 d, but is still quite reasonable and gives a double-wave curve. However, there are only 19 observations – far too few for a reliable period estimation.

#### $\eta$ Cen

A description of recent emission-line activity in the shell star  $\eta$  Cen is given by Dachs *et al.* (1986). They conclude that at times the internal self-absorption by the shell is so

severe that it may obscure a certain fraction of the light from the photosphere. Mennickent and Vogt (1988) observed the star during 1987. The lower order Balmer lines showed weak variable emission wings and variable shell absorption cores.

Baade (1983) discovered significant line-profile variations with a time scale of hours. During a search for rapid spectroscopic variations, Ghosh *et al.*, (1988) discovered a continuum flux variability with a time scale of hours.

Our results clearly show large short-period variability. A nightly range in brightness of nearly 0.1 mag is not uncommon. The Fourier periodogram shows a very strong peak at  $1.56 \text{ d}^{-1}$ , but the resulting light curve has a large scatter. The PDM periodogram shows, in addition, large peaks at one-half and one-third this frequency. An examination of the phase plots show that a better fit is obtained by assuming a double-wave light curve with  $f = 0.78 \text{ d}^{-1}$ , but even then the scatter is large. By far the best phase plot is obtained with  $f = 0.52 \text{ d}^{-1}$  which results in a triple-wave light curve. While double-wave light curves are common, only one case of a triple-wave light curve ( $\alpha$  Eri, Balona *et al.*, 1987) is known.

This surprising result prompted us to re-observe  $\eta$  Cen for a further season in 1988 (SAAO observations only). Our analysis of these data showed quite convincingly that  $f = 0.52 \text{ d}^{-1}$  is indeed present once again and gives a light curve with the least scatter. On this basis we feel that a triple-wave light curve with period of 1.927 d is likely to be the correct interpretation (Fig. 3). The peculiar light curve and the very large amplitude (0.14 mag in 1987, 0.10 mag in 1988) makes this star unique among the periodic Be variables. Line profile observations of  $\eta$  Cen are potentially of great importance in view of the large amplitude. If NRP is involved, the geometrical distortions and/or temperature variation must be exceedingly severe to give rise to such a large amplitude for an  $\ell = |m| = 3$  mode.

#### HD 137518

This star was included in our program at the suggestion of C. Waelkens who found it to be a large-amplitude variable. Very little is known about this star. It has been classified as a luminous blue supergiant (B5Ia) by McConnell and Bidelman (1976). On the other hand, Garrison *et al.*, (1977) obtain a classification B1IIIInep with a comment that the HeI lines are strongly veiled and the suspicion that it is a double-line spectroscopic binary. In the *Michigan Spectral Catalogue* (Houk, 1978), it is given the classification B1/2(I/IIIn), while Mermilliod (1987) tabulates it as O9Vn.

The photometric variability was discovered by Strohmeier *et al.* (1964) who found a range of 0.5 mag from sky patrol plates of the Bamberg Southern Station. The variability was later confirmed by Kozok (1985).

Our 1987 observations show a light range of 0.5 mag, but we could not find a definite periodicity. The Fourier periodogram shows a strong peak at  $f = 0.11 \text{ d}^{-1}$  or its

one-day alias at  $0.89 \text{ d}^{-1}$ . Smaller peaks are present at  $f = 0.36$  and  $1.29 \text{ d}^{-1}$ . The PDM periodogram is similar except that  $f = 0.12 \text{ d}^{-1}$  is by far the strongest. We examined the light curve at many different frequencies given by the Fourier and PDM periodograms, but no single choice produced a reasonable light curve. We could assume multiperiodicity of course. If we prewhiten the data by removing a sinusoid with  $f_1 = 0.115 \text{ d}^{-1}$ , we obtain a noisy Fourier periodogram in which the highest peak is at  $f_2 = 0.55 \text{ d}^{-1}$ . A multiperiodic Fourier fit with  $f_1$  and  $f_2$  leads to a rms scatter of 0.08 mag which is clearly unacceptable. Removing these two frequencies leads to a strong peak at  $f_3 = 0.35 \text{ d}^{-1}$ , which is the third harmonic of  $f_1$ . A Fourier fit with  $f_1$ ,  $f_2$  and  $f_3$  gives an rms scatter of 0.044 which is still very much higher than the expected observational error. Further prewhitening just gives a periodogram in which the noise level steadily increases towards low frequencies.

It is clear that no direct evidence for coherent multiperiodicity exists as even a solution with three frequencies is quite insufficient to describe the data. A solution can always be found by including more and more frequencies until the resulting rms scatter is sufficiently small, but there is no guarantee that the resulting solution will bear any resemblance to the true physics of the problem. We prefer to believe that the only likely periodicity in HD137518 is  $P = 8.70 \text{ d}$  ( $f = 0.115 \text{ d}^{-1}$ ) and its harmonics (Fig. 4). Prewhitening by a Fourier curve leads to a further possible periodicity of 0.685 d, but the reality of this period is far from convincing. The residual variability can be attributed to random fluctuations.

This star is a very puzzling object. Without detailed spectroscopic observations it seems impossible to determine its nature. The large light amplitude suggests some type of eclipsing phenomenon.

#### $\kappa^1$ Aps

Slettebak (1982) describes how two spectra, taken one day apart, show striking differences in the widths of the absorption lines (from  $v \sin i = 250 \text{ km s}^{-1}$  to  $350 \text{ km s}^{-1}$ ), suggesting that this star may be a double-line spectroscopic binary. Mennickent and Vogt (1988) observed this star during 1988 and found the spectrum characterized by narrow absorption cores flanked by weak emission at H $\beta$ .

The Fourier periodogram of this star shows it to be another periodic variable with a frequency of  $1.61 \text{ d}^{-1}$ . The PDM periodogram also gives a strong signal at half this frequency, indicating a double-wave light curve. However, the phase plot shows that this effect arises from only five data points which deepens the second minimum; apart from this there is no strong reason to believe that the light curve is double-wave.

Observations at SAAO during 1988 give much the same results, but this time the double-wave nature is more convincing. This time the amplitude difference in the two waves is based on more observations and there are also distinct differences in their shapes. By combining the data

from the two seasons we find the best double-wave period to be 1.238 d (Fig. 5). There has been relatively little change in overall amplitude between the two seasons.

#### $\eta^1$ TrA

This star turned out to be one of the few constant Be stars in our sample. No sign of periodicity is evident in either the Fourier or PDM periodograms. The rms scatter of all observations is 3 millimag – the value expected for a constant star. It was not observed in 1988.

#### 48 Lib

This star has a long history of documented spectroscopic observations. It is well known for the long-lasting presence of numerous sharp shell absorption lines and for strong regular variations of Balmer emission line profiles and radial velocities of shell absorption lines with a quasi-period of about ten years (cf., e.g. Aydin and Faraggiana, 1978). Dachs *et al.* (1986) has described the emission-line history in recent years. Mennickent and Vogt (1988) found the spectrum of 48 Lib in 1987 to be characterized by many faint FeII shell lines, some of them surrounded by emission. H $\beta$  and H $\gamma$  had deep central absorption cores with emission flanks ( $V < R$ ). The  $V/R$  ratio was variable.

The 1987 data show a very strong peak at 2.49 d $^{-1}$  in the Fourier periodogram. The PDM periodogram shows, in addition, smaller peaks at one-half and one-third of this frequency. Examination of the phase plots gives us no reason to suspect that the any of the lower frequencies are physically real. There is no appreciable reduction in scatter or differences in the waves. The phase plot for the single-wave period of 0.40 d shows a rather asymmetrical light curve with a gradual decline and sharp rise (Fig. 6).

To confirm the period, we re-observed the star during 1988. We obtained the same results. The best period from the combined data is 0.4017 d. This is the shortest period we have found among the periodic Be stars.

Ringuelet-Kaswalder (1963) reported periodic radial velocity variations with  $P = 0.1154295$  d suggesting a pulsation similar to that of the  $\beta$  Cep variables. Such a variation has never been confirmed, though rapid changes in spectral appearance during one night is not unknown (Aydin and Faraggiana, 1978). We phased Ringuelet-Kaswalder's data on our photometric period, but were unable to produce a satisfactory velocity curve.

#### $\chi$ Oph

Recent variations in the emission lines in this well-observed Be star are discussed by Dachs *et al.* (1986). The considerable number of radial velocities in the literature were analyzed by Harmanec (1987b) who tentatively proposed that  $\chi$  Oph is a spectroscopic binary with a period of 34.12 d in a highly eccentric orbit ( $e = 0.699$ ).

Balona and Engelbrecht (1987) observed this star photometrically during 1985 and suggested a possible period of 0.935 d. The available data were insufficient to discriminate between this period and its one-day alias at 14.3 d.

The shorter period was considered the more likely owing to the considerable brightness variations during the course of a night. In neither case is the resulting light curve particularly convincing.

Our 1987 data show much the same behaviour. Again, the strongest frequencies in both the Fourier and PDM periodograms occur at either 0.935 d or 14.3 d, but this time the longer period has a much higher peak. The resulting phase plot shows much less scatter for 14.3 d than for 0.935 d. Thus, by combining data from two separate sites we eliminate the alias problem and conclude that the true period is 14.3 d. Combining the 1985 and 1987 data yields a best period of 13.774 d (Fig. 7). A small number of observations were obtained in 1988.

This is the longest period so far detected for a periodic Be star. The question arises as to whether it represents an orbital variation. We looked for, but could not find, evidence for a radial velocity variation with this period from the data collected by Harmanec (1987b).

In spite of the fact that the 13.774 d period fits the photometric data, it is clear that there are variations on a much shorter time scale. However, these cannot be periodic as they would leave a signature in the periodogram. We are forced to conclude that this is one more example of the "flickering" phenomenon so common amongst these stars.

#### $\zeta$ Oph

$\zeta$  Oph (O9.5Ve) is one of the most rapidly rotating stars known. Emission lines at H $\alpha$  and HeI were observed between July 1973 and April 1974 (Irvine, 1974; Niemela and Mendez, 1974; Barker and Brown, 1974). During 1979 there was no hint of emission at H $\alpha$ , but by March 1980 H $\alpha$  appeared as a well-developed double-emission feature quite similar to that observed during 1974 (Ebbets 1981).

Walker *et al.* (1979) were the first to discover the by now well-known phenomenon of "moving bumps" in the line profile of B stars. It was in this star,  $\zeta$  Oph, that they were first detected. At that time they were interpreted as nonuniformities in the stellar photosphere carried across the line of sight by rotation. In this way a rotational period of 21.7 hours was deduced, leading to an estimated  $v \sin i = 560$  km s $^{-1}$ . This value is considerably higher than observed ( $v \sin i = 370$  km s $^{-1}$ ), but it is sensitive to the adopted stellar radius. Harmanec (1989) has shown that a somewhat smaller radius, which is still consistent with the physical parameters for this star, will bring the expected projected rotational velocity into good agreement with the observed value.

Walker *et al.* (1981) presented further observations of this new phenomenon and proposed an alternative model in which the moving bumps arise from obscurations in the circumstellar material. They also suggested NRP as a possible explanation.

Vogt and Penrod (1983) obtained extensive high-resolution spectra of  $\zeta$  Oph and confirmed the moving bumps seen

by the earlier workers. They rejected the interpretation in terms of rotational modulation or obscuration by circumstellar material. While these models give rise to moving bumps in agreement with the observations, the expected light variations are not seen. They favour the NRP model. Harmanec (1989) has criticized this conclusion as it is based on only one night of photometry which does indeed show some variability. Nevertheless, the NRP model has been accepted by most groups as the most probable explanation in this and other Be stars.

It is clear that intensive photometric observations should shed considerable light on the nature of moving bumps. According to Vogt and Penrod (1983) the NRP interpretation should give rise to periodic light variations of amplitude 0.02 mag with a period of a few hours. Such periodic variations are easily detectable.

We started our photometric campaign on this star in 1985. The results of this first season were briefly reported in Balona and Engelbrecht (1987). During 1985  $\zeta$  Oph showed a clear indication of pulsation with a period of 0.193 d and amplitude of 0.02 mag. There was also a longer period of 1.075 d with amplitude 0.03 mag (Fig. 8a). The short-period pulsations were interpreted as NRP, but not as a new phenomenon. Since  $\zeta$  Oph lies on the hot end of the  $\beta$  Cep instability strip, it seemed reasonable to classify the star as a  $\beta$  Cep variable. However, the presence of the longer period in a  $\beta$  Cep star is most unusual ; rather it is indicative of a periodic Be star. It is also of particular interest to note that if the light curve is phased on the long period of 1.075 d the short-period variation is easily visible as a sinusoidal modulation of the light curve (Fig. 8b). It looks very much as if the short period is practically one-sixth of the long period to within observational errors. This seems to confirm Harmanec's (1987a) suspicion that whenever a short-period is present it turns out to be a sub-multiple of the rotation period.

We observed  $\zeta$  Oph during 1987 expecting to confirm the provisional results announced by Balona and Engelbrecht (1987). The Fourier and PDM periodograms both show some power at frequencies of 1.00 or 2.00  $d^{-1}$  with a peak-to-peak amplitude of 0.01 mag. There is no sign of the short period variation seen in 1985. While the long period is close to the value seen in 1985, it is practically the same as one cycle per day and, as such, must give rise to a grave suspicion that this is an artifact. Indeed, since we had to observe  $\zeta$  Oph at SAAO using a neutral density filter it would not be entirely surprising if some slight systematic error was present in the data. Under these conditions we can understand the results of the periodogram and we do not consider the one day period to be real. Indeed, the rms error of one observation from both sites is only 4 millimags., i.e. close to the expected rms error of a constant star. We conclude that  $\zeta$  Oph was constant in light during June/July 1987.

Some observations at SAAO during early 1988 again

indicated a constant light. But in June of that year a dramatic change took place : our results of four photometric nights show marked photometric activity with night-to-night variations exceeding 0.01 mag. It would be very interesting to discover if this is associated with the development of emission at  $H\alpha$ . Further photometric monitoring is planned for 1989.

#### $\iota$ Ara

Our photometric observations of this star now cover four seasons (1985 - 1988). The 1985 light variations were particularly difficult to interpret in spite of the fact that there were no long-term trends or other difficulties. The problem is the very large flickering associated with this star which tends to mask the underlying variation almost completely. Balona and Engelbrecht (1987) obtained a double-wave light curve with period 0.515 d as the most likely solution for this season. During 1986 the amplitude appeared to increase steadily during the observing period of one month, but a period of 0.56 d could be extracted. This is the same as the period obtained in 1985 within the observational error and results in a single-wave light curve.

Our 1987 data showed that once more the star was active with a short time scale. The strongest peak in the Fourier periodogram occurs at 0.265 d, but the PDM periodogram indicates that a double-wave solution is better as the strongest peak occurs at 0.53 d. The 1988 observations are too few to derive a period, but a solution with  $P = 0.56$  d fits the data well.

The conclusion is that in spite of the low signal-to-noise ratio, a period close to 0.55 d is derived independently from each observing season. On this account we feel confident that this is indeed the correct period. Combining all the data gives a best solution of  $P = 0.5565$  d (Fig. 9). The flickering in this star is one of the largest we have encountered. As a consequence the periodicity in this star is somewhat blurred, but periodicity implies coherence over many cycles and this period was the only one where such a coherence was observed. The rms deviation from the light curve is as large as 0.02 mag though the amplitude is as large as 0.10 mag.

#### $\alpha$ Ara

Mennickent and Vogt (1988) found the spectrum in 1986-7 to show double emission in  $H\beta$  and  $H\gamma$  ( $V \simeq R$ ) with small variations in the  $V/R$  ratio on a short time scale.

This star was immediately recognized as a large-amplitude short-period variable during the first observing run in 1985. The Fourier periodogram shows strong peaks at frequencies of 2.04 or 3.04  $d^{-1}$  and its one day aliases. These two frequencies are also the strongest in the PDM periodogram, but the double-wave solution at 1.52  $d^{-1}$  is nearly as strong. An examination of the light curve does indicate that there is a significant difference in amplitude between the two waves in the double-wave solution. Thus we adopted a period of 0.658 d ( $f = 1.52$   $d^{-1}$ , Balona and Engelbrecht, 1987).

In 1986 the highest peak in the periodograms occurs at

$2.04 \text{ d}^{-1}$ ; the solution for 1985 thus appears to be a one-day alias of the true period. That a wrong choice was made is not surprising in view of the closeness of the frequency to a half a day and the severe aliasing problem.

Our 1987 data shows that the ambiguity is completely resolved thanks to the different longitudes of the observing sites. The true period is either a single-wave with  $P = 0.49 \text{ d}$  or a double-wave with  $P = 0.98 \text{ d}$ . It is very difficult to be sure of the double-wave solution since the two waves are of nearly equal amplitude, but it is a distinct possibility. Unfortunately, the period is so close to one day that we cannot use the other observations (all made at SAAO) to settle this problem. By combining all the data we find a best period of  $0.9807 \text{ d}$  for the double-wave solution (Fig. 10). Results from 1988 indicate that this period fits the data well.

### 66 Oph

During 1987 this star showed double emission at  $\text{H}\beta$  and  $\text{H}\gamma$  (Mennickent and Vogt, 1988).

We observed this star for one season only (1987). Apart from a gradual increase in brightness, there is no indication of periodicity. Nevertheless, the star is variable as the scatter (0.01 mag) is much larger than expected. It appears that we are seeing only the flickering component.

### V986 Oph

This star is not known as a Be star, but we have included it here owing to its great interest in connection with the role of NRP in the Be phenomenon. V986 Oph has a long history of photometric variability. Lynds (1959) determined a period of  $0.289 \text{ d}$  and classified it as a  $\beta$  Cep variable. Jerzykiewicz (1975) confirmed this period from seven nights of photometry and deduced a peak-to-peak amplitude of 0.014 mag. However, not all the data could be described by a simple sinusoidal variation with this period. Pike and Lloyd (1979) failed to detect any radial velocity variations with the photometric period, though a systematic decrease in nightly mean velocities was observed.

To clarify the nature of this star, C. D. Pike organized an international campaign for July 1980. Fullerton *et al.* (1985) analyzed both the photometry and radial velocities obtained during the campaign and found a period of  $0.325 \text{ d}$  with a light amplitude of 0.01 mag in  $B$ . No significant radial velocity variations were found, but the systematic variation in mean nightly velocity was confirmed and orbital elements derived ( $P = 25.56 \text{ d}$ ,  $2K = 35 \text{ km s}^{-1}$ ,  $e = 0.23$ ). They also obtained an extensive set of high dispersion line-profile observations and discovered high-order profile variations progressing from blue to red with a period of about  $0.3 \text{ d}$ . They deduced  $\ell = 6$ , but at times  $\ell = 4$  and  $\ell = 8$  are excited.

This confusing picture of V986 Oph prompted us to include it in our observing project. Our photometry for 1985 was obtained over 10 nights and showed light variations at a low level. Fourier periodogram analysis showed peaks at  $1.74$  and  $1.16 \text{ d}^{-1}$  and their one-day aliases (Fig. 11a - top panel). An unique period or periods could not be

determined from these data alone. Combining these data with all available photometry showed the most likely period to be  $0.30 \text{ d}$  or  $0.23 \text{ d}$ , its alias.

Our 1987 data again shows micro-variability, but owing to the spacing in longitude the aliasing problem largely disappears. The Fourier periodogram shows three strong peaks at  $f_1 = 3.30$ ,  $f_2 = 1.41$  and  $f_3 = 0.77 \text{ d}^{-1}$  with semi-amplitudes of 4.7, 4.6 and 3.8 millimags respectively (Fig. 11a - middle panel). Prewhitenning by  $f_1$  leaves  $f_2$  and  $f_3$  with  $f_2$  slightly stronger (Fig. 11a - bottom panel). Further prewhitening by either  $f_2$  or  $f_3$  leads to a pure noise spectrum. Our conclusion is that there are two periods present : the short period is well determined to be  $0.303 \text{ d}$  (Fig. 11b), but the long period is uncertain, it is either  $1.299 \text{ d}$  or  $0.709 \text{ d}$ .

The short period is the same as the one originally suggested by Lynds and Jerzykiewicz and also found in our 1985 data. We cannot confirm the period found by Fullerton *et al.* (1985); indeed there is no indication of a  $0.325 \text{ d}$  period in our data at all. We note, however, that the  $0.303 \text{ d}$  period is the same as their estimate of the period of the high-order line-profile variations. The low photometric amplitude is consistent with a high-order NRP mode. The nature of the long period is uncertain. It could be another NRP mode (either rotationally split  $p$ -mode or a  $g$ -mode), but the period is close to the expected period of rotation so that some kind of rotational modulation is not ruled out.

There appears to be no compelling reason to classify this star as anything other than a  $\beta$  Cep variable as originally suggested by Lynds (1959). Admittedly, it has the longest known period of a star in this class, but this is not surprising in view of its early spectral type and luminosity. It does apparently oscillate with somewhat higher order modes than the majority of other  $\beta$  Cep variables, but this is not a criterion for reclassification. The only reasonable criterion for classification is a definite indication of a different pulsation mechanism. Since there is no reason to believe that the pulsation mechanism in this star is any different from other  $\beta$  Cep variables, we feel that reclassification serves merely to confuse the issue.

### $\epsilon$ Cap

Mennickent and Vogt (1988) observed this star in 1986 and found deep central absorption with no traces of emission at  $\text{H}\beta$ . Dachs *et al.* (1986) found  $\text{H}\alpha$  to have a central absorption and emission wings in 1982.

Our 1985 photometry showed that  $\epsilon$  Cap underwent a sharp drop in brightness of nearly 0.2 mag in just three days followed by a rapid recovery. Excluding this period, the periodogram shows a set of peaks with  $f = 2.60 \text{ d}^{-1}$  and its one-day aliases, but a double-wave light curve with half this frequency is possible as the waves have distinctly different amplitudes. This solution ( $P = 0.769 \text{ d}$ ) was proposed by Balona and Engelbrecht (1987).

The 1986 photometry again showed indications of peri-

odicity, but we could not confirm the solution derived for 1985. Instead a frequency close to two cycles per day was indicated. The severe aliasing problem prevented a meaningful solution.

The combined ESO and SAAO data of 1987 showed a slow brightening over the observing run. Removing this trend showed a possible period very close to one day. By combining all available data we determined the best value to be 1.030 d (Fig. 12). Because it is so close to one day and because of the discrepancy obtained in 1985, this period must be regarded as very uncertain.

#### $\eta$ PsA

Mennickent and Vogt (1988) find H $\beta$  to be very diffuse with weak double emission ( $V \approx R$ ) around a broad absorption core.

We were unable to detect any clear periodicity in  $\eta$  PsA during 1985, but the star was clearly a low-amplitude variable. A Fourier periodogram suggested a possible period of 0.774 d, which is the value quoted by Balona and Engelbrecht (1987). During 1986, the low-level variability (flickering) was again evident. No reliable period could be extracted. The same is true for the combined ESO and SAAO data of 1987.

#### HR 8408

Slettebak (1982) discovered HR 8408 to be a Be star showing faint double emission ( $V \approx R$ ) at H $\beta$  surrounding an absorption core. Mennickent and Vogt (1988) did not see emission during 1986, but suspect incipient emission at H $\beta$ . Buscombe and Morris (1961) found a radial velocity range of 40 km s $^{-1}$ .

This star was originally used as a photometric standard before we discovered its variability. The 1985 SAAO photometry shows clear indications of periodicity with a frequency  $f = 2.53$  d $^{-1}$ , though a double-wave solution with half this frequency shows the two waves to have significantly different amplitudes.

The 1986 data produced the highest peak in the Fourier periodogram at  $f = 1.53$  d $^{-1}$ , i.e. the one-day alias of the solution adopted for 1985. Again, a double-wave solution appeared to be significant.

The combined ESO and SAAO data of 1987 removed the aliasing problem. The correct frequency was found to be  $f = 1.53$  d $^{-1}$  or the double-wave solution at  $f = 0.765$  d $^{-1}$ . By combining all available data we refined the period to 1.3106 if we adopt the double-wave solution (Fig. 13). This value (or possibly the single-wave period of 0.6553 d) seems to be well established.

#### $\text{o Aqr}$

The emission features in this star have been relatively static since 1975. The spectrum shows double emission in the Balmer lines ( $V \approx R$ ) and a central shell absorption feature (Mennickent and Vogt 1988).

Photometry during 1985 showed a steady brightening of the star over a period of three months with a large excursion

in magnitude on one occasion. These complications made period extraction very difficult. A subset of the data indicated a possible double-wave light curve with  $P = 1.449$  d (Balona and Engelbrecht 1987). During 1986 a much larger amount of data was obtained and the star was more quiescent. Both the Fourier and PDM periodograms show a strong signal at  $f = 1.40$  d $^{-1}$  ( $P = 0.71$  d) - half the period found in 1985. However, a double-wave light curve with the period found in 1985 also described the 1986 data well.

The combined ESO and SAAO data for 1987 again shows a strong signal with  $f = 1.39$  d $^{-1}$  or the double-wave equivalent at  $f = 0.70$  d $^{-1}$ . This time the amplitude difference in the waves did not appear to be significant. Considering the good agreement of the period determinations of three seasons, we can state with little doubt that the period of this star is  $P = 1.4325$  d from all available data (Fig. 14). This gives a double-wave light curve with distinctly different amplitudes for the two waves in 1985 and 1986. An increase in amplitude from 0.01 mag in 1986 to 0.02 mag in 1987 seems to have occurred.

One of the aims of this project was to determine the temperature range at which periodic Be stars could be found.  $\text{o Aqr}$  is one of the coolest stars (B7IVe) which has been found to be periodic. Our impression is that periodicity is most likely to be found amongst earlier spectral types.

#### $\psi^2$ Aqr

This star is not known as an emission-line star. We decided to include it in our observing program owing to its very rapid rotation ( $v \sin i = 332$  km s $^{-1}$ ) and to test the hypothesis that rapidly-rotating non-Be stars do not show low-order profile variations (which are expected to give rise to periodic light variations). Abt and Levy (1978) regard the star as a probable spectroscopic binary.

The SAAO data for 1986 showed a strong peak at  $f = 1.87$  d $^{-1}$  in the Fourier periodogram. The PDM periodogram suggested that a double-wave solution with half this frequency is more appropriate. The ESO and SAAO data of 1987 confirmed these conclusions. We have adopted a double-wave solution with  $P = 1.073$  d obtained from an analysis of all existing photometry. A large decrease in amplitude seems to have occurred between the two seasons (Fig. 15).

