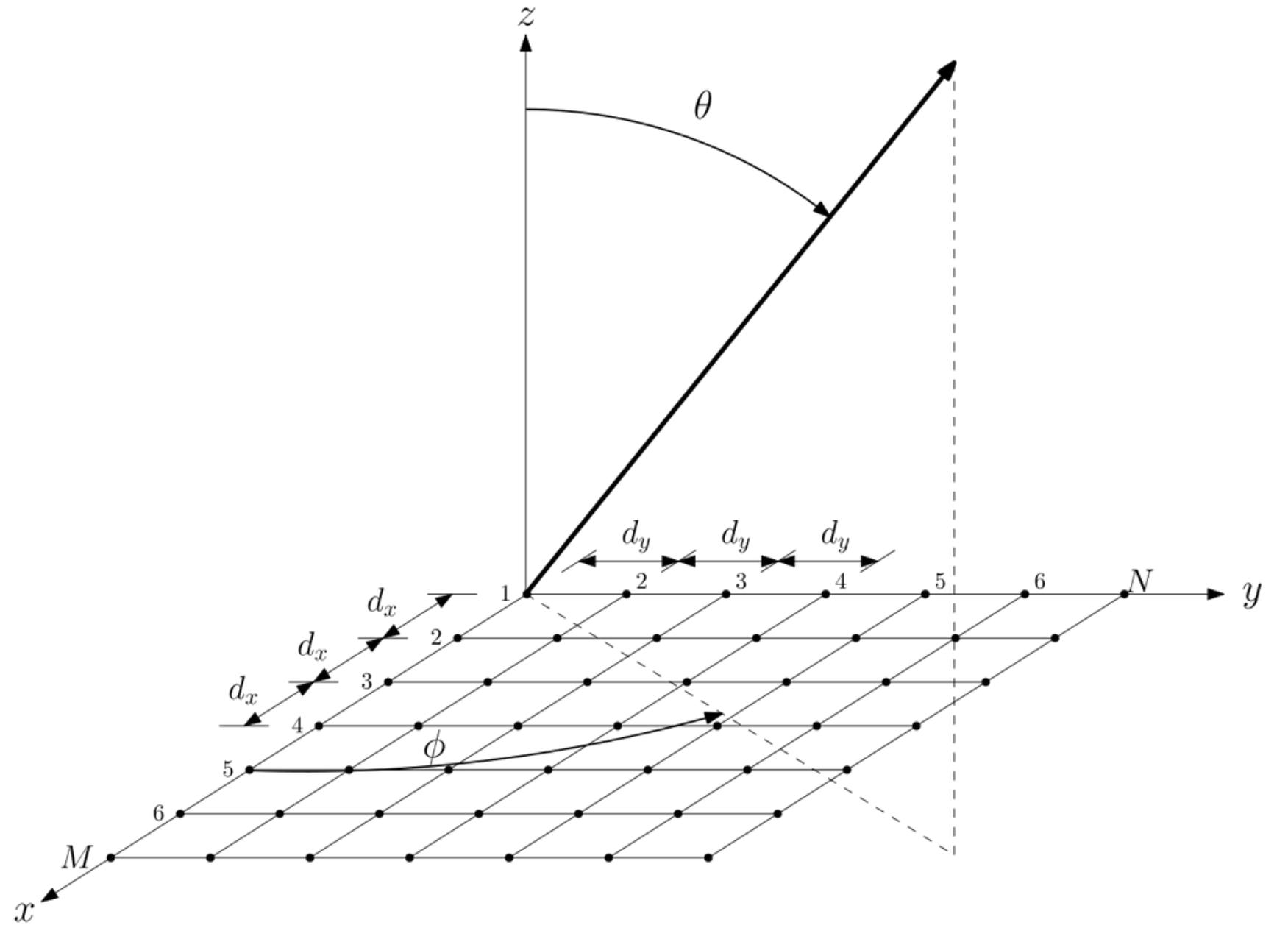


# Phased array antenna simulations

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# Phased array antennas



Array factors

$$\mathbf{E}_{array}(\theta, \phi) = AF(\theta, \phi) \mathbf{E}_{element}(\theta, \phi)$$

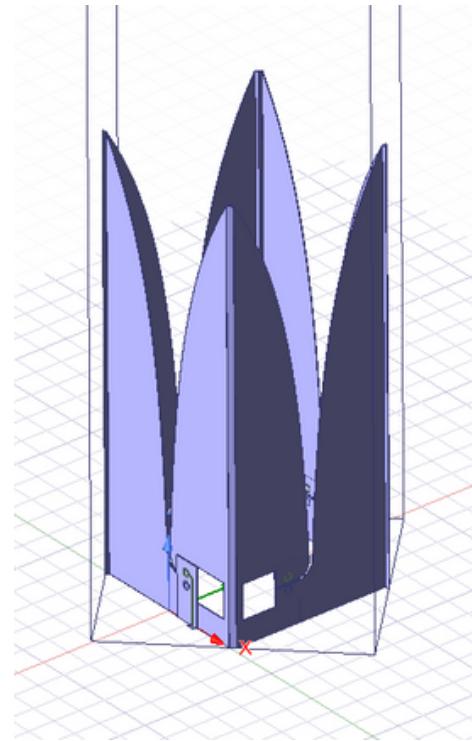
$$AF(\theta, \phi) = \sum_n^N e^{j k \mathbf{r}_n \cdot (\hat{r}(\theta, \phi) - \hat{r}_S)}$$

Phase shift

$$\psi_n = -k \mathbf{r}_n \cdot \hat{r}_S$$

# The technology

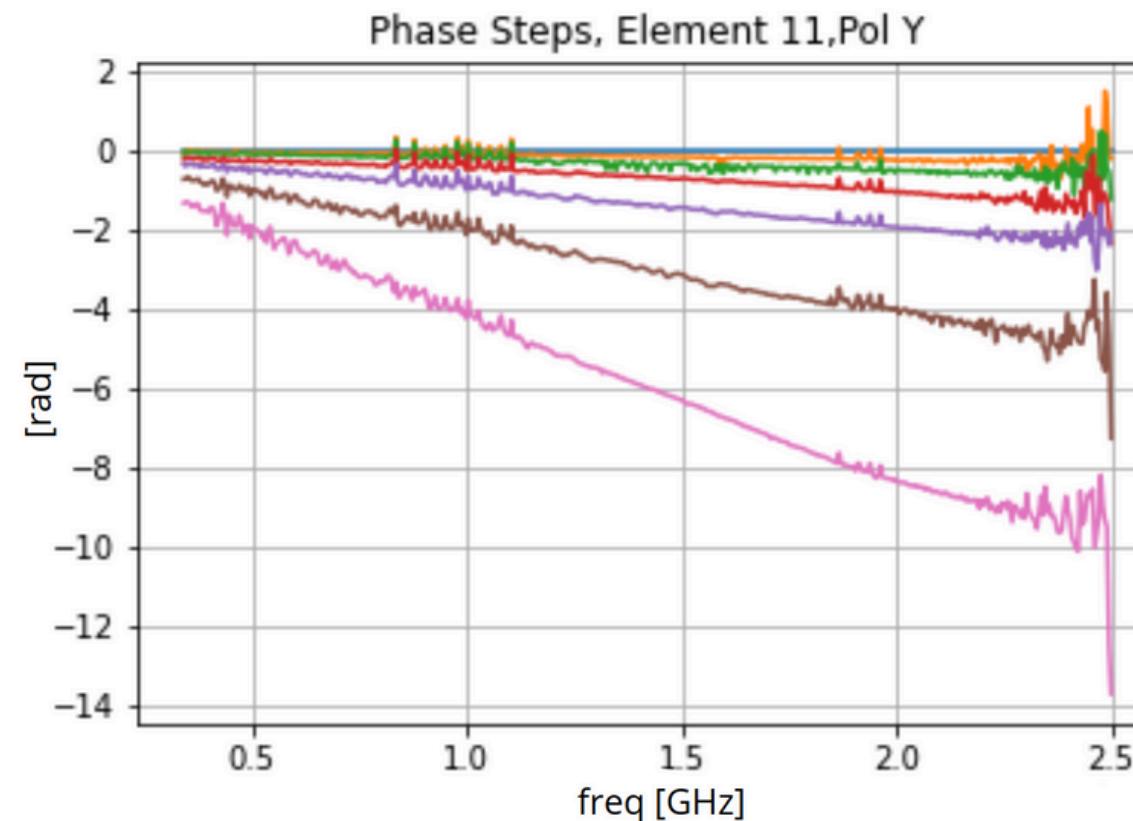
- **Vivaldi antenna:** This tapering design helps provide a wide bandwidth and directional radiation pattern.
- **Beamforming:** performed in two steps: the first is analog on the tile and the second is digital. The digital step form multiple directed beams at the same time.
- **True-time delay:** exhibit constant phase slope over frequency.  $\psi_n = -k\mathbf{r}_n \cdot \hat{\mathbf{r}}_S$



Vivaldi antenna



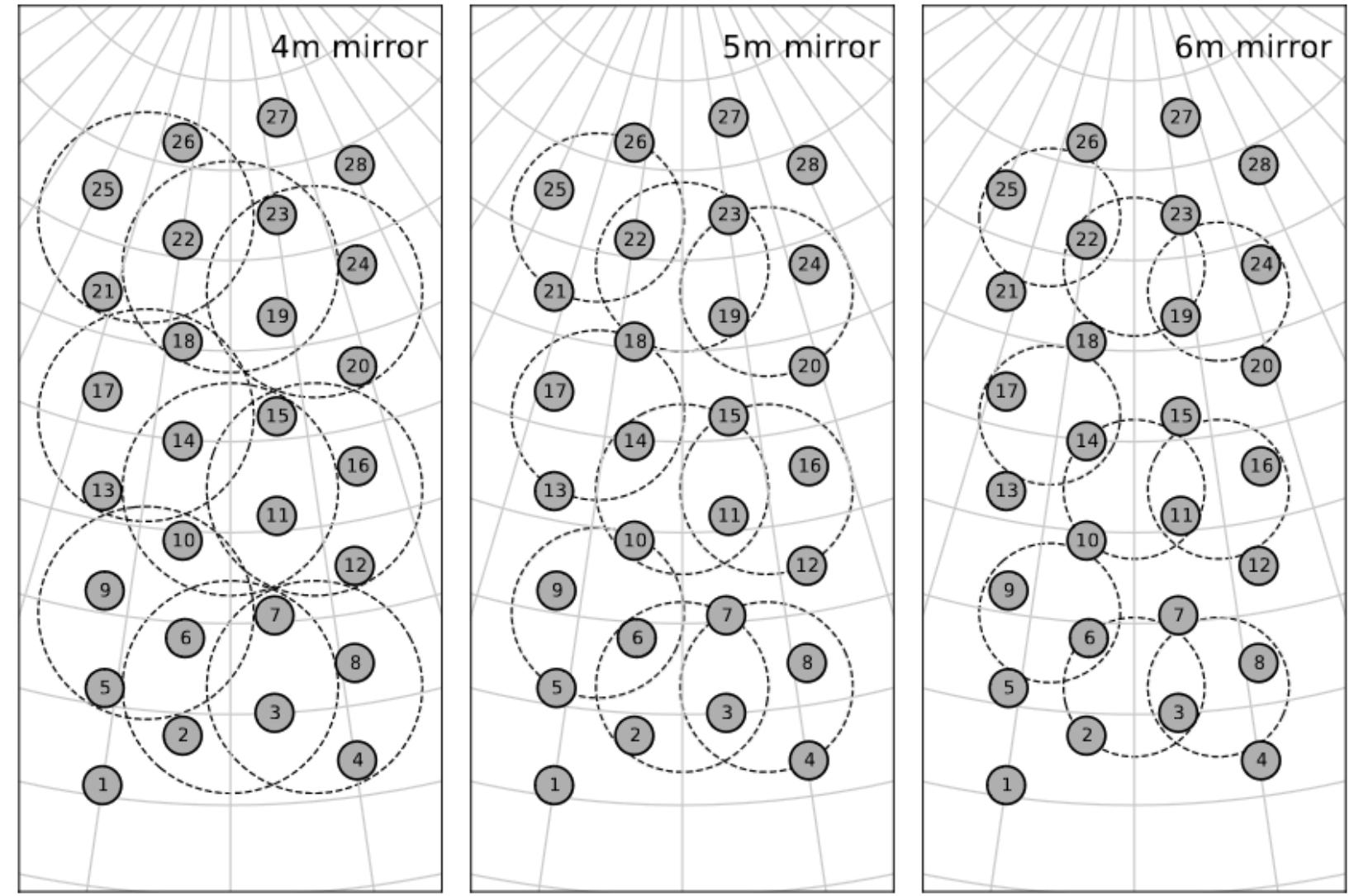
Tile



# Science case: FRB

- BINGO IX paper (Santos et al) proposes **single-dished outriggers** for the BINGO Interferometry System (BIS).
- However, one **phased-array (PA) outrigger** can cover the entire instantaneous area of the BINGO main telescope.
- We explore PA stations with similar effective area as in Santos et al.
- The minimal flux density of the sources will also depend the system temperature (*to be measured*) by

$$S_{\min}^{\text{intf}} = \frac{1}{\sqrt{2n_p \Delta\nu \tau}} \left[ \sum_{\substack{i, j, \\ i < j}} \frac{G_i P_{n,i}(\mathbf{n})}{K_i T_{\text{sys},i}} \frac{G_j P_{n,j}(\mathbf{n})}{K_j T_{\text{sys},j}} \right]^{-1/2}.$$



