



1. DATA

3. Results & Conclusion

2. Methodology

4. NEXT STEPS





Simulation with thermal noise and foreground

21 cm map<sup>1</sup> + foreground map<sup>2</sup>





+ thermal noise<sup>3</sup>



 $\rightarrow$ 

foreground cleaning<sup>4</sup>



Simulates the observed data

Simulates the instrumental contamination

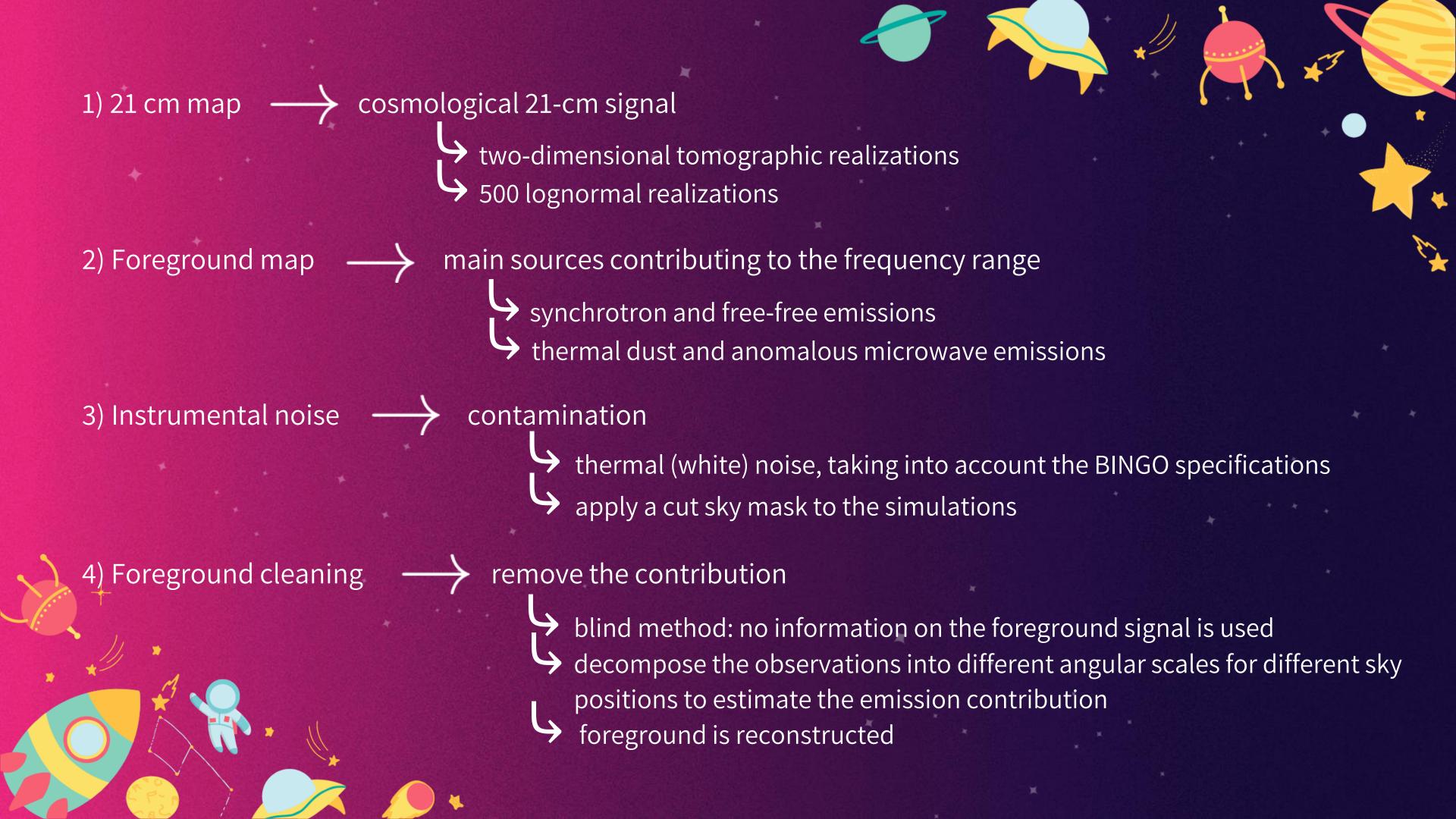
final simulation: data I used

CMB

Planck 2018

BAO

SDSS, SDSS III and 6dFGS







• 3 dark sector interaction models: <a href="https://arxiv.org/abs/2308.05807">https://arxiv.org/abs/2308.05807</a>

Parameter		Prior		
$\Omega_b h^2$		[0.005, 0.04]		Baryonic matter density
$\Omega_c h^2$		[0.001, 0.5]		Dark matter density
$100\theta_{s}$		[1.03, 1.05]		$\rightarrow$ BAO's angular size in the recombination sky