The light variation in this star is in every respect similar to those in Be stars. It would be of great+ value to look for possible emission at H $\alpha$  in this star. If it can be shown that emission is definitely not present but that short-period light variations are present, it could offer an important clue to the relationship between the short-period variations and the Be phenomenon.

#### 5. Colour variations.

There are very few observations of periodic Be stars observed in more than one colour. Van Vuuren *et al.* (1988) found that the colour variations in periodic Be stars in the

open cluster NGC 3766 were very small. The evidence indicates that the star is bluest (hotest) when brightest. The ESO observations produced simultaneous *uvby* colours. In figures 16 - 20 we show light and colour curves for some of the periodic Be stars observed at ESO. In general, the conclusions of van Vuuren *et al.*, (1988) are supported. The best evidence comes from observations of 48 Lib (Fig. 17). There is a small phase shift between *b* and *u-b*, but the star is brightest when bluest.

## 6. Conclusions.

A detailed interpretation of the results found in this paper and others in the series will be presented elsewhere. Here we summarize the main findings.

We found periodic light variations in almost all Be stars. Of the 17 candidates, only three stars were constant in light, three others had uncertain periods. The Be star HD137518 has a very unusual light curve which may not be strictly periodic or else could be another example of a triple-wave light curve as found in  $\eta$  Cen and  $\alpha$  Eri.

In most cases the light curves are double waves with unequal maxima and minima. Sometimes the waves are of nearly equal amplitude, so it is quite possible that the periods of some apparently single-wave variables may be in error by a factor of two. Such a situation was found for two Be stars in the cluster NGC3766 (Ahmed 1 and Ahmed 15) where an apparently single-wave light curve became a double-wave light curve with twice the period in the course of a year (van Vuuren *et al.* 1988). The triple-wave light curve of  $\eta$  Cen is remarkable for its large amplitude, but has a counterpart in the bright star  $\alpha$  Eri (Balona *et al.*, 1987).

It has become increasingly clear, both in this work and in previous studies of the light variability of periodic Be stars, that the light variations are never adequately described

by purely periodic variations. There is always a residual scatter which is often many times larger than the expected observational error. We have called these random, short-term variations *flickering*. It seems that flickering is closely associated with periodic variability. The constant Be stars are really constant and do not show this flickering (or at least not to the same extent).

In all cases we found the period of light variation to be consistent with the expected period of rotation of the star. There is no clear evidence for multiperiodicity in any of the stars. In the case of  $\zeta$  Oph we found a  $\beta$  Cep - like variation with a period of 0.193 d superimposed on a longer period of 1.075 d which is practically six times the length of the shorter period and could be the rotation period. This behaviour was only seen for one season. This star is clearly an exception to the rule and deserves closer study. For  $\chi$  Oph (and possibly HD137518) we found evidence for a period longer than the expected period of rotation. It is not very clear whether this variation is intrinsic to the star (as is the case in general) or a result of a close companion or circumstellar material.

The results of this work clearly show the importance of multi-site observations for these short-period variables. The aliasing problem that still existed in some stars was completely resolved by combining results from the two sites.

## Acknowledgements.

At the time that these results were obtained and for part of the duration of the reductions, J.C. was at the *Astronomisch Instituut van de Katholieke Universiteit Leuven* under grant OT/86/28 of the *Onderzoeksfonds van de Katholieke Universiteit Leuven*. This support is gratefully acknowledged. We are also grateful to Dr. C. Waelkens for his role in organizing the observing campaign and for his stimulating comments.

## References

- ABT H. A. and LEVY S. G. : 1978, *Astrophys. J. Suppl. Ser.* **36**, 241.
- ASHOKA B. N., SURENDIRANATH R. and KAMESWARA RAO N. : 1985, *Acta Astron.* **35**, 395.
- AYDIN C. and FARAGGIANA R. : 1978, *Astron. Astrophys. Suppl. Ser.* **34**, 51.
- BAADE D. : 1983, *Astron. Astrophys.* **124**, 283.
- BAADE D. : 1984, *Astron. Astrophys.* **135**, 101.
- BAADE D. : 1987a, in *IAU Symp. No. 132*, The Impact of very high S/N Spectroscopy on Stellar Physics, G. Cayrell and M. Spite Eds. (Kluwer Academic Publ., Dordrecht) p. 193.
- BAADE D. : 1987b, *Inf. Bull. var. Stars* **3124**.
- BAADE D., DACHS J., VAN DE WEYGAERT R. and STEEMAN F. : 1988, *Astron. Astrophys.* **198**, 211.
- BALONA L. A. and ENGELBRECHT C. A. : 1986, *Mon. Not. R. Astron. Soc.* **219**, 131.
- BALONA L. A., ENGELBRECHT C. A. and MARANG F. : 1987, *Mon. Not. R. Astron. Soc.* **227**, 123.
- BALONA L. A. and ENGELBRECHT C. A. : 1987, in *IAU Coll. 92*, The Physics of Be Stars A. Slettebak and T. P. Snow Eds. (Cambridge

- University Press, Cambridge) p. 87.
- BARKER P. K. and BROWN T. : 1974, *Astrophys. J. Lett.* **192**, L11.
- BUSCOMBE W. and MORRIS P. M. : 1961, *Mon. Not. R. Astron. Soc.* **123**, 233.
- DACHS J., HANUSCHIK R., KAISER D., BALLEREAU D., BOUCHET P., KIEHLING R., KOZOK J., RUDOLPH R. and SCHLOSSER W. : 1986, *Astron. Astrophys. Suppl. Ser.* **63**, 87.
- EBBETS D. : 1981, *Publ. Astron. Soc. Pac.* **93**, 119.
- FULLERTON A. W., BOLTON C. T. and PENROD G. D. : 1985, *J. R. Astron. Soc. Canada* **79**, 236.
- GARRISON R. F., HILTNER W. A. and SCHILD R. E. : 1977, *Astrophys. J. Suppl. Ser.* **35**, 111.
- GHOSH K. K., VELU C., KUPPUSWAMY K., JAYKUMAR K. and ROSARIO M. J. : 1987, *Inf. Bull. Var. Stars* **3056**.
- GHOSH K. K., PUCALENTHI S. and JAYKUMAR K. : 1988, *Be star Newslett.* **18**, 7.
- HARMANEC P. : 1987a, *Inf. Bull. Var. Stars* **3097**.
- HARMANEC P. : 1987b, *Bull. Astron. Inst. Czech.* **38**, 283.
- HARMANEC P. : 1989, *Bull. Astron. Inst. Czech.*, in press.
- HENDRY E. M. and BAHNG J.D.R. : 1981, *J. Astron. Astrophys.* **2**, 141.
- HOUK N. : 1978, Michigan catalogue of two-dimensional spectral types for HD stars, **2**.
- IRVINE N. J. : 1974, *Astrophys. J. Lett.* **188**, L19.
- JERZYKIEWICZ M. : 1975, *Acta Astron.* **25**, 81.
- KOZOK J. R. : 1985, *Astron. Astrophys. Suppl. Ser.* **62**, 7.
- KUBIAK M. and SEGGEWISS W. : 1982, *Acta Astron.* **32**, 371.
- LYNDS C. R. : 1959, *Astrophys. J.* **130**, 577.
- MCCONNELL D. J. and BIDELMAN W. P. : 1976, *Astron. J.* **81**, 225.
- MENNICKENT R. I. and VOGT N. : 1988, *Astron. Astrophys. Suppl. Ser.* **74**, 497.
- MERMILLIOD J. C. : 1987, *Astron. Astrophys. Suppl. Ser.* **71**, 119.
- NIEMELA V. S. and MENDEZ R. H. : 1974, *Astrophys. J. Lett.* **187**, L23.
- PALMER H. K. : 1906, *Lick Obs. Bull.* **4**, 97.
- PENROD G. D. : 1987, in *IAU Coll. 92*, The Physics of Be Stars A. Slettebak and T.P. Snow Eds. (Cambridge University Press, Cambridge) p. 463.
- PERCY J. R. : 1987, in *IAU Coll. 92*, The Physics of Be stars A. Slettebak and T. P. Snow Eds. (Cambridge University Press, Cambridge) p. 49.
- PIKE C. D. and LLOYD C. : 1979, *Inf. Bull. Var. Stars* **1716**.
- RAJAMOHAN R. : 1977, *Kodaikanal Obs. Bull. Ser. A* **2**, 6.
- RINGUELET-KASWALDER A. : 1963, *Astrophys. J.* **137**, 1310.
- SLETTEBAK A. : 1982, *Astrophys. J. Suppl. Ser.* **50**, 55.
- STELLINGWERF R. F. : 1978, *Astrophys. J.* **224**, 953.
- STROHMEIER W., KNIGGE R. and OTT H. : 1964, *Inf. Bull. Var. Stars* **62**.
- SWINGLER D. N. : 1989, *Astron. J.* **97**, 280.
- VAN VUUREN G. W., BALONA L. A. and MARANG F. : 1988, *Mon. Not. R. Astron. Soc.* **234**, 373.
- VOGT S. S. and PENROD G. D. : 1983, *Astrophys. J.* **275**, 661.
- WAELKENS C. and RUFENER F. : 1983, *Astron. Astrophys.* **121**, 45.
- WALKER G. A. H., YANG S. and FAHLMAN G. G. : 1979, *Astrophys. J.* **233**, 199.
- WALKER G. A. H., YANG S. and FAHLMAN G. G. : 1981, in Proc. Workshop on Pulsating B Stars G.E.V.O.N. and C. Sterken Eds. (Nice Obs.) p. 261.
- WALKER G. A. H. and YANG S. : 1987, Observational astrophysics with very high precision data (Univ. Liege) p. 337.
- WILSON R. E. : 1914, *Lick Obs. Bull.* **8**, 130.

TABLE I.—*A summary of stars observed. Notes are given for program stars ; the others are comparison stars. The abbreviation 2W and 3W stands for double-wave and triple -wave light curve respectively. The grouping includes stars which are in close proximity and for which the same transparency corrections were applied. The projected rotational velocity, the mean Strömgren b magnitude and the total number of observations is shown.*

HR	HD	Name	MK type	vsini	$\langle b \rangle$	N	Notes
HR 5190	HD 120307	$\nu$ Cen	B2IV	91	3.33	279	P = 2.621 d.
HR 5193	HD 120324	$\mu$ Cen	B2IV-Ve	175	3.39	360	P = 2.1017 d, 2W.
HR 5439	HD 127971		B7V		5.872	202	
HR 5440	HD 127972	$\eta$ Cen	B1.5Vne	333	2.32	352	P = 1.927 d, 3W.
HR 5668	HD 135348		B3IV		6.026	178	
	HD 137518		B1IIIInep		7.94	183	P = 8.696 d?
HR 5730	HD 137387	$\kappa^1$ Aps	B1pne		5.39	308	P = 1.238 d, 2W.
HR 5786	HD 138867		B9V		5.950	246	
HR 6172	HD 149671	$\eta^1$ TrA	B7IVe		5.89	202	Constant.
HR 6233	HD 151441		B8II-III		6.150	215	
HR 5902	HD 142096		B2.5V	197	5.065	230	
HR 5941	HD 142983	48 Lib	B5IIIpe	393	4.81	232	P = 0.4017 d.
HR 5993	HD 144470		B1V	142	3.986	401	
HR 6112	HD 147933		B2V	303	4.781	134	
HR 6118	HD 148184	$\chi$ Oph	B2IV:pe	134	4.61	346	P = 13.774 d.
HR 6175	HD 149757	$\zeta$ Oph	O9.5Vne	379	2.67	313	P <sub>1</sub> = 1.075, P <sub>2</sub> = 0.193 d.
HR 6224	HD 151133		B9.5III	100	6.050	131	
HR 6451	HD 157042	$\iota$ Ara	B2IIIIne	369	5.25	542	P = 0.556 d, 2W.
HR 6460	HD 157243		B7III	150	5.120	684	
HR 6475	HD 157599		B8-9V		6.190	405	
HR 6510	HD 158427	$\alpha$ Ara	B2IIIIne	298	2.81	545	P = 0.9807 d, 2W.
HR 6712	HD 164284	66 Oph	B2Ve	221	4.77	160	Constant.
HR 6719	HD 164432		B2IV		6.352	287	
HR 6732	HD 164716		B9V		6.878	66	
HR 6747	HD 165174	V986 Oph	B0IIIIn	434	6.20	274	P <sub>1</sub> =0.303, P <sub>2</sub> =1.3/0.7 d.
HR 8260	HD 205637	$\epsilon$ Cap	B2.5Vpe	293	4.36	567	P = 1.03 d?
HR 8293	HD 206546		A3m	47	6.336	581	
HR 8386	HD 209014	$\eta$ PsA	B8Ve		5.39	598	Constant.
HR 8408	HD 209522		B4IVne	300	5.90	597	P = 1.3106 d, 2W.
HR 8446	HD 210300		A5V		6.500	482	
HR 8402	HD 209409	$\circ$ Aqr	B7IVe	227	4.67	526	P = 1.433 d, 2W.
HR 8451	HD 210419		A1Vnn	254	6.270	455	
HR 8533	HD 212404		A0V	65	5.750	406	
HR 8840	HD 219402		A3V		5.561	389	
HR 8858	HD 219688	$\psi^2$ Aqr	B5V	332	4.32	388	P = 1.073 d, 2W.

TABLE IIa. — *Strömgren-b* observations for periodic Be stars observed at SAAO. The heliocentric Julian day is with respect to JD2440000.000.

HR 5190	HJD	b	HJD	b	HJD	b	HR 5190		
	6981.2317	3.329	7241.5998	3.301	7250.6334	3.302			
HJD	b		7241.6214	3.301	7251.3468	3.287	HJD	b	
6959.2227	3.326	6981.3039	3.331	7241.6288	3.295	7251.3967	3.301	7266.6034	3.302
6959.2679	3.314	6981.3763	3.336	7242.2955	3.301	7251.4350	3.304	7267.2784	3.305
6959.3058	3.315	6982.2613	3.322	7242.3293	3.295	7251.4798	3.305	7267.3253	3.299
6959.3447	3.322	6982.3033	3.324	7242.3639	3.303	7251.5175	3.306	7268.2740	3.298
6959.3817	3.329	6982.3941	3.317	7242.4022	3.303	7251.5553	3.298	7268.3204	3.301
6959.4169	3.331	6983.2248	3.337	7242.4351	3.301	7251.6264	3.299	7268.3712	3.297
6959.4671	3.328	6983.2623	3.333	7242.4691	3.302	7252.3409	3.297	7268.4254	3.295
6960.4225	3.341	6983.3049	3.340	7242.5016	3.301	7252.3767	3.300	7268.4744	3.302
6961.2177	3.324	6983.3515	3.327	7242.5813	3.304	7252.4118	3.296	7270.2753	3.302
6961.2457	3.322	7236.3839	3.297	7242.6059	3.301	7252.4502	3.293	7270.3249	3.300
6961.2717	3.320	7236.4252	3.292	7242.6217	3.302	7252.4933	3.293	7270.3770	3.299
6961.3051	3.321	7236.4608	3.306	7243.3035	3.313	7252.5343	3.290	7270.4316	3.297
6961.3383	3.319	7236.5013	3.296	7243.3364	3.308	7252.5718	3.288	7270.4852	3.302
6961.3701	3.321	7236.5363	3.300	7243.3866	3.308	7252.6081	3.293	7270.5198	3.296
6961.4039	3.323	7236.5711	3.297	7243.4018	3.306	7252.6444	3.288	7270.5547	3.299
6961.4493	3.320	7236.6065	3.300	7243.4337	3.305	7252.6809	3.290	7270.5862	3.301
6962.2118	3.307	7236.6294	3.294	7243.4663	3.303	7252.7082	3.287	7270.6178	3.296
6962.2566	3.319	7237.2946	3.304	7243.5014	3.306	7252.7421	3.311	7274.2888	3.309
6962.2845	3.324	7237.3298	3.305	7243.5338	3.306	7252.7812	3.310	7274.3473	3.297
6962.3180	3.324	7237.3642	3.306	7243.5645	3.307	7252.8143	3.309	7274.4088	3.298
6963.2381	3.363	7237.3980	3.301	7243.5939	3.307	7252.8496	3.305	7274.4663	3.300
6963.3844	3.343	7237.4322	3.306	7243.6132	3.311	7253.0154	3.308	7274.5045	3.302
6963.4322	3.341	7237.4664	3.300	7243.6426	3.311	7253.0464	3.306	7274.5426	3.305
6964.2222	3.319	7238.2952	3.310	7245.3135	3.302	7253.2964	3.298	7274.5828	3.305
6966.2255	3.318	7238.3271	3.311	7246.3011	3.312	7253.3359	3.297	7274.5761	3.301
6966.2510	3.323	7238.3598	3.305	7246.3329	3.305	7253.4091	3.296	7321.4137	3.297
6966.3001	3.323	7238.3944	3.300	7246.3650	3.304	7253.4427	3.295	7322.2234	3.308
6966.3609	3.328	7238.4292	3.303	7246.3962	3.307	7253.4803	3.300	7322.2834	3.305
6966.3923	3.328	7238.4620	3.307	7246.4297	3.308	7253.5154	3.297	7322.3407	3.306
6966.4498	3.321	7240.4967	3.303	7246.4662	3.303	7253.5524	3.290	7322.3973	3.303
6967.2222	3.336	7238.5291	3.300	7246.4984	3.305	7253.5871	3.295	7322.4569	3.303
6967.2512	3.333	7238.5612	3.304	7246.5204	3.304	7253.6210	3.300	7323.2288	3.295
6967.2836	3.335	7238.5956	3.305	7246.5628	3.308	7253.6294	3.301	7323.2858	3.291
6967.3203	3.337	7238.6240	3.304	7246.6060	3.300	7253.6333	3.300	7323.3409	3.311
6967.3495	3.346	7238.6599	3.305	7246.6331	3.301	7253.6368	3.299	7323.4231	3.293
6967.3813	3.341	7240.2931	3.311	7246.6495	3.304	7253.6481	3.299	7323.4773	3.296
6967.4118	3.334	7240.3257	3.309	7247.3082	3.300	7253.6811	3.299	7328.2484	3.298
6967.4628	3.353	7240.3615	3.311	7247.3452	3.296	7253.7173	3.297	7328.3338	3.298
6968.1219	3.341	7240.3959	3.309	7247.3791	3.296	7253.7504	3.302	7329.2485	3.302
6968.2458	3.332	7240.4288	3.307	7247.4125	3.298	7253.7912	3.302	7329.3097	3.302
6968.2735	3.339	7240.4636	3.306	7247.4498	3.300	7253.8211	3.294	7329.3602	3.302
6968.3018	3.337	7240.4964	3.309	7247.4839	3.298	7253.8624	3.305	7329.4219	3.307
6968.3318	3.357	7240.5293	3.310	7247.5213	3.299	7253.9068	3.298	7330.2522	3.300
6972.2623	3.327	7240.5626	3.304	7247.5558	3.299	7253.9413	3.294	7330.3110	3.305
6972.3074	3.330	7240.5956	3.307	7247.5891	3.302	7253.9881	3.299	7330.3578	3.303
6972.3585	3.340	7240.6216	3.309	7247.6141	3.301	7254.0188	3.297	7330.4208	3.307
6978.2279	3.330	7240.6403	3.312	7247.6308	3.298	7254.5552	3.294	7331.2508	3.292
6978.2657	3.32	7241.2953	3.307	7248.3134	3.310	7255.5912	3.306	7331.3115	3.297
6978.3111	3.341	7241.3289	3.303	7248.4083	3.304	7256.2761	3.304	7331.3640	3.294
6978.3564	3.344	7241.3624	3.298	7248.5753	3.315	7256.3096	3.308	7331.4328	3.299
6978.4156	3.334	7241.3961	3.303	7250.3382	3.284	7256.3422	3.304		
6980.2167	3.342	7241.4290	3.300	7250.3889	3.290	7256.3743	3.308		
6980.2988	3.328	7241.4622	3.304	7250.4330	3.288	7256.4403	3.304		
6980.3467	3.342	7241.4971	3.299	7250.4904	3.298	7256.4977	3.306		
6980.3858	3.334	7241.5310	3.300	7250.5435	3.307	7256.5316	3.307		
6980.4386	3.340	7241.5663	3.300	7250.5967	3.318	7256.5724	3.306		
HR 5193	HJD	b	HJD	b	HJD	b	HR 5193		
	6981.2666	3.415	7241.5012	3.376	7250.6345	3.398			
HJD	b		7241.5234	3.378	7251.3480	3.371	HJD	b	
6959.2439	3.407	6981.3775	3.421	7241.5204	3.373	7251.3808	3.386	7267.2791	3.392
6959.2692	3.393	6982.2629	3.392	7242.2566	3.379	7251.4261	3.386	7267.3262	3.389
6959.3072	3.389	6982.3046	3.397	7242.3204	3.377	7251.4809	3.384	7268.2751	3.384
6959.3459	3.382	6982.3457	3.397	7242.3651	3.380	7251.5184	3.384	7268.3214	3.391
6959.3829	3.394	6982.3953	3.386	7242.4032	3.382	7251.5545	3.380	7268.3721	3.385
6959.4181	3.391	6982.2261	3.408	7242.4361	3.382	7251.6274	3.376	7268.4261	3.381
6959.4683	3.395	6983.2637	3.409	7242.4700	3.389	7252.3419	3.386		
6960.4242	3.406	7240.3062	3.406	7242.5023	3.387	7252.3777	3.389		
6961.2147	3.408	7240.3528	3.393	7242.5826	3.381	7252.4126	3.386		
6961.2467	3.408	7240.3852	3.350	7242.6071	3.373	7252.4515	3.380		
6961.2727	3.408	7240.4266	3.356	7242.6231	3.377	7252.4930	3.382		
6961.3062	3.411	7236.4618	3.361	7243.3044	3.378	7252.5354	3.380		
6961.3393	3.406	7236.5024	3.359	7243.3376	3.381	7252.5723	3.384		
6961.3711	3.403	7236.5374	3.359	7243.3697	3.381	7252.6093	3.386		
6961.4056	3.406	7236.5720	3.356	7243.4027	3.383	7252.6462	3.383		
6961.4504	3.402	7236.6083	3.360	7243.4347	3.377	7260.3107	3.381		
6962.2124	3.386	7236.6303	3.359	7243.4686	3.373	7260.3830	3.380		
6962.2577	3.396	7238.2956	3.363	7243.5024	3.369	7261.4763	3.381		
6962.2856	3.404	7237.3308	3.363	7243.5347	3.369	7261.5142	3.375		
6962.3198	3.401	7237.3654	3.365	7243.5675	3.370	7261.5492	3.375		
6962.3495	3.401	7237.3988	3.361	7243.5950	3.376	7261.5818	3.369		
6965.2792	3.401	7237.4331	3.363	7243.6146	3.382	7261.6165	3.371		
6965.3859	3.392	7237.4657	3.364	7243.6306	3.380	7261.6473	3.369		
6965.4132	3.386	7238.5300	3.376	7245.2971	3.366	7262.2974	3.377		
6966.2237	3.357	7238.3280	3.367	7246.3021	3.398	7262.3378	3.381		
6966.2546	3.358	7238.3608	3.366	7246.3339	3.394	7262.3741	3.379		
6966.2833	3.367	7238.3954	3.368	7246.3660	3.396	7262.4101	3.376		
6966.3131	3.361	7238.4302	3.372	7246.3971	3.394	7262.4433	3.375		
6966.3619	3.358	7238.4640	3.374	7246.4206	3.397	7262.4815	3.381		
6966.3932	3.362	7238.4976	3.373	7246.4674	3.397	7262.5168	3.370		
6966.4249	3.356	7238.5300	3.376	7246.4993	3.393	7262.5535	3.377		
6967.2241	3.366	7238.5652	3.375	7246.5314	3.391	7262.5882	3.379		
6967.2552	3.357	7238.5976	3.375	72					

## PHOTOMETRY OF BE STARS

TABLE IIa (*continued*)