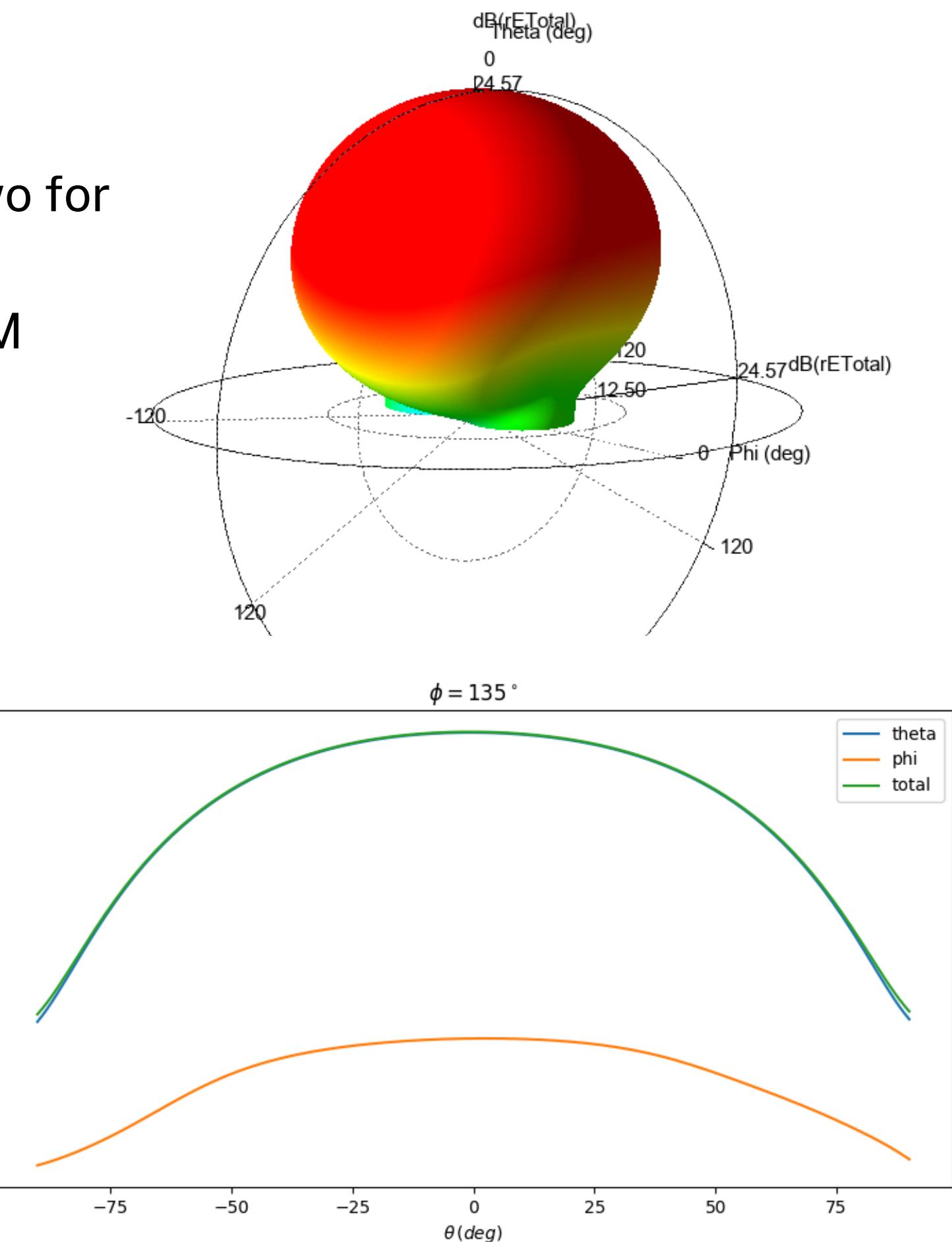
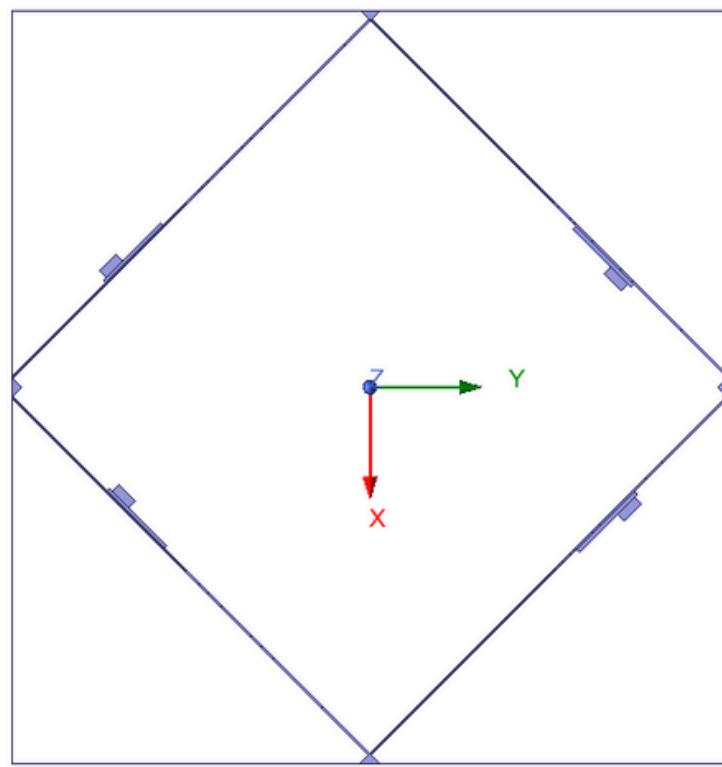
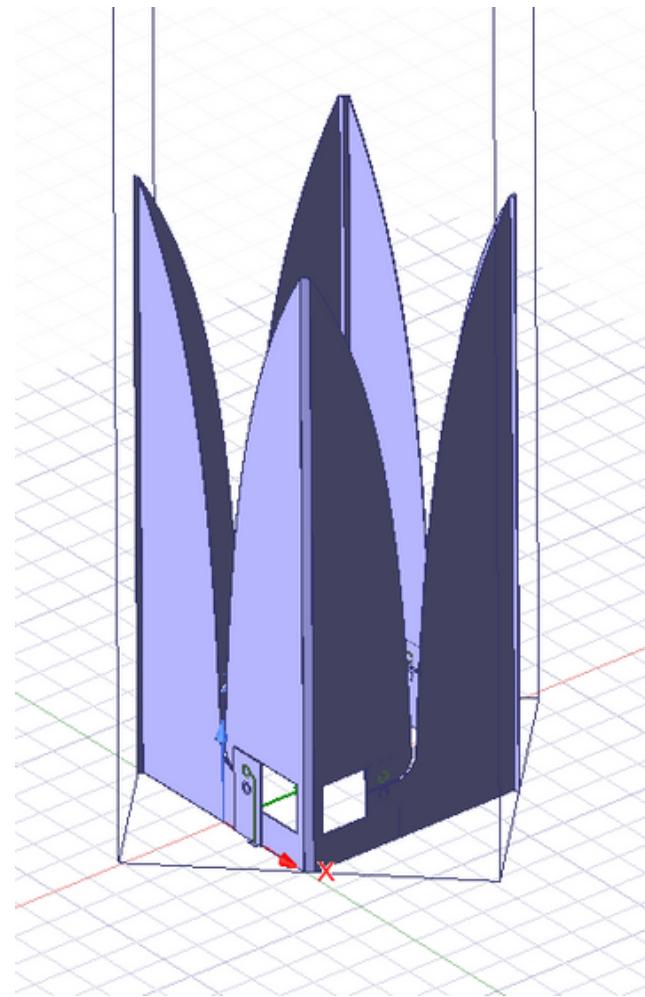
Santos et al

- Preliminary results points that BIS detection with one PA outrigger of 6m of diameter is 10 times higher than with nine single-dished outriggers of 6m.

# **Results**

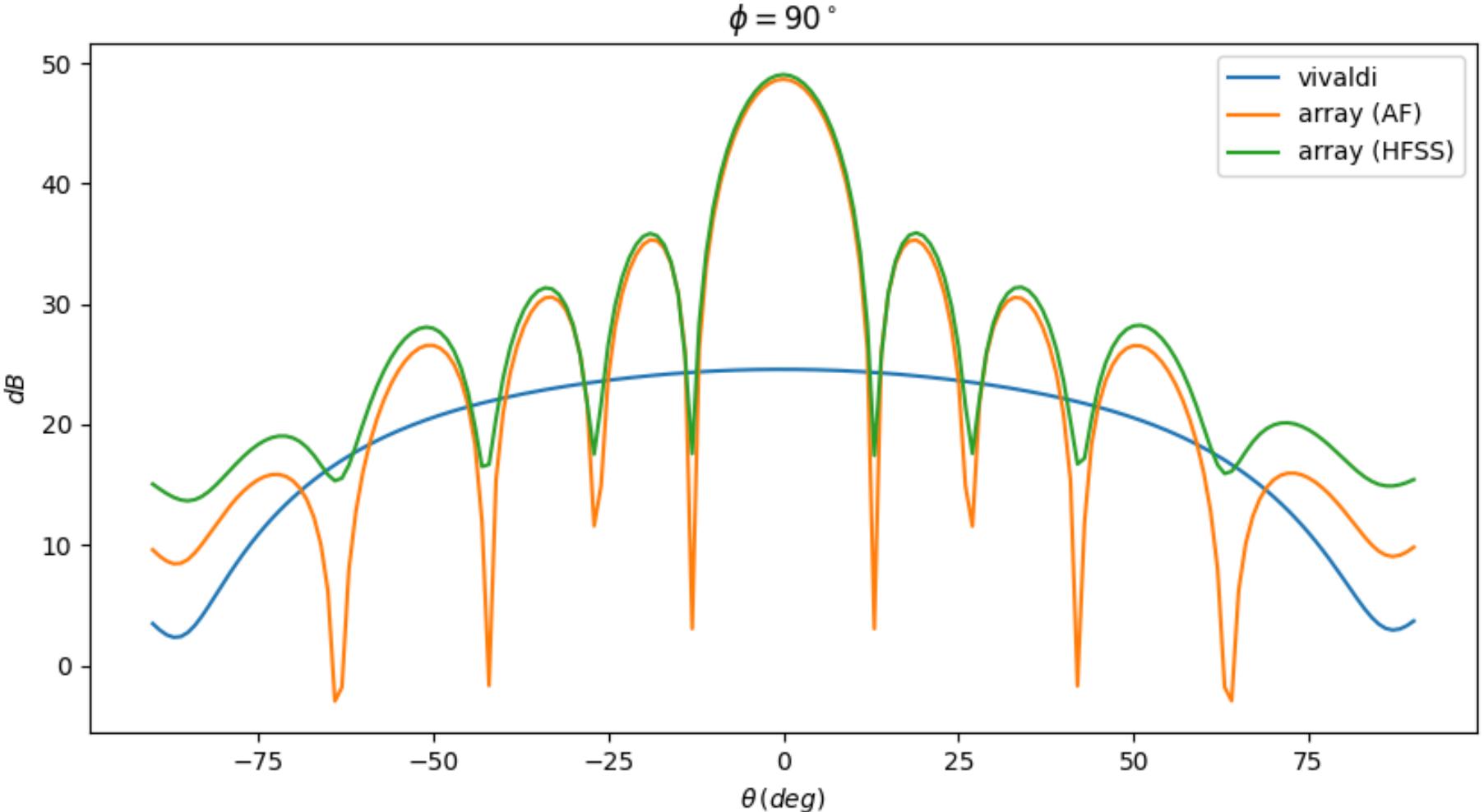
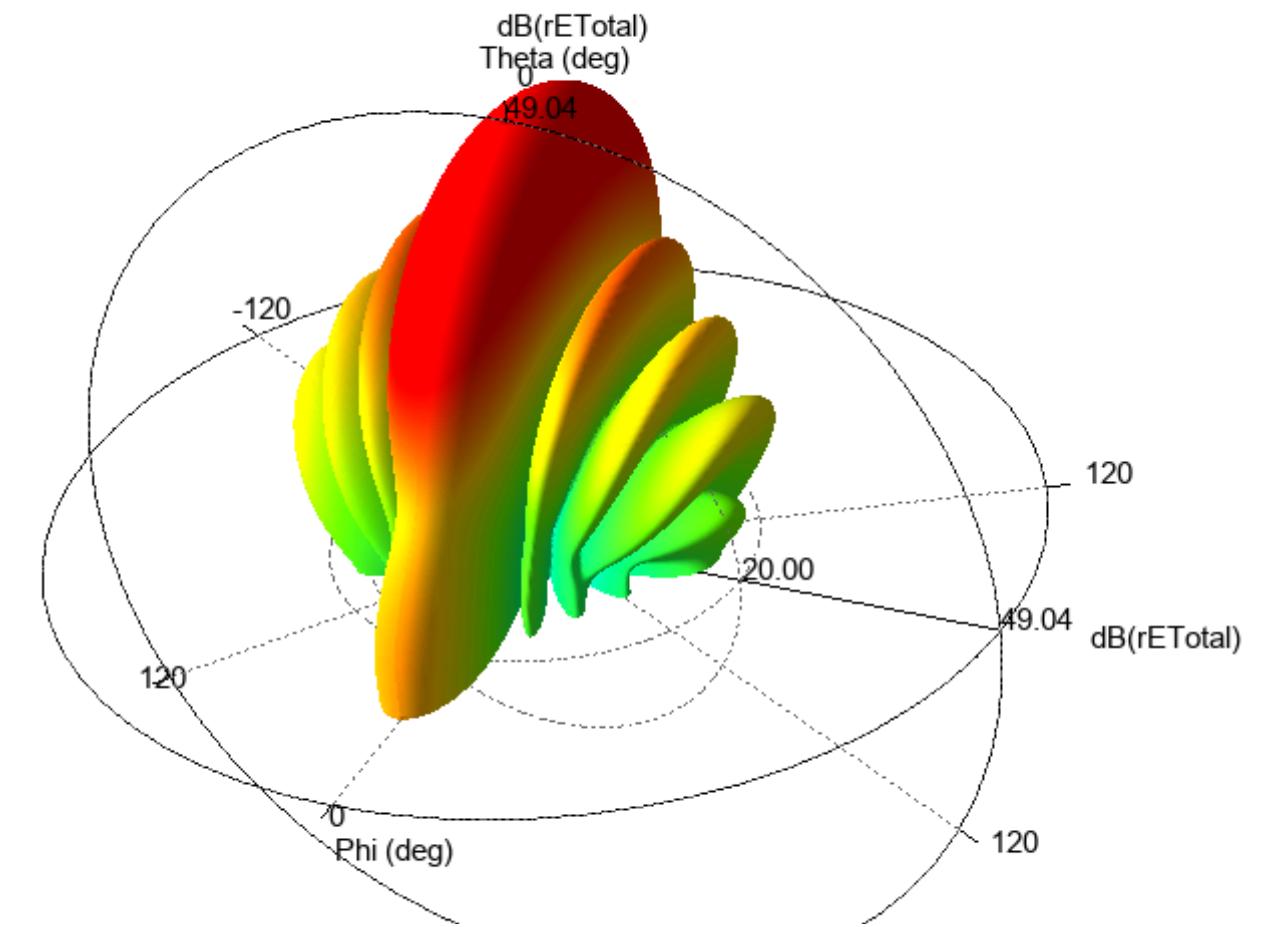
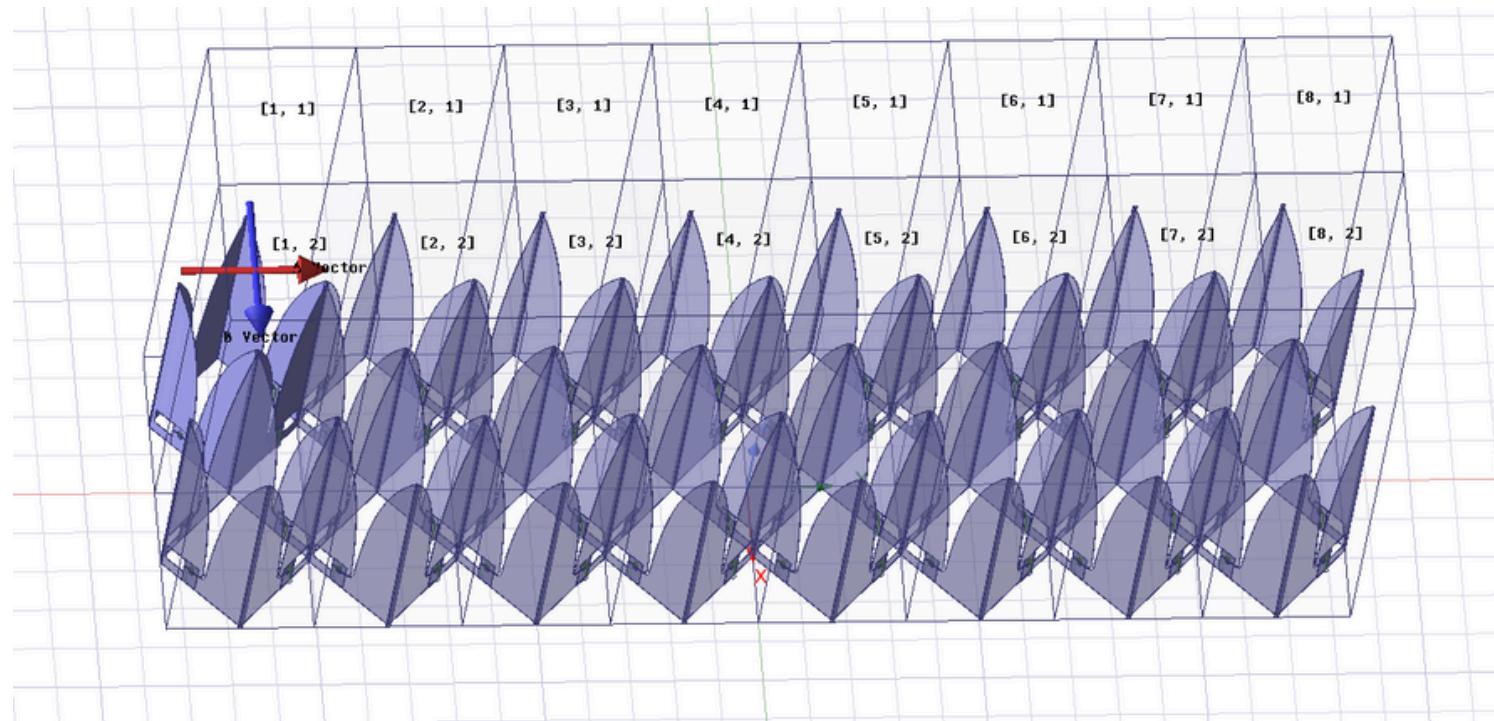
# Unit Cell simulation

