

WORK UPDATE

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UNIVERSIDADE DE SÃO PAULO | AUGUST 14, 2024



OVERVIEW

1 - Study of covariant derivatives

2 - Study of spinors

3 - Study of the tetrad

4 - Study of spin connection

5 - Study of Clifford algebra





LAGRANGIAN

Old

$$L = \frac{1}{2}\partial^{\mu}\phi\partial_{\mu} - V(\phi) + \bar{\psi}(i\overleftarrow{\partial} - M)\psi + \phi\bar{\psi}\psi$$

New

$$L = \sqrt{-g} \{ -V(\phi) \sqrt{1 - \alpha \partial^{\mu} \phi \partial_{\mu} \phi} + \frac{i}{2} [\bar{\psi} \gamma^{\mu} \nabla_{\mu} \psi - \bar{\psi} \overleftarrow{\nabla}_{\mu} \gamma^{\mu} \psi] - (M - \beta \phi) \bar{\psi} \psi \}$$





TACHYONIC FIELD

$$L = \sqrt{-g} \{ -V(\phi) \sqrt{1 - \alpha \partial^{\mu} \phi \partial_{\mu} \phi} \}$$

Energy density and the pressure density of the tachyon scalar field model:

$$\rho = \frac{V(\phi)}{\sqrt{1 - \dot{\phi}^2}}$$

$$P = -V(\phi)\sqrt{1 - \dot{\phi}^2}$$





FERMIONIC FIELD

$$\frac{i}{2} [\bar{\psi}\gamma^{\mu}\nabla_{\mu}\psi - \bar{\psi}\overleftarrow{\nabla}_{\mu}\gamma^{\mu}\psi]$$

$$C^{\mu}_{\kappa\lambda} = -4\pi G \epsilon_{abcd} e^a_{\lambda} e^b_{\kappa} e^{c\mu} (\bar{\psi} \gamma_5 \gamma^d \psi)$$

$$K^{\lambda}_{\nu\mu} = -2\pi G \epsilon_{abcd} e^a_{\mu} e^b_{\nu} e^{c\lambda} (\bar{\psi} \gamma_5 \gamma^d \psi)$$

Introducing the symmetric Ricci tensor: $R_{\mu\nu}=\widetilde{R}_{\mu\nu}+\widetilde{\nabla}_{\lambda}K^{\lambda}_{\mu\nu}-\widetilde{\nabla}_{\nu}K^{\lambda}_{\mu\lambda}+K^{\lambda}_{\theta\lambda}K^{\theta}_{\mu\nu}-K^{\lambda}_{\theta\nu}K^{\theta}_{\mu\lambda}$

$$\widetilde{R}_{\mu\nu} - \frac{1}{2}\widetilde{R}g_{\mu\nu} = 8\pi G(\widetilde{T}_{\mu\nu} - \frac{3}{2}\pi Gg_{\mu\nu}\sigma^2)$$