$\ln(10^{10}A_S)$		[2.7, 4.0]		Primordial spectrum amplitude
$n_{s}$		[0.9, 1.07]		$\longrightarrow$ Spectral index of the primordial space
	Model 1	Model 2	Model 3	
$\omega_0$	[-3.0, -0.3]	[-3.0, -1.0]	[-3.0, -1.0]	Dark energy equation of state
$\lambda_{1(2)}$	[-1.5, 1.5]	[0.0, 0.04]	[0.0, 0.04]	-> Coupling constant

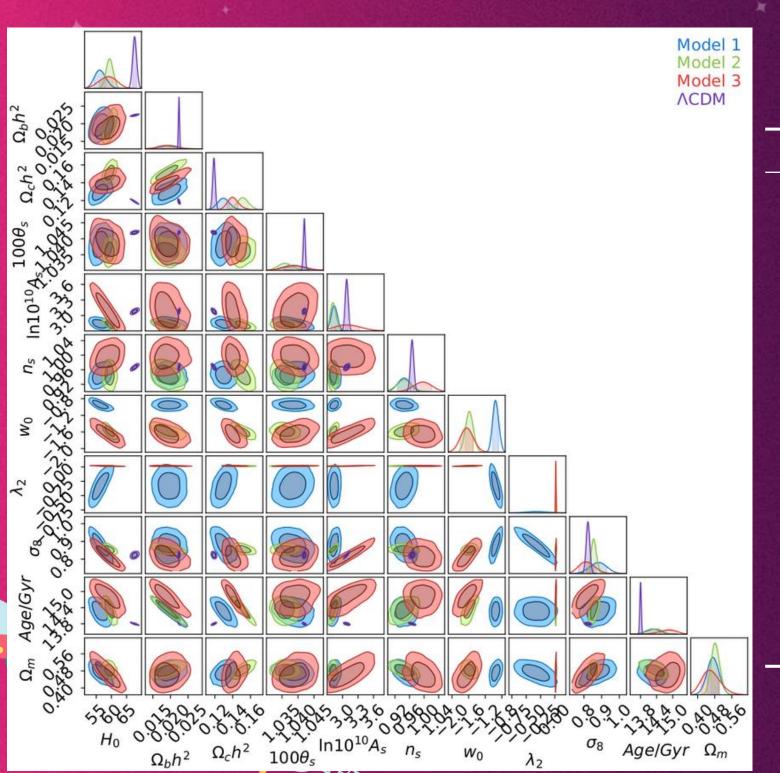
Depends on the dark energy Depends on the dark energy Depends on the dark matter

	Model	Q	ω	λ
$\leftarrow$		$3\lambda_2 H  ho_d$	$-1 < \omega < 0$	$\lambda_2 < 0$
$\leftarrow$	- II	$3\lambda_2 \ H ho_d$	$\omega < -1$	$0 < \lambda_2 < -2\omega\Omega_c$
$\leftarrow$	- 111	3λ <sub>1</sub> Ηρ <sub>c</sub>	$\omega < -1$	$0 < \lambda_1 < -\frac{\omega}{4}$