- One unit cell is assumed with four antenna elements, two for each polarization.
- We excite two elements of same polarization and run EM simulations on HFSS software.



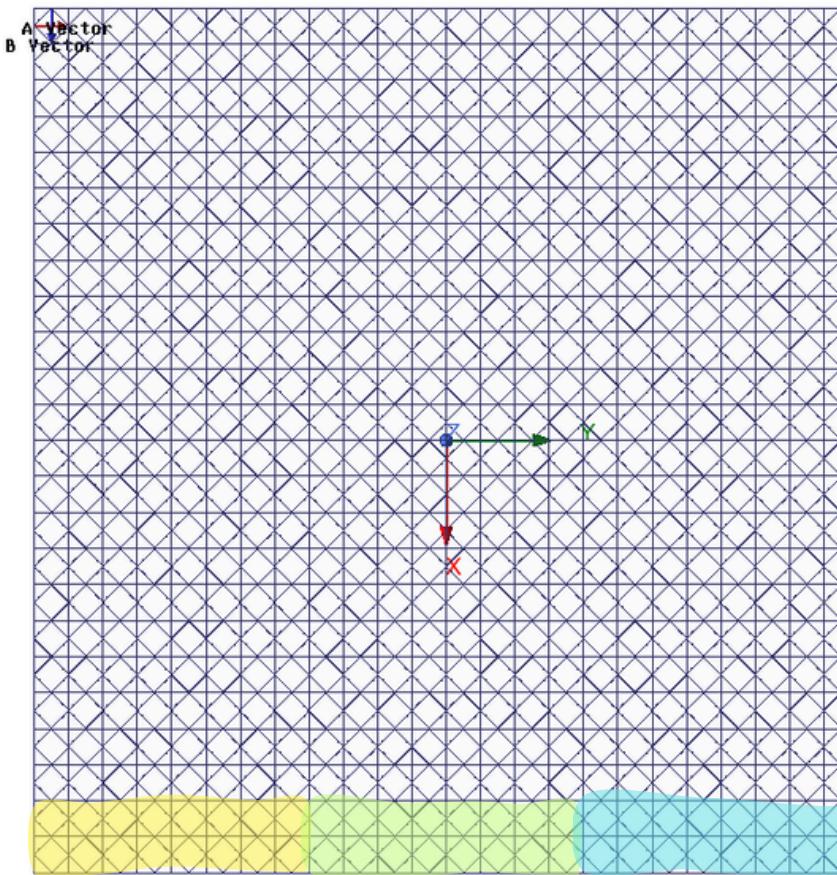
# Vivaldi Tile simulation

- We simulated a 2x8 array. It correspond to a 35cm x 134 cm tile that will be used on ABDUS project.

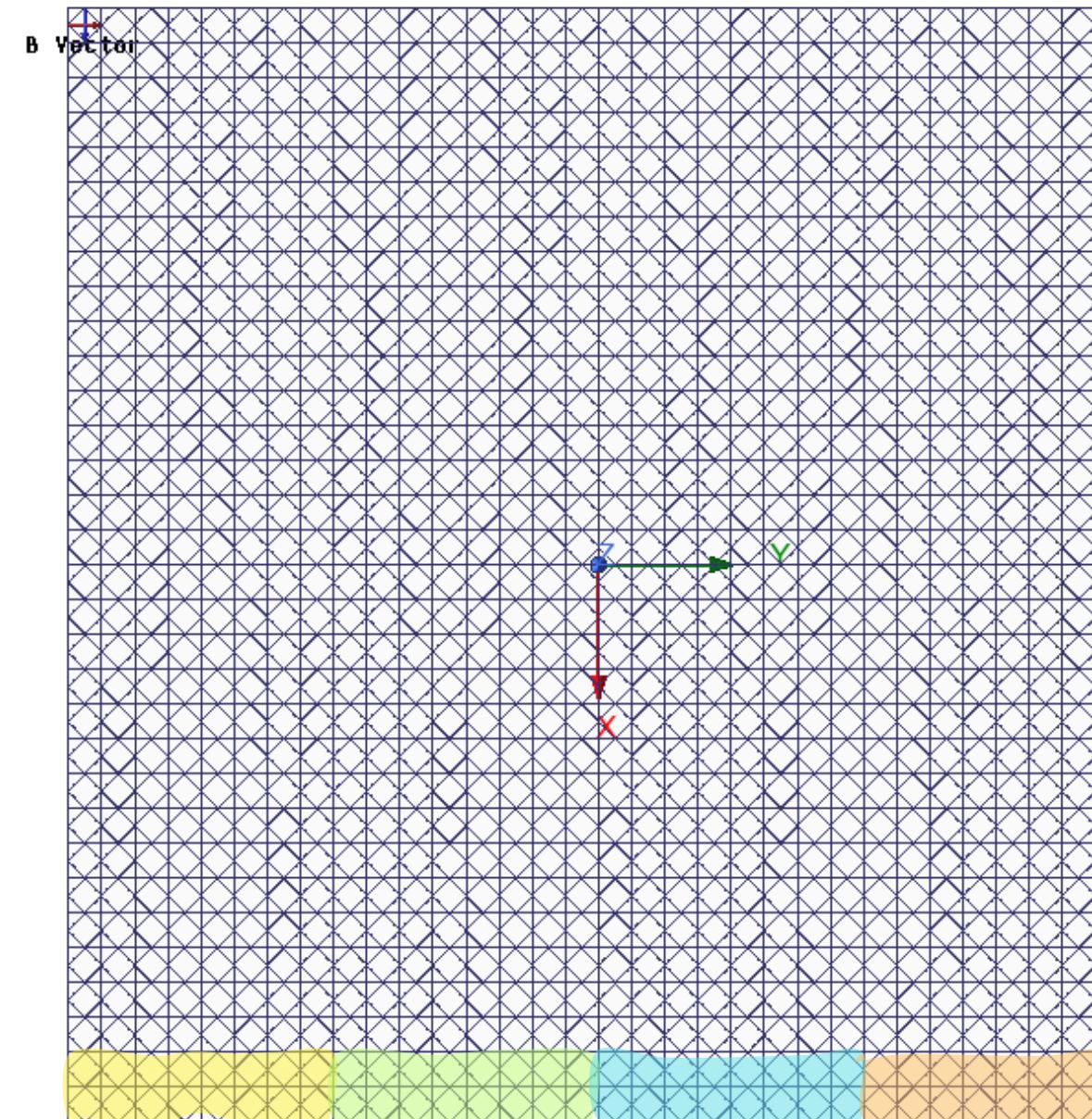


# **Simulating stations**

# Stations comparison

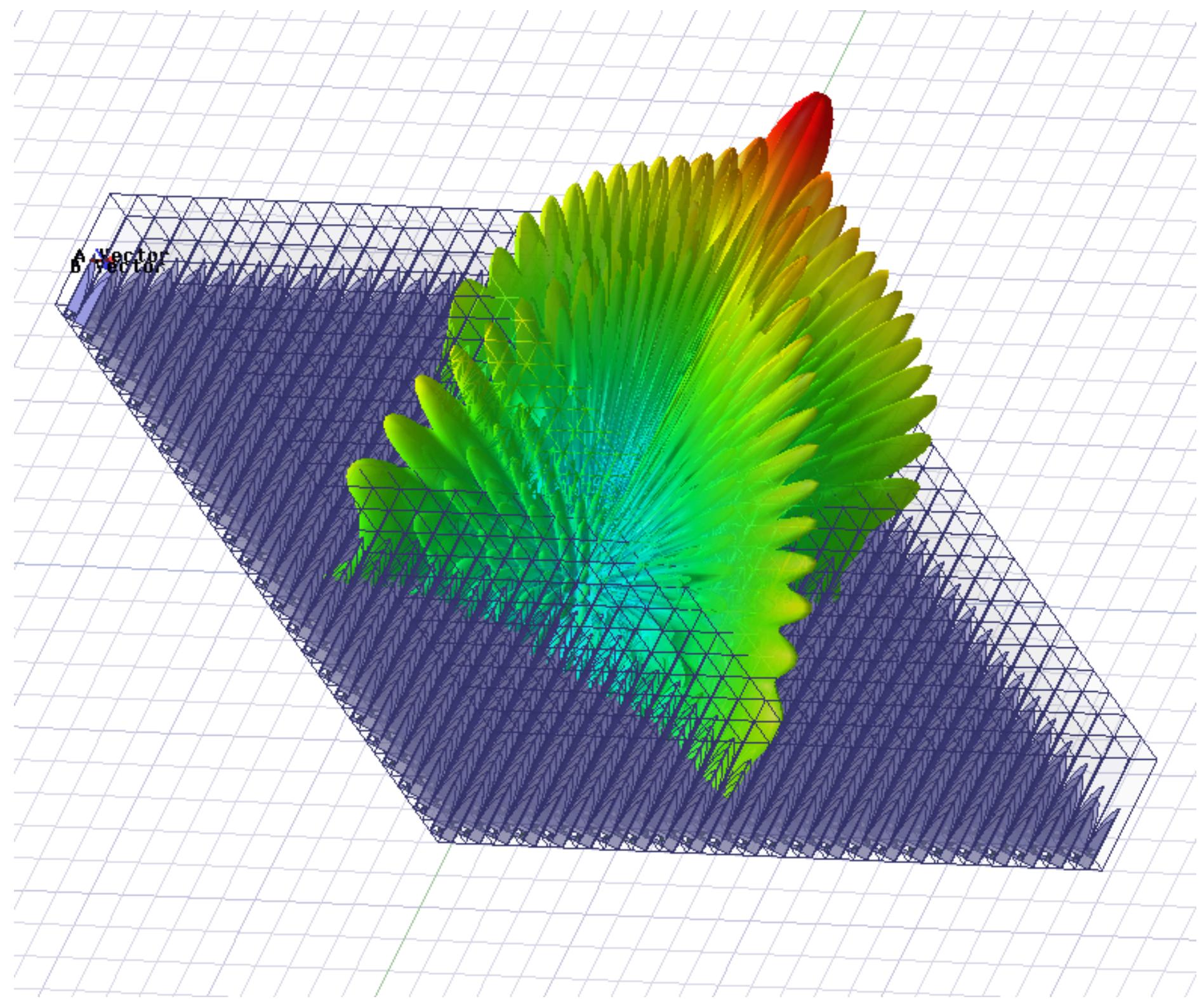


- 24x24 unit cells
- 36 tiles
- 422cm x 402cm
- Physical area: 17m<sup>2</sup>
- Effective area: 16.24m<sup>2</sup>



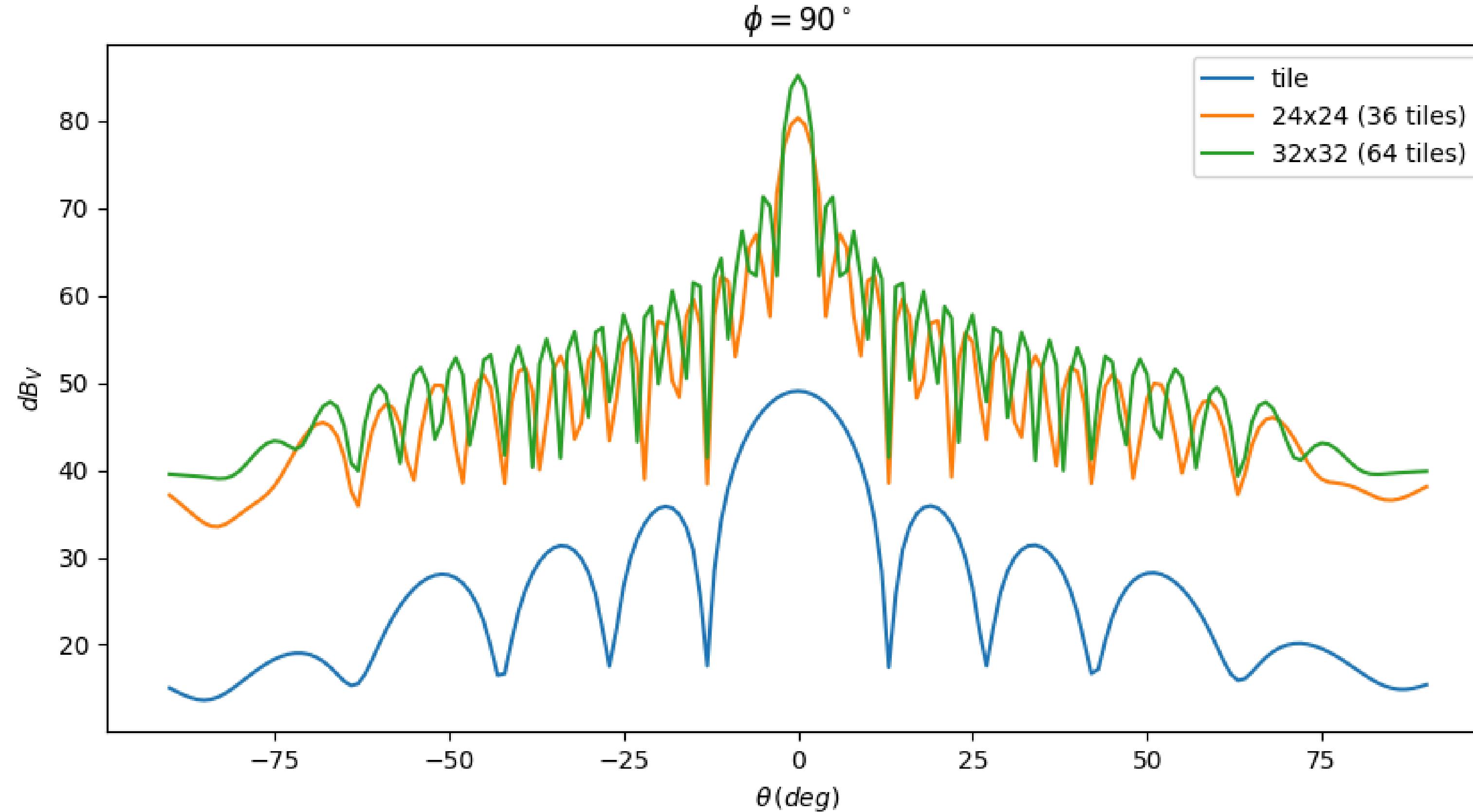
- 32x32 unit cells
- 64 tiles
- 563cm x 536cm
- Physical area: 30.21m<sup>2</sup>
- Effective area: 28.84m<sup>2</sup>