HR 5440	HJD	b	HJD	b	HJD	b	HR 5440
HJD	b	6980.4417 2.296	7241.6265 2.193	7251.4028 2.180	HJD	b	
6959.2277 2.298	6981.2354 2.311	7241.6340 2.193	7251.4390 2.188	7268.2777 2.213			
6959.2714 2.287	6981.2666 2.310	7242.2992 2.206	7251.4844 2.180	7268.3242 2.182			
6959.3088 2.290	6981.3074 2.311	7242.3327 2.187	7251.5215 2.178	7268.3755 2.195			
6959.3477 2.292	6981.3793 2.313	7242.3685 2.181	7251.5571 2.186	7268.4290 2.210			
6959.3845 2.298	6982.2655 2.275	7242.4058 2.161	7251.6304 2.201	7268.4777 2.222			
6959.4197 2.306	6982.3064 2.287	7242.4386 2.150	7252.3449 2.206	7270.2797 2.205			
6959.4699 2.313	6982.3475 2.294	7242.4726 2.147	7252.3804 2.220	7270.3281 2.198			
6960.4272 2.286	6982.3973 2.298	7242.5056 2.149	7252.4153 2.227	7270.3818 2.200			
6961.2201 2.298	6982.4503 2.300	7242.5846 2.172	7252.4984 2.204	7270.4355 2.198			
6961.2483 2.289	6982.2286 2.312	7242.6094 2.181	7252.5385 2.193	7270.4890 2.203			
6961.2744 2.306	6982.2659 2.310	7242.6260 2.188	7252.5747 2.184	7270.5234 2.193			
6961.3079 2.312	6982.3080 2.316	7243.3069 2.192	7252.6119 2.178	7270.5590 2.186			
6961.3407 2.316	6982.3561 2.318	7243.3401 2.199	7252.6493 2.174	7270.5912 2.173			
6961.3725 2.331	7236.3883 2.213	7243.3720 2.211	7260.3150 2.200	7270.6222 2.176			
6961.4063 2.333	7236.4291 2.226	7243.4050 2.204	7260.3875 2.192	7274.2955 2.227			
6961.4522 2.341	7236.4642 2.225	7243.4376 2.193	7261.4797 2.235	7274.3519 2.210			
6962.2143 2.319	7236.5032 2.221	7243.4711 2.194	7261.5173 2.224	7274.4132 2.183			
6962.2594 2.302	7236.5400 2.216	7243.5047 2.191	7261.5517 2.217	7274.4716 2.160			
6962.2873 2.290	7236.5744 2.196	7243.5367 2.197	7261.5842 2.203	7274.5091 2.162			
6962.3212 2.281	7236.6124 2.175	7243.5699 2.202	7261.6192 2.202	7274.5469 2.160			
6962.3541 2.280	7236.6339 2.184	7243.5975 2.200	7261.6500 2.199	7274.5803 2.163			
6963.2412 2.342	7237.3331 2.166	7243.6181 2.204	7262.3017 2.200	7321.4189 2.197			
6963.2739 2.345	7237.3679 2.171	7243.6334 2.198	7262.3409 2.201	7322.2281 2.180			
6963.3874 2.372	7237.4019 2.178	7245.3170 2.211	7262.3767 2.194	7322.2885 2.198			
6963.4348 2.341	7237.4352 2.181	7246.3043 2.163	7262.4473 2.187	7322.3454 2.215			
6963.4784 2.296	7237.4680 2.192	7246.3362 2.150	7262.4844 2.188	7322.4024 2.219			
6966.2254 2.269	7238.3308 2.213	7246.3481 2.155	7262.5196 2.188	7322.4621 2.215			
6966.2563 2.272	7238.3634 2.218	7246.3991 2.157	7262.5562 2.194	7323.2344 2.215			
6966.2850 2.278	7238.3984 2.213	7246.4333 2.169	7262.5928 2.192	7323.2919 2.205			
6966.3328 2.298	7238.4326 2.202	7246.4696 2.187	7262.6250 2.201	7323.3463 2.187			
6966.3634 2.309	7238.4667 2.207	7246.5019 2.204	7263.3001 2.218	7323.4277 2.166			
6966.3948 2.317	7238.4999 2.195	7246.5336 2.217	7263.3377 2.234	7323.4827 2.195			
6966.4265 2.316	7238.5322 2.183	7246.5663 2.215	7263.3704 2.242	7328.2554 2.219			
6966.4665 2.318	7238.5672 2.178	7246.6095 2.217	7263.4057 2.243	7328.3394 2.226			
6967.2257 2.372	7238.5999 2.173	7246.6370 2.216	7263.4421 2.239	7328.4166 2.210			
6967.2540 2.384	7238.6289 2.171	7246.6534 2.215	7263.4772 2.228	7329.2549 2.166			
6967.2863 2.397	7238.6440 2.169	7247.3115 2.204	7263.5120 2.214	7329.3145 2.182			
6967.3227 2.393	7240.3299 2.221	7247.3485 2.195	7263.5489 2.204	7329.3651 2.201			
6967.3524 2.379	7240.3648 2.232	7247.3825 2.200	7263.5861 2.191	7329.4265 2.202			
6967.3838 2.350	7240.3996 2.219	7247.4157 2.193	7264.2809 2.180	7330.2594 2.206			
6967.4241 2.319	7240.4326 2.207	7247.4531 2.194	7265.3107 2.227	7330.3160 2.201			
6967.4651 2.287	7240.4673 2.189	7247.4879 2.184	7265.4204 2.203	7330.3629 2.194			
6968.2220 2.304	7240.5002 2.170	7247.5245 2.176	7265.4715 2.176	7330.4257 2.181			
6968.2485 2.312	7240.5323 2.155	7247.5590 2.173	7265.5225 2.173	7331.2561 2.181			
6968.2758 2.319	7240.5658 2.148	7247.5922 2.167	7265.5585 2.172	7331.2581 2.186			
6968.3042 2.328	7240.5988 2.151	7247.6180 2.169	7266.2798 2.179	7331.3168 2.209			
6968.3342 2.356	7240.6261 2.159	7247.6349 2.170	7266.3131 2.184	7331.3690 2.214			
6972.2683 2.337	7240.6448 2.164	7248.3164 2.169	7266.3458 2.186	7331.4383 2.217			
6972.3107 2.344	7241.3325 2.183	7248.4118 2.205	7266.3783 2.194				
6978.2321 2.319	7241.3660 2.198	7248.5807 2.233	7266.4461 2.182				
6978.2691 2.316	7241.3999 2.203	7250.3453 2.203	7266.5018 2.194				
6978.3143 2.311	7241.4324 2.211	7250.3942 2.211	7266.5411 2.212				
6978.3607 2.293	7241.4655 2.211	7250.4387 2.217	7266.5763 2.221				
6978.4192 2.292	7241.5005 2.198	7250.4966 2.210	7266.6071 2.218				
6980.3026 2.278	7241.5348 2.184	7250.5483 2.204	7267.2821 2.210				
6980.3499 2.282	7241.5719 2.187	7250.6010 2.203	7267.3288 2.197				

HR 137518	HJD	b	HR 5730	HJD	b	HJD	b	HJD	b	v
6978.3178 2.056	6981.2392 2.056	7247.5922 2.167	6978.2766 5.412	7261.5444 5.377	7270.6582 5.353					
6978.3642 8.019	6978.4229 8.054	7248.3126 5.411	6978.3216 5.411	7261.5768 5.373	7274.2851 5.388					
6959.2314 8.019	6959.2753 8.003	7249.2340 5.373	6978.3673 5.412	7261.6112 5.376	7274.3435 5.376					
6959.2753 8.003	6978.4865 8.061	7249.2775 5.367	6978.4254 5.390	7261.6421 5.378	7274.4049 5.362					
6959.3122 8.025	6980.3063 7.978	7249.3148 5.371	6978.4894 5.380	7262.3316 5.395	7274.4636 5.365					
6959.3510 8.014	6980.3533 8.003	7249.3528 5.378	6980.3096 5.386	7262.3860 5.396	7274.5006 5.371					
6959.3880 8.000	6980.3933 8.022	7249.3983 5.373	6980.3567 5.376	7262.4067 5.424	7274.5373 5.374					
6959.4228 7.998	6980.4505 8.032	7249.4245 5.384	6980.3960 5.376	7262.4388 5.436	7274.5722 5.380					
6959.4733 7.999	6981.2392 7.879	7249.4757 5.383	6980.4534 5.381	7262.4762 5.437	7321.4212 5.368					
6959.5259 7.989	6981.2722 7.893	7249.5280 5.389	6981.2425 5.405	7262.5115 5.427	7322.2306 5.363					
6960.4309 7.989	6981.3109 7.914	7249.5620 5.397	6981.2749 5.413	7262.5487 5.406	7322.2909 5.373					
6960.4879 7.975	6981.3826 7.971	7249.6089 5.375	6981.3136 5.423	7262.5830 5.394	7322.3475 5.371					
6960.5329 7.960	6981.4322 7.999	7249.6535 5.368	6981.3586 5.441	7262.6465 5.386	7322.4045 5.384					
6961.2228 7.984	6982.2690 7.889	7249.6917 5.370	6981.4114 5.400	7263.2922 5.361	7322.4646 5.399					
6961.2508 8.007	6982.3109 7.851	7249.7256 5.378	6981.4567 5.416	7263.2922 5.361	7322.4364 5.373					
6961.2775 8.024	6982.3507 7.827	7249.7794 5.377	6982.3334 5.387	7263.3299 5.361	7323.2364 5.373					
6961.3107 8.046	6982.4009 7.824	7249.8131 5.381	6982.3533 5.391	7263.3632 5.372	7323.2938 5.353					
6961.3435 8.065	6982.4534 7.825	7249.8571 5.388	6982.4048 5.397	7263.4344 5.373	7323.3486 5.346					
6961.3752 8.080	6983.2326 7.657	7249.8971 5.398	6982.4574 5.406	7263.4693 5.385	7328.2584 5.337					
6961.4093 8.101	6983.2695 7.644	7249.9356 5.398	6983.2356 5.431	7263.5040 5.386	7328.3443 5.347					
6961.4548 8.101	6983.3116 7.627	7249.9736 5.398	6983.2752 5.427	7263.5413 5.386	7329.2574 5.375					
6962.2177 8.135	6983.3839 7.667	7249.1926 5.398	6983.2196 5.408	7263.5783 5.395	7329.3165 5.367					
6962.2625 8.139	6983.3909 7.673	7249.2364 5.383	6983.2644 5.408	7263.6151 5.379	7330.4835 5.371					
6962.2908 8.141	6983.4050 7.681	7249.2803 5.387	6983.3033 5.371	7263.6515 5.376	7331.2617 5.360					
6962.3239 8.136			7249.3257 5.358	7264.4128 5.384	7329.4842 5.333					
6965.2447 7.866			7249.3627 5.357	7265.4662 5.407	7330.2624 5.366					
6965.2777 7.856			7249.3808 5.359	7265.4858 5.413	7330.3183 5.374					
6965.3907 7.801			7249.3848 5.359	7266.3706 5.348	7331.4412 5.363					
6965.4376 7.783										

TABLE IIa (*continued*)

HR 5941	HJD													
	b	b	b	b	b	b	b	b	b	b	b	b	b	b
6959.2407 4.840	6978.4322 4.803	6978.3751 4.800	6959.2407 4.803	6980.3161 4.791	6959.2407 4.803	6980.3628 4.801	6959.2407 4.803	6980.4019 4.810	6959.2407 4.803	6980.4593 4.815	6959.2407 4.803	6980.4593 4.815	6959.2407 4.803	6981.2490 4.827
6959.3966 4.807	6981.2828 4.827	6981.3750 4.830	6959.3966 4.807	6981.3917 4.836	6959.3966 4.807	6981.4450 4.815	6959.3966 4.807	6982.2794 4.798	6959.3966 4.807	6982.3199 4.793	6959.3966 4.807	6982.3199 4.793	6959.3966 4.807	6982.4117 4.809
6959.4211 4.793	6981.2828 4.827	6981.3750 4.830	6959.4211 4.793	6981.4639 4.822	6959.4211 4.793	6981.4639 4.822	6959.4211 4.793	6982.4117 4.809	6959.4211 4.793	6982.4117 4.809	6959.4211 4.793	6982.4117 4.809	6959.4211 4.793	6982.4117 4.809
6959.4869 4.804	6981.3250 4.830	6959.4869 4.804	6959.5330 4.805	6981.3917 4.836	6959.5330 4.805	6981.4382 4.816	6959.5330 4.805	6982.2794 4.798	6959.5330 4.805	6982.3199 4.793	6959.5330 4.805	6982.3199 4.793	6959.5330 4.805	6982.4117 4.827
6960.4963 4.844	6982.2794 4.798	6982.3199 4.793	6960.4963 4.844	6982.3199 4.793	6961.2307 4.830	6982.3592 4.801	6960.4963 4.844	6982.3199 4.793	6961.2307 4.830	6982.3592 4.801	6961.2307 4.830	6982.3592 4.801	6961.2307 4.830	6982.4117 4.809
6961.2507 4.830	6982.3592 4.801	6961.2507 4.830	6961.3833 4.806	6983.2813 4.822	6961.3833 4.806	6983.2813 4.822	6961.3833 4.806	6982.2794 4.798	6961.3833 4.806	6982.2794 4.798	6961.3833 4.806	6982.2794 4.798	6961.3833 4.806	6982.2794 4.798
6961.4430 4.801	6982.2794 4.798	6961.4430 4.801	6962.2273 4.791	7321.5036 4.804	6961.4430 4.801	7322.2370 4.804	6961.4430 4.801	7322.2974 4.804	6962.2273 4.791	7322.2974 4.804	6962.2273 4.791	7322.2974 4.804	6962.2273 4.791	7322.2974 4.804
6962.2273 4.791	7322.2974 4.804	6962.2707 4.804	6962.3316 4.813	7322.4103 4.831	6962.2707 4.804	7322.4103 4.831	6962.3316 4.813	7322.4103 4.831	6962.3316 4.813	7322.4103 4.831	6962.3316 4.813	7322.4103 4.831	6962.3316 4.813	7322.4103 4.831
6962.2855 4.830	6982.4639 4.822	6962.3316 4.813	6962.3316 4.813	6981.3917 4.836	6962.3316 4.813	6981.4382 4.816	6962.3316 4.813	6982.2794 4.798	6962.3316 4.813	6982.3199 4.793	6962.3316 4.813	6982.3199 4.793	6962.3316 4.813	6982.4117 4.809
6962.3190 4.809	6981.2423 4.809	6962.3190 4.809	6961.3512 4.817	6983.2813 4.822	6961.3512 4.817	6983.2813 4.822	6961.3512 4.817	6982.2794 4.798	6961.3512 4.817	6982.2794 4.798	6961.3512 4.817	6982.2794 4.798	6961.3512 4.817	6982.2794 4.798
6962.4036 4.814	6981.3917 4.836	6962.4036 4.814	6963.2853 4.817	7323.2424 4.827	6962.4036 4.814	7323.2424 4.827	6963.2853 4.817	7323.2424 4.827	6964.3986 4.802	7323.2998 4.820	6964.3986 4.802	7323.2998 4.820	6964.3986 4.802	7323.2998 4.820
6965.2853 4.817	7323.2424 4.827	6965.2853 4.817	6965.3896 4.802	7329.3224 4.827	6965.2853 4.817	7329.3224 4.827	6965.3896 4.802	7329.3224 4.827	6966.4457 4.806	7323.3256 4.804	6966.4457 4.806	7323.3256 4.804	6966.4457 4.806	7323.3256 4.804
6966.2855 4.830	6982.4639 4.822	6966.2855 4.830	6967.2984 4.785	7330.4337 4.840	6966.2855 4.830	7330.4337 4.840	6967.2984 4.785	7330.4337 4.840	6968.4981 4.803	7329.4342 4.803	6968.4981 4.803	7329.4342 4.803	6968.4981 4.803	7329.4342 4.803
6968.2888 4.801	6981.3917 4.836	6968.2888 4.801	6969.3146 4.813	7331.2669 4.856	6968.2888 4.801	7331.2669 4.856	6969.3146 4.813	7331.2669 4.856	6970.3168 4.801	7331.3252 4.832	6970.3168 4.801	7331.3252 4.832	6970.3168 4.801	7331.3252 4.832
6970.2841 4.795	7331.3252 4.832	6970.2841 4.795	6970.2847 4.810	7331.4470 4.810	6970.2841 4.795	7331.4470 4.810	6970.2847 4.810	7331.4470 4.810	6971.5175 4.805	7331.5073 4.814	6971.5175 4.805	7331.5073 4.814	6971.5175 4.805	7331.5073 4.814
6972.2529 4.802	6980.3600 4.804	6972.2529 4.802	6972.3146 4.813	7331.2669 4.856	6972.2529 4.802	7331.2669 4.856	6972.3146 4.813	7331.2669 4.856	6973.3965 4.811	7331.3256 4.832	6973.3965 4.811	7331.3256 4.832	6973.3965 4.811	7331.3256 4.832
6974.2457 4.816	6982.2846 4.802	6974.2457 4.816	6975.3146 4.814	7331.3252 4.832	6974.2457 4.816	7331.3252 4.832	6975.3146 4.814	7331.3252 4.832	6976.3756 4.814	7331.3256 4.832	6976.3756 4.814	7331.3256 4.832	6976.3756 4.814	7331.3256 4.832
6977.3372 4.669	6959.4935 2.658	6977.3372 4.669	6978.3738 2.676	6959.4935 2.658	6977.3372 4.669	6959.4935 2.658	6978.3738 2.676	6959.4935 2.658	6979.3372 4.669	6959.4935 2.658	6979.3372 4.669	6959.4935 2.658	6979.3372 4.669	6959.4935 2.658

HR 6175	HJD													
	b	b	b	b	b	b	b	b	b	b	b	b	b	b
6277.3874 2.666	6959.5413 2.670	6277.4133 2.663	6960.5019 2.656	6277.4258 2.662	6960.5470 2.650	6277.4258 2.662	6960.5470 2.650	6278.3056 2.656	6962.3655 2.664	6278.4187 2.659	6962.4187 2.659	6278.4187 2.659	6962.4187 2.659	6278.4187 2.659
6277.4240 2.670	6982.4639 4.822	6277.4240 2.670	6961.4430 4.814	6277.4240 2.670	6961.4430 4.814	6277.4240 2.670	6961.4430 4.814	6277.4240 2.670	6961.4430 4.814	6277.4240 2.670	6961.4430 4.814	6277.4240 2.670	6961.4430 4.814	6277.4240 2.670
6277.4257 2.668	6982.3592 4.801	6277.4257 2.668	6961.4430 4.814	6277.4257 2.668	6961.4430 4.814	6277.4257 2.668	6961.4430 4.814	6277.4257 2.668	6961.4430 4.814	6277.4257 2.668	6961.4430 4.814	6277.4257 2.668	6961.4430 4.814	6277.4257 2.668
6277.2877 2.668	6982.3592 4.801	6277.2877 2.668	6961.4430 4.814	6277.2877 2.668	6961.4430 4.814	6277.2877 2.668	6961.4430 4.814	6277.2877 2.668	6961.4430 4.814	6277.2877 2.668	6961.4430 4.814	6277.2877 2.668	6961.4430 4.814	6277.2877 2.668
6277.3039 2.670	6982.3592 4.801	6277.3039 2.670	6961.4430 4.814	6277.3039 2.670	6961.4430 4.814	6277.3039 2.670	6961.4430 4.814	6277.3039 2.670	6961.4430 4.814	6277.3039 2.670	6961.4430 4.814	6277.3039 2.670	6961.4430 4.814	6277.3039 2.670
6277.3182 2.670	6982.3592 4.801	6277.3182 2.670	6961.4430 4.814	6277.3182 2.670	6961.4430 4.814	6277.3182 2.670	6961.4430 4.814	6277.3182 2.670	6961.4430 4.814	6277.3182 2.670	6961.4430 4.814	6277.3182 2.670	6961.4430 4.814	6277.3182 2.670
6277.3130 2.675	6982.3592 4.801	6277.3130 2.675	6961.4430 4.814	6277.3130 2.675	6961.4430 4.814	6277.3130 2.675	6961.4430 4.814	6277.3130 2.675	6961.4430 4.814	6277.3130 2.675	6961.4430 4.814	6277.3130 2.675	6961.4430 4.814	6277.3130 2.675
6277.3475 2.677	6982.3592 4.801	6277.3475 2.677	6961.4430 4.814	6277.3475 2.677	6961.4430 4.814	6277.3475 2.677	6961.4430 4.814	6277.3475 2.677	6961.4430 4.814	6277.3475 2.677	6961.4430 4.814	6277.3475 2.677	6961.4430 4.814	6277.3475 2.677
6277.3874 2.676	6982.3592 4.801	6277.3874 2.676	6961.4430 4.814	6277.3874 2.676	6961.4430 4.814	6277.3874 2.676	6961.4430 4.814	6277.3874 2.676	6961.4430 4.814	6277.3874 2.676	6961.4430 4.814	6277.3874 2.676	6961.4430 4.814	6277.3874 2.676
6277.4065 2.676	6982.3592 4.801	6277.4065 2.676	6961.4430 4.814	6277.4065 2.676	6961.4430 4.814	6277.4065 2.676	6961.4430 4.814	6277.4065 2.676	6961.4430 4.814	6277.4065 2.676	6961.4430 4.814	6277.4065 2.676	6961.4430 4.814	6277.4065 2.676
6277.4247 2.676	6982.3592 4.801	6277.4247 2.676	6961.4430 4.814	6277.4247 2.676	6961.4430 4.814	6277.4247 2.676	6961.4430 4.814	6277.4247 2.676	6961.4430 4.814	6277.4247 2.676	6961.4430 4.814	6277.4247 2.676	6961.4430 4.814	6277.4247 2.676
6277.4257 2.676	6982.3592 4.801	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676
6277.4257 2.676	6982.3592 4.801	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676	6961.4430 4.814	6277.4257 2.676
6277.4262 2.676	6982.3592 4.801	6277.4262 2.676	6961.4430 4.814	6277.4262 2.676	6961.4430 4.814	6277.4262 2.676	6961.4430 4.814	6277.4262 2.676	6961.4430 4.814	6277.4262 2.676	6961.4430 4.814	6277.4262 2.676	6961.4430 4.814	6277.4262 2.676
6277.4271 2.676	6982.3592 4.801	6277.4271 2.676	6961.4430 4.814	6277.4271 2.676	6961.4430 4.814	6277.4271 2.676	696							

TABLE IIa (*continued*)

HR 451	HJD	b	HJD	b	HR 6510	HJD	b	HJD	b	HR 6747	HJD	b	HJD	b	HR 6510	HJD	b	HJD	b			
	6972.3387	5.280	7328.5936	5.207		6277.3967	2.833	6676.3866	3.124		6691.3507	3.089		6676.4035	3.130		6691.3742	3.093		6676.4035	3.130	
HJD	b		HJD	b	HJD	b		HJD	b	HJD	b		HJD	b	HJD	b		HJD	b			
6961.2414	5.258	6972.5434	5.247	7329.3419	5.203	6271.2252	2.818	6277.4214	2.826	6677.3404	3.110	6692.2260	3.064		6677.3599	3.117	6693.2259	3.049		6677.3599	3.117	
6961.2673	5.267	6972.5646	5.248	7329.3918	5.273	6271.2295	2.836	6277.4285	2.846	6677.3830	3.124	6693.2492	3.042		6677.4044	3.121	6693.2721	3.034		6677.4044	3.121	
6961.3277	5.235	6973.4105	5.268	7329.5100	5.254	6271.3131	2.836	6278.2639	2.839		6680.3172	3.104	6693.2934	3.037		6680.3288	3.104	6693.3185	3.046		6680.3288	3.104
6961.3303	5.251	6973.5615	5.249	7329.5476	5.236	6271.3288	2.839	6278.3624	2.832		6680.3456	3.107	6693.3423	3.039		6680.3587	2.843	6693.3581	3.046		6680.3587	2.843
6961.3992	5.235	6978.2589	5.238	7329.6029	5.201	6271.3423	2.845	6278.3753	2.830		6680.3686	3.107	6693.3481	3.039		6680.3753	2.830	6693.3481	3.039		6680.3753	2.830
6961.4267	5.240	6978.2973	5.238	7330.2900	5.219	6271.3740	2.841	6278.4019	2.828		6680.3885	3.107	6693.3681	3.041		6680.3885	2.828	6693.3681	3.041		6680.3885	2.828
6961.4728	5.248	6978.3389	5.241	7330.3417	5.201	6271.3883	2.841	6278.4145	2.822		6680.3951	3.118	6694.2309	3.005		6680.3951	2.822	6694.2309	3.005		6680.3951	2.822
6962.2798	5.228	6978.3888	5.249	7330.3900	5.199	6271.4018	2.836	6278.4268	2.820		6680.3954	3.116	6694.2543	3.043		6680.3954	2.820	6694.2543	3.043		6680.3954	2.820
6962.3082	5.224	6978.4448	5.267	7330.4527	5.202	6271.4158	2.835	6278.4357	2.818		6680.3956	3.118	6694.2778	3.054		6680.3956	2.818	6694.2778	3.054		6680.3956	2.818
6962.3411	5.241	6978.4971	5.268	7330.5089	5.170	6271.4285	2.832	6279.3406	2.813		6680.3958	3.122	6694.3016	3.153		6680.3958	2.813	6694.3016	3.153		6680.3958	2.813
6962.5534	5.229	6978.5164	5.227	7330.5453	5.186	6271.4424	2.827	6279.3546	2.829		6680.3960	3.124	6694.3309	3.058		6680.3960	2.829	6694.3309	3.058		6680.3960	2.829
6962.5807	5.228	6978.5564	5.249	7330.5817	5.190	6271.3523	2.844	6279.3676	2.823		6680.3962	3.122	6694.3346	3.066		6680.3962	2.823	6694.3346	3.066		6680.3962	2.823
6962.6106	5.226	6978.5755	5.244	7331.2879	5.242	6271.3704	2.841	6279.3897	2.824		6680.3964	3.122	6694.3348	3.066		6680.3964	2.824	6694.3348	3.066		6680.3964	2.824
6965.2246	5.220	6980.3287	5.238	7331.3451	5.250	6271.3862	2.841	6279.3967	2.824		6680.3966	3.116	6694.3351	3.067		6680.3966	2.824	6694.3351	3.067		6680.3966	2.824
6965.2953	5.223	6980.3741	5.237	7331.4013	5.229	6271.3982	2.841	6279.4145	2.822		6680.3968	3.118	6694.3353	3.067		6680.3968	2.822	6694.3353	3.067		6680.3968	2.822
6965.4084	5.263	6980.4126	5.242	7331.4665	5.226	6271.4022	2.838	6279.4226	2.820		6680.3970	3.117	6694.2762	3.023		6680.3970	2.820	6694.2762	3.023		6680.3970	2.820
6965.4352	5.276	6980.4713	5.259	7331.4671	5.223	6271.4022	2.838	6279.4254	2.820		6680.3972	3.104	6694.2993	3.042		6680.3972	2.820	6694.2993	3.042		6680.3972	2.820
6965.4990	5.273	6980.5113	5.263	7331.4677	5.221	6271.4022	2.838	6279.4259	2.820		6680.3974	3.105	6694.3351	3.057		6680.3974	2.820	6694.3351	3.057		6680.3974	2.820
6965.5461	5.249	6980.5421	5.238	7331.4682	5.218	6271.3366	2.850	6281.2660	2.841		6680.3976	3.105	6694.2496	3.051		6680.3976	2.841	6694.2496	3.051		6680.3976	2.841
6965.5914	5.239	6980.5607	5.251	7331.4687	5.221	6271.3523	2.844	6281.2799	2.836		6680.3978	3.106	6694.2731	3.025		6680.3978	2.836	6694.2731	3.025		6680.3978	2.836
6965.6115	5.249	6980.5783	5.247	7331.4692	5.221	6271.3677	2.837	6281.2955	2.829		6680.3980	3.107	6694.2942	3.068		6680.3980	2.829	6694.2942	3.068		6680.3980	2.829
6965.6244	5.259	6981.2610	5.255	7331.5250	5.225	6271.3862	2.831	6281.3107	2.821		6680.3982	3.076	6694.3172	3.086		6680.3982	2.821	6694.3172	3.086		6680.3982	2.821
6966.2777	5.263	6981.2937	5.271	7331.5665	5.218	6271.3862	2.831	6281.3257	2.819		6680.3984	3.076	6694.3429	3.083		6680.3984	2.819	6694.3429	3.083		6680.3984	2.819
6966.3121	5.258	6981.3532	5.268	7332.5677	5.243	6271.4125	2.826	6281.3394	2.818		6680.3986	3.094	6694.2391	3.073		6680.3986	2.818	6694.2391	3.073		6680.3986	2.818
6966.3830	5.239	6981.4029	5.239	7332.5689	5.240	6271.4206	2.825	6281.3536	2.815		6680.3988	3.095	6694.2609	3.069		6680.3988	2.815	6694.2609	3.069		6680.3988	2.815
6966.4142	5.229	6981.4985	5.228	7332.6038	5.240	6271.4217	2.825	6281.3681	2.812		6680.3990	3.096	6697.3071	3.076		6680.3990	2.812	6697.3071	3.076		6680.3990	2.812
6966.4430	5.233	6981.5313	5.239			6271.4263	2.825	6281.3967	2.812		6680.3992	3.097	6697.3317	3.088		6680.3992	2.812	6697.3317	3.088		6680.3992	2.812
6966.4856	5.256	6981.5542	5.263			6271.4319	2.825	6281.4097	2.812		6680.3994	3.098	6697.3355	3.089		6680.3994	2.812	6697.3355	3.089		6680.3994	2.812
6966.5254	5.257	6982.2914	5.222			6271.3862	2.830	6281.4229	2.817		6680.3996	3.098	6698.2447	3.047		6680.3996	2.817	6698.2447	3.047		6680.3996	2.817
6966.5659	5.255	6982.3316	5.246			6271.3862	2.830	6281.4324	2.819		6680.3998	3.098	6698.2658	3.063		6680.3998	2.819	6698.2658	3.063		6680.3998	2.819
6966.5866	5.255	6982.3702	5.282			6271.4073	2.824	6282.3377	2.820		6680.3999	3.098	6698.2782	3.074		6680.3999	2.820	6698.2782	3.074		6680.3999	2.820
6966.6136	5.258	6982.4745	5.279			6271.3928	2.826	6282.3501	2.820		6680.3999	3.099	6698.3325	3.095		6680.3999	2.820	6698.3325	3.095		6680.3999	2.820
6967.2647	5.265	6983.2569	5.268			6271.3705	2.850	6282.3741	2.820		6680.3999	3.103	6699.2414	3.065		6680.3999	2.820	6699.2414	3.065		6680.3999	2.820
6967.2747	5.265	6983.2925	5.258			6271.3243	2.853	6282.4009	2.822		6680.3999	3.104	6699.2462	3.066		6680.3999	2.822	6699.2462	3.066		6680.3999	2.822
6967.3113	5.261	6983.3381	5.249			6271.3384	2.852	6282.4130	2.824		6680.3999	3.104	6699.2532	3.072		6680.3999	2.852	6699.2532	3.072		6680.3999	2.852
6967.3411	5.249	6983.4237	5.257			6271.3564	2.843	6282.4245	2.826		6680.3999	3.104	6700.2415	3.072		6680.3999	2.843	6700.2415	3.072		6680.3999	2.843
6967.3772	5.249	6983.4397	5.240			6271.3678	2.850	6282.4287	2.821		6680.3999	3.105	6700.2861	3.066		6680.3999	2.821	6700.2861	3.066		6680.3999	2.821
6967.4039	5.249	6983.4801	5.240			6271.3809	2.851	6282.4309	2.824		6680.3999	3.106	6703.3059	3.089		6680.3999	2.851	6703.3059	3.089		6680.3999	2.851
6967.4320	5.249	6983.5001	5.216			6271.3822	2.850	6282.4387	2.824		6680.3999	3.106	6703.3286	3.089		6680.3999	2.850	6703.3286	3.089		6680.3999	2.850
6967.4620	5.249	6983.5301	5.216			6271.3962	2.851	6282.4406	2.820		6680.3999	3.106	6703.3460	3.084		6680.3999	2.851	6703.3460	3.084		6680.3999	2.851
6967.4981	5.249	6983.5512	5.216			6271.3992	2.851	6282.4406	2.820		6680.3999	3.106	6703.3592	3.089		6680.3999	2.851	6703.3592	3.089		6680.3999	2.851
6967.5252	5.249	6983.5832	5.216			6271.3720	2.852	6282.4406	2.820		6680.3999	3.106	6703.3794	3.089		6680.3999	2.852	6703.3794	3.089		6680.3999	2.852
6967.5520	5.249	6983.6034	5.216			6271.3776	2.852	6282.4406	2.820		6680.3999	3.106	6703.3974	3.089		6680.3999	2.852	6703.3974	3.089		6680.3999	2.852
6967.5828	5.249	6983.6314	5.216			6271.3878	2.852	6282.4406	2.820		6680.3999	3.106	6703.4150	3.089		6680.3999	2.852	6703.4150	3.089		6680.3999	2.852
6967.6137	5.249	6983.6764	5.216			6271.3932	2.852	6282.4406	2.820		6680.3999	3.106	6703.4349	3.089		6680.3999	2.852	6703.4349	3.089		6680.3999	2.852
6967.6457	5.249	6983.7001	5.216			6271.3992	2.852	6282.4406	2.820		6680.3999	3.106	6703.4523	3.089		6680.3999	2.852	6703.4523	3.089		6680.3999	2.852
6967.6747	5.249	6983.7314	5.216			6271.4050	2.852	6282.4406	2.820	</												