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NEXT STEPS

Tensor pertubation for fermionic field

CLASS

if you	prefer output in CAMR/He	ealPix/LensPix units/orde	r. set 'format' to 'ca	mb' in input file				
		header, set 'headers' to		iii iiipac iice				
		re C_l^phi-phi for the le						
Remembe:	r the conversion factors	s:						
C_l^dd	$(deflection) = l(l+1) C_{-}$	_l^phi-phi						
C_l^gg	(shear/convergence) = 1/	/4 (l(l+1))^2 C_l^phi-phi						
2:	7:TPhi		4:TE	5:BB	6:phiphi			
	1.404305087108e-10	8:Ephi 5.182392798948e-15	3.897398229419e-13	0.000000000000e+00	8.320634			
7e-09	4.654920130753e-10	-2.059561357077e-12		0.0000000000000000000000000000000000000	6.320034			
	1.325290292153e-10	7.446442600614e-15	4.589184267630e-13	0.00000000000e+00	5.084366			
0e-09	3.293216034626e-10	-1.667994027860e-12		0.0000000000000000000000000000000000000	3.004300			
	1.250005333600e-10	7.281585906855e-15	4.472542531455e-13	0.000000000000e+00	3.520716			
3e-09	2.491049265376e-10	-1.194672552031e-12		3.0000000000000000000000000000000000000	3.020120			
	1.192731114781e-10	5.531188719197e-15	3.955420357531e-13	0.000000000000e+00	2.619506			
6e-09	1.966181163637e-10	-7.625969446411e-13						
	1.152516403912e-10	3.464613201530e-15	3.292652694992e-13	0.00000000000e+00	2.040065			
9e-09	1.598901093992e-10	-4.116504341388e-13						
	1.126136674560e-10	1.924330768169e-15	2.644077824186e-13	0.00000000000e+00	1.641136			
7e-09	1.329384335627e-10	-1.550999202817e-13						
	1.109583013716e-10	1.105162215578e-15	2.094454657941e-13	0.00000000000e+00	1.352895			
1e-09	1.124470687839e-10	9.641968930672e-15						
	1.100985272189e-10	7.911648734672e-16	1.680373319566e-13	0.00000000000e+00	1.136915			
1e-09	9.648263094699e-11	9.511561168800e-14						
	1.098448341413e-10	6.852895378100e-16	1.406432042403e-13	0.00000000000e+00	9.702839			
4e-10	8.381531595207e-11	1.196818470011e-13						
	1.100070012772e-10	6.074803561535e-16	1.258009852578e-13	0.00000000000e+00	8.310422973254e-10	7.427411362106e-11	1.037537755311e-13	
	1.105527993374e-10	5.186527643866e-16	1.208969832918e-13	0.00000000000e+00	7.260472361994e-10	6.578358317214e-11	6.701046202050e-14	
	1.113495809980e-10	4.481728358072e-16	1.228519060080e-13	0.00000000000e+00	6.398566843841e-10	5.867813063624e-11	2.487788372596e-14	
	1.123367738049e-10	4.210819271608e-16	1.289086186422e-13	0.00000000000e+00	5.681510908433e-10	5.264550828855e-11	-1.155258667909e-14	
	1.135201843317e-10	4.327036319947e-16	1.369381816230e-13	0.0000000000e+00	5.078020536465e-10	4.755812491006e-11	-3.603790485079e-14	
	1.148459736594e-10	4.628738248884e-16	1.454631601526e-13	0.00000000000e+00	4.564951732576e-10	4.320355929953e-11	-4.698476168460e-14	
	1.162966147486e-10	4.984478020560e-16	1.536776020916e-13	0.00000000000e+00	4.124856414927e-10	3.937188827107e-11	-4.626103396763e-14	
	1.178464650474e-10	5.421957444395e-16	1.613412121429e-13	0.000000000000e+00	3.742953717545e-10	3.599780567649e-11	-3.765644180635e-14	
	1.194976360677e-10	6.035395360864e-16	1.686553790887e-13	0.000000000000e+00	3.413048797651e-10	3.304216498065e-11	-2.505115332398e-14	
	1.212093586417e-10	6.924899532311e-16	1.759792102925e-13	0.00000000000e+00	3.122545768496e-10	3.039362385799e-11	-1.298257237758e-14	
	1.229912121585e-10 1.248376386126e-10	8.095022392855e-16 9.508917277756e-16	1.836258776879e-13 1.918588269560e-13	0.00000000000e+00 0.00000000000e+00	2.866858975902e-10 2.639815602939e-10	2.805197678598e-11 2.600972435387e-11	-3.905563030686e-15 5.856652247094e-16	
	1.248376386126e-10 1.267467793450e-10	1.114547629876e-15	1.918588269560e-13 2.006482837858e-13	0.000000000000e+00	2.639815602939e-10 2.438365174244e-10	2.421389448061e-11	6.212074529831e-16	
	1.286973377485e-10	1.114547629876e-15 1.298620017681e-15	2.000482837858E-13 2.097914151764E-13	0.0000000000000000000000000000000000000	2.438365174244e-10 2.257897356727e-10	2.421389448061e-11 2.257527031265e-11	-2.724501481498e-15	
	1.306904397564e-10	1.503464353781e-15	2.189494134169e-13	0.00000000000e+00	2.095956052707e-10	2.107378512239e-11	-7.866577522241e-15	
	1.327210446523e-10	1.729857597048e-15	2.276991279793e-13	0.00000000000e+00	1.949363796914e-10	1.969914926530e-11	-1.292614950737e-14	
	1.347983236592e-10	1.977628447008e-15	2.357464443728e-13	0.00000000000e+00	1.817355427814e-10	1.845280244592e-11	-1.695555873977e-14	
	1.369195034590e-10	2.245986339246e-15	2.427941079296e-13	0.00000000000e+00	1.698325814232e-10	1.732638233194e-11	-1.93333873977e-14 -1.918435417597e-14	
	1.390675972263e-10	2.534361385785e-15	2.485470445627e-13	0.00000000000e+00	1.589377846305e-10	1.629914848414e-11	-1.896735160698e-14	
			2.529357545324e-13	0.00000000000e+00	1.489907339104e-10	1.536383928368e-11	-1.674580148639e-14	
	1.412463149111e-10	2.044930091000e-15						
)	1.412463149111e-10 1.434553205707e-10	2.844938691068e-15 3.181182287947e-15	2.559839693731e-13	0.00000000000e+00	1.399172750641e-10	1.450844577207e-11	-1.349742672400e-14	