# Simulated station - 3d

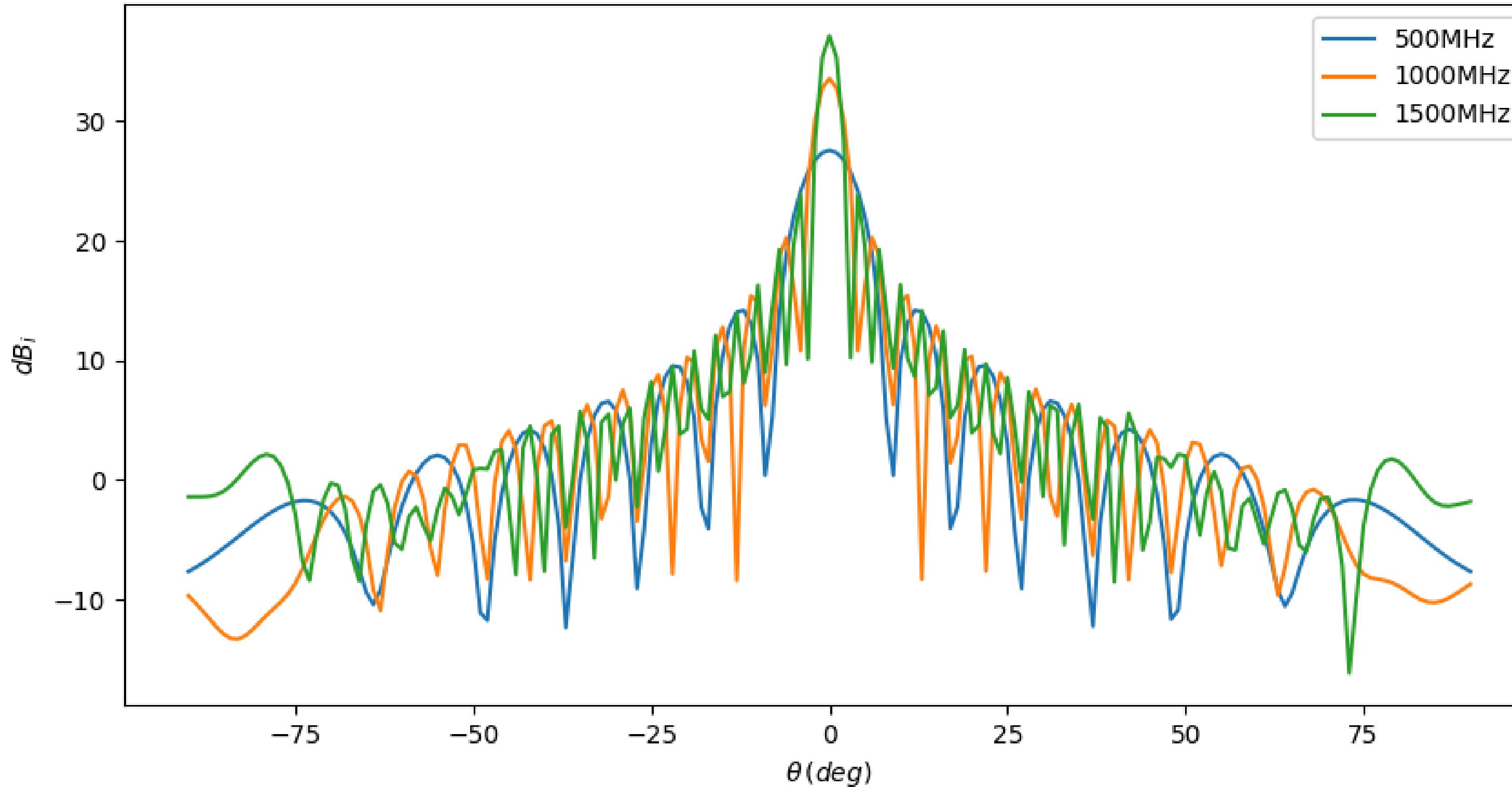


dB

# Simulated station - 36 vs 64 tiles



# Simulated station - Frequencies



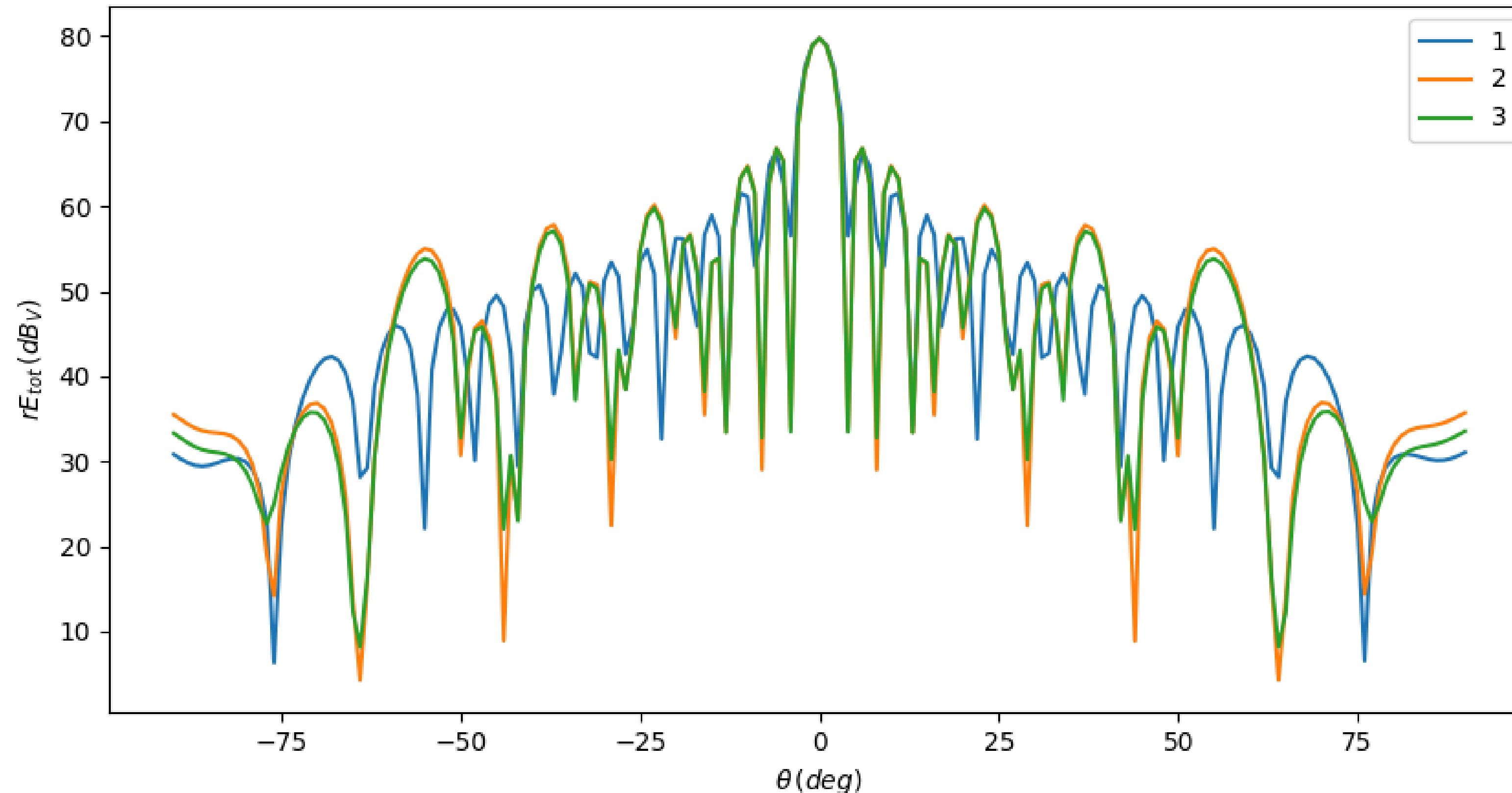
	01 tile	36 tiles	64 tiles
Size (m x m)	0.35 x 1.34	4.22 x 4.02	5.63 x 5.36
Gain (dBi)	5.02	33.6	36.1
Phis. area (m <sup>2</sup> )	0.47	17	30.2
Eff. area (m <sup>2</sup> )	0.45	16.2	28.8
FWHM (deg)	10 (42)	4	2
First side lobe (deg)	-13.3	-13.3	-13.9

**Table 1.** Technical data for the tile and for the 36-tile and 64-tile station. All the quantities except the size and physical area were computed from the HFSS simulations.