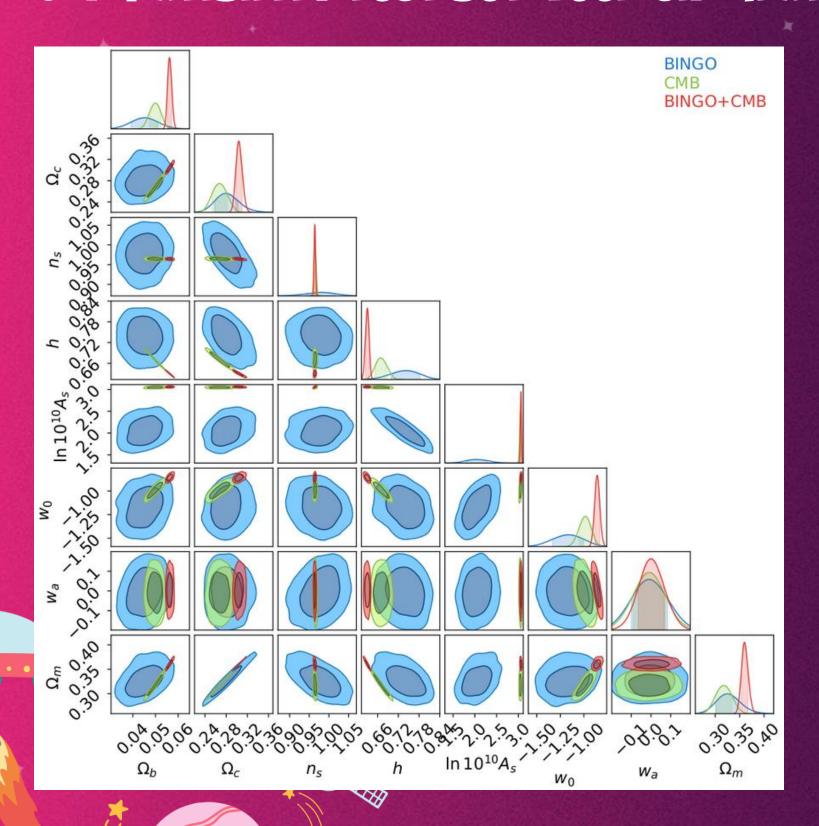


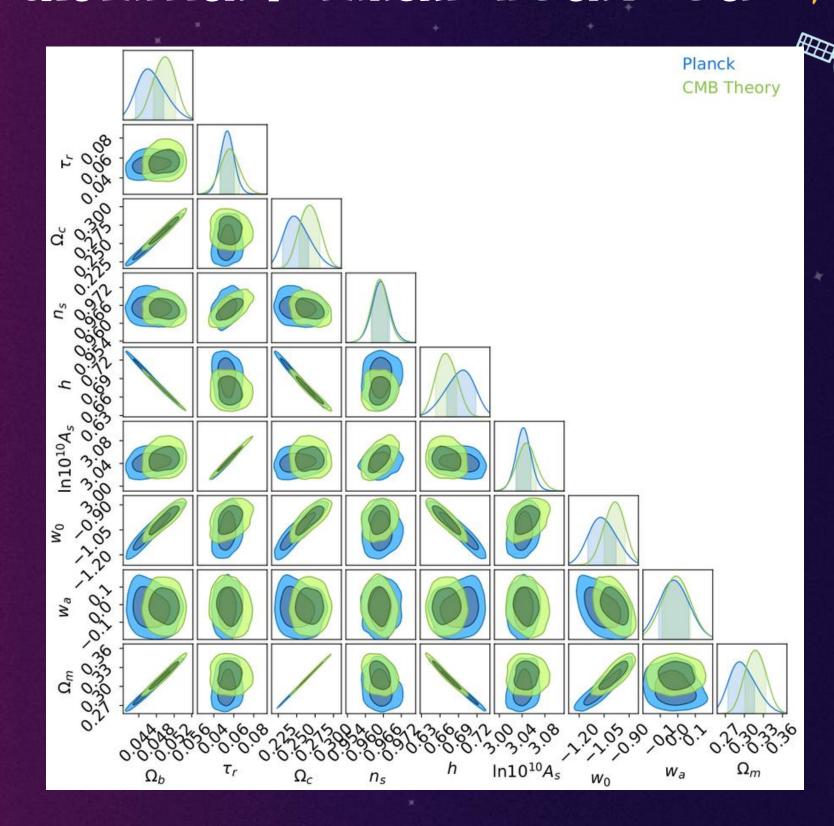
## COSMOLOGICAL PARAMETERS - BINGO



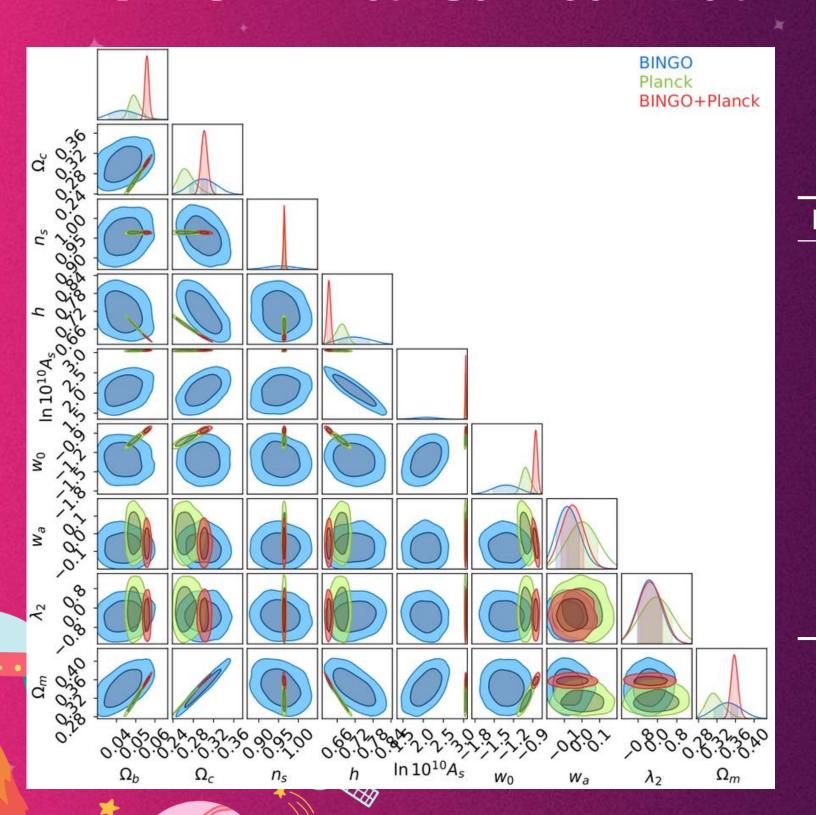
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Parameter	Model /ICDM	Model 1	Model 2	Model 3
$\Omega_b h^2$	$0.0224 \pm 0.0002$	$0.0196^{+0.0024}_{-0.0023}$	$0.0187^{+0.0025}_{-0.0023}$	$0.0185^{+0.0028}_{-0.0024}$
$\Omega_c h^2$	$0.1189^{+0.0014}_{-0.0015}$	$0.1294^{+0.0067}_{-0.0064}$	$0.1510^{+0.0066}_{-0.0060}$	$0.1396^{+0.0065}_{-0.0060}$
$100 heta_{\scriptscriptstyle S}$	$1.0420 \pm 0.0003$	$1.0384^{+0.0030}_{-0.0031}$	$1.0359 \pm 0.022$	$1.0392^{+0.0033}_{-0.0035}$
$\ln(10^{10}A_s)$	$3.080 \pm 0.029$	$2.838^{+0.061}_{-0.059}$	$2.814^{+0.055}_{-0.051}$	$3.07^{+0.26}_{-0.23}$