J. Cuypers *et al.*

TABLE IIa (*continued*)

HR 8260	HJD	b	HJD	b	HJD	b	HR 8260	HJD	b	HJD	b	HJD	b
6278.6244	4.414	6363.3383	4.423	6682.4108	4.434	6696.3946	4.439	6962.5653	4.369	6978.5594	4.353		
6278.6377	4.413	6363.3541	4.423	6682.4259	4.430	6696.4163	4.432	6962.5880	4.369	6978.5786	4.353		
6271.4769	4.431	6378.6490	4.425	6683.3699	4.419	6691.2644	4.424	6962.4976	4.369	6978.5939	4.356		
6271.4901	4.432	6281.4625	4.427	6683.3847	4.421	6691.2874	4.427	6967.2471	4.425	6980.4281	4.352		
6271.5035	4.423	6281.4756	4.428	6363.4002	4.423	6682.4736	4.435	6967.2690	4.426	6980.4863	4.350		
6271.5166	4.429	6281.4870	4.426	6364.2602	4.421	6682.4886	4.433	6691.3333	4.428	6980.5245	4.350		
6272.5344	4.428	6281.4989	4.426	6364.2740	4.419	6682.5042	4.436	6691.3585	4.432	6980.5453	4.351		
6272.5557	4.431	6281.5119	4.422	6367.2909	4.446	6682.5199	4.435	6691.3818	4.432	6980.5630	4.349		
6272.5697	4.432	6281.5236	4.421	6367.3075	4.453	6682.5357	4.437	6691.4031	4.432	6980.5810	4.345		
6272.5832	4.430	6281.5368	4.418	6676.3762	4.437	6682.5536	4.436	6691.4214	4.432	6980.5956	4.343		
6274.5166	4.438	6281.5499	4.417	6676.3932	4.437	6684.3218	4.424	6691.4387	4.434	6980.6085	4.341		
6274.5292	4.439	6281.5619	4.417	6676.4102	4.437	6684.3414	4.423	6691.4624	4.435	6980.6229	4.340		
6274.5401	4.437	6281.5750	4.417	6676.4275	4.440	6684.3629	4.427	6691.4803	4.436	6980.6362	4.341		
6274.5515	4.435	6281.5861	4.412	6676.4440	4.443	6684.3850	4.430	6691.4969	4.438	6980.6483	4.342		
6276.4542	4.417	6281.5973	4.411	6678.4662	4.442	6684.4088	4.427	6691.5144	4.434	6980.6860	4.341		
6276.4677	4.415	6281.6084	4.409	6678.4820	4.447	6684.4260	4.433	6692.4233	4.438	6980.6999	4.346		
6276.4794	4.414	6281.6205	4.407	6676.3037	4.451	6684.4416	4.431	6692.4465	4.436	6980.6962	4.347		
6276.4907	4.414	6281.6313	4.410	6676.3202	4.448	6684.4585	4.414	6692.4671	4.434	6980.7166	4.346		
6276.5022	4.417	6282.4729	4.411	6676.3571	4.447	6684.5027	4.454	6692.4892	4.438	6980.7273	4.343		
6276.5141	4.417	6282.4844	4.408	6676.3556	4.448	6684.5179	4.436	6692.5119	4.438	6700.2490	4.431		
6276.5252	4.415	6282.4960	4.411	6677.3473	4.414	6686.2367	4.422	6693.2342	4.423	6700.2720	4.430		
6276.5377	4.415	6282.5070	4.413	6677.3667	4.434	6686.2546	4.425	6693.2578	4.424	6700.2946	4.433		
6276.5503	4.417	6282.5181	4.417	6677.3898	4.436	6686.2722	4.424	6693.2803	4.425	6700.3190	4.433		
6276.5633	4.414	6282.5294	4.419	6677.4109	4.435	6686.2895	4.428	6693.3030	4.433	6700.3453	4.422		
6276.5745	4.415	6282.5400	4.421	6677.4228	4.441	6686.3068	4.427	6693.3262	4.434	6700.4444	4.432		
6276.5848	4.417	6282.5508	4.422	6677.4454	4.448	6686.3234	4.430	6693.3524	4.436	6700.4665	4.431		
6276.5973	4.413	6282.5673	4.421	6677.4616	4.443	6686.3406	4.432	6693.3766	4.437	6702.4865	4.435		
6276.6082	4.409	6282.5800	4.424	6677.4783	4.447	6686.3599	4.429	6693.3996	4.438	6703.2703	4.435		
6276.6183	4.411	6282.5926	4.422	6677.4977	4.453	6686.3786	4.429	6693.4194	4.437	6703.2910	4.426		
6276.6287	4.412	6282.6047	4.424	6677.5132	4.451	6686.3970	4.433	6693.4389	4.438	6703.3147	4.424		
6276.6391	4.413	6282.6166	4.421	6677.5323	4.440	6686.4127	4.435	6693.4650	4.444	6703.3733	4.433		
6276.6494	4.414	6282.6286	4.420	6677.5538	4.462	6686.4279	4.436	6693.4876	4.445	6703.3569	4.433		
6277.4611	4.413	6282.6401	4.419	6680.2267	4.427	6686.4432	4.436	6693.5092	4.456	6703.3779	4.434		
6277.4735	4.412	6334.3637	4.563	6680.2489	4.429	6686.4578	4.439	6694.2395	4.419	6703.3984	4.434		
6277.4864	4.415	6334.3843	4.563	6680.2654	4.430	6686.4735	4.442	6694.2624	4.421	6703.4184	4.436		
6277.4961	4.417	6334.4028	4.563	6680.2845	4.429	6686.4890	4.439	6694.2857	4.434	6703.4393	4.433		
6277.5100	4.417	6334.4203	4.560	6680.3027	4.430	6686.5066	4.442	6694.3094	4.418	6703.4722	4.433		
6277.5190	4.417	6334.4368	4.554	6680.3205	4.436	6686.5228	4.441	6694.3387	4.432	6703.4944	4.438		
6277.5305	4.422	6334.4543	4.555	6680.3388	4.436	6686.5397	4.444	6694.3664	4.429	6703.5130	4.437		
6277.5416	4.419	6334.4702	4.549	6680.3560	4.434	6686.3208	4.401	6694.3982	4.431	6703.5570	4.436		
6277.5560	4.415	6334.4874	4.546	6680.3777	4.436	6686.2540	4.430	6695.3092	4.456	6703.5779	4.438		
6277.6151	4.421	6338.3944	4.490	6680.3996	4.437	6686.2716	4.415	6695.4356	4.434	6703.5984	4.438		
6277.6267	4.416	6338.4120	4.493	6680.4287	4.441	6686.2896	4.417	6695.2616	4.424	6703.6183	4.437		
6277.6372	4.420	6338.4271	4.493	6680.4388	4.442	6686.3076	4.421	6695.2841	4.444	6703.6395	4.437		
6277.6475	4.419	6338.4433	4.483	6680.4500	4.441	6686.3267	4.423	6695.3071	4.442	6703.6608	4.436		
6277.6552	4.410	6338.4580	4.488	6680.4736	4.442	6690.2412	4.418	6695.3143	4.436	6703.6820	4.436		
6278.4497	4.423	6338.4731	4.481	6680.4975	4.446	6690.2719	4.430	6695.3364	4.429	6703.7037	4.448		
6278.4872	4.419	6339.3485	4.467	6690.5145	4.445	6690.2971	4.428	6695.3833	4.435	6703.7222	4.442		
6278.5007	4.411	6339.3567	4.467	6692.2295	4.418	6690.3220	4.433	6695.4048	4.438	6703.7467	4.442		
6278.5137	4.413	6339.3813	4.468	6692.2506	4.418	6690.3587	4.430	6695.4267	4.426	6703.7657	4.442		
6278.5262	4.416	6339.4034	4.471	6692.2683	4.417	6690.3862	4.428	6695.4505	4.440	6703.6241	4.433		
6278.5389	4.408	6339.4191	4.471	6692.2882	4.421	6690.4145	4.434	6695.5063	4.434	6703.6471	4.434		
6278.5513	4.412	6339.4343	4.471	6692.3012	4.422	6690.4378	4.435	6696.2577	4.430	6703.6689	4.431		
6278.5664	4.410	6339.4561	4.465	6692.3187	4.425	6690.4664	4.436	6695.3605	4.435	6703.6904	4.432		
6278.5768	4.407	6363.2663	4.424	6692.3336	4.422	6690.4852	4.432	6695.3894	4.429	6703.7123	4.432		
6278.5890	4.407	6363.2835	4.423	6692.3565	4.426	6690.5067	4.427	6695.3949	4.429	6703.7341	4.432		
6278.6000	4.401	6363.3067	4.423	6692.3739	4.427	6690.5271	4.425	6695.3510	4.433	6703.7558	4.432		
6278.6106	4.401	6363.3226	4.424	6692.3924	4.427	6690.5475	4.425	6695.3765	4.430	6703.7768	4.432		

HR	8402	HJD		HJD		HJD			
		b	b	b	b	b	b		
		HJD	b	HJD	b	HJD	b		
		6278.	6184	6.710	6363.	3791	4.677		
7331.	4200	4.521		6363.	3940	4.675	6682. 5099	4.685	
7331.	4868	4.526		6364.	2530	4.673	6682. 5250	4.684	
7331.	5534	4.540		6364.	2703	4.671	6684. 5410	4.681	
7332.	5856	4.527		6361.	2850	4.676	6684. 3468	4.688	
		6271.	4495	4.708	6361.	4815	4.717	6684. 3641	4.680
		6271.	4840	4.709	6361.	4952	4.718	6684. 3271	4.687
		6271.	4968	4.706	6361.	4699	4.718	6684. 3271	4.687
		6271.	5103	4.707	6361.	4815	4.717	6684. 3011	4.673
		6271.	5128	4.706	6361.	4932	4.715	6684. 3097	4.669
		6272.	5376	4.697	6361.	5053	4.717	6684. 4311	4.676
		6272.	5502	4.697	6361.	5179	4.716	6676. 3984	4.690
		6272.	5619	4.696	6361.	5311	4.716	6676. 4153	4.686
		6274.	5097	4.706	6361.	5429	4.716	6676. 4321	4.688
		6274.	5229	4.706	6361.	5558	4.711	6676. 4503	4.687
		6274.	5348	4.707	6361.	5676	4.711	6676. 4677	4.687
		6274.	5463	4.705	6361.	5805	4.711	6676. 4895	4.688
		6276.	4472	4.714	6361.	5921	4.710	6676. 5090	4.684
		6276.	4622	4.715	6361.	6030	4.706	6676. 5256	4.689
		6276.	4743	4.709	6361.	6141	4.710	6676. 5424	4.686
		6276.	4854	4.710	6361.	6260	4.709	6676. 3571	4.679
		6276.	4967	4.710	6362.	4668	4.701	6677. 3713	4.685
		6276.	5086	4.711	6362.	4792	4.701	6677. 3943	4.682
		6276.	5198	4.705	6362.	4909	4.700	6677. 4159	4.683
		6276.	5324	4.706	6362.	5021	4.701	6677. 4337	4.691
		6276.	5447	4.711	6362.	5131	4.700	6677. 4501	4.685
		6276.	5569	4.710	6362.	5241	4.699	6677. 4665	4.683
		6276.	5691	4.711	6362.	5352	4.699	6677. 4833	4.686
		6276.	5796	4.707	6362.	5455	4.697	6677. 5024	4.686
		6276.	5923	4.706	6362.	5617	4.697	6677. 5185	4.681
		6276.	6025	4.708	6362.	5750	4.695	6677. 5373	4.677
		6276.	6133	4.711	6362.	5877	4.696	6680. 2341	4.679
		6276.	6236	4.706	6362.	5999	4.697	6680. 2543	4.688
		6276.	6339	4.704	6362.	6119	4.697	6680. 2705	4.688
		6276.	6442	4.712	6362.	6238	4.696	6680. 2897	4.686
		6277.	4547	4.700	6334.	3772	4.715	6680. 3075	4.686
		6277.	4672	4.702	6334.	3944	4.715	6680. 3258	4.684
		6277.	4792	4.698	6334.	4134	4.716	6680. 3433	4.685
		6277.	4910	4.702	6334.	4301	4.714	6680. 3612	4.687
		6277.	5017	4.697	6334.	4465	4.709	6680. 3825	4.686
		6277.	5130	4.699	6334.	4633	4.713	6680. 4047	4.685
		6277.	5252	4.700	6334.	4801	4.710	6680. 4230	4.683
		6277.	5364	4.696	6338.	3871	4.668	6680. 4431	4.683
		6277.	5494	4.697	6338.	4050	4.671	6680. 4860	4.684
		6277.	6101	4.698	6338.	4209	4.672	6680. 5026	4.683
		6277.	6205	4.698	6338.	4370	4.672	6680. 5202	4.684
		6277.	6322	4.699	6338.	4518	4.671	6682. 2367	4.679
		6277.	6430	4.699	6338.	4667	4.675	6682. 2559	4.676
		6278.	4425	4.707	6339.	3413	4.681	6682. 2731	4.687
		6278.	4568	4.709	6339.	3596	4.693	6682. 3064	4.684
		6278.	4815	4.712	6339.	3747	4.689	6682. 3235	4.684
		6278.	4939	4.711	6339.	3953	4.684	6682. 3401	4.686
		6278.	5073	4.716	6339.	4128	4.685	6682. 3590	4.683
		6278.	5203	4.717	6339.	4277	4.680	6682. 3787	4.682
		6278.	5325	4.716	6363.	2576	4.672	6682. 3975	4.682
		6278.	5450	4.717	6363.	2774	4.674	6682. 4154	4.691
		6278.	5575	4.713	6363.	3001	4.673	6682. 4303	4.686
		6278.	5706	4.719	6363.	3167	4.674	6682. 4464	4.685
		6278.	5828	4.710	6363.	3323	4.677	6682. 4624	4.688
		6278.	5950	4.711	6363.	3479	4.677	6682. 4788	4.683
		6278.	6067	4.709	6363.	3637	4.677	6682. 4933	4.682

## PHOTOMETRY OF BE STARS

TABLE IIa (*continued*)

HR 8402	HJD	b	HJD	b	HJD	b	HR 8408	HJD	b	HJD	b	HJD	b
	6698.3691	4.692	6966.6191	4.659	6981.5856	4.665		6278.5217	5.895	6334.4154	5.896	6680.2855	5.884
HJD	b		6699.2565	4.693	6966.6404	4.660	HJD	b		6334.4320	5.897	6680.3037	5.887
6691.5027	4.679	6700.2559	4.678	6966.6558	4.659	6981.6224	4.668	6278.5341	5.898	6334.4482	5.899	6680.3215	5.890
6691.5201	4.683	6700.2782	4.677	6967.4561	4.669	6981.6369	4.664	6271.4856	5.897	6334.4653	5.901	6680.3394	5.894
6692.2414	4.687	6700.3009	4.680	6967.4997	4.668	6981.6522	4.664	6271.4993	5.896	6334.4823	5.901	6680.3569	5.897
6692.4314	4.686	6700.3255	4.679	6967.5492	4.677	6981.6646	4.661	6271.5121	5.900	6338.3888	5.901	6680.3786	5.897
6692.4528	4.681	6700.3517	4.680	6967.5729	4.677	6981.5965	4.673	6271.5246	5.900	6338.4068	5.899	6680.4007	5.895
6692.4738	4.684	6702.4504	4.676	6967.5882	4.675	6983.6142	4.677	6272.5392	5.893	6278.6083	5.904	6338.4227	5.896
6692.4963	4.685	6702.4734	4.672	6967.6051	4.679	6983.6356	4.671	6272.5518	5.895	6278.6203	5.908	6338.4387	5.896
6692.5195	4.689	6703.2767	4.688	6967.6230	4.676	6983.6486	4.669	6272.5634	5.897	6278.6335	5.907	6338.4535	5.897
6693.2409	4.681	6703.2985	4.689	6967.6353	4.677	6983.6601	4.669	6272.5791	5.896	6278.6444	5.910	6338.4682	5.892
6693.2641	4.686	6703.3208	4.689	6967.6460	4.678	7322.5149	4.669	6274.5119	5.885	6278.6569	5.914	6680.4747	5.881
6693.2867	4.671	6703.3435	4.690	6967.6562	4.679	7322.5751	4.664	6274.5245	5.887	6278.6635	5.910	6680.4988	5.877
6693.3090	4.688	6703.3630	4.692	6970.5504	4.682	7322.6182	4.667	6274.5364	5.886	6278.6691	5.904	6339.3163	5.907
6693.3231	4.691	6703.3839	4.694	6970.5769	4.681	7327.6153	4.670	6274.6572	5.909	6339.4144	5.887	6682.2516	5.876
6693.3582	4.689	6703.4044	4.695	6970.6098	4.682	7328.5772	4.674	6276.4492	5.884	6339.4292	5.879	6682.2693	5.876
6693.3829	4.690	6703.4247	4.696	6970.6338	4.680	7328.6177	4.673	6276.4637	5.887	6281.4712	5.897	6682.2857	5.879
6693.4058	4.692	6703.4592	4.692	6970.6501	4.687	7328.6623	4.668	6276.4757	5.883	6281.4829	5.892	6682.3021	5.879
6693.4258	4.686	6703.4783	4.689	6972.5303	4.674	7329.5888	4.667	6276.4870	5.886	6281.4947	5.890	6363.2789	5.906
6693.4447	4.689	6959.4507	4.654	6972.5521	4.676	7329.6242	4.672	6276.4983	5.885	6281.5080	5.893	6363.3019	5.904
6693.4731	4.700	6959.5155	4.672	6972.5749	4.674	7329.6704	4.669	6276.5103	5.889	6281.5194	5.893	6363.3182	5.901
6693.4939	4.704	6959.5324	4.675	6972.5941	4.678	7329.6818	4.668	6276.5216	5.889	6281.5322	5.891	6363.3237	5.899
6693.5153	4.705	6959.5942	4.665	6972.6075	4.675	7330.5685	4.670	6276.5458	5.893	6363.3494	5.893	6682.3933	5.879
6694.2694	4.680	6959.6164	4.671	6972.6206	4.674	7330.6022	4.667	6276.5463	5.893	6363.3580	5.886	6682.4267	5.880
6694.2686	4.684	6959.6358	4.668	6972.6340	4.677	7332.5907	4.671	6276.5706	5.901	6281.5820	5.893	6363.3954	5.883
6694.2927	4.688	6959.6519	4.671	6972.6463	4.679	6975.5469	4.684	6276.5756	5.902	6281.5933	5.892	6363.4252	5.882
6694.3164	4.682	6959.6642	4.688	6972.6590	4.678	6975.5850	4.688	6276.5813	5.900	6281.5935	5.893	6363.4575	5.877
6694.3452	4.679	6959.6770	4.669	6972.6742	4.675	6975.6166	4.676	6276.5957	5.897	6281.6045	5.902	6363.4779	5.883
6694.3711	4.695	6960.2520	4.680	6972.6845	4.678	6975.6463	4.673	6276.6040	5.902	6281.6135	5.906	6367.2865	5.892
6694.3996	4.691	6960.3567	4.691	6972.6970	4.676	6975.6666	4.678	6276.6147	5.905	6281.6223	5.894	6682.5057	5.871
6694.4177	4.701	6960.5930	4.674	6975.5713	4.680	6975.7231	4.680	6276.6251	5.904	6281.6374	5.898	6682.5208	5.871
6695.3003	4.685	6960.6104	4.679	6975.5887	4.679	6980.6442	4.670	6276.6355	5.905	6281.6433	5.895	6682.5365	5.871
6695.3670	4.686	6960.6282	4.684	6975.6034	4.677	6980.6610	4.674	6276.6456	5.902	6281.6495	5.895	6682.5544	5.873
6695.3984	4.684	6960.6410	4.677	6975.6166	4.674	6980.6781	4.672	6276.6554	5.895	6281.6554	5.892	6682.5730	5.870
6695.4112	4.686	6960.6523	4.674	6975.6332	4.676	6980.6962	4.674	6276.6597	5.906	6281.6613	5.894	6684.3227	5.894
6695.4332	4.688	6960.6648	4.676	6975.6472	4.675	6980.7049	4.676	6277.4562	5.903	6281.6673	5.896	6686.4344	5.893
6695.4369	4.692	6960.6766	4.663	6975.6613	4.678	6980.7123	4.674	6277.4696	5.904	6281.6742	5.887	6686.4369	5.889
6695.2534	4.682	6961.4431	4.687	6978.4735	4.684	6980.7321	4.675	6277.4805	5.903	6281.6807	5.890	6686.3988	5.889
6695.3081	4.690	6961.4860	4.680	6978.5244	4.679	6980.7531	4.674	6277.4925	5.902	6281.6922	5.892	6686.4100	5.887
6695.3310	4.694	6962.5431	4.667	6978.5453	4.677	6980.7662	4.670	6277.5031	5.901	6281.6997	5.895	6686.4270	5.886
6695.3573	4.692	6962.5707	4.666	6978.5647	4.677	6980.7782	4.671	6277.5145	5.900	6281.6915	5.892	6686.4425	5.875
6696.0008	4.664	6962.6235	4.671	6978.6002	4.674	6980.7973	4.674	6277.5268	5.902	6281.5257	5.899	6686.4595	5.857
6696.4122	4.693	6962.6440	4.678	6978.6490	4.674	6980.8114	4.676	6277.5597	5.902	6281.6404	5.898	6686.4744	5.872
6696.4436	4.685	6963.2430	4.680	6978.6960	4.681	6980.8383	4.671	6277.5698	5.907	6281.6943	5.891	6686.4994	5.872
6695.4661	5.884	6963.3554	5.884	6978.7580	5.887	6980.8512	4.681	6277.5825	5.908	6281.6987	5.893	6686.5191	5.862
6695.4999	4.676	6965.4699	4.676	6980.5499	4.682	6980.9203	4.676	6277.5947	5.909	6282.0407	5.890	6686.2558	5.881
6695.5189	4.682	6968.5698	4.681	6980.5869	4.681	6980.9339	4.679	6277.6057	5.909	6282.0528	5.897	6686.2732	5.883
6695.5292	4.687	6965.5617	4.672	6980.5880	4.682	6980.9455	4.678	6277.6143	5.905	6282.0630	5.895	6686.2905	5.886
6696.3570	5.879	6965.6202	4.683	6980.6009	4.680	6980.9636	4.678	6277.6266	5.902	6282.0743	5.892	6686.3373	5.873
6697.3224	4.687	6965.6266	4.677	6980.6138	4.681	6980.9764	4.678	6277.6388	5.902	6282.0852	5.892	6686.3543	5.872
6697.3712	4.677	6965.6423	4.678	6980.6284	4.684	6980.9890	4.674	6277.6507	5.908	6282.0971	5.898	6686.3745	5.872
6697.3925	4.678	6965.6570	4.679	6980.6410	4.679	6980.9998	4.671	6277.6625	5.902	6282.1080	5.897	6686.3957	5.872
6697.4133	4.675	6966.4584	4.672	6980.6532	4.677	6980.9991	4.677	6277.6716	5.897	6282.1199	5.896	6686.4136	5.880
6697.4254	4.685	6965.4843	4.680	6980.6830	4.680	6980.9998	4.678	6277.6829	5.906	6282.1318	5.897	6686.4289	5.881
6697.4393	5.874	6965.3083	5.876	6980.6919	5.903	6980.9995	5.907	6277.6946	5.907	6282.1437	5.897	6686.4367	5.882
6698.2733	5.888	6965.3354	5.885	6980.6918	5.913	6980.9998	5.905	6278.5143	5.901	6282.1557	5.897	6686.4553	5.884
6698.2884	5.882	6965.3620	5.885	6980.6947	5.905	6980.9998	5.904	6278.5284	5.902	6282.1677	5.896	6686.4744	5.883
6698.3258	5.892	6965.3844	5.886	6980.7019	5.890	6980.9998	5.906	6278.5466	5.905	6282.1796	5.895	6686.4923	5.882
6698.3578	5.882	6965.4060	5.888	6980.7139	5.908	6980.9998	5.906	6278.5667	5.905	6282.1915	5.894	6686.5113	5.882
6698.3975	5.878	6965.4203	5.882	6980.7260	5.904	6980.9998	5.906	6278.5886	5.905	6282.1934	5.894	6686.5303	5.882
6698.4384	5.878	6965.4517	5.874	6980.7364	5.905	6980.9998	5.905	6278.6004	5.905	6282.2052	5.894	6686.5395	5.881
6698.4495	5.872	6965.4517	5.875	6980.7494	5.905	6980.9998	5.905	6278.6122	5.904	6282.2171	5.894	6686.5587	5.880
6698.4593	5.873	6965.5080	5.879	6980.7610	5.905	6980.9998	5.905	6278.6240	5.904	6282.2293	5.894	6686.5774	5.881
6698.4874	5												