**Stations with randomly  
spaced tiles**

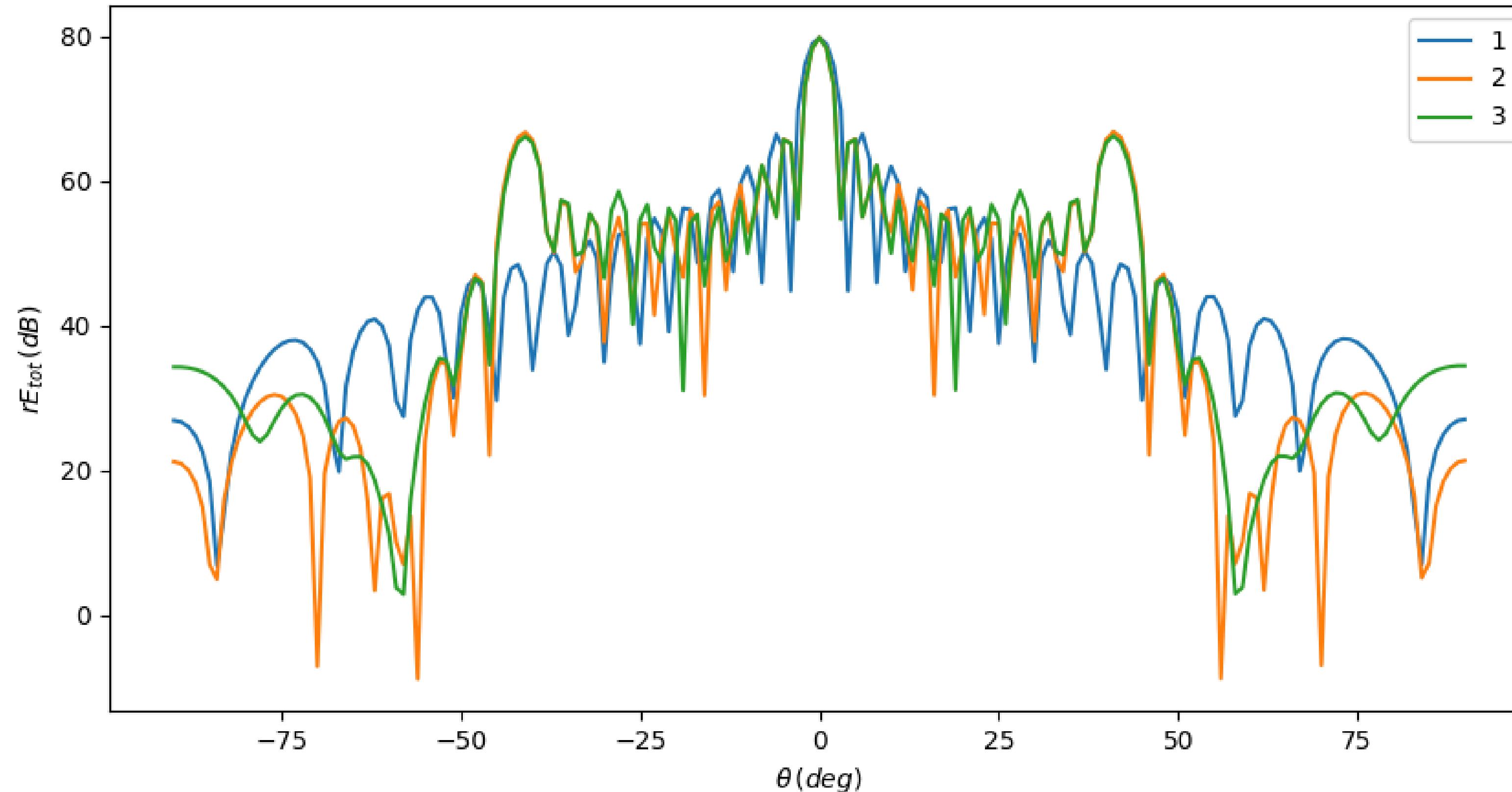
# 10cm spaced station

$\phi = 90^\circ$   $s = 10\text{cm}$

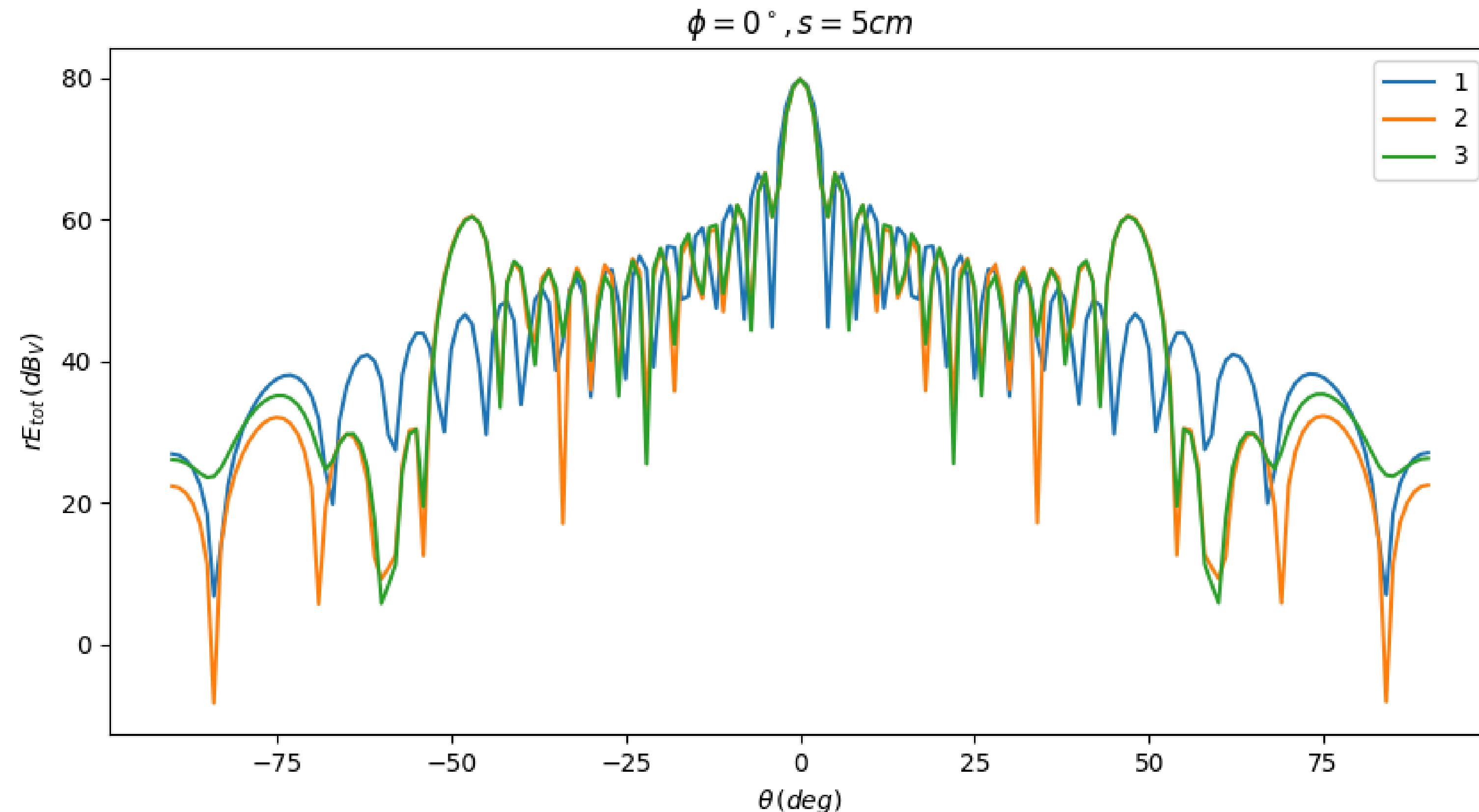


# 10cm spaced station

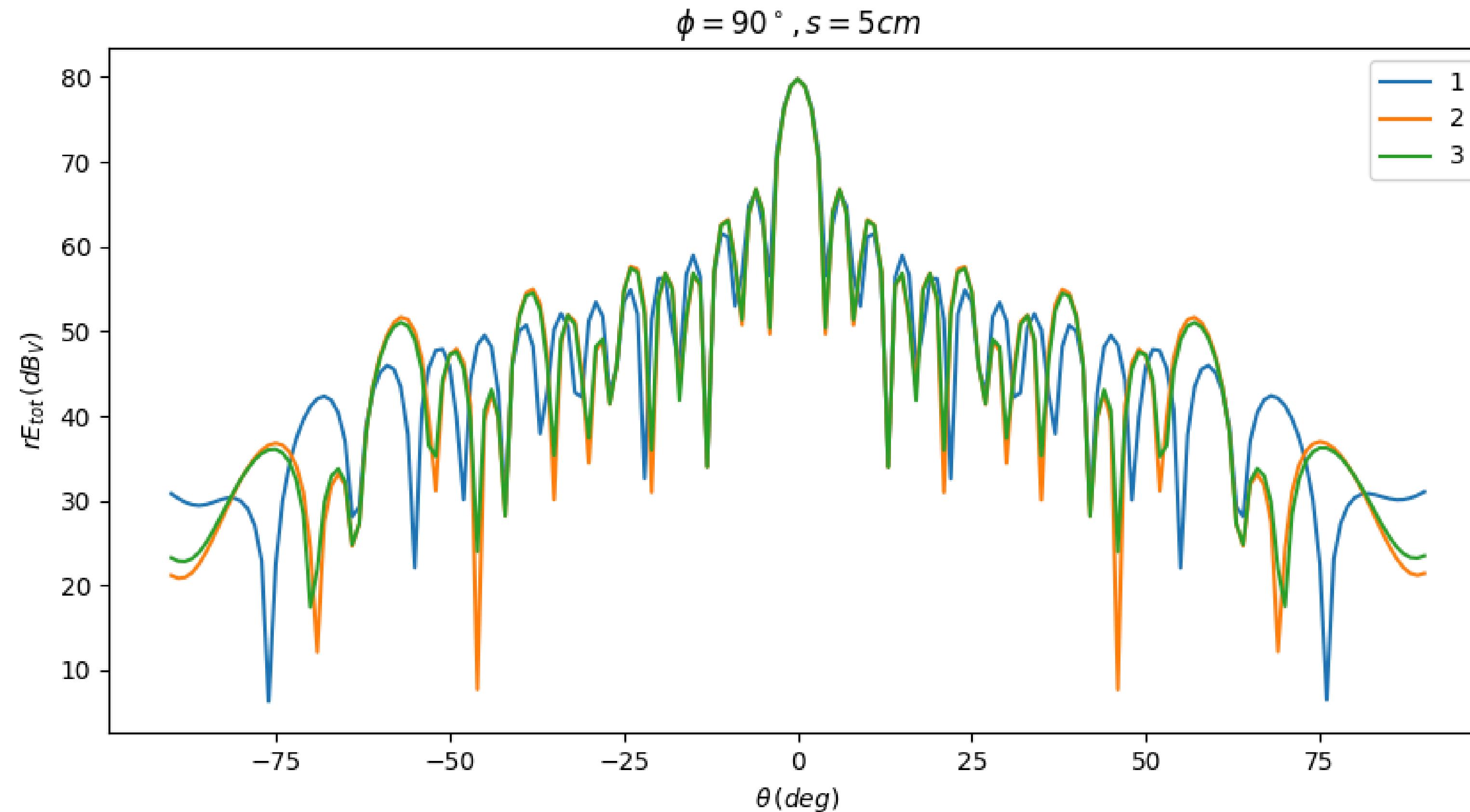
$\phi = 0^\circ, s = 10\text{cm}$



# 5cm spaced station

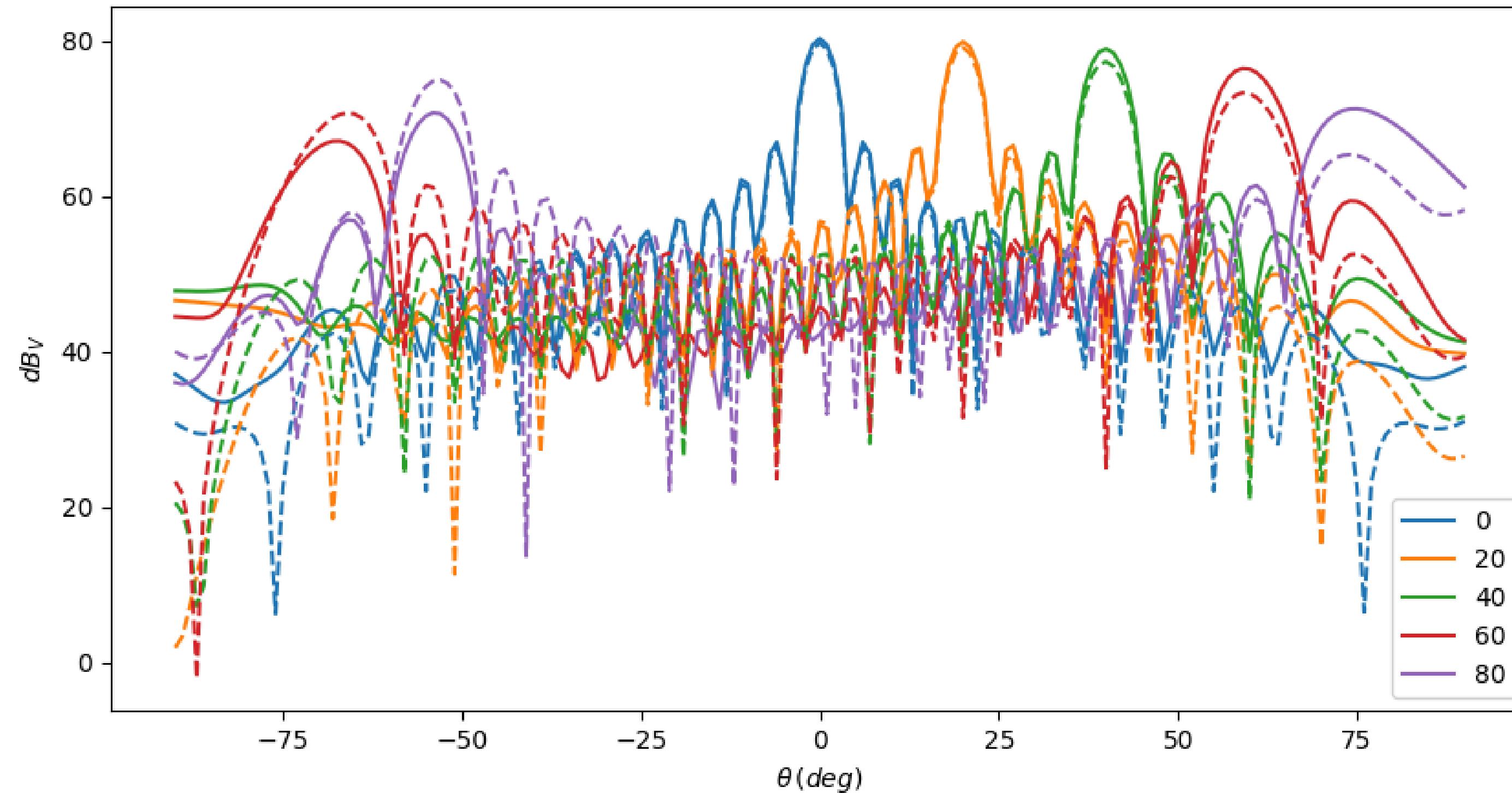