$n_{\scriptscriptstyle S}$	0.967 ± 0.005	$0.945^{+0.017}_{-0.018}$	$0.946 \pm 0.020$	$0.997^{+0.24}_{-0.25}$
$\omega_0$		$-0.953^{\pm 0.075}_{-0.071}$	$-1.58^{+0.09}_{-0.10}$	$-1.65 \pm 0.16$
$\lambda_{(1;2)}$		$-0.33 \pm 0.18$	$0.0156^{+0.0050}_{-0.0049}$	$0.0161^{+0.0047}_{-0.0045}$
$H_0$	$67.7 \pm 0.7$	$56.1^{+1.9}_{-1.8}$	59.4+1.1	$59.1^{+2.6}_{-3.1}$
$\sigma_8$	$0.822^{+0.11}_{-0.10}$	$0.883^{+0.045}_{-0.049}$	$0.854^{+0.016}_{-0.015}$	$0.810^{+0.047}_{-0.041}$
Age/Gyr	$13.80 \pm 0.03$	$14.28^{+0.25}_{-0.27}$	$14.37^{+0.24}_{-0.27}$	$14.87^{+0.29}_{-0.35}$
$\Omega_m$	*	$0.471^{+0.029}_{-0.027}$	$0.482^{+0.021}_{-0.020}$	$0.455^{+0.047}_{-0.40}$