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WORK UPDATE

Alcides Vicente de Mello

UNIVERSIDADE DE SÃO PAULO | SEPTEMBER II, 2024





- tCl (Temperature Power Spectrum)
- pCl (Polarization Power Spectrum)
- ICI (Lensing Power Spectrum)
- As (Amplitude of the Primordial Spectrum)
- mPk (Power Spectrum of Matter)
- pCl (Polarization Power Spectrum)





```
#number_count_contributions = # nCl contributions
#(density,lensing,rsd,gr) -> (density, lensing, rsd+doppler, all others)
selection=gaussian
                                 # nCl window function type
#selection_mean=1.0,1.25,2.0,3.5 # Mean redshifts of nCl window functions
#selection_width = 0.1
                          # Widths of nCl window functions
#selection_bias =
                                 # Biases of nCl window functions
#selection_magnification_bias = # Biases of lensing of nCl
#non_diagonal=3
                                  # Number of non-diagonal terms
l_max_scalars = 3000
                                 # lmax of CMB for scalar mode
                                 # lmax of CMB for tensor mode
#l_max_tensors = 500
\#l_{max_lss} = 300
                                  # lmax of nCl
P_k_{max_h/Mpc} = 2.
                                 # Maximum k for P(k) in 1/Mpc
\#P_k_max_1/Mpc = 0.7
                                  # Maximum k for P(k) in h/Mpc
z_{pk} = 0.13
                                     # Redshifts of P(k,z)
```





```
# Dimensionless reduced Hubble parameter (H_0 / (100km/s/Mpc))
h = 0.67810
#H0 = 67.810
                                  # Hubble parameter in km/s/Mpc
                                  # Angular size of the sound horizon, exactly 100(ds_dec/da_dec)
#100*theta_s = 1.041783
                                  # with decoupling time given by maximum of visibility function
                                  # (different from theta_MC of CosmoMC and
                                  # slightly different from theta_* of CAMB)
T_{cmb} = 2.7255
                                  # CMB temperature
                                  # Reduced baryon density (Omega*h^2)
omega_b = 0.02238280
\#0mega_b =
                                  # Baryon density
omega\_cdm = 0.1201075
                                  # Reduced cold dark matter density (Omega*h^2)
#Omega_cdm =
                                  # CDM density
omega\_dcdmdr = 0.0
                                  # Reduced decaying dark matter density (Omega*h^2)
#Omega_dcdmdr =
                                  # DCDM density
\#Gamma_dcdm = 0.0
                                  # Decay constant of DCDM in km/s/Mpc
Omega_k = 0.
                                  # Curvature density
Omega_fld = 0
                                  # Dark Energy as Fluid density
Omega\_scf = 0
                                  # Dark Energy as Scalar field density
```