# 5cm spaced station

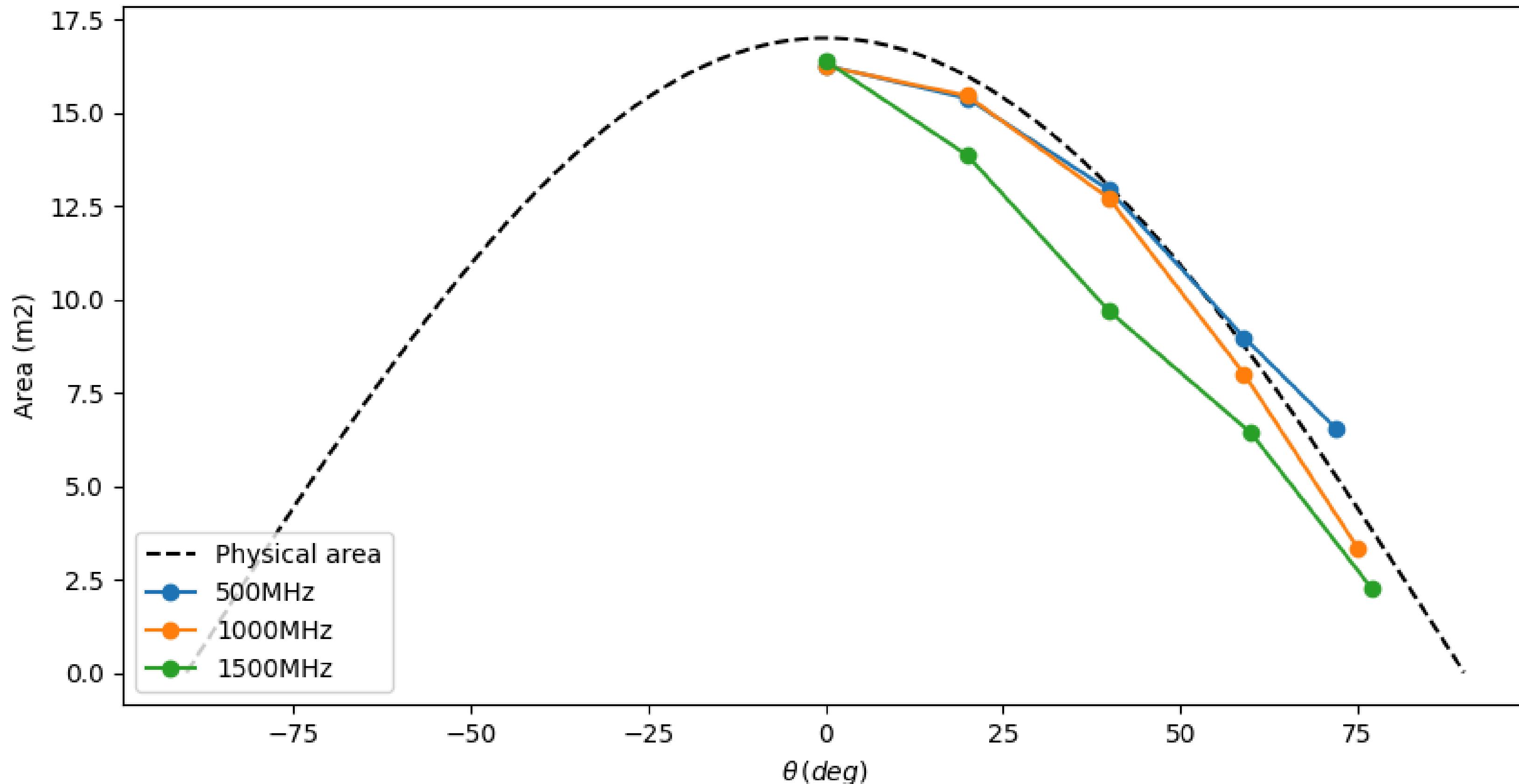


# **Scan angle and effective area**

# Beam steering



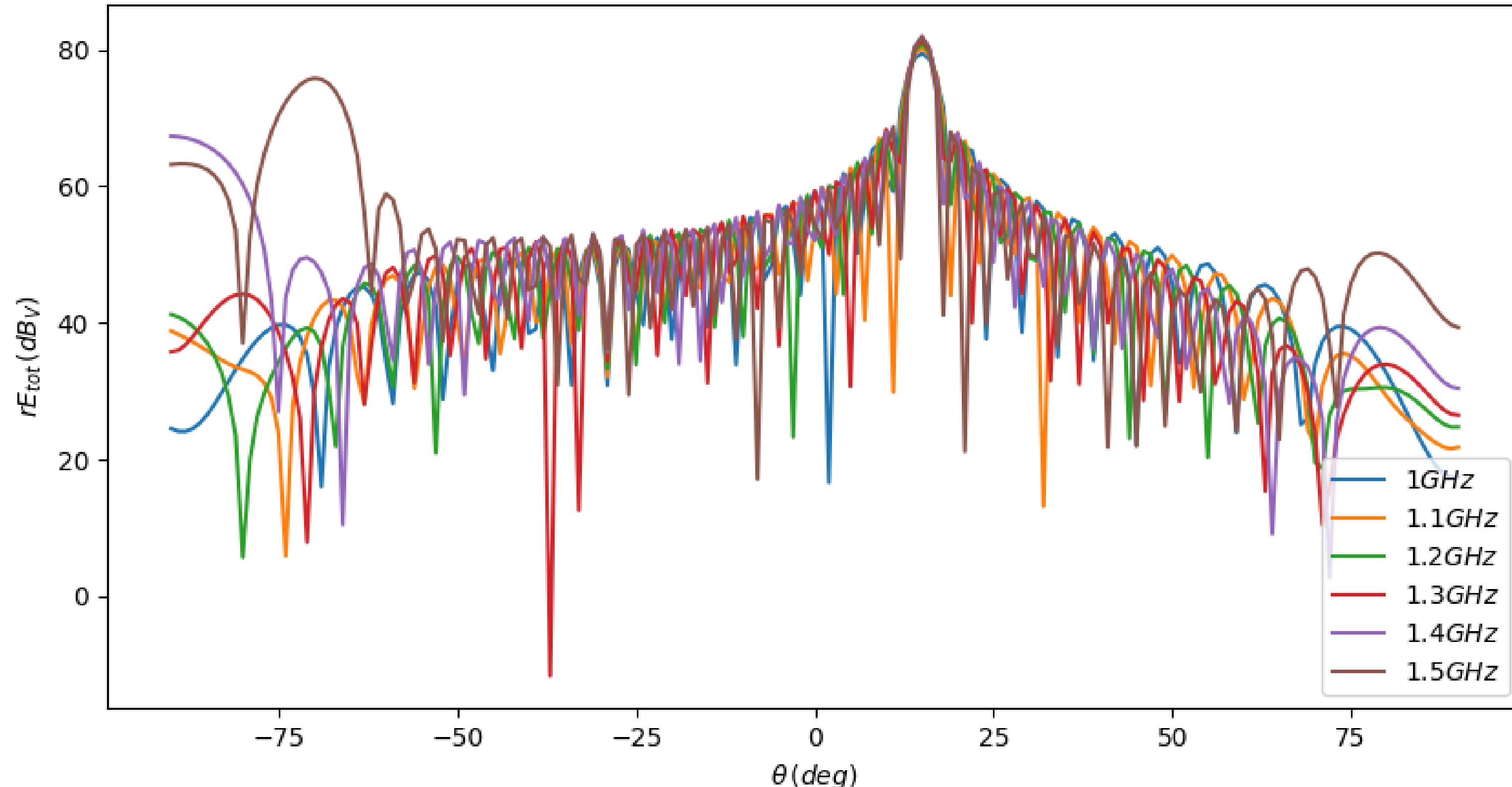
# Effective area



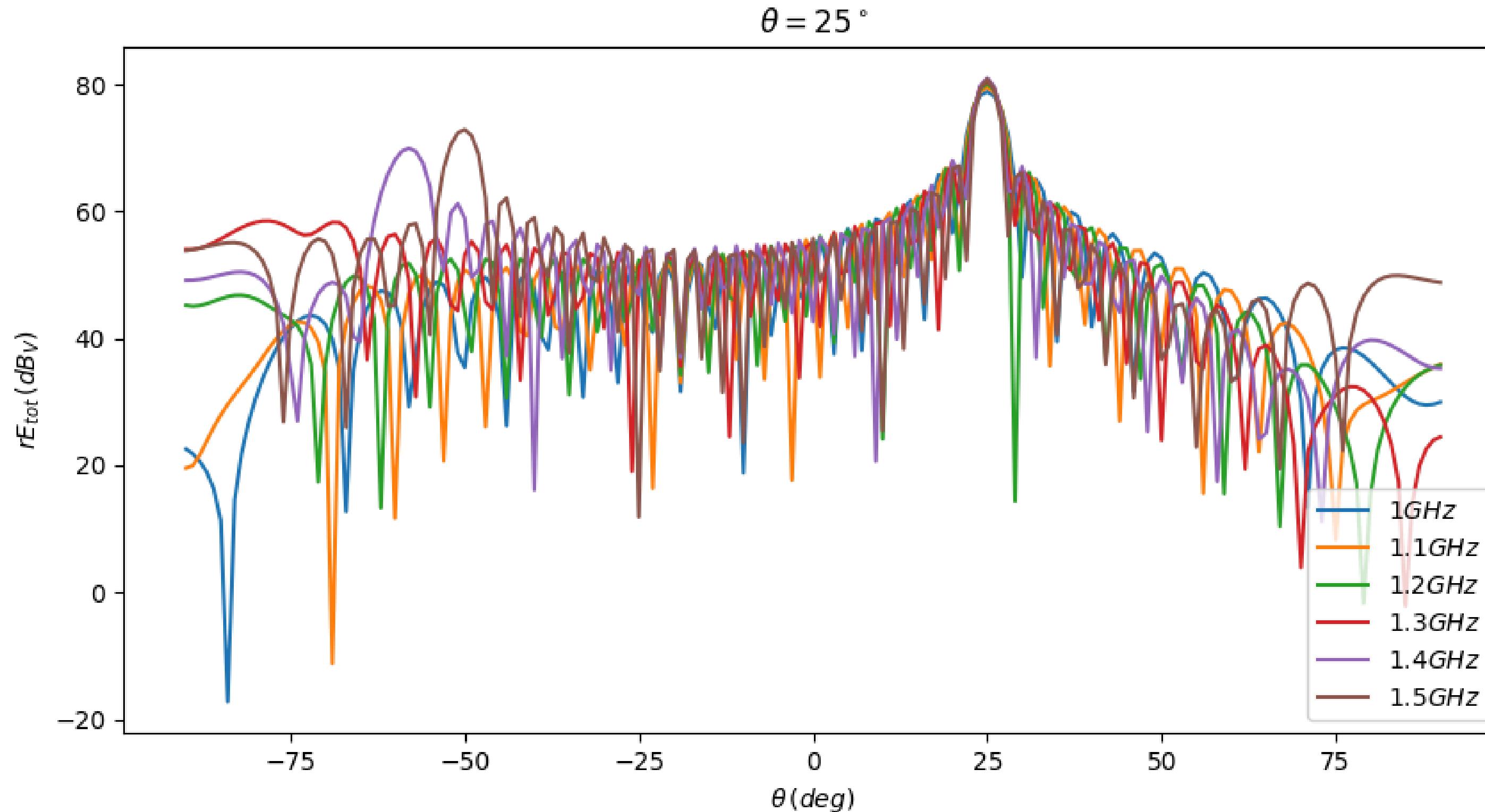
# **Scan angle and grating lobes**

# Sweeping frequency at $15^\circ$

$\theta = 15^\circ$

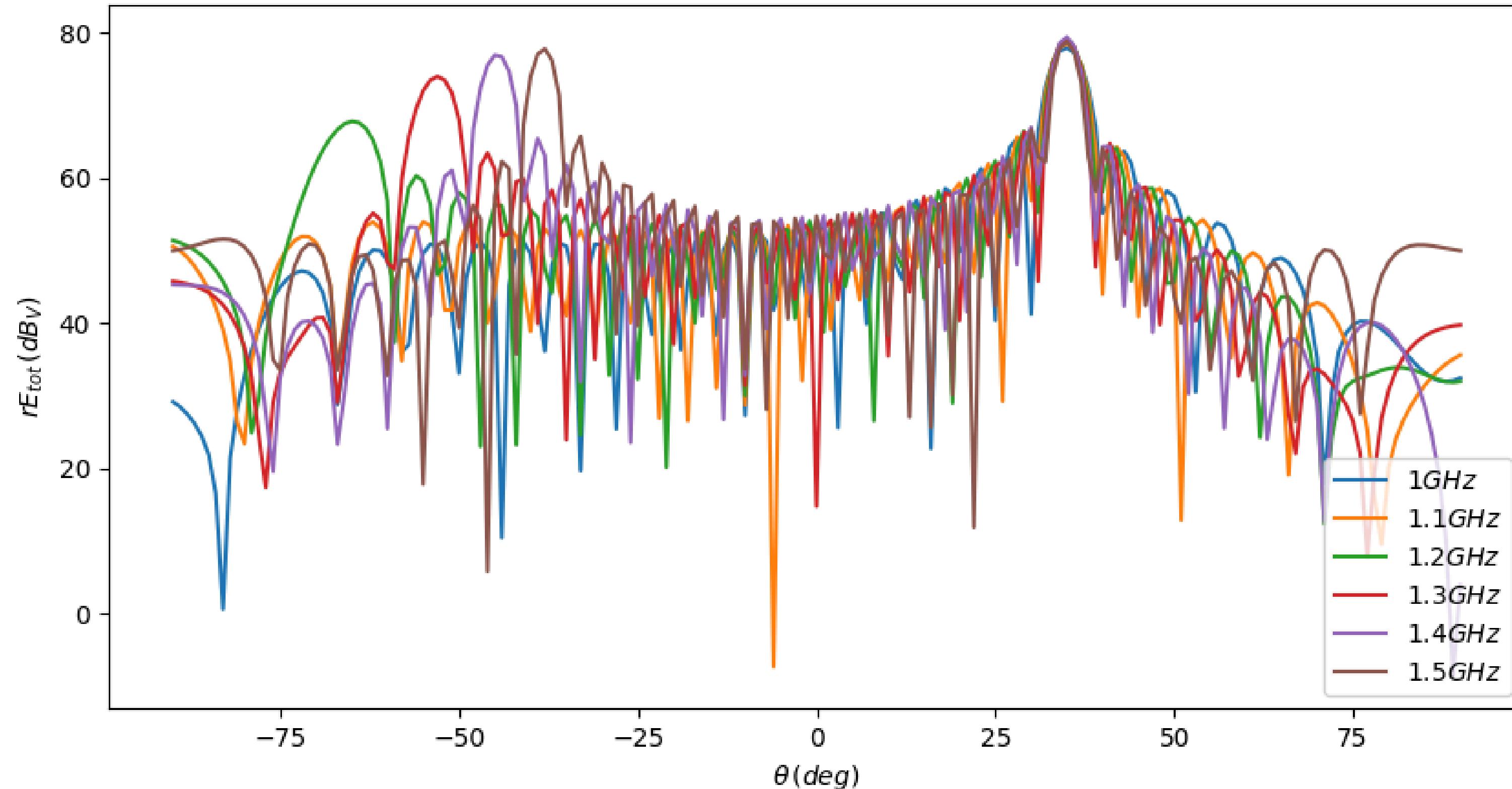


# Sweeping frequency at $25^\circ$



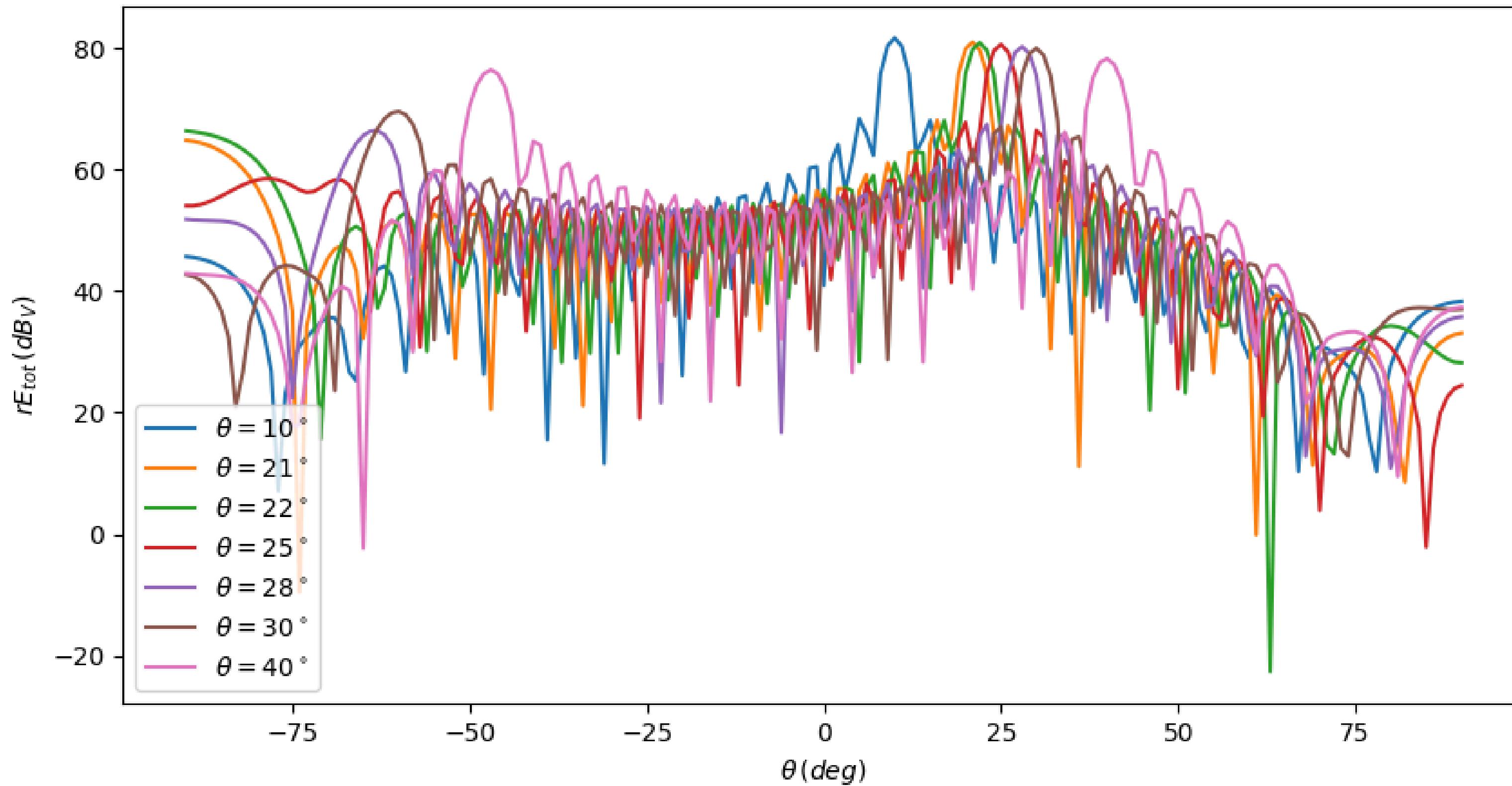