## COMPATIBILITY TEST BETWEEN SIMULATIONS - CHEVALLIER-POLARSKI-LINDER MODEL





## COMPATIBILITY TEST BETWEEN THE SIMULATION AND REAL DATA - CPL MODEL



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Parameter	Theoretical	CMB - Planck	BINGO	CMB + BINGO
$\Omega_b$	0.0493	$0.0460^{+0.0036}_{-0.0028}$	$0.0473^{+0.0080}_{-0.0076}$	$0.0558 \pm 0.0011$
$\Omega_c$	0.2645	$0.246^{+0.020}_{-0.016}$	$0.292^{+0.031}_{-0.030}$	$0.302^{+0.007}_{-0.006}$
$ au_r$	0.0544	$0.0534^{+0.0071}_{-0.0070}$		$0.0574^{+0.0047}_{-0.0046}$
$n_{\scriptscriptstyle S}$	0.9649	$0.9650^{+0.0030}_{-0.0029}$	$0.958^{+0.0038}_{-0.0042}$	$0.9643 \pm 0.0020$
h	0.6736	$0.697^{+0.021}_{-0.027}$	$0.740^{+0.048}_{-0.050}$	$0.632 \pm 0.006$
$\ln(10^{10}A_S)$	3.044	$3.042^{+0.014}_{-0.013}$	$1.99^{+0.027}_{-0.024}$	3.051 ± 0.009
$\omega_0$	-1.0	$-1.07^{+0.09}_{-0.08}$	$-1.31^{+0.02}_{-0.01}$	$-0.85 \pm 0.03$
$\omega_a$	0.0	$-0.024^{+0.087}_{-0.085}$	$-0.036^{+0.086}_{-0.084}$	$-0.054^{+0.063}_{-0.065}$
$\Omega_m$	0.3138	$0.292^{+0.024}_{-0.018}$	$0.336^{+0.037}_{-0.034}$	$0.358 \pm 0.008$
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- Use new simulation for BINGO;
- Combine BINGO simulation with other data (CMB and BAO) for the three interaction models;
- Repeat the comparison with ΛCDM model to analyze the viability of the proposed models.