TABLE IIa (*continued*)

HR 8858	HJD	b	HJD	b	HJD	b	HR 8858	HJD	b
	6682.4665 4.346		6692.5017 4.339		6700.2841 4.344			6983.6001 4.328	
HJD	b		6682.4823 4.343		6692.5257 4.346		HJD	b	
6676.3843 4.324			6682.4971 4.341		6693.2454 4.346		6983.6067 4.320		
6676.4019 4.329			6682.5135 4.343		6693.2693 4.345		6983.6253 4.323		
6676.4192 4.330			6682.5289 4.344		6693.2913 4.354		6983.6310 4.322		
6676.4355 4.333			6682.5453 4.346		6693.3143 4.352		6983.6477 4.319		
6676.4539 4.337			6682.5615 4.347		6693.3366 4.356		6983.6520 4.328		
6676.4716 4.339			6682.5776 4.347		6693.3627 4.349		6983.6567 4.337		
6676.4939 4.343			6682.5908 4.344		6693.3874 4.353		6983.6601 4.321		
6676.5127 4.342			6682.6036 4.344		6693.4107 4.346		6983.6619 4.324		
6676.5274 4.349			6684.3513 4.330		6693.4303 4.348		6983.6638 4.324		
6676.5464 4.343			6684.3726 4.329		6693.4502 4.339		6983.6650 4.322		
6676.5623 4.342			6684.3962 4.322		6693.4755 4.335		6983.6665 4.322		
6676.5747 4.347			6684.4185 4.331		6693.4987 4.335		6983.6689 4.324		
6676.5898 4.341			6684.4352 4.329		6693.5202 4.325		6983.6722 4.327		
6676.6024 4.341			6684.4512 4.338		6693.5372 4.316		6983.6732 4.326		
6677.3567 4.344			6684.4962 4.339		6693.5524 4.312		6983.6742 4.325		
6677.3754 4.355			6686.2812 4.343		6693.5647 4.309		6983.6760 4.324		
6677.3980 4.328			6686.2987 4.343		6694.2514 4.320		6983.6770 4.323		
6677.4197 4.327			6686.3157 4.342		6694.2981 4.333		6983.6780 4.322		
6677.4380 4.319			6686.3317 4.345		6694.3221 4.338		6983.6790 4.321		
6677.4543 4.324			6686.3492 4.347		6694.3513 4.341		6983.6808 4.320		
6677.4705 4.310			6686.3680 4.347		6694.3779 4.324		6983.6826 4.319		
6677.4879 4.322			6686.3874 4.351		6694.3993 4.334		6983.6845 4.318		
6677.5059 4.319			6686.4054 4.349		6694.4226 4.336		6983.6864 4.317		
6677.5223 4.317			6686.4219 4.352		6694.4471 4.327		6983.6883 4.316		
6677.5412 4.333			6686.4365 4.353		6695.2731 4.336		6983.6902 4.315		
6677.5623 4.332			6686.4519 4.354		6695.3186 4.330		6983.6921 4.314		
6677.5792 4.326			6686.4668 4.352		6695.3453 4.320		6983.6940 4.313		
6680.2589 4.324			6686.4823 4.348		6695.3725 4.345		6983.6959 4.312		
6680.2748 4.322			6686.4994 4.348		6695.3940 4.347		6983.6978 4.311		
6680.2937 4.334			6686.5167 4.339		6696.4162 4.349		6983.6997 4.310		
6680.3115 4.334			6686.5325 4.337		6695.4384 4.349		6983.7016 4.309		
6680.3296 4.329			6686.5485 4.332		6695.5187 4.359		6983.7035 4.308		
6680.3475 4.330			6686.5592 4.333		6695.5354 4.370		6983.7054 4.307		
6680.3650 4.338			6686.5702 4.328		6696.2699 4.332		6983.7073 4.306		
6680.3865 4.343			6686.5829 4.330		6696.2919 4.333		6983.7092 4.305		
6680.4088 4.348			6686.5947 4.329		6696.3132 4.342		6983.7111 4.304		
6680.4312 4.352			6688.2442 4.306		6696.3362 4.344		6983.7130 4.303		
6680.4471 4.362			6688.2636 4.314		6696.3629 4.341		6983.7149 4.302		
6680.4639 4.359			6688.2995 4.336		6696.3845 4.341		6983.7168 4.301		
6680.4906 4.353			6688.3175 4.331		6696.4058 4.336		6983.7187 4.300		
6680.5070 4.354			6688.3360 4.335		6696.4280 4.345		6983.7206 4.299		
6680.5412 4.353			6690.3101 4.330		6696.4491 4.354		6983.7225 4.298		
6680.5580 4.350			6690.3343 4.336		6697.2588 4.334		6983.7244 4.297		
6680.5804 4.366			6690.3702 4.323		6697.2817 4.331		6983.7263 4.296		
6682.2413 4.346			6690.3996 4.335		6697.3046 4.332		6983.7282 4.295		
6682.2603 4.342			6690.4272 4.327		6697.3287 4.335		6983.7301 4.294		
6682.2768 4.337			6690.4497 4.327		6697.3509 4.334		6983.7320 4.293		
6682.2937 4.339			6690.4753 4.340		6697.3769 4.334		6983.7339 4.292		
6682.3106 4.335			6690.4962 4.335		6697.3983 4.342		6983.7358 4.291		
6682.3271 4.335			6690.5167 4.342		6697.4195 4.339		6983.7377 4.290		
6682.3439 4.338			6690.5384 4.340		6698.2647 4.349		6983.7396 4.289		
6682.3614 4.316			6690.5549 4.349		6698.2863 4.349		6983.7415 4.288		
6682.4011 4.328			6690.5787 4.353		6698.3081 4.344		6983.7434 4.287		
6682.4189 4.352			6692.4370 4.346		6698.3303 4.340		6983.7453 4.286		
6682.4346 4.338			6692.4575 4.344		6698.3537 4.340		6983.7472 4.285		
6682.4501 4.339			6692.4789 4.342		6700.2636 4.353		6983.7491 4.284		
							6983.6569 4.325		

TABLE IIb. — *Four-colour Strömgren photometry for periodic Be stars observed at ESO. Observations between JD2446991 and 2447000 were obtained at SAAO. The heliocentric Julian day is with respect to JD 2440000.000.*

HR 5193	HJD	u	v	b	y	HR 5440	HJD	u	v	b	y			
6961.6036	3.954	3.687	3.418	3.444	6973.5925	3.874	3.661	3.365	3.387	6979.4842	2.839	2.599	2.326	2.348
6961.6330	3.953	3.685	3.416	3.444	6973.6791	3.869	3.634	3.358	3.378	6979.5299	2.835	2.603	2.326	2.350
6961.6522	3.952	3.685	3.417	3.446	6975.5980	3.942	3.674	3.399	3.430	6979.5722	2.822	2.592	2.318	2.342
6961.6780	3.950	3.684	3.416	3.444	6975.6156	3.938	3.693	3.429	3.438	6979.6114	2.800	2.569	2.294	2.321
6961.7008	3.948	3.679	3.410	3.438	6975.6553	3.933	3.661	3.406	3.417	6979.6440	2.790	2.561	2.287	2.314
6961.7379	3.947	3.680	3.412	3.442	6979.4812	3.859	3.688	3.413	3.443	6980.4842	2.823	2.575	2.311	2.341
6961.5516	3.926	3.682	3.400	3.434	6979.3282	3.855	3.685	3.408	3.439	6980.5227	2.832	2.582	2.316	2.344
6961.5675	3.929	3.665	3.394	3.420	6979.5701	3.957	3.689	3.411	3.441	6980.5680	2.834	2.583	2.315	2.344
6961.5856	3.929	3.668	3.395	3.421	6979.6095	3.957	3.685	3.408	3.441	6980.6065	2.835	2.582	2.315	2.343
6961.6102	3.929	3.670	3.397	3.423	6979.6422	3.957	3.683	3.406	3.438	6980.6690	2.819	2.584	2.317	2.343
6961.5185	3.933	3.668	3.395	3.422	6980.4793	3.938	3.676	3.409	3.443	6980.7219	2.83	2.587	2.316	2.347
6961.6633	3.933	3.670	3.395	3.422	6980.4823	3.939	3.680	3.413	3.446	6981.4893	2.816	2.584	2.316	2.341
6961.6956	3.935	3.673	3.401	3.427	6980.5207	3.957	3.677	3.410	3.442	6981.5468	2.783	2.558	2.289	2.315
6961.7329	3.935	3.672	3.398	3.425	6980.5654	3.962	3.682	3.414	3.444	6981.5860	2.780	2.544	2.277	2.303
6961.5115	3.918	3.654	3.384	3.410	6980.6022	3.960	3.682	3.412	3.443	6981.6266	2.782	2.553	2.283	2.311
6961.5525	3.918	3.650	3.378	3.405	6980.6627	3.958	3.679	3.409	3.444	6981.6719	2.797	2.563	2.293	2.325
6961.5902	3.907	3.649	3.378	3.404	6981.4869	3.974	3.694	3.424	3.453	6982.6656	2.848	2.593	2.327	2.362
6961.6421	3.907	3.647	3.376	3.402	6981.5446	3.969	3.691	3.419	3.450	6983.5214	2.931	2.637	2.400	2.406
6961.6895	3.895	3.644	3.370	3.400	6981.5840	3.969	3.691	3.420	3.451	6988.5932	2.948	2.662	2.408	2.429
6961.7197	3.900	3.646	3.375	3.401	6981.6247	3.969	3.691	3.417	3.450	6988.6292	2.976	2.689	2.427	2.442
6961.4974	3.882	3.647	3.370	3.391	6981.6698	3.968	3.690	3.416	3.455	6988.6308	2.945	2.688	2.411	2.431
6961.5349	3.887	3.670	3.385	3.412	6982.6634	3.977	3.688	3.416	3.451	6988.6322	2.946	2.683	2.411	2.432
6961.5629	3.883	3.645	3.369	3.388	6988.6272	3.920	3.645	3.386	3.428	6988.6664	2.920	2.638	2.388	2.412
6961.5947	3.872	3.642	3.365	3.386	6988.6716	3.872	3.640	3.384	3.420	6988.7123	2.857	2.601	2.333	2.360
6961.6236	3.873	3.642	3.363	3.382	6981.5807	2.873	2.605	2.335	2.357	6989.6183	2.881	2.613	2.344	2.365
6961.6716	3.872	3.640	3.364	3.380	6989.5889	3.876	3.642	3.365	3.383	6990.5870	2.874	2.609	2.341	2.361
6961.6851	3.934	3.677	3.402	3.433	6990.5207	3.876	3.677	3.410	3.442	6990.5872	2.874	2.605	2.334	2.357
6961.6917	3.928	3.673	3.399	3.428	6990.5656	3.922	3.682	3.414	3.444	6990.6227	2.845	2.597	2.324	2.346
6961.6773	3.948	3.675	3.406	3.433	6991.5211	3.917	3.658	3.386	3.418	6990.6642	2.863	2.585	2.315	2.341
6970.6210	3.936	3.665	3.392	3.422	6991.5750	3.973	3.671	3.412	3.443	6990.6875	2.865	2.579	2.307	2.334
6970.6626	3.929	3.657	3.381	3.413	6991.5210	3.917	3.658	3.386	3.418	6991.5238	2.917	2.626	2.358	2.382
6970.6953	3.919	3.652	3.375	3.404	6991.5790	3.971	3.671	3.412	3.443	6991.5932	2.948	2.662	2.408	2.429
6971.4822	3.893	3.638	3.366	3.389	6991.5517	3.947	3.671	3.412	3.443	6998.6292	2.976	2.689	2.427	2.442
6971.5597	3.872	3.642	3.365	3.386	6998.6157	3.947	3.671	3.412	3.443	6999.6183	2.948	2.662	2.408	2.429
6971.6236	3.873	3.642	3.363	3.382	6998.6634	3.977	3.671	3.412	3.443	6999.6560	2.944	2.675	2.400	2.422
6971.6716	3.872	3.640	3.364	3.380	6999.6093	3.947	3.671	3.412	3.443	6999.6888	2.948	2.662	2.408	2.429
6971.6899	3.876	3.642	3.365	3.383	7000.6075	3.937	3.674	3.412	3.443	7000.6552	2.853	2.597	2.326	2.351
6971.6851	3.934	3.677	3.402	3.433	7001.6091	5.951	5.704	5.394	5.365	7001.5579	6.007	5.757	5.447	5.416
6971.6491	3.947	3.673	3.403	3.430	7001.6363	5.951	5.704	5.391	5.366	7001.5980	5.965	5.713	5.404	5.376
6971.6773	3.948	3.675	3.406	3.433	7001.6796	5.940	5.695	5.385	5.350	7001.6360	5.951	5.699	5.389	5.360
6971.7197	3.900	3.646	3.375	3.401	7001.7176	5.917	5.692	5.380	5.348	7001.6796	5.940	5.695	5.383	5.350
6971.6974	3.882	3.647	3.370	3.401	7001.7207	5.931	5.688	5.376	5.348	7001.7207	5.931	5.688	5.376	5.348
6971.7157	3.902	3.649	3.369	3.379	7001.7204	5.917	5.682	5.373	5.348	7001.7704	5.925	5.681	5.371	5.343
6971.7137	3.902	3.649	3.369	3.379	7001.7213	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7157	3.902	3.649	3.369	3.379	7001.7223	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7197	3.902	3.649	3.369	3.379	7001.7233	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7217	3.902	3.649	3.369	3.379	7001.7243	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7237	3.902	3.649	3.369	3.379	7001.7253	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7257	3.902	3.649	3.369	3.379	7001.7263	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7277	3.902	3.649	3.369	3.379	7001.7273	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7297	3.902	3.649	3.369	3.379	7001.7283	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7317	3.902	3.649	3.369	3.379	7001.7293	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7337	3.902	3.649	3.369	3.379	7001.7303	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7357	3.902	3.649	3.369	3.379	7001.7313	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7377	3.902	3.649	3.369	3.379	7001.7323	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7397	3.902	3.649	3.369	3.379	7001.7333	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7417	3.902	3.649	3.369	3.379	7001.7343	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7437	3.902	3.649	3.369	3.379	7001.7353	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7457	3.902	3.649	3.369	3.379	7001.7363	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7477	3.902	3.649	3.369	3.379	7001.7373	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7497	3.902	3.649	3.369	3.379	7001.7383	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7517	3.902	3.649	3.369	3.379	7001.7393	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7537	3.902	3.649	3.369	3.379	7001.7403	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7557	3.902	3.649	3.369	3.379	7001.7413	5.917	5.682	5.372	5.348	7001.7872	5.926	5.683	5.371	5.341
6971.7577	3.902	3.649	3.369	3.379	7001.7423	5.917	5.682	5.372	5.348	7001.7872				

TABLE IIb (*continued*)

HR 5941	HJD	u	v	b	y	HR 5941	HJD	u	v	b	y			
6960.5071	5.862	5.185	4.836	4.793	6970.7108	5.806	5.147	4.796	4.766	6970.7496	5.832	5.161	4.808	4.776
6960.5455	5.839	5.171	4.824	4.787	6971.5018	5.817	5.149	4.797	4.770	6971.5357	5.853	5.183	4.830	4.802
6960.5681	5.826	5.162	4.814	4.780	6971.5669	5.841	5.162	4.812	4.782	6971.5669	5.845	5.164	4.811	4.781
6960.5956	5.818	5.150	4.805	4.772	6971.6061	5.845	5.164	4.811	4.781	6971.6437	5.857	5.182	4.819	4.789
6960.6264	5.810	5.140	4.792	4.763	6971.6869	5.897	5.197	4.843	4.807	6971.7301	5.913	5.213	4.859	4.819
6960.6358	5.824	5.145	4.800	4.768	6971.7792	5.880	5.185	4.834	4.794	6972.3022	5.901	5.203	4.850	4.818
6960.7196	5.839	5.160	4.813	4.783	6972.5372	5.909	5.210	4.856	4.820	6972.5649	5.958	5.203	4.851	4.815
6960.7407	5.843	5.159	4.812	4.783	6972.6129	5.869	5.178	4.827	4.794	6972.6496	5.829	5.150	4.800	4.770
6960.7848	5.856	5.166	4.817	4.784	6972.6878	5.816	5.140	4.790	4.761	6972.7350	5.829	5.167	4.846	4.768
6960.8104	5.858	5.173	4.822	4.788	6972.7682	5.842	5.158	4.805	4.774	6973.4952	5.817	5.158	4.844	4.764
6962.5198	5.876	5.192	4.842	4.805	6973.5274	5.827	5.151	4.793	4.771	6973.5660	5.842	5.162	4.812	4.782
6962.5211	5.868	5.176	4.829	4.791	6973.6106	5.818	5.164	4.813	4.783	6973.6496	5.829	5.150	4.800	4.770
6962.5231	5.858	5.178	4.826	4.791	6973.6925	5.896	5.102	4.850	4.816	6973.7482	5.837	5.160	4.812	4.778
6962.5424	5.850	5.167	4.818	4.785	6973.6935	5.839	5.163	4.809	4.779	6973.7495	5.817	5.154	4.814	4.764
6962.5562	5.841	5.157	4.808	4.776	6975.6670	5.846	5.168	4.815	4.784	6975.7052	5.801	5.133	4.780	4.753
6962.5799	5.843	5.158	4.809	4.779	6980.5439	5.880	5.196	4.835	4.810	6980.5803	5.877	5.190	4.837	4.803
6962.5951	5.836	5.149	4.802	4.772	6980.5885	5.853	5.184	4.837	4.804	6980.6282	5.851	5.181	4.829	4.793
6962.6103	5.836	5.150	4.803	4.774	6980.6833	5.873	5.171	4.824	4.792	6980.7362	5.851	5.154	4.804	4.773
6962.6279	5.832	5.149	4.801	4.771	6980.7628	5.847	5.159	4.808	4.770	6980.7654	5.849	5.160	4.808	4.779
6962.6484	5.836	5.150	4.803	4.773	6980.7795	5.854	5.162	4.826	4.814	6980.8021	5.886	5.182	4.839	4.773
6962.7013	5.847	5.162	4.814	4.784	6981.5244	5.817	5.143	4.793	4.764	6981.5604	5.826	5.146	4.796	4.766
6962.7228	5.853	5.166	4.818	4.787	6981.5630	5.826	5.146	4.796	4.766	6981.5844	5.849	5.160	4.808	4.779
6962.7444	5.858	5.156	4.821	4.790	6981.6003	5.839	5.161	4.808	4.779	6981.6435	5.839	5.163	4.809	4.779
6962.7629	5.867	5.172	4.825	4.790	6981.6853	5.879	5.183	4.828	4.801	6981.7052	5.801	5.133	4.780	4.753
6962.8030	5.868	5.178	4.827	4.788	6981.7886	5.832	5.133	4.799	4.773	6981.8086	5.846	5.178	4.820	4.793
6965.5312	5.836	5.154	4.805	4.774	6981.8228	5.847	5.159	4.808	4.770	6981.8446	5.856	5.161	4.813	4.783
6965.5609	5.848	5.166	4.813	4.784	6981.8454	5.849	5.162	4.820	4.814	6981.8677	5.847	5.159	4.808	4.770
6965.6061	5.855	5.172	4.820	4.788	6981.8795	5.849	5.160	4.808	4.779	6981.8934	5.849	5.161	4.811	4.814
6965.6448	5.851	5.170	4.819	4.785	6981.9135	5.849	5.162	4.827	4.814	6981.9395	5.845	5.163	4.818	4.815
6965.7327	5.844	5.160	4.810	4.777	6982.0321	5.884	5.173	4.824	4.796	6982.0552	5.801	5.132	4.804	4.780
6965.7808	5.837	5.158	4.810	4.775	6982.0889	5.829	5.156	4.804	4.776	6982.1206	5.839	5.164	4.812	4.811
6967.5179	5.843	5.158	4.809	4.782	6982.1765	5.829	5.156	4.804	4.769	6982.1877	5.831	5.157	4.812	4.811
6967.5483	5.835	5.154	4.804	4.776	6982.2141	5.837	5.158	4.810	4.783	6982.2446	5.836	5.159	4.808	4.779
6967.5573	5.840	5.159	4.806	4.779	6982.2888	5.847	5.159	4.808	4.770	6982.3026	5.848	5.159	4.809	4.773
6968.5673	5.864	5.184	4.829	4.799	6982.3444	5.851	5.164	4.816	4.800	6982.3593	5.851	5.165	4.817	4.823
6968.6057	5.841	5.160	4.807	4.778	6982.3602	5.847	5.164	4.807	4.800	6982.3750	5.847	5.165	4.811	4.814
6968.6480	5.843	5.160	4.802	4.773	6982.3955	5.845	5.166	4.808	4.804	6982.4106	5.848	5.167	4.812	4.814
6968.6767	5.834	5.151	4.800	4.772	6982.4399	5.850	5.164	4.808	4.823	6982.4631	5.844	5.165	4.811	4.829
6968.7244	5.851	5.164	4.811	4.781	6982.4931	5.845	5.164	4.804	4.829	6982.5374	5.846	5.165	4.812	4.829
6968.7363	5.857	5.166	4.814	4.783	6982.5378	5.843	5.165	4.811	4.829	6982.5846	5.846	5.166	4.813	4.835
6969.5975	5.852	5.167	4.812	4.785	6982.7788	5.842	5.168	4.807	4.827	6982.8086	5.845	5.169	4.815	4.833
6969.6331	5.861	5.164	4.810	4.784	6982.8123	5.849	5.166	4.814	4.807	6982.8466	5.851	5.167	4.815	4.827
6969.6622	5.880	5.182	4.829	4.799	6994.3660	5.402	4.860	4.792	4.799	6994.3966	5.403	4.864	4.800	4.816
6969.6988	5.888	5.196	4.842	4.809	6994.4104	5.430	4.875	4.834	4.817	6994.4404	5.435	4.879	4.845	4.821
6969.7277	5.892	5.215	4.859	4.818	6995.2274	5.392	4.828	4.776	4.789	6995.2507	4.897	4.800	4.569	4.271
6970.6364	5.817	5.153	4.800	4.771	6995.2720	5.398	4.848	4.792	4.802	6995.2951	4.880	4.793	4.561	4.255
6970.6766	5.793	5.135	4.782	4.754	6995.3222	5.410	4.853	4.795	4.807	6995.3393	4.886	4.800	4.572	4.263
6961.6140	5.265	5.142	4.608	4.270	6971.5051	5.266	5.138	4.612	4.277	6971.5397	5.281	5.155	4.625	4.289
6961.6423	5.261	5.161	4.608	4.267	6971.5684	5.262	5.130	4.601	4.264	6971.6099	5.259	5.126	4.597	4.257
6961.6634	5.260	5.141	4.606	4.264	6971.6475	5.258	5.126	4.594	4.254	6971.6747	5.250	5.127	4.595	4.254
6961.6870	5.261	5.161	4.608	4.263	6971.6903	5.263	5.125	4.593	4.250	6971.7337	5.253	5.124	4.594	4.250
6961.7093	5.259	5.140	4.606	4.261	6971.7377	5.253	5.124	4.595	4.250	6971.7837	5.254	5.125	4.596	4.250
6961.7488	5.258	5.137	4.604	4.258	6971.7877	5.254	5.125	4.596	4.250	6972.0561	5.253	5.120	4.590	4.249
6961.7676	5.255	5.138	4.603	4.258	6972.0864	5.254	5.118	4.586	4.249	6972.1668	5.249	5.118	4.583	4.246
6961.7994	5.253	5.135	4.602	4.257	6972.5378	5.249	5.119	4.584	4.246	6972.6089	5.249	5.118	4.584	4.246
6961.8193	5.256	5.139	4.604	4.258	6972.6333	5.249	5.117	4.581	4.243	6972.6713	5.245	5.119	4.589	4.246
6963.5617	5.272	5.156	4.621	4.283	6972.6337	5.246	5.117	4.581	4.243	6972.6960	5.243	5.117	4.582	4.243
6963.5794	5.274	5.155	4.623	4.282	6972.7316	5.251	5.112	4.585	4.246	6972.7352	5.251	5.112	4.586	4.246
6963.5965	5.277	5.155	4.623	4.283	6972.7808	5.244	5.113	4.584	4.243	6972.7830	5.233	5.114	4.582	4.243
6963.6250	5.274	5.158	4.622	4.281	6972.7837	5.233	5.114	4.582	4.243	6972.7872	5.233	5.114	4.582	4.243
6963.6493	5.274	5.157	4.623	4.282	6972.7878	5.234	5.117	4.587	4.244	6973.6467	5.255	5.113	4.611	4.267
6965.5720	5.273	5.159	4.622	4.287	6973.6713	5.253	5.119	4.607	4.263	6973.6755	5.256	5.115	4.615	4.276
6965.6092	5.275	5.159	4.630	4.287	6978.7465	5.269	5.115	4.624	4.284	6978.7465	5.269	5.115	4.624	4.284
6965.6479	5.271	5.159	4.624	4.284	6979.5120	5.275	5.116	4.633	4.293	6979.5496	5.276	5.116	4.635	4.293
6965.7351	5.275	5.159	4.631	4.286	6979.5594	5.276	5.116	4.635	4.293	6979.5924	5.275	5.116	4.632	4.292
6965.7832	5.272	5.15												