# Sweeping frequency at $35^\circ$

$\theta = 35^\circ$

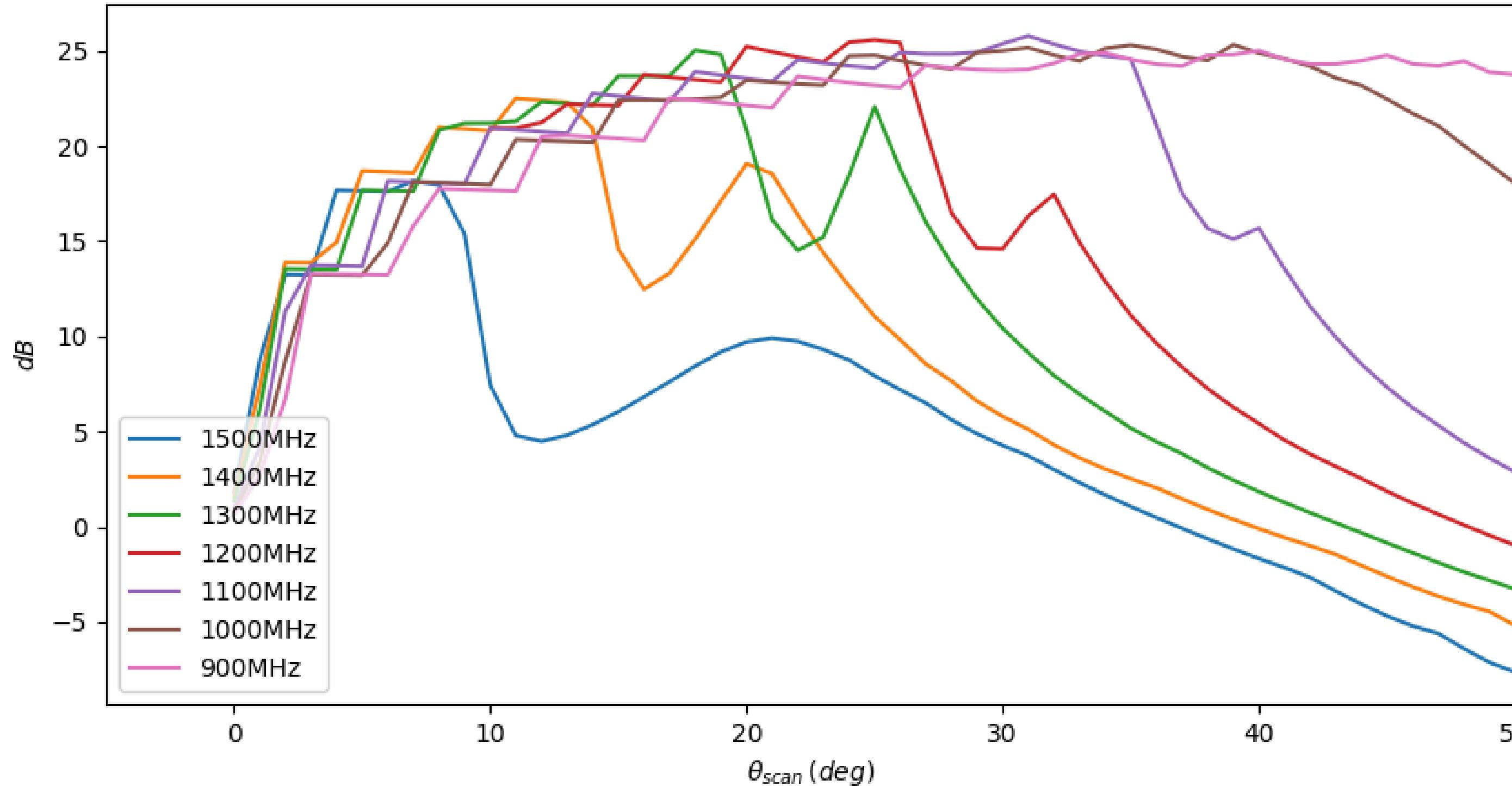


# Sweeping theta at 1.3GHz

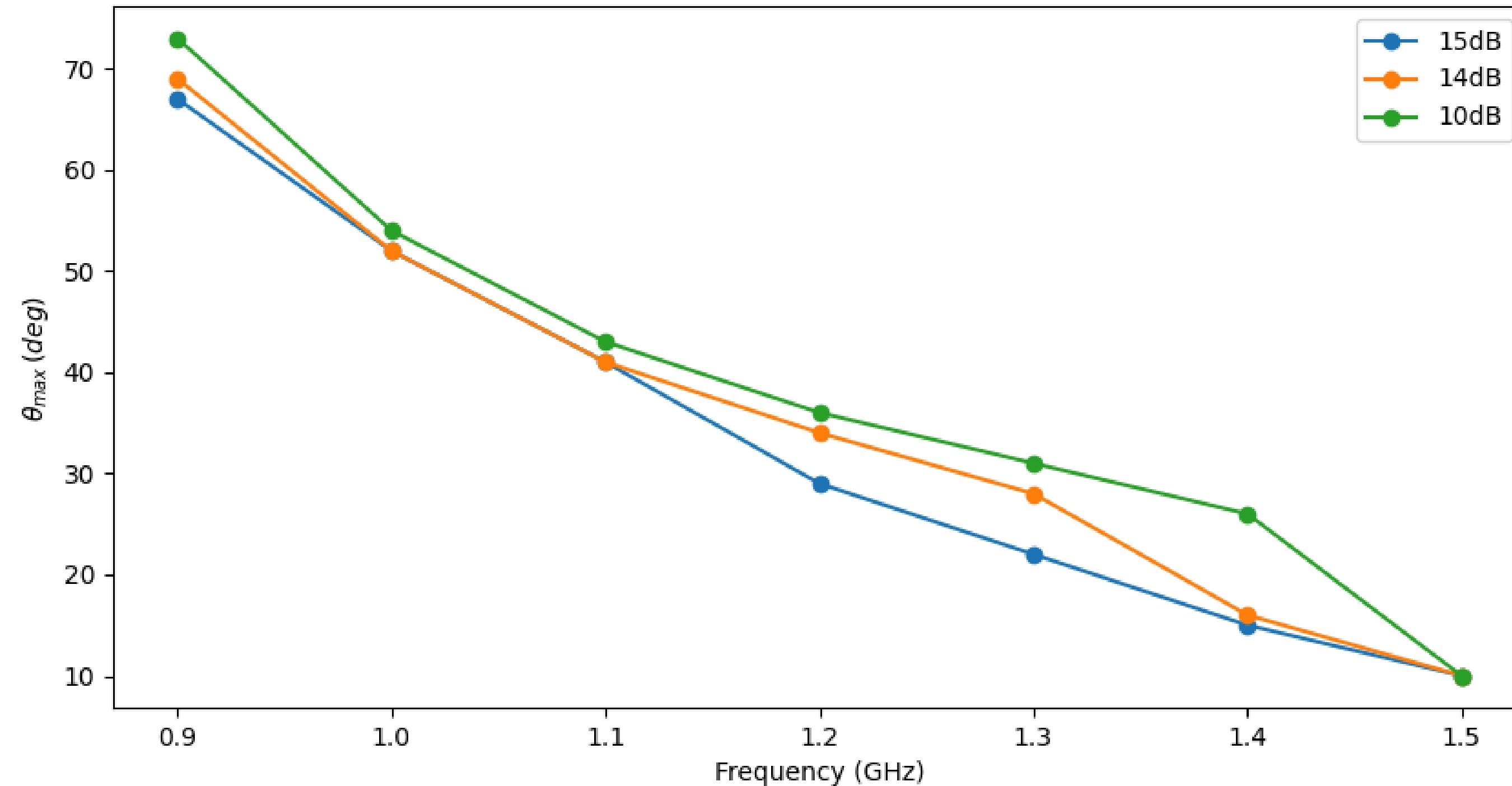
1.3GHz



# Difference between main lobe and maximum value at negative theta



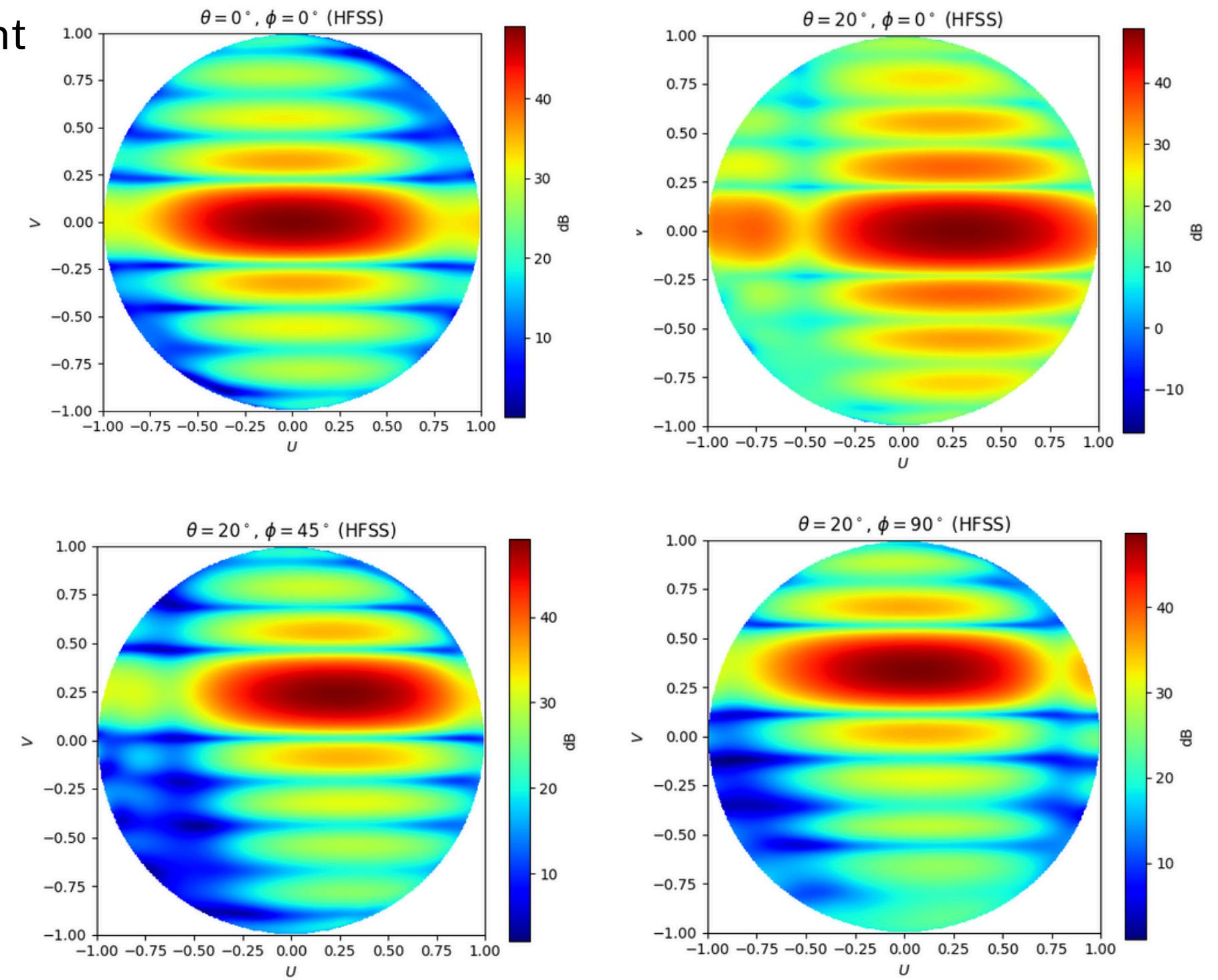
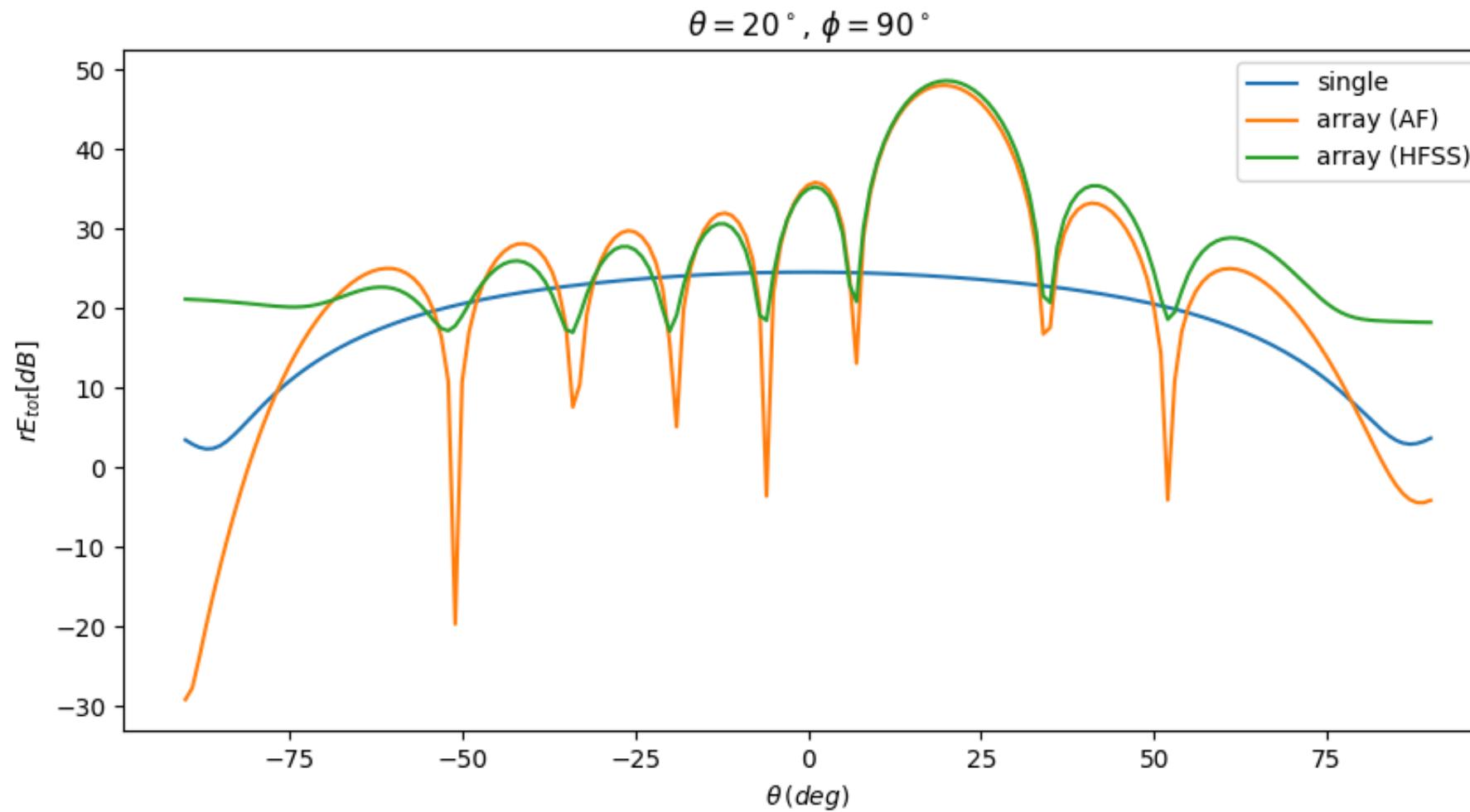
# Maximum scan angle