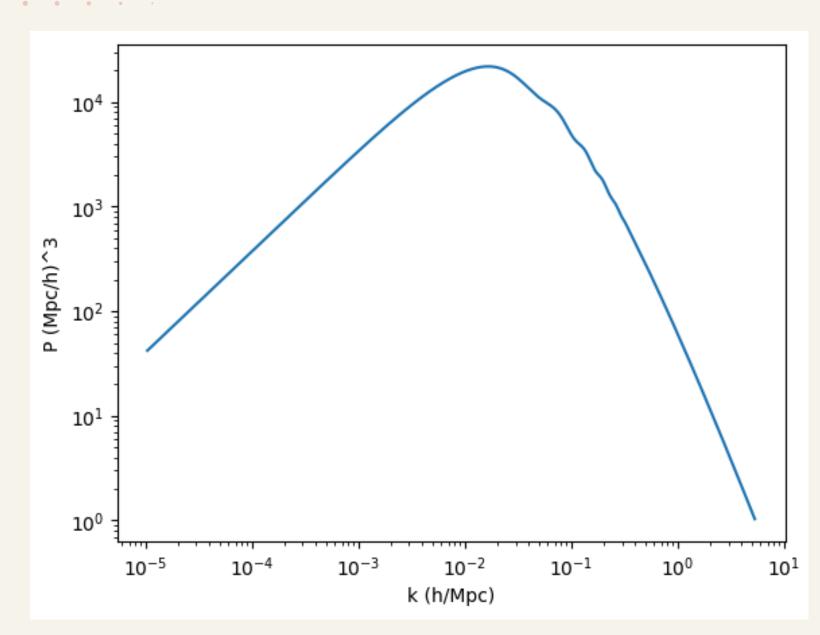


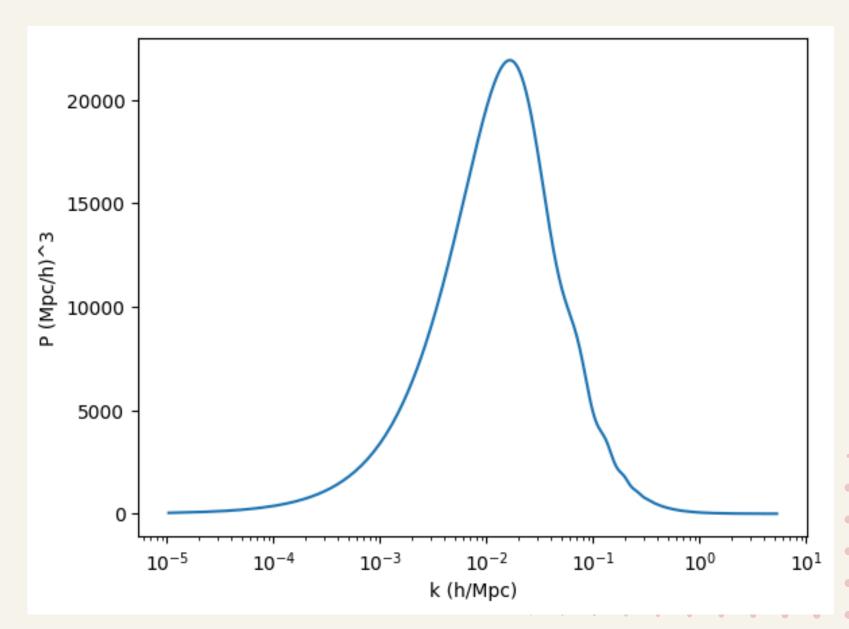


```
----> Primordial parameters:
P_k_ini type = analytic_Pk
                                  # Select primordial spectrum
#('analytic_Pk','inflation_V','inflation_H','inflation_V_end','two scales','external_Pk')
k_{pivot} = 0.05
                                  # Pivot scale for A_s,n_s
A_s = 2.100549e-09
                                  # Amplitude of prim spectrum
\#\ln 10^{10}A_s = 3.0980
                                  # In Amplitude of prim spectrum
\# sigma8 = 0.848365
                                  # Final density averaged over 8 Mpc
n_s = 0.9660499
                                  # Spectrum tilt
alpha_s = 0.
                                  # Spectrum running of tilt
\#r = 1.
                                  # If tensors are activated
# See explanatory.ini for more information about all the different primordial spectra
```

















- r (Tensor-to-Scalar Ratio)
- Change more parameters







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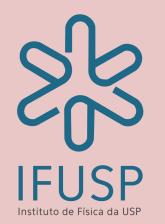
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WORK UPDATE

Alcides Vicente de Mello



UNIVERSIDADE DE SÃO PAULO | 03 de Dezembro, 2024

O CLASS (Cosmic Linear Anisotropy Solving System) é um código computacional amplamente utilizado em cosmologia para calcular perturbações lineares e prever observáveis cosmológicos, como o espectro de potência da radiação cósmica de fundo (CMB), o espectro de potência da matéria, e outros. Ele é uma ferramenta essencial para estudos de estrutura em larga escala do universo e análises de dados cosmológicos.





- 1. Versatilidade: É utilizado para diversos modelos cosmológicos, desde o modelo padrão (ΛCDM) até extensões que consideram energia escura dinâmica, gravidade modificada, neutrinos massivos, entre outros.
- 2. Eficiência: É projetado para ser rápido e eficiente, adequado tanto para simulações grandes quanto para ajustes de parâmetros em observações cosmológicas.
- 3. Modularidade: Possui uma estrutura modular que facilita a personalização e a adição de novos modelos ou extensões cosmológicas.







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