## PHOTOMETRY OF BE STARS

HR 6175	HJD	u	v	b	y	HR 6451	HJD	u	v	b	y			
6965.5464	3.179	3.057	2.668	2.540	6979.5250	3.185	3.059	2.669	2.543	6968.8680	5.786	5.588	5.266	5.236
6965.5803	3.175	3.051	2.662	2.535	6979.5625	3.183	3.061	2.670	2.546	6968.8879	5.775	5.581	5.257	5.231
6965.6241	3.174	3.051	2.662	2.536	6979.6048	3.180	3.061	2.670	2.546	6969.6081	5.761	5.568	5.243	5.225
6965.6656	3.173	3.053	2.662	2.535	6979.6400	3.180	3.057	2.666	2.542	6969.6413	5.753	5.568	5.243	5.226
6965.7590	3.184	3.061	2.672	2.543	6980.5166	3.191	3.063	2.677	2.532	6969.6705	5.755	5.566	5.241	5.223
6965.8076	3.181	3.057	2.669	2.534	6980.5625	3.182	3.060	2.671	2.546	6969.6978	5.741	5.551	5.225	5.207
6967.5305	3.186	3.065	2.672	2.550	6980.5986	3.179	3.057	2.667	2.541	6969.7373	5.733	5.542	5.217	5.195
6967.5578	3.179	3.059	2.665	2.542	6980.6479	3.186	3.059	2.671	2.543	6970.6139	5.765	5.562	5.235	5.218
6967.5911	3.179	3.059	2.667	2.543	6980.7042	3.184	3.059	2.670	2.543	6970.6469	5.769	5.572	5.246	5.227
6967.6281	3.181	3.058	2.666	2.543	6980.7538	3.179	3.056	2.665	2.539	6970.6872	5.793	5.594	5.269	5.248
6967.6663	3.177	3.052	2.662	2.537	6981.5406	3.184	3.082	2.679	2.560	6970.7202	5.799	5.593	5.270	5.248
6967.6969	3.179	3.054	2.665	2.539	6981.5805	3.181	3.057	2.666	2.541	6970.7602	5.773	5.572	5.250	5.230
6967.7238	3.181	3.057	2.667	2.541	6981.6178	3.189	3.063	2.672	2.546	6970.7981	5.758	5.552	5.235	5.214
6967.7659	3.181	3.057	2.666	2.538	6981.6534	3.191	3.063	2.675	2.548	6970.8201	5.756	5.553	5.232	5.210
6967.8187	3.187	3.062	2.672	2.550	6982.6475	3.181	3.054	2.665	2.541	6970.8382	5.744	5.549	5.224	5.205
6968.5135	3.188	3.066	2.672	2.544	6982.7029	3.190	3.063	2.674	2.547	6971.5072	5.805	5.599	5.256	5.236
6968.5465	3.184	3.066	2.672	2.544	6991.2889	2.789	2.743	2.652	2.568	6971.5427	5.815	5.611	5.265	5.265
6968.5835	3.179	3.057	2.667	2.543	6991.3312	2.778	2.752	2.654	2.567	6971.5809	5.800	5.599	5.266	5.243
6968.6242	3.181	3.058	2.666	2.543	6991.3824	2.785	2.764	2.665	2.574	6971.6191	5.790	5.588	5.235	5.228
6968.6612	3.179	3.058	2.669	2.542	6991.4262	2.775	2.767	2.668	2.576	6971.6555	5.788	5.572	5.230	5.224
6968.6889	3.181	3.059	2.669	2.542	6991.4671	2.788	2.763	2.666	2.580	6971.7497	5.812	5.609	5.265	5.254
6968.7373	3.194	3.072	2.679	2.548	6991.5284	2.804	2.756	2.680	2.567	6971.7920	5.822	5.617	5.274	5.264
6968.7753	3.192	3.054	2.665	2.536	6994.2640	2.798	2.766	2.662	2.570	6971.8262	5.809	5.601	5.259	5.248
6969.6138	3.178	3.052	2.662	2.536	6994.3079	2.805	2.774	2.677	2.587	6971.8523	5.799	5.582	5.241	5.233
6970.7424	3.181	3.059	2.669	2.543	6994.3521	2.794	2.783	2.684	2.592	6971.8684	5.784	5.569	5.229	5.218
6970.7654	3.185	3.062	2.670	2.544	6994.3969	2.804	2.775	2.669	2.572	6972.0575	5.765	5.563	5.240	5.226
6971.5132	3.181	3.070	2.670	2.553	6994.4398	2.781	2.764	2.678	2.591	6972.5429	5.770	5.573	5.249	5.237
6971.5423	3.185	3.063	2.671	2.544	6995.2071	2.812	2.783	2.681	2.592	6972.5779	5.767	5.570	5.247	5.235
6970.6177	3.182	3.061	2.669	2.547	6995.3568	2.807	2.787	2.682	2.594	6972.6249	5.771	5.571	5.247	5.233
6970.6518	3.181	3.060	2.666	2.541	6998.3034	2.777	2.767	2.665	2.579	6972.6592	5.774	5.570	5.246	5.233
6970.6911	3.181	3.058	2.668	2.541	6998.3524	2.777	2.766	2.664	2.576	6972.6980	5.768	5.565	5.240	5.225
6970.7654	3.185	3.062	2.670	2.544	6998.4011	2.788	2.764	2.672	2.588	6972.7432	5.770	5.564	5.242	5.227
6971.5132	3.181	3.070	2.670	2.553	6998.4470	2.784	2.762	2.678	2.590	6972.7864	5.782	5.583	5.257	5.242
6971.5474	3.201	3.080	2.688	2.564	6999.2550	2.772	2.779	2.675	2.588	6973.0211	5.792	5.577	5.253	5.234
6971.5843	3.189	3.066	2.675	2.549	6999.3035	2.794	2.782	2.690	2.596	6973.5345	5.789	5.580	5.255	5.242
6971.6243	3.183	3.062	2.670	2.544	6999.3479	2.785	2.772	2.675	2.589	6973.5773	5.773	5.573	5.251	5.234
6971.6624	3.192	3.066	2.674	2.547	6999.3909	2.776	2.763	2.669	2.580	6973.6215	5.762	5.560	5.239	5.221
6971.7004	3.189	3.061	2.671	2.541	6999.4336	2.772	2.763	2.664	2.577	6973.7013	5.774	5.574	5.248	5.228
6971.7540	3.183	3.057	2.666	2.536	7000.2527	2.775	2.765	2.666	2.583	6973.7582	5.777	5.575	5.251	5.231
6971.8015	3.180	3.056	2.669	2.532	7000.2970	2.775	2.761	2.665	2.579	6973.8027	5.791	5.589	5.264	5.240
6972.5122	3.188	3.059	2.670	2.541	7000.3403	2.778	2.761	2.663	2.575	6975.6125	5.775	5.568	5.241	5.223
6972.5479	3.185	3.056	2.667	2.540	7000.3967	2.790	2.762	2.671	2.577	6975.6280	5.772	5.566	5.241	5.220
6972.5881	3.181	3.060	2.666	2.539	7000.4406	2.792	2.763	2.672	2.577	6975.6480	5.777	5.573	5.246	5.229
6972.6294	3.179	3.057	2.665	2.539	6999.3035	2.794	2.782	2.670	2.596	6977.7319	5.777	5.571	5.248	5.226
6972.6651	3.179	3.057	2.666	2.541	6999.3568	2.776	2.763	2.668	2.582	6977.7633	5.771	5.567	5.243	5.222
6972.7020	3.176	3.056	2.664	2.538	6999.4006	2.775	2.767	2.665	2.581	6977.7930	5.769	5.563	5.240	5.218
6972.7476	3.184	3.058	2.670	2.540	6999.4521	2.781	2.768	2.670	2.576	6977.8219	5.766	5.566	5.242	5.222
6972.7911	3.181	3.057	2.666	2.538	6999.5016	2.784	2.763	2.668	2.578	6978.7573	5.777	5.571	5.249	5.228
6973.5398	3.188	3.059	2.668	2.541	6999.5474	2.785	2.764	2.670	2.581	6978.8111	5.767	5.566	5.238	5.222
6973.5813	3.179	3.055	2.668	2.542	6998.6848	5.759	5.566	5.235	5.215	6979.5186	5.779	5.578	5.250	5.234
6973.6260	3.173	3.053	2.666	2.541	6998.7327	5.769	5.572	5.246	5.217	6979.5569	5.787	5.582	5.257	5.239
6973.6704	3.170	3.053	2.660	2.534	6998.7713	5.771	5.573	5.248	5.220	6979.5999	5.784	5.583	5.258	5.236
6973.7626	3.157	3.042	2.654	2.527	6998.8162	5.816	5.579	5.255	5.228	6979.6347	5.787	5.581	5.254	5.235
6977.7374	3.173	3.057	2.657	2.539	6998.8419	5.789	5.593	5.269	5.242	6979.6647	5.781	5.578	5.251	5.232
6979.3745	5.394	5.294	5.266	5.295	6999.3811	3.130	2.800	2.782	2.792	6979.7359	3.195	3.148	2.816	2.788
6994.3344	5.401	5.314	5.266	5.285	6999.4229	3.130	2.801	2.781	2.793	6980.7741	3.178	3.134	2.804	2.777
6994.4229	5.409	5.303	5.269	5.288	6999.5155	3.130	2.801	2.781	2.794	6980.8160	3.164	3.123	2.792	2.767
6994.5155	5.371	5.251	5.223	5.246	6999.5622	3.130	2.801	2.781	2.795	6980.8453	3.176	3.136	2.808	2.784
6995.2405	5.434	5.325	5.295	5.322	6999.6145	3.139	2.801	2.781	2.797	6980.8717	3.137	3.122	2.801	2.781
6995.2848	5.435	5.320	5.294	5.316	6999.6651	3.139	2.802	2.781	2.795	6980.8893	3.172	3.131	2.801	2.775
6995.3344	5.385	5.276	5.245	5.277	6999.7008	3.137	2.803	2.782	2.796	6980.9100	3.174	3.122	2.801	2.792
6998.2708	5.381	5.273	5.252	5.276	6999.7439	3.139	2.803	2.782	2.795	6980.9409	3.141	3.158	2.809	2.800
6999.3340	5.393	5.303	5.273	5.301	6999.7661	3.139	2.803	2.782	2.796	6980.9707	3.182	3.140	2.807	2.785
6999.3833	5.372	5.280	5.256	5.282	6999.8094	3.138	2.802	2.789	2.791	6980.9245	3.172	3.126	2.798	2.779
6999.4289	5.372	5.268	5.240	5.264	6999.8592	3.138	2.803	2.784	2.798	6980.9413	3.171	3.125	2.811	2.806
6999.4640	5.364	5.262	5.233	5.256	6999.8832</td									

TABLE IIb (*continued*)

HR	HJD						HJD								
	u	v	b	y	u	v	b	y	u	v	b	y			
6510	7000.2380	2.978	2.839	2.810	2.829	7000.2826	2.983	2.839	2.806	2.837	6979.7881	6.622	6.565	6.192	6.096
6979.5612	3.379	3.137	2.808	2.789	7000.3259	2.984	2.834	2.809	2.833	6971.8227	6.617	6.563	6.189	6.092	
6979.6038	3.376	3.132	2.803	2.785	7000.3787	2.986	2.843	2.815	2.841	6972.5718	6.640	6.582	6.206	6.110	
6979.6389	3.368	3.126	2.797	2.778	7000.4304	2.997	2.850	2.820	2.843	6972.6196	6.638	6.581	6.206	6.110	
6979.6691	3.372	3.129	2.800	2.780	7000.4828	2.994	2.861	2.826	2.863	6972.6567	6.629	6.574	6.200	6.107	
6979.7630	3.382	3.142	2.814	2.792	6979.7849	3.380	3.139	2.810	2.791	6972.6952	6.625	6.573	6.234	6.198	
6979.7805	3.388	3.145	2.816	2.797	6979.8229	3.381	3.138	2.807	2.789	6972.7409	6.628	6.575	6.202	6.108	
6980.5152	3.390	3.146	2.818	2.803	6980.5152	3.390	3.146	2.818	2.803	6972.7828	6.628	6.574	6.203	6.108	
6980.5613	3.373	3.128	2.801	2.784	6980.5973	3.368	3.121	2.792	2.773	6972.8149	6.633	6.581	6.203	6.110	
6980.6465	3.363	3.119	2.790	2.771	6980.6465	3.363	3.119	2.790	2.771	6973.5728	6.653	6.591	6.217	6.111	
6980.7029	3.374	3.134	2.804	2.782	6980.7524	3.385	3.142	2.814	2.793	6973.5728	6.653	6.591	6.217	6.111	
6980.7886	3.386	3.144	2.812	2.793	6980.8206	3.387	3.142	2.811	2.796	6973.5728	6.653	6.591	6.217	6.111	
6980.8422	3.392	3.148	2.814	2.797	6980.8589	3.393	3.151	2.814	2.798	6973.6990	6.660	6.594	6.222	6.122	
6981.5389	3.374	3.132	2.801	2.785	6981.5389	3.374	3.132	2.801	2.785	6973.7755	6.668	6.599	6.233	6.135	
6981.5793	3.361	3.119	2.783	2.775	6981.6156	3.363	3.121	2.792	2.775	6973.7997	6.668	6.609	6.234	6.135	
6981.6555	3.371	3.127	2.797	2.776	6981.7019	3.381	3.141	2.807	2.791	6975.6747	6.648	6.584	6.212	6.114	
6982.6456	3.385	3.137	2.807	2.788	6982.7014	3.390	3.145	2.814	2.798	6976.6157	6.627	6.568	6.194	6.101	
6991.2734	2.996	2.847	2.814	2.838	6991.3172	2.984	2.835	2.810	2.826	6976.6157	6.627	6.568	6.194	6.101	
6991.3672	2.976	2.826	2.795	2.819	6991.4106	2.967	2.827	2.802	2.823	6976.6157	6.627	6.568	6.194	6.101	
6991.4586	2.980	2.840	2.811	2.835	6991.5105	2.994	2.844	2.813	2.836	6976.6157	6.640	6.586	6.209	6.115	
6991.5463	3.002	2.836	2.810	2.837	6991.5944	2.986	2.844	2.811	2.837	6976.6157	6.640	6.586	6.209	6.115	
6994.2482	2.996	2.836	2.812	2.833	6994.2929	2.982	2.839	2.801	2.826	6976.6157	6.640	6.586	6.209	6.115	
6994.3365	3.004	2.857	2.805	2.829	6994.3805	2.973	2.836	2.804	2.825	6976.6157	6.640	6.586	6.209	6.115	
6994.4248	2.993	2.834	2.810	2.829	6994.4712	3.002	2.846	2.805	2.836	6976.6157	6.640	6.586	6.209	6.115	
6994.5173	3.003	2.846	2.804	2.836	6995.2419	2.997	2.836	2.805	2.833	6976.6157	6.640	6.586	6.209	6.115	
6995.2867	3.005	2.853	2.827	2.845	6995.2914	2.983	2.835	2.810	2.829	6976.6157	6.640	6.586	6.209	6.115	
6995.3362	2.991	2.846	2.814	2.848	6995.3404	3.002	2.844	2.804	2.835	6976.6157	6.640	6.586	6.209	6.115	
6996.2810	2.994	2.829	2.811	2.835	6996.3360	2.996	2.844	2.823	2.847	6976.6157	6.640	6.586	6.209	6.115	
6996.3853	2.994	2.844	2.813	2.836	6996.4308	3.002	2.843	2.816	2.837	6976.6157	6.640	6.586	6.209	6.115	
6996.4461	3.003	2.853	2.821	2.843	6996.5166	3.027	2.873	2.843	2.862	6976.6157	6.640	6.586	6.209	6.115	
6999.2389	2.984	2.840	2.811	2.833	6999.2880	2.987	2.846	2.818	2.839	6970.7570	6.647	6.592	6.217	6.119	
6999.3330	3.002	2.853	2.819	2.850	6999.3764	3.005	2.849	2.821	2.844	6970.7808	6.644	6.587	6.216	6.118	
6999.4189	2.988	2.858	2.826	2.850	6999.4672	3.000	2.863	2.835	2.852	6971.6523	6.639	6.580	6.207	6.113	
6999.5172	3.021	2.865	2.833	2.844	6999.5172	3.021	2.865	2.833	2.844	6971.7369	6.630	6.572	6.198	6.103	
8260	HJD						HJD								
HR	u	v	b	y	u	v	b	y	u	v	b	y			
	6977.7445	4.940	4.652	4.331	4.398	6977.7797	4.941	4.655	4.334	4.398	6977.7125	5.794	5.036	4.659	4.685
6963.7250	4.978	4.674	4.355	4.421	6977.8097	4.941	4.654	4.333	4.398	6977.7483	5.813	5.055	4.676	4.704	
6963.7493	4.975	4.673	4.355	4.420	6977.8324	4.941	4.655	4.335	4.398	6977.7834	5.813	5.054	4.675	4.704	
6963.7710	4.979	4.675	4.358	4.423	6977.8447	4.946	4.658	4.337	4.403	6977.8132	5.813	5.054	4.676	4.704	
6963.8057	4.980	4.676	4.360	4.425	6977.8605	4.948	4.657	4.338	4.403	6977.8406	5.812	5.050	4.673	4.698	
6963.8321	4.984	4.677	4.362	4.425	6977.8815	4.950	4.662	4.342	4.407	6976.8815	5.812	5.048	4.670	4.698	
6963.8511	4.981	4.678	4.362	4.424	6977.8815	4.955	4.667	4.347	4.411	6976.8815	5.812	5.049	4.670	4.698	
6963.8612	4.981	4.677	4.362	4.426	6978.7653	4.948	4.656	4.336	4.399	6976.8815	5.812	5.049	4.669	4.698	
6963.8825	4.977	4.677	4.361	4.422	6978.7964	4.947	4.658	4.339	4.402	6976.8815	5.812	5.049	4.671	4.698	
6963.8836	4.981	4.674	4.361	4.422	6978.8194	4.951	4.662	4.342	4.407	6976.8815	5.812	5.049	4.671	4.698	
6963.8935	4.987	4.675	4.363	4.423	6979.6766	4.948	4.661	4.344	4.405	6976.8815	5.812	5.049	4.672	4.698	
6963.8983	4.978	4.674	4.363	4.423	6979.7671	4.952	4.659	4.345	4.406	6976.8815	5.812	5.049	4.672	4.698	
6965.8120	4.978	4.674	4.357	4.420	6979.7709	4.951	4.664	4.347	4.410	6976.8815	5.812	5.049	4.672	4.698	
6965.8390	4.972	4.671	4.355	4.419	6979.8106	4.956	4.661	4.347	4.407	6976.8815	5.812	5.049	4.672	4.698	
6965.8556	4.973	4.671	4.355	4.417	6979.8285	4.951	4.661	4.344	4.406	6976.8815	5.812	5.049	4.672	4.698	
6965.8878	4.971	4.668	4.353	4.416	6979.8386	4.953	4.660	4.344	4.407	6976.8815	5.812	5.049	4.672	4.698	
6965.8922	4.965	4.663	4.347	4.411	6979.8553	4.959	4.662	4.347	4.410	6976.8815	5.812	5.049	4.672	4.698	
6967.7308	4.976	4.662	4.361	4.426	6979.8712	4.952	4.664	4.348	4.407	6976.8815	5.812	5.049	4.672	4.698	
6967.7659	4.979	4.661	4.362	4.425	6979.8872	4.949	4.659	4.341	4.405	6976.8815	5.812	5.049	4.672	4.698	
6967.8095	4.990	4.671	4.364	4.426	6979.9017	4.957	4.663	4.347	4.408	6976.8815	5.812	5.049	4.672	4.698	
6967.8422	4.974	4.674	4.357	4.420	6979.9166	4.946	4.652	4.332	4.391	6976.8815	5.812	5.049	4.672	4.698	
6967.8642	4.972	4.672	4.354	4.417	6979.9280	4.949	4.659	4.341	4.403	6976.8815	5.812	5.049	4.672	4.698	
6967.8947	4.971	4.671	4.358	4.421	6979.9380	4.951	4.664	4.344	4.404	6976.8815	5.812	5.049	4.672	4.698	
6967.9291	4.972	4.676	4.358	4.421	6979.9561	4.970	4.664	4.331	4.393	6976.8815	5.812	5.049	4.672	4.698	
6970.7297	4.972	4.676	4.358	4.421	6979.9691	4.951	4.674	4.340	4.404	6976.8815	5.812	5.049	4.672	4.698	
6970.7670	4.972	4.675	4.358	4.421	6979.9851	4.954	4.668	4.330	4.398	6976.8815	5.812	5.049	4.672	4.698	
6970.8042	4.967	4.669	4.353	4.415	6979.9911	4.951	4.674	4.347	4.402	6976.8815	5.812	5.049	4.672	4.698	
6970.8265	4.953	4.658	4.340	4.402	6980.0004	4.954	4.668	4.336	4.398	6976.8815	5.812	5.049	4.672	4.698	
6970.8521	4.971	4.678	4.360	4.424	6980.0268	4.954	4.668	4.336	4.398	6976.8815	5.812	5.049	4.672	4.698	
6971.8571	4.971	4.672	4.357	4.421	6980.0521	4.954	4.667	4.337	4.398	6976.8815	5.812	5.049	4.672	4.698	
6971.8791	4.973	4.670	4.353	4.416	6980.0806	4.954	4.666	4.336	4.398	6976.8815	5.812	5.049	4.672	4.698	
6971.8907	4.973	4.671	4.353	4.417	6980.1080	4.954	4.667	4.337	4.398	6976.8815	5.812	5.049	4.672	4.698	
6971.9033	4.973	4.667	4.351	4.414	6980.1356	4.954	4.668	4.338	4.398	6976.8815	5.812	5.049	4.672	4.698	
6972.6695	4.979	4.674	4.358	4.422	6980.1631	4.955	4.669	4.339	4.398	6976.8815	5.812	5.049	4.672	4.698	
6972.7063	4.979	4.674	4.359	4.422	6980.1906	4.955	4.670	4.338	4.398	6976.8815	5.812	5.049	4.672	4.698	
6972.7512	4.979	4.675	4.359	4.420	6980.2180	4.955	4.671	4.337	4.398	6976.8815	5.812	5.049	4.6		

TABLE IIb (*continued*)

HR 8408	HJD	u	v	b	y	HR 8858	HJD	u	v	b	y
6960.7686	6.688	6.225	5.891	5.942	5.977	6.7457	6.687	6.220	5.886	5.943	5.977
6960.7697	6.689	6.220	5.888	5.939	5.977	7.809	6.697	6.232	5.896	5.952	5.977
6960.8037	6.699	6.232	5.900	5.951	5.977	8.109	6.701	6.236	5.901	5.952	5.977
6960.8110	6.699	6.231	5.900	5.952	5.977	8.336	6.696	6.233	5.899	5.953	5.977
6960.8678	6.695	6.230	5.900	5.951	5.977	8.498	6.697	6.231	5.899	5.953	5.977
6960.8785	6.697	6.231	5.901	5.953	5.977	8.618	6.693	6.230	5.895	5.950	5.977
6960.8869	6.692	6.229	5.899	5.951	5.977	8.829	6.693	6.229	5.893	5.947	5.977
6960.8946	6.693	6.224	5.899	5.949	5.978	7.668	6.687	6.220	5.884	5.939	5.961
6960.9011	6.693	6.229	5.899	5.949	5.978	7.797	6.680	6.220	5.889	5.940	5.961
6960.9084	6.696	6.229	5.901	5.950	5.978	8.208	6.681	6.218	5.886	5.939	5.962
6960.9141	6.697	6.230	5.894	5.951	5.979	7.685	6.697	6.232	5.903	5.953	5.962
6960.9203	6.697	6.230	5.901	5.951	5.979	7.918	6.696	6.231	5.900	5.947	5.962
6965.8142	6.673	6.213	5.882	5.936	5.979	8.120	6.693	6.231	5.902	5.951	5.962
6965.8165	6.679	6.217	5.886	5.938	5.979	8.272	6.696	6.228	5.899	5.948	5.962
6965.8397	6.676	6.216	5.886	5.939	5.979	8.358	6.698	6.235	5.904	5.953	5.962
6965.8571	6.675	6.215	5.888	5.936	5.979	8.591	6.693	6.232	5.901	5.950	5.963
6965.8799	6.671	6.216	5.888	5.936	5.979	8.727	6.696	6.233	5.901	5.952	5.963
6965.8839	6.674	6.217	5.883	5.938	5.979	8.891	6.694	6.232	5.901	5.950	5.963
6967.7317	6.680	6.216	5.881	5.937	5.979	9.039	6.697	6.235	5.905	5.956	5.967
6967.7711	6.682	6.222	5.887	5.942	5.980	6.571	6.694	6.227	5.892	5.946	5.980
6967.8108	6.684	6.228	5.892	5.946	5.980	7.127	6.682	6.213	5.880	5.936	5.980
6967.8441	6.679	6.223	5.888	5.941	5.980	7.941	6.679	6.212	5.880	5.932	5.980
6967.8674	6.676	6.219	5.886	5.935	5.980	8.254	6.680	6.214	5.882	5.935	5.980
6968.7034	6.679	6.215	5.882	5.939	5.980	8.467	6.682	6.217	5.886	5.938	5.980
6968.7423	6.697	6.222	5.897	5.953	5.980	8.670	6.683	6.219	5.886	5.939	5.980
6968.7800	6.692	6.223	5.890	5.946	5.980	8.896	6.679	6.219	5.886	5.940	5.980
6968.8194	6.687	6.222	5.887	5.942	5.980	9.005	6.680	6.218	5.883	5.938	5.980
6968.8485	6.697	6.230	5.897	5.949	5.981	6.654	6.692	6.222	5.887	5.936	5.981
6968.8753	6.685	6.217	5.884	5.937	5.982	6.586	6.721	6.225	5.894	5.943	5.982
6968.8973	6.686	6.220	5.887	5.940	5.982	7.129	6.728	6.234	5.906	5.959	5.982
6969.7085	6.685	6.223	5.891	5.943	5.982	7.584	6.701	6.208	5.881	5.932	5.982
6969.7479	6.690	6.221	5.891	5.942	5.982	8.526	6.700	6.199	5.904	5.950	5.983
6970.7308	6.689	6.224	5.892	5.944	5.982	6.348	6.297	5.956	5.936	6.003	5.983
6970.7703	6.696	6.218	5.892	5.946	5.982	8.485	6.267	5.921	5.908	5.983	5.983
6970.8057	6.698	6.221	5.888	5.941	5.982	5.573	6.268	5.909	5.893	5.963	5.983
6970.8280	6.685	6.222	5.888	5.942	5.982	5.598	6.280	5.919	5.905	5.972	5.982
6970.8571	6.686	6.221	5.890	5.944	5.982	6.524	6.286	5.922	5.905	5.981	5.982
6970.8669	6.690	6.227	5.898	5.950	5.982	6.879	6.285	5.925	5.904	5.975	5.982
6970.8780	6.690	6.227	5.895	5.949	5.982	7.005	6.285	5.927	5.904	5.982	5.982
6970.8933	6.689	6.227	5.895	5.948	5.982	6.013	6.294	5.928	5.904	5.975	5.982
6971.7058	6.695	6.224	5.888	5.946	5.982	6.545	6.281	5.935	5.907	5.986	5.982
6971.7605	6.691	6.214	5.884	5.942	5.982	7.000	5.610	6.275	5.905	5.911	5.962
6971.8049	6.689	6.219	5.885	5.942	5.982	7.293	6.252	5.903	5.887	5.958	5.982
6971.8247	6.696	6.220	5.889	5.945	5.982	7.568	6.262	5.915	5.898	5.963	5.982
6971.8582	6.694	6.222	5.890	5.945	5.982	7.618	6.265	5.917	5.894	5.965	5.982
6971.8813	6.698	6.226	5.893	5.947	5.982	7.781	6.214	5.886	5.937	5.982	5.982
6971.8926	6.700	6.227	5.893	5.947	5.982	7.809	6.215	5.886	5.937	5.982	5.982
6971.9044	6.698	6.225	5.892	5.949	5.982	7.836	6.216	5.887	5.937	5.982	5.982
6972.6706	6.701	6.224	5.894	5.942	5.982	7.722	6.275	5.899	5.693	4.344	4.402
6972.7523	6.689	6.216	5.888	5.940	5.982	7.843	6.238	4.679	4.324	4.384	5.977
6972.7981	6.686	6.214	5.886	5.937	5.982	8.139	5.235	4.679	4.323	4.385	5.977
6972.8133	6.694	6.220	5.890	5.943	5.982	8.367	5.239	4.677	4.322	4.385	5.977
6972.8698	6.695	6.224	5.895	5.943	5.982	8.534	5.235	4.673	4.320	4.382	5.977
6972.8952	6.699	6.230	5.899	5.950	5.982	8.660	5.228	4.673	4.318	4.381	5.977
6973.6677	6.673	6.209	5.886	5.934	5.982	8.859	5.233	4.674	4.320	4.381	5.977

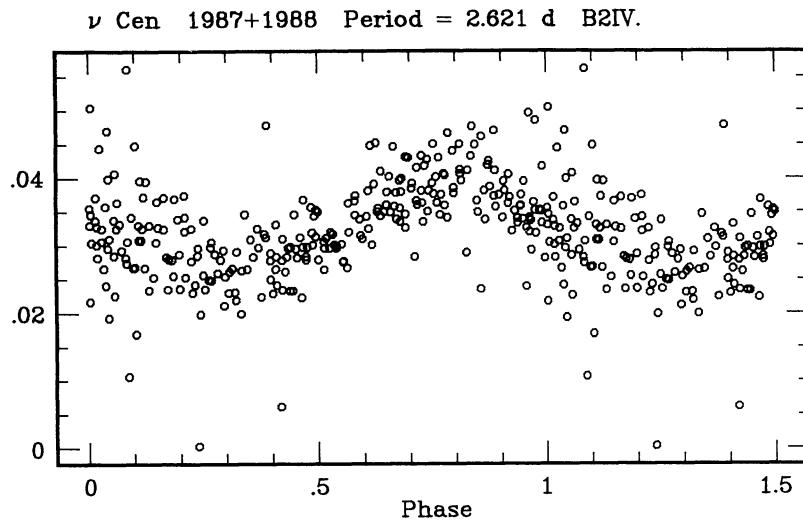


FIGURE 1. — Light curve of  $\nu$  Cen ( $P = 2.621$  d). In this and subsequent figures, the scale is in magnitudes and the epoch of phase zero is JD2446000.000.

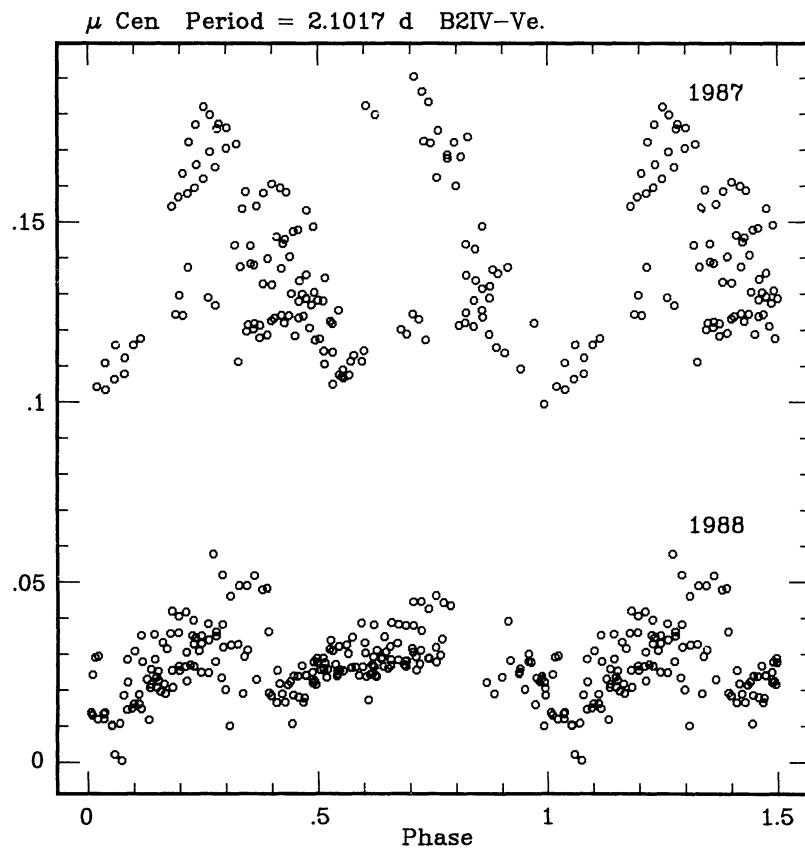
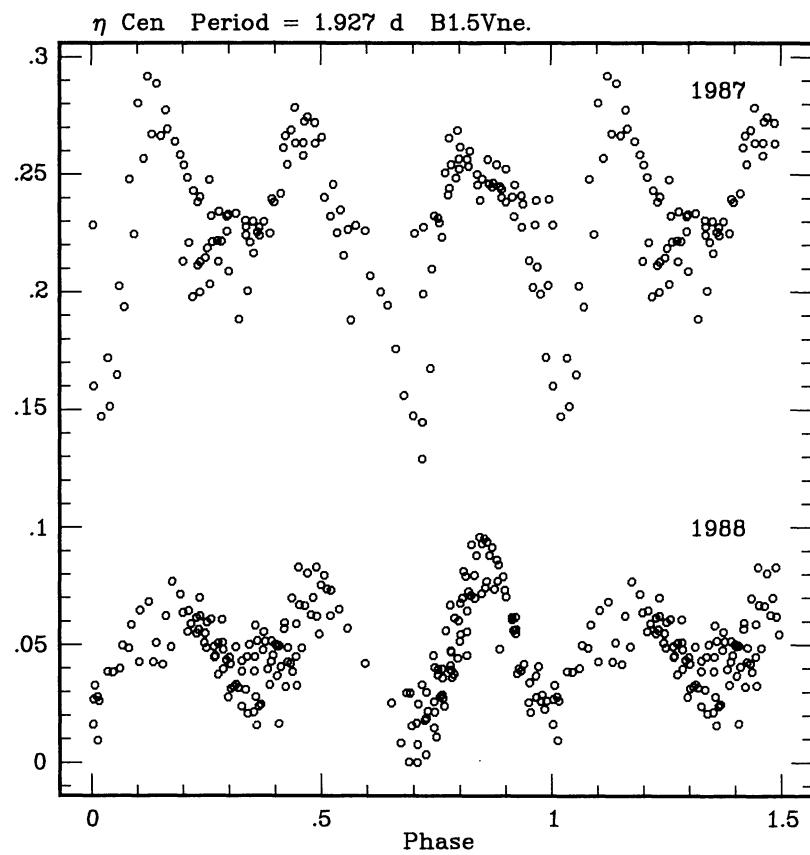
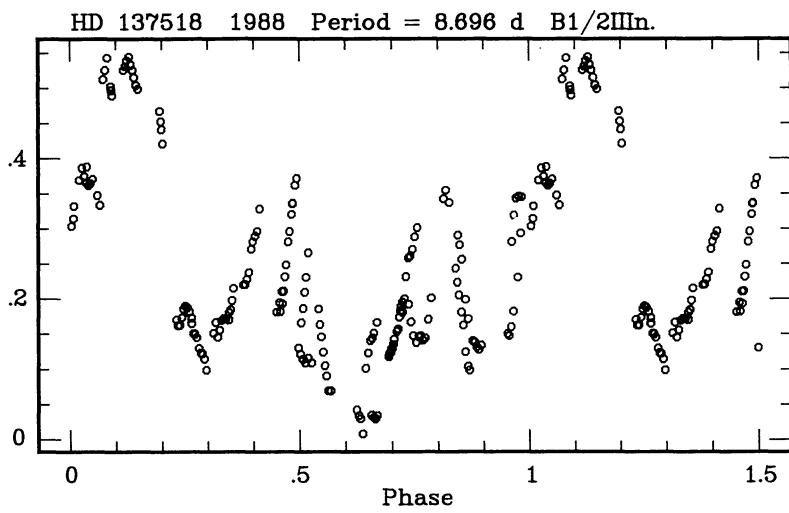
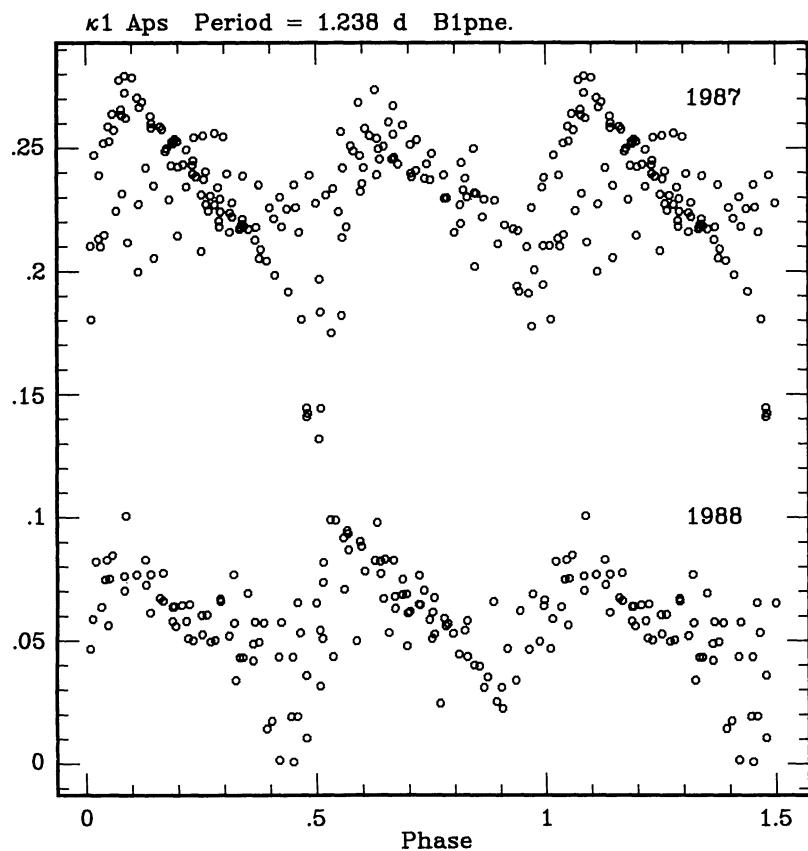
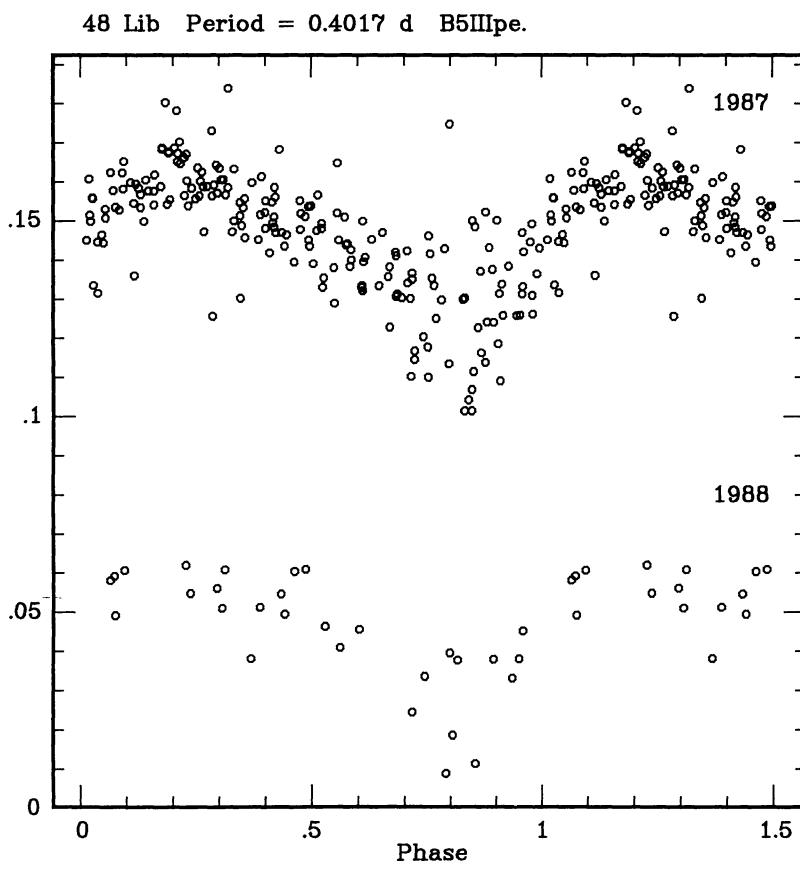
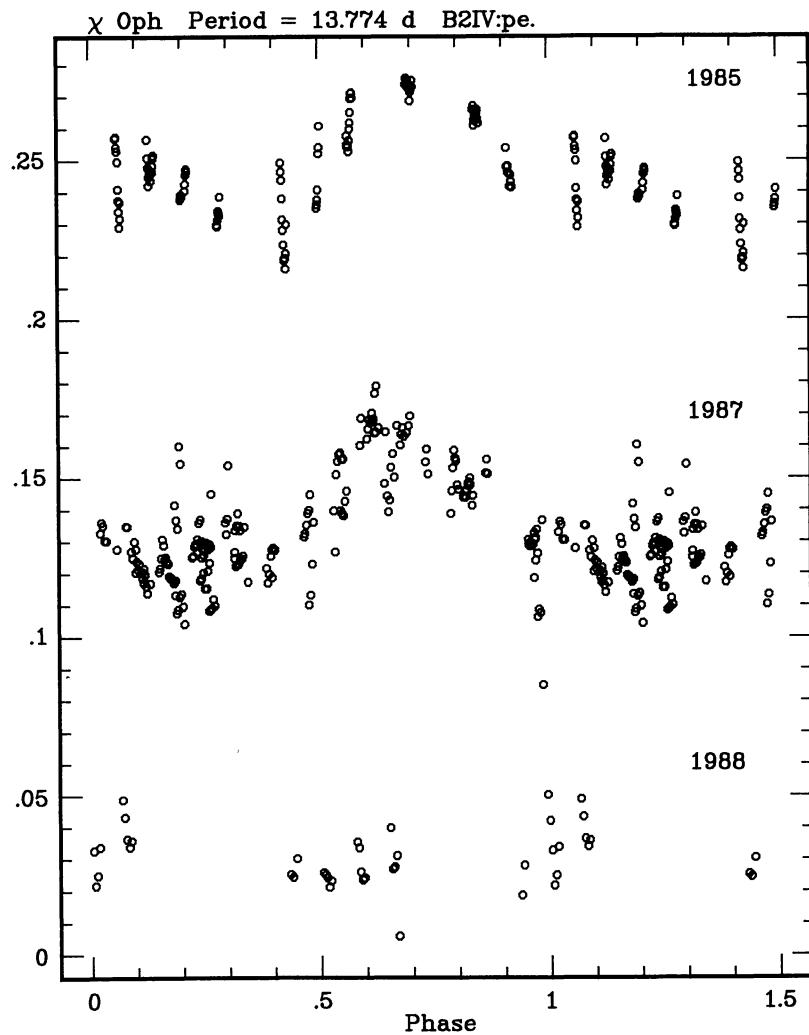
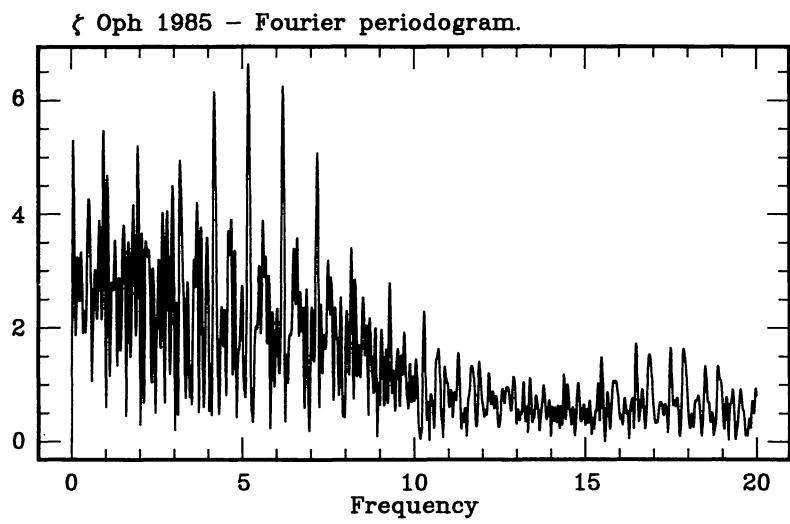
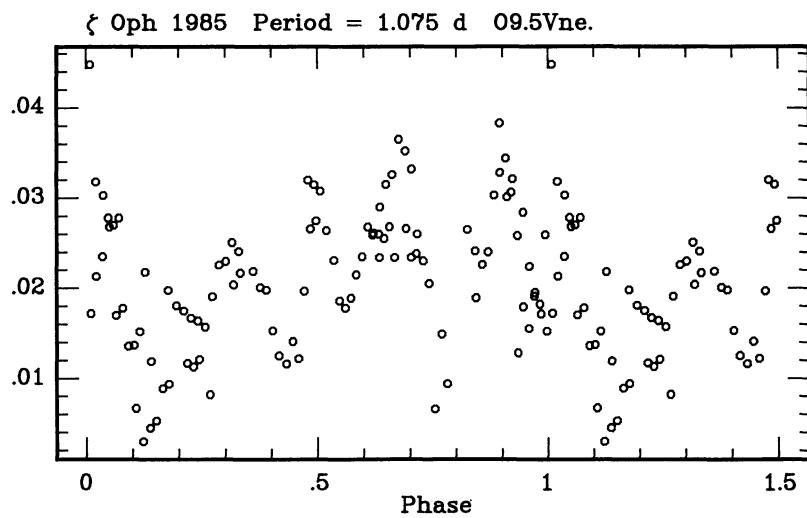
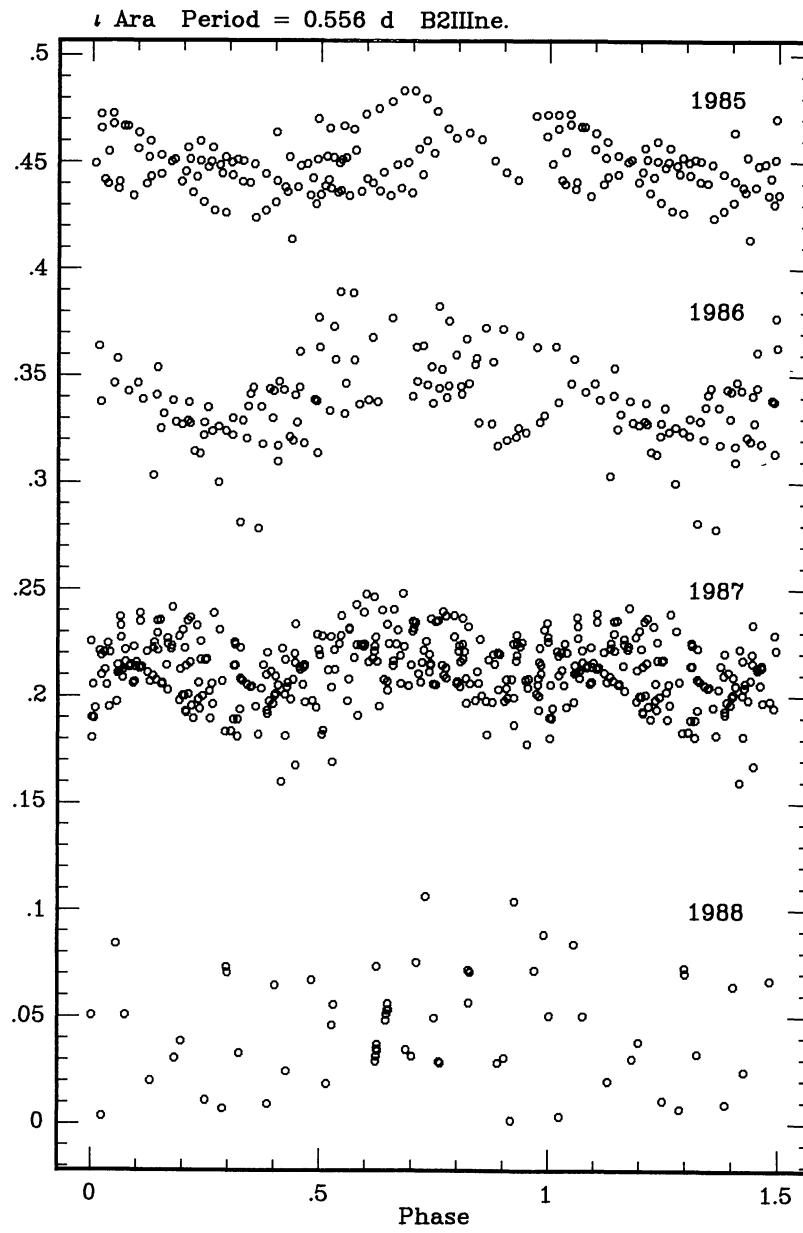


FIGURE 2. — Light curve of  $\mu$  Cen ( $P = 2.1017$  d).

FIGURE 3. — Light curve of  $\eta$  Cen ( $P = 1.927$  d).FIGURE 4. — Light curve of HD137518 ( $P = 8.696$  d).

FIGURE 5. — Light curve of  $\kappa^1$  Aps ( $P = 1.238$  d).FIGURE 6. — Light curve of 48 Lib ( $P = 0.4017$  d).

FIGURE 7. — Light curve of  $\chi$  Oph ( $P = 13.774$  d).FIGURE 8a. — Fourier periodogram for 1985 data of  $\zeta$  Oph. The frequency is in cycles  $d^{-1}$  and the semi-amplitude in millimagnitudes.

FIGURE 8b. — Light curve of  $\zeta$  Oph for 1985 ( $P = 1.075$  d).FIGURE 9. — Light curve of  $\iota$  Ara ( $P = 0.556$  d).

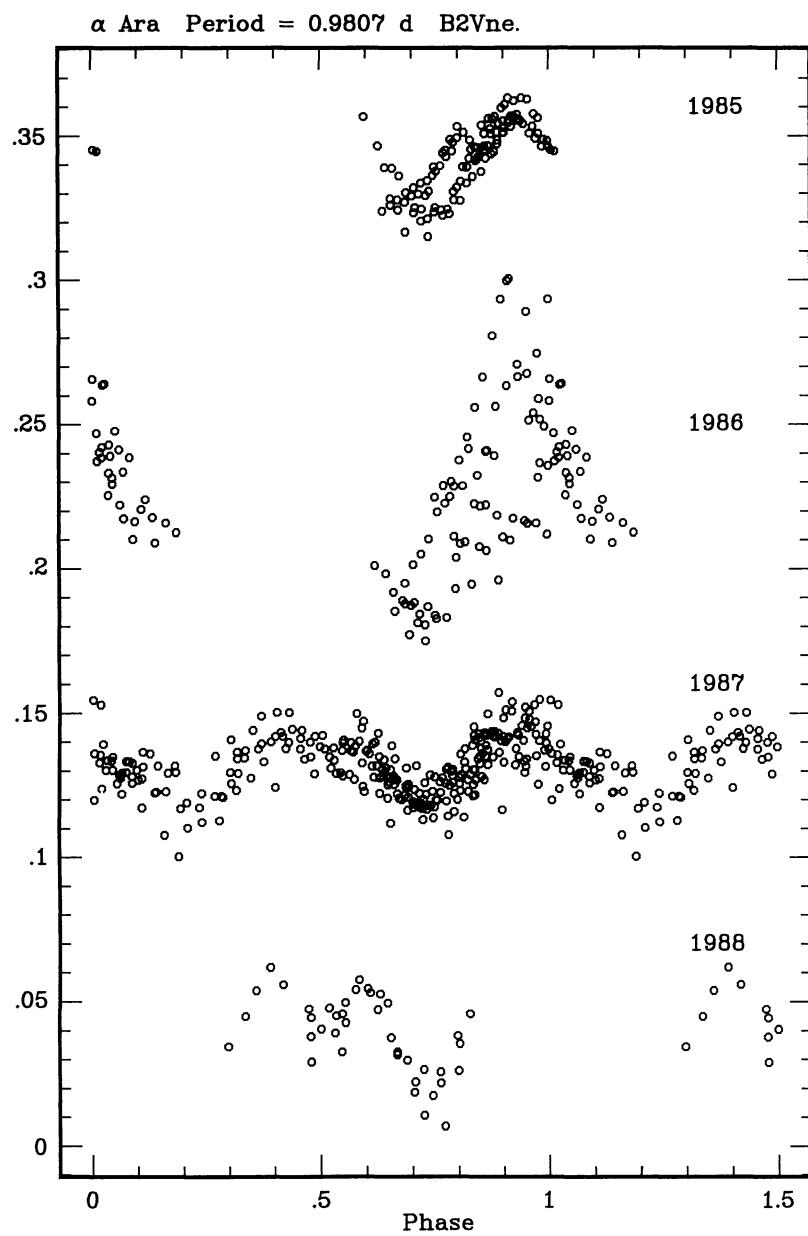


FIGURE 10. — Light curve of  $\alpha$  Ara ( $P = 0.9807$  d).

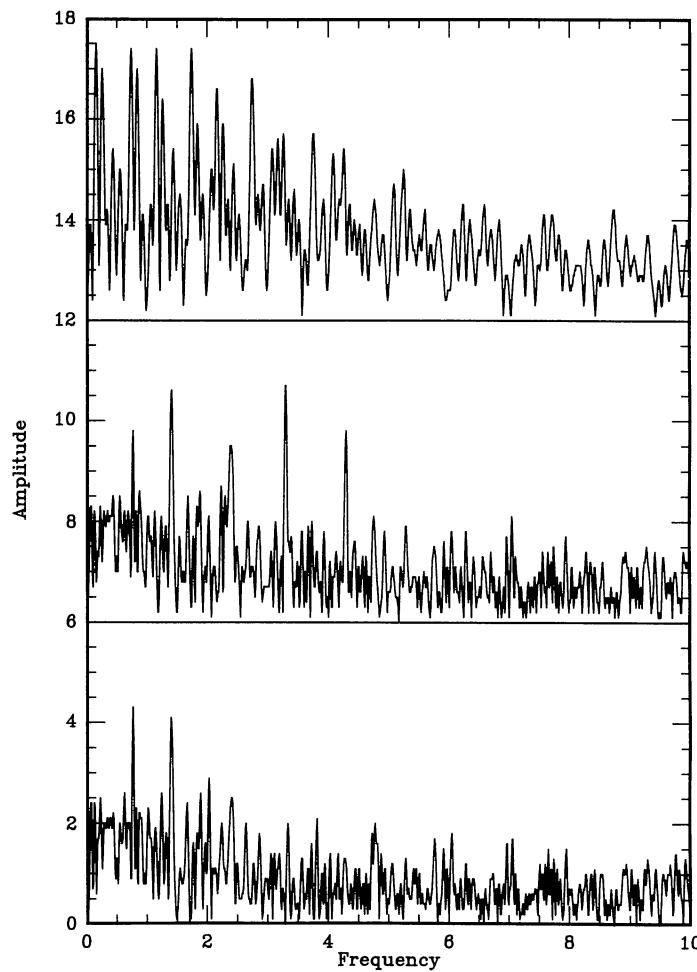


FIGURE 11a. — Fourier periodograms for V986 Oph. Top panel - 1985 data ; middle panel - 1987 data ; bottom panel - 1987 data prewhitened by  $3.30 \text{ d}^{-1}$ . The frequency is in cycles  $\text{d}^{-1}$  and the semi-amplitude in millimagnitudes.

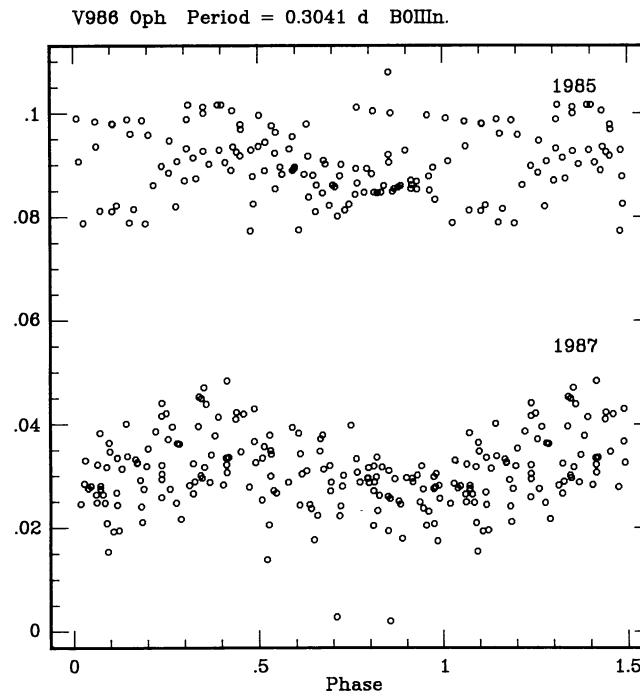
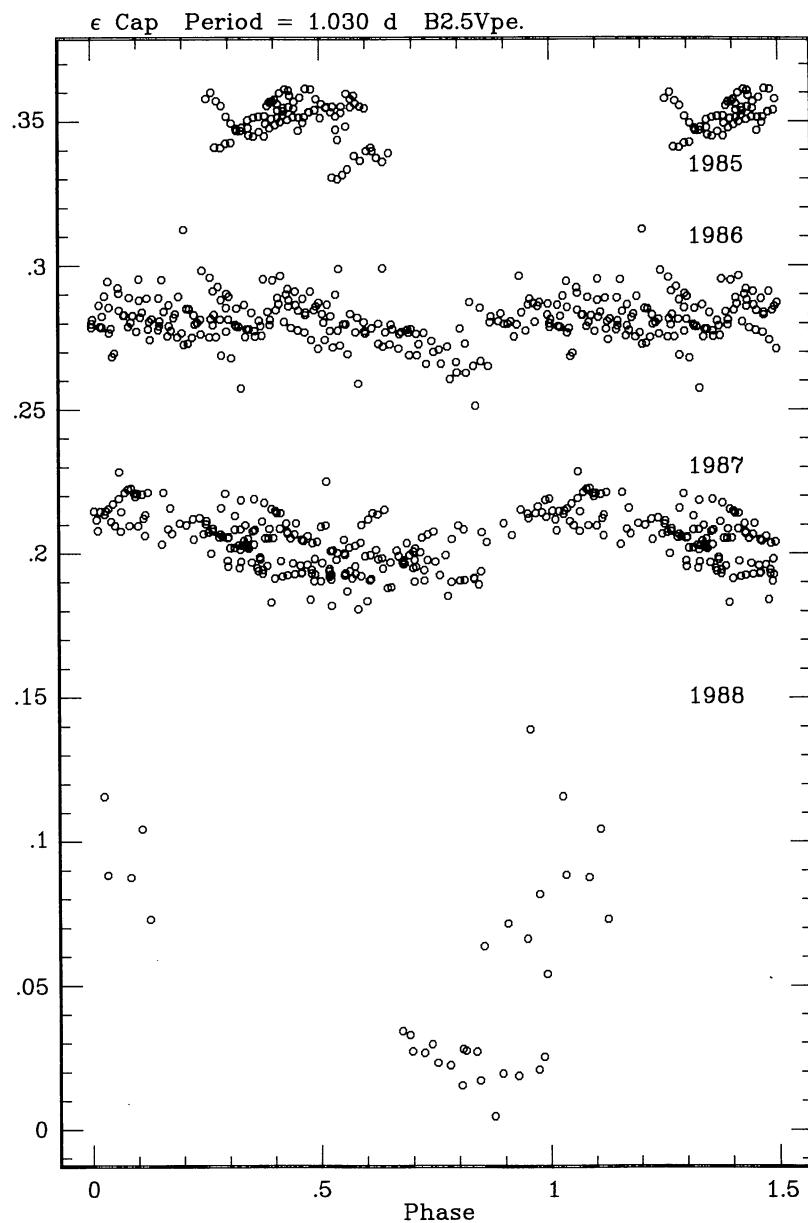


FIGURE 11b. — Light curve of V986 Oph ( $P = 0.3041 \text{ d}$ ).

FIGURE 12. — Light curve of  $\epsilon$  Cap ( $P = 1.030$  d).

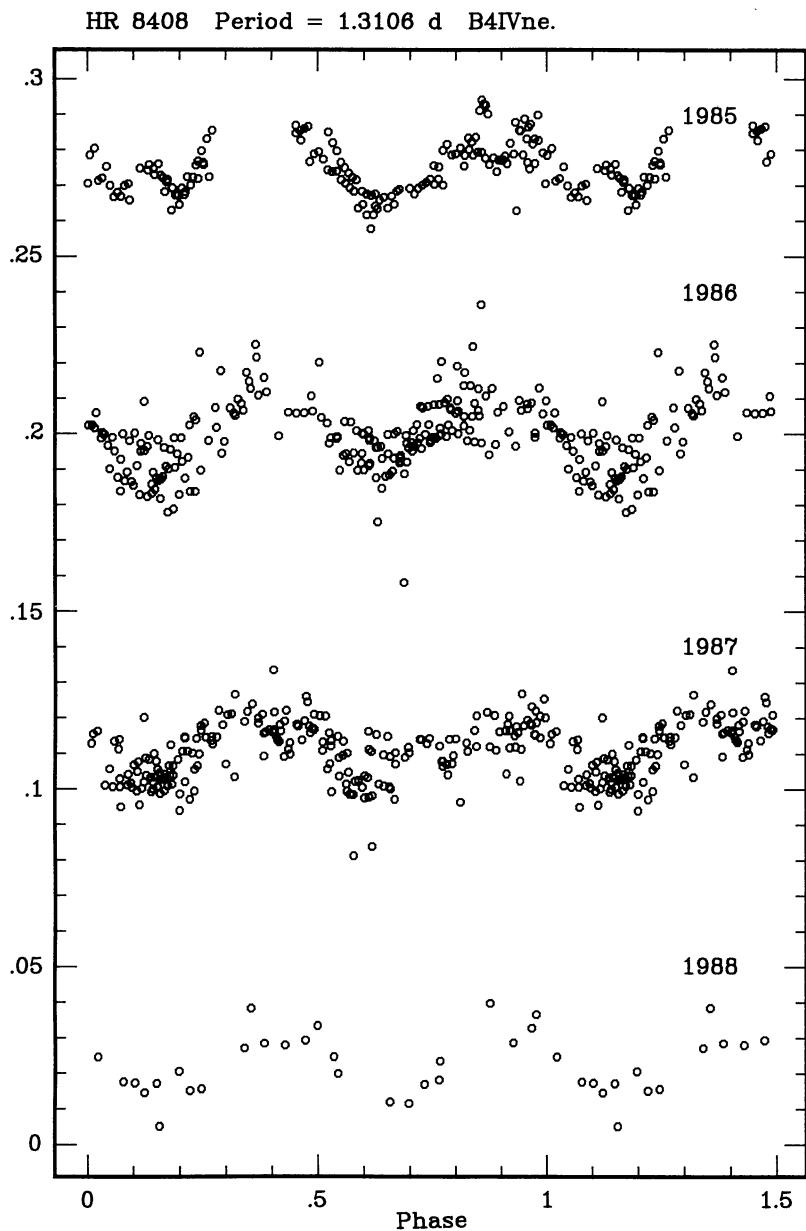
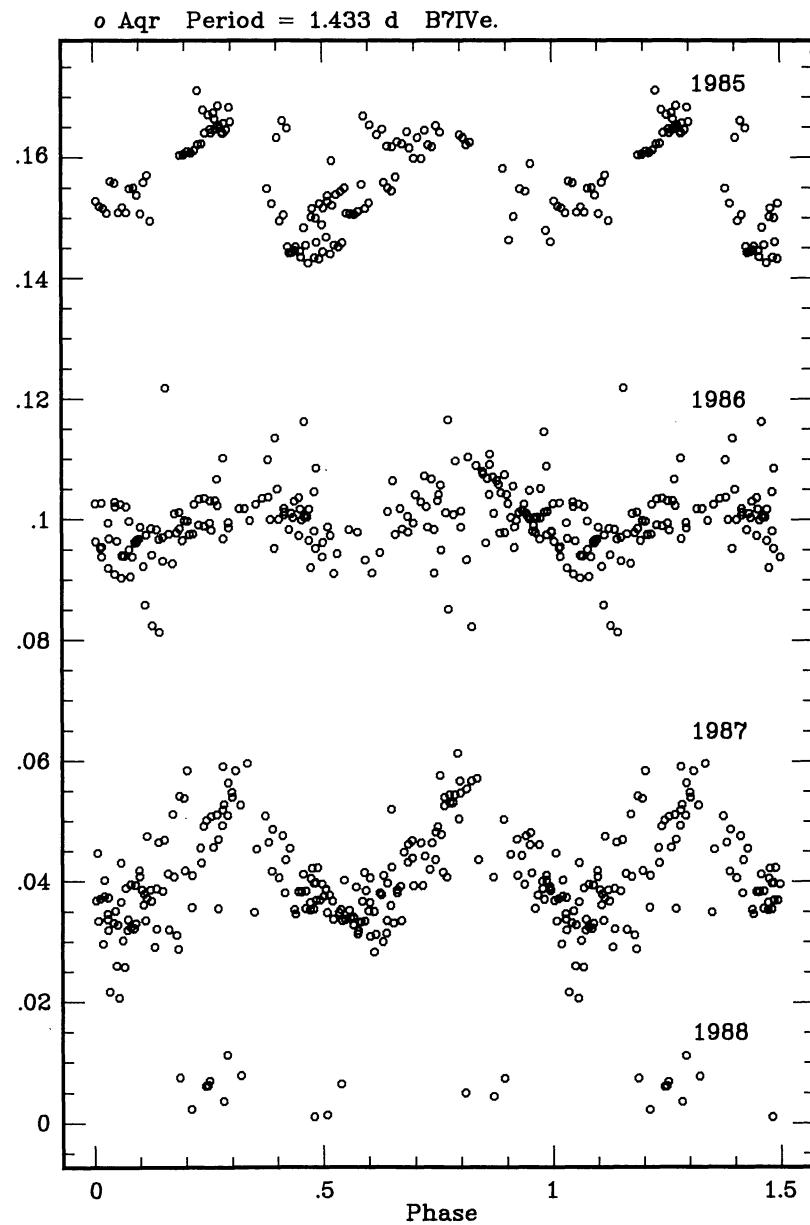
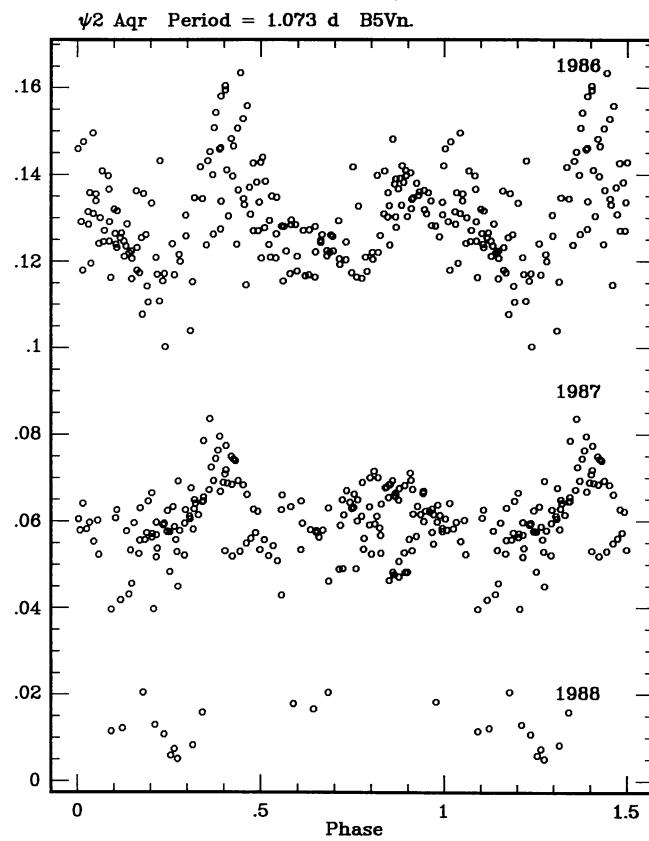
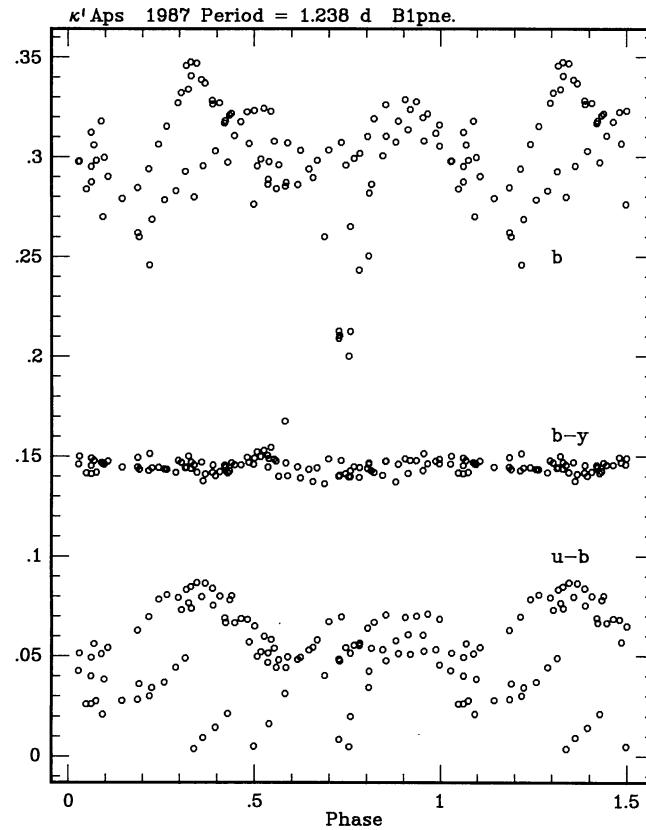
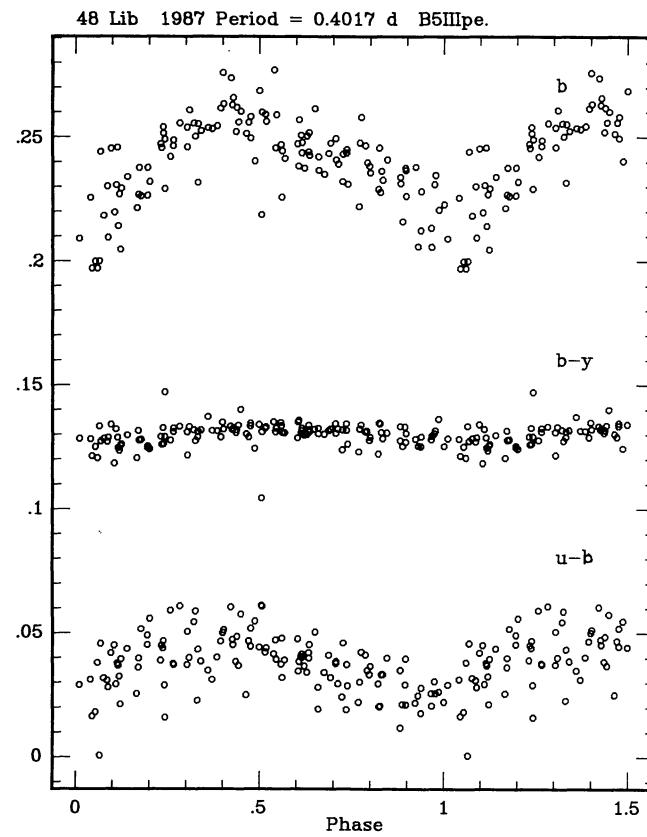
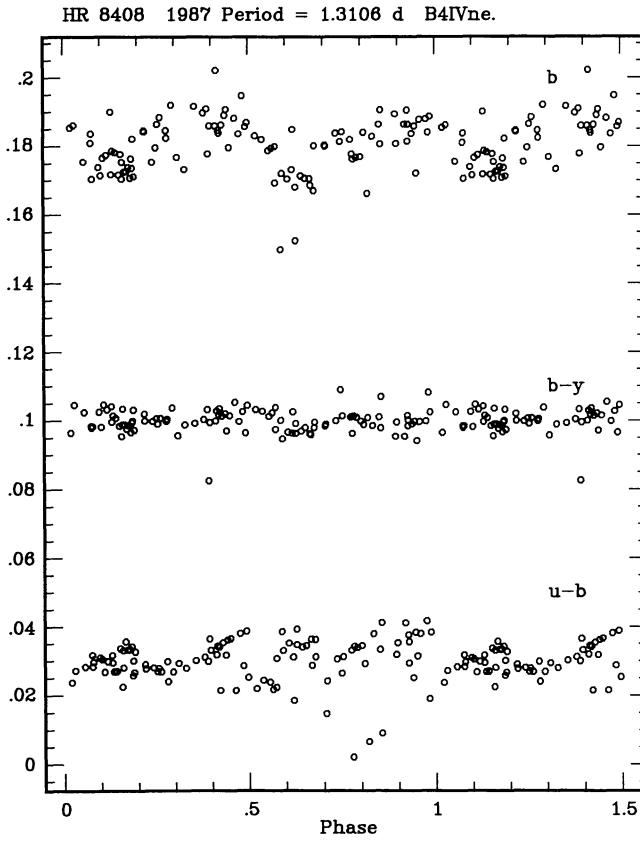
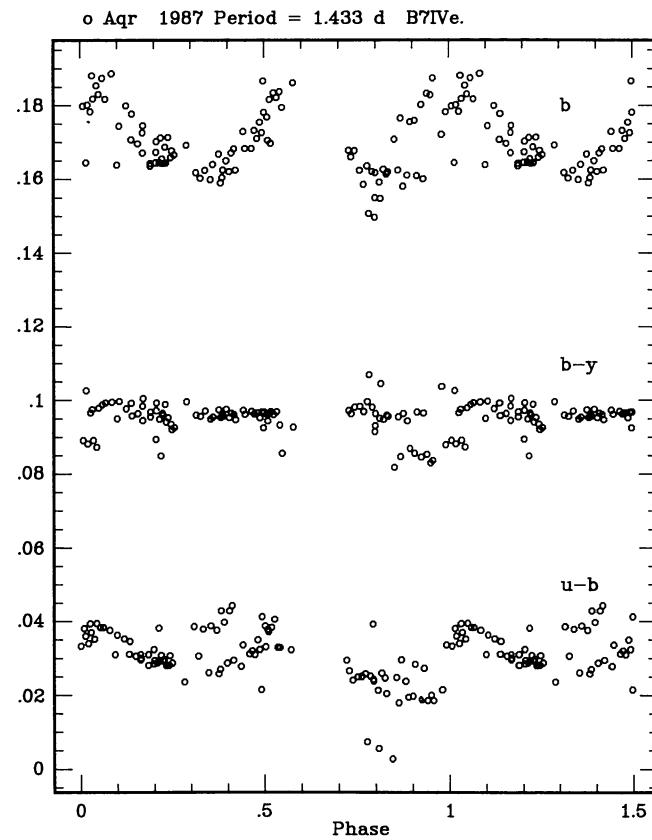
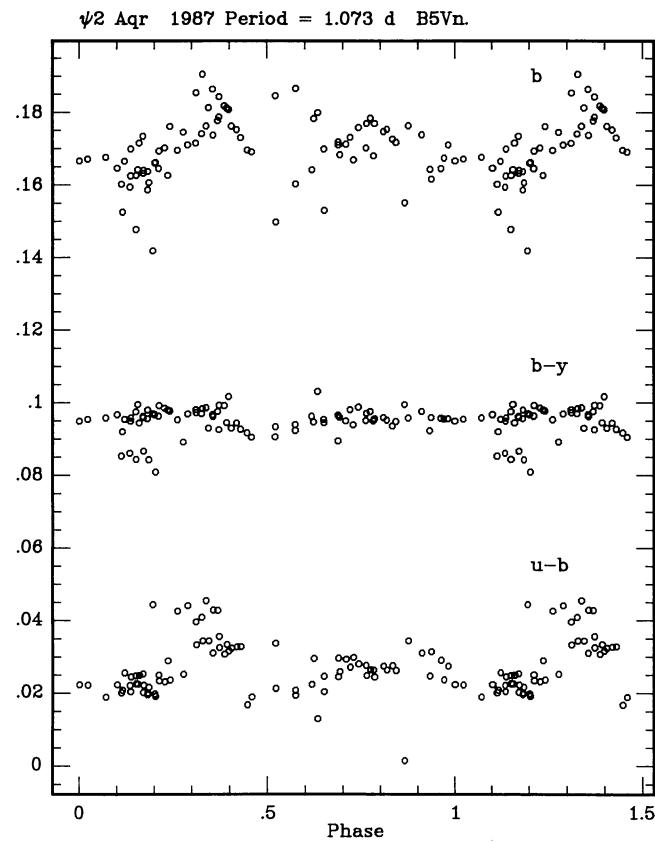


FIGURE 13. — Light curve of HR 8408 ( $P = 1.3106$  d).

FIGURE 14. — Light curve of *o* Aqr ( $P = 1.433$  d).

FIGURE 15. — Light curve of  $\psi^2$  Aqr ( $P = 1.073$  d).FIGURE 16. — Light and colour variations in  $\kappa^1$  Aps ( $P = 1.238$  d). In this and subsequent figures the colour and light variations are plotted according to the astronomical convention, i.e. with more negative values (bluer colour) upwards.

FIGURE 17. — Light and colour variations in 48 Lib ( $P = 0.4017$  d).FIGURE 18. — Light and colour variations in HR 8408 ( $P = 1.3106$  d).

FIGURE 19. — Light and colour variations in o Aqr ( $P = 1.433$  d).FIGURE 20. — Light and colour variations in  $\psi^2$  Aqr ( $P = 1.073$  d).