**End**

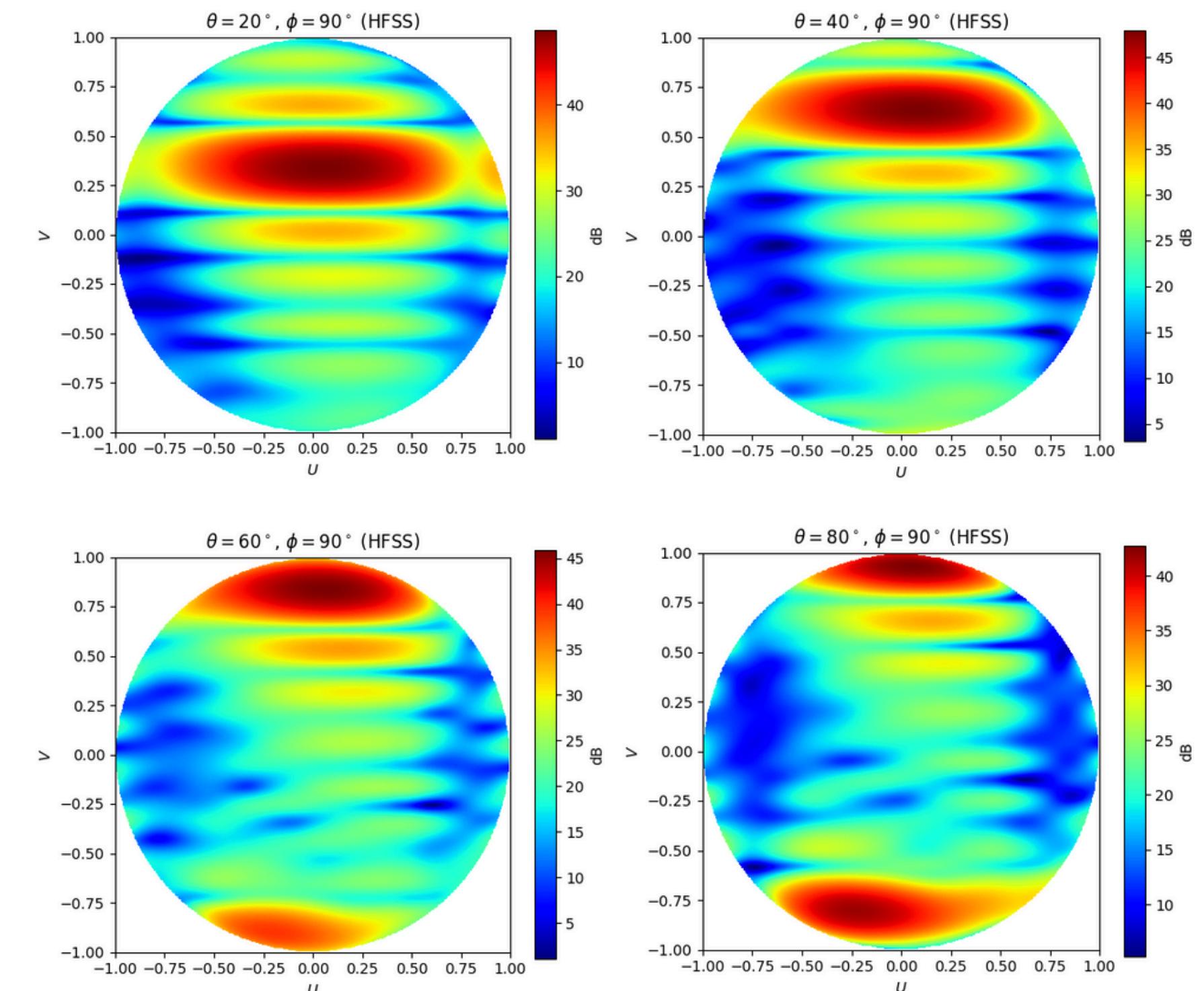
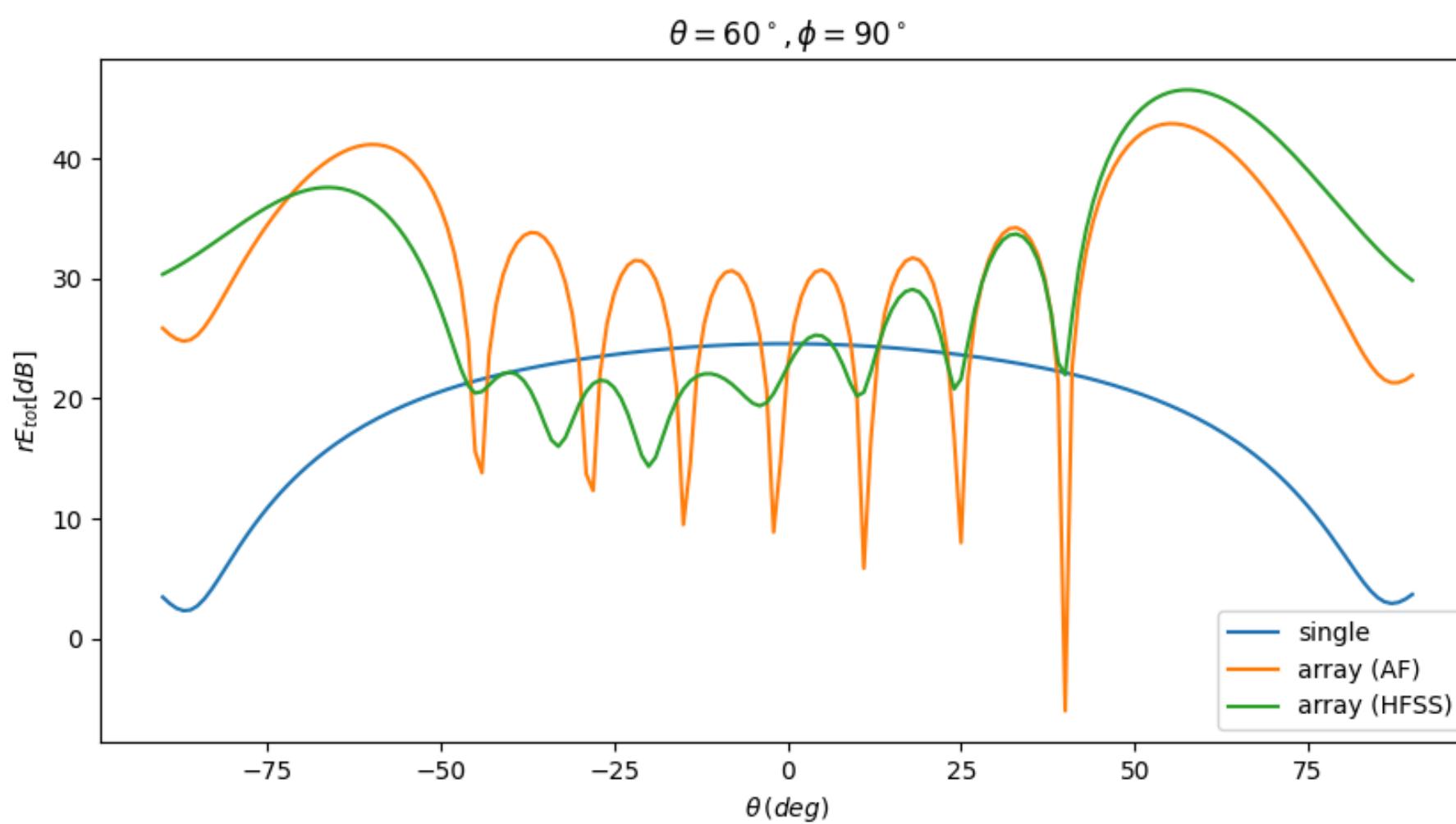
# Vivaldi Tile simulation: scanning

- We then present simulations steering in some different directions.

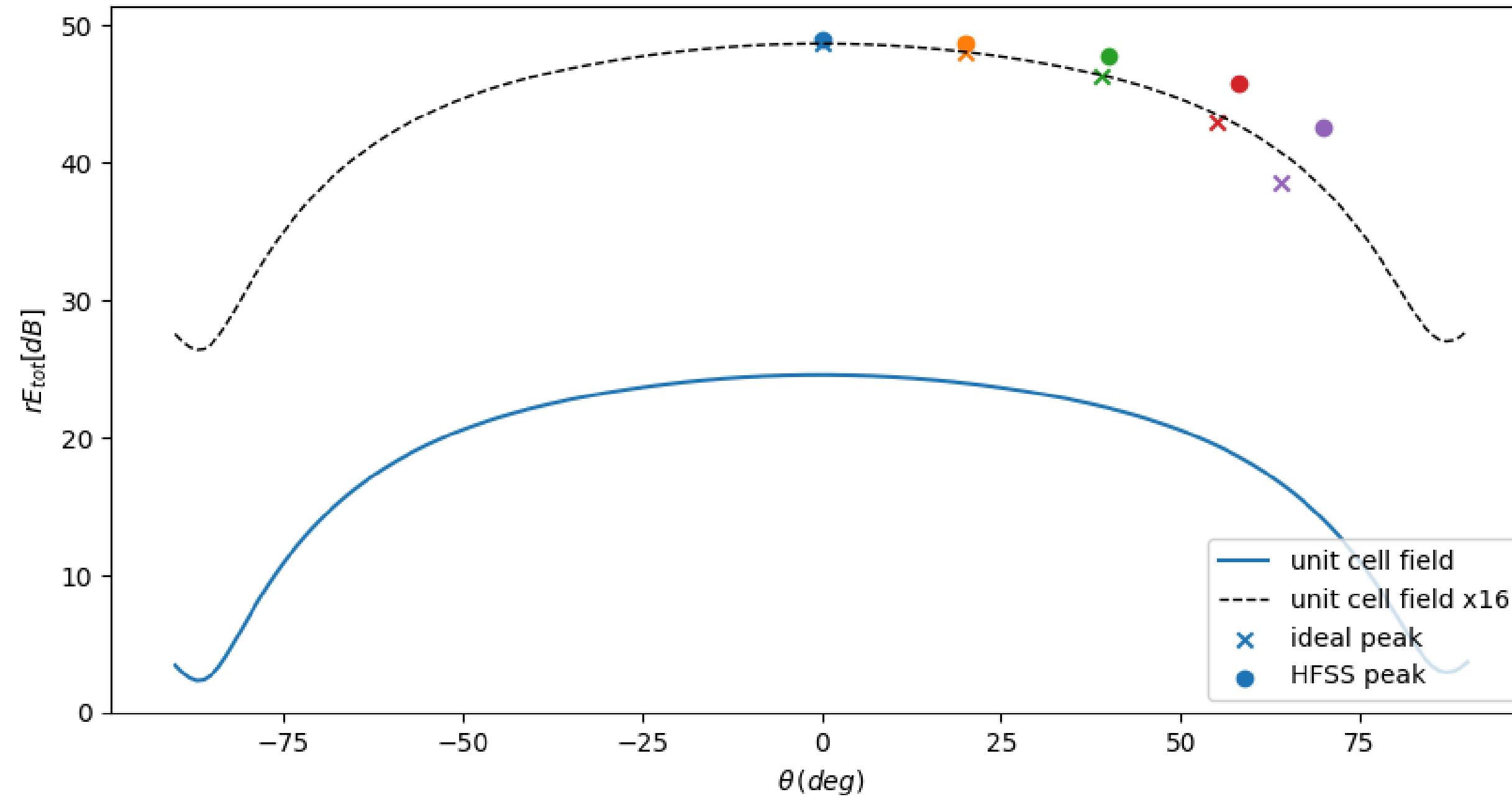


# Vivaldi Tile simulation: scanning II

- Now we scan the theta angle and keep the phi fixed (90 deg).
- We found grating lobes for large theta

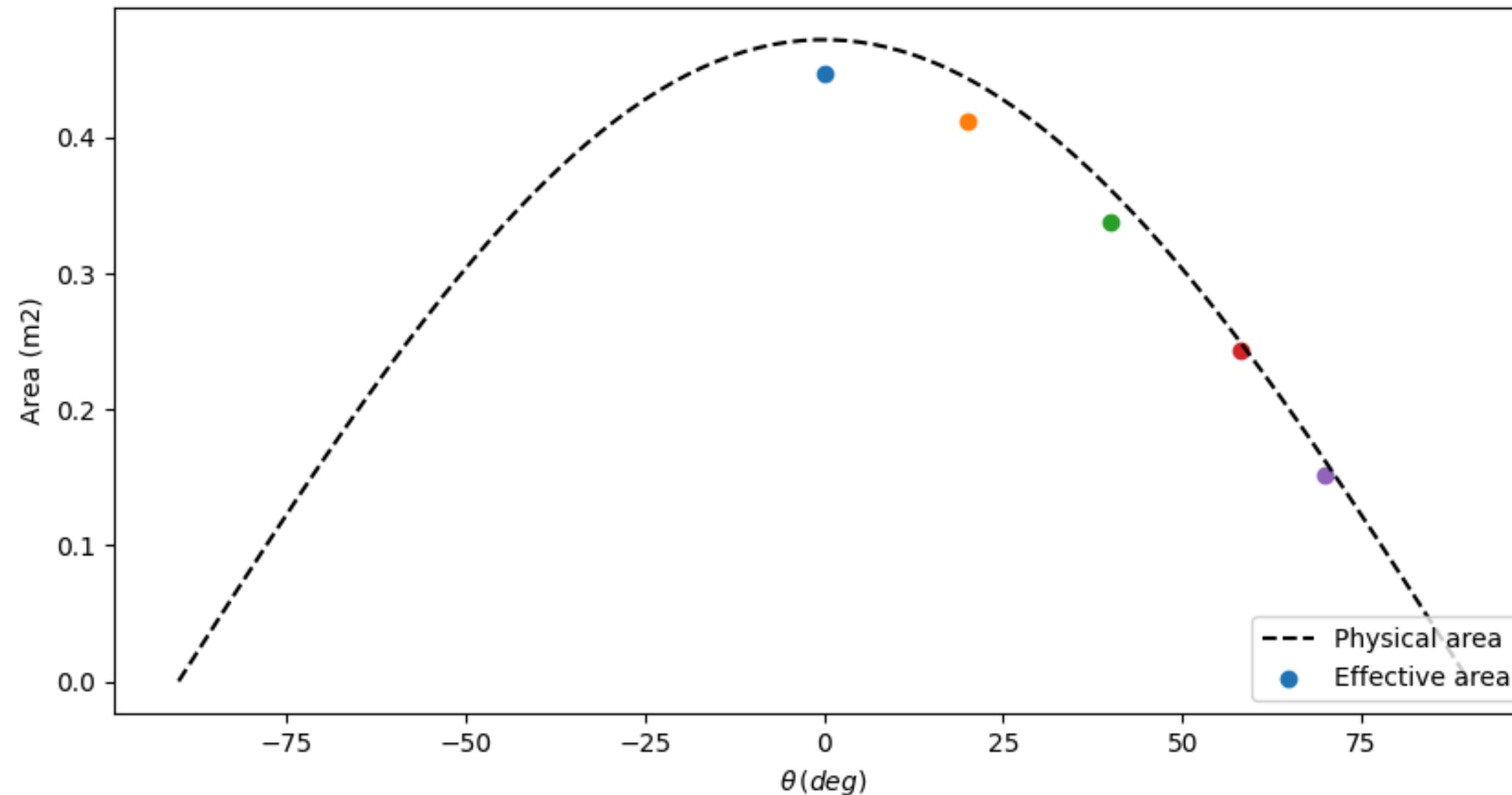


# Maximum value of the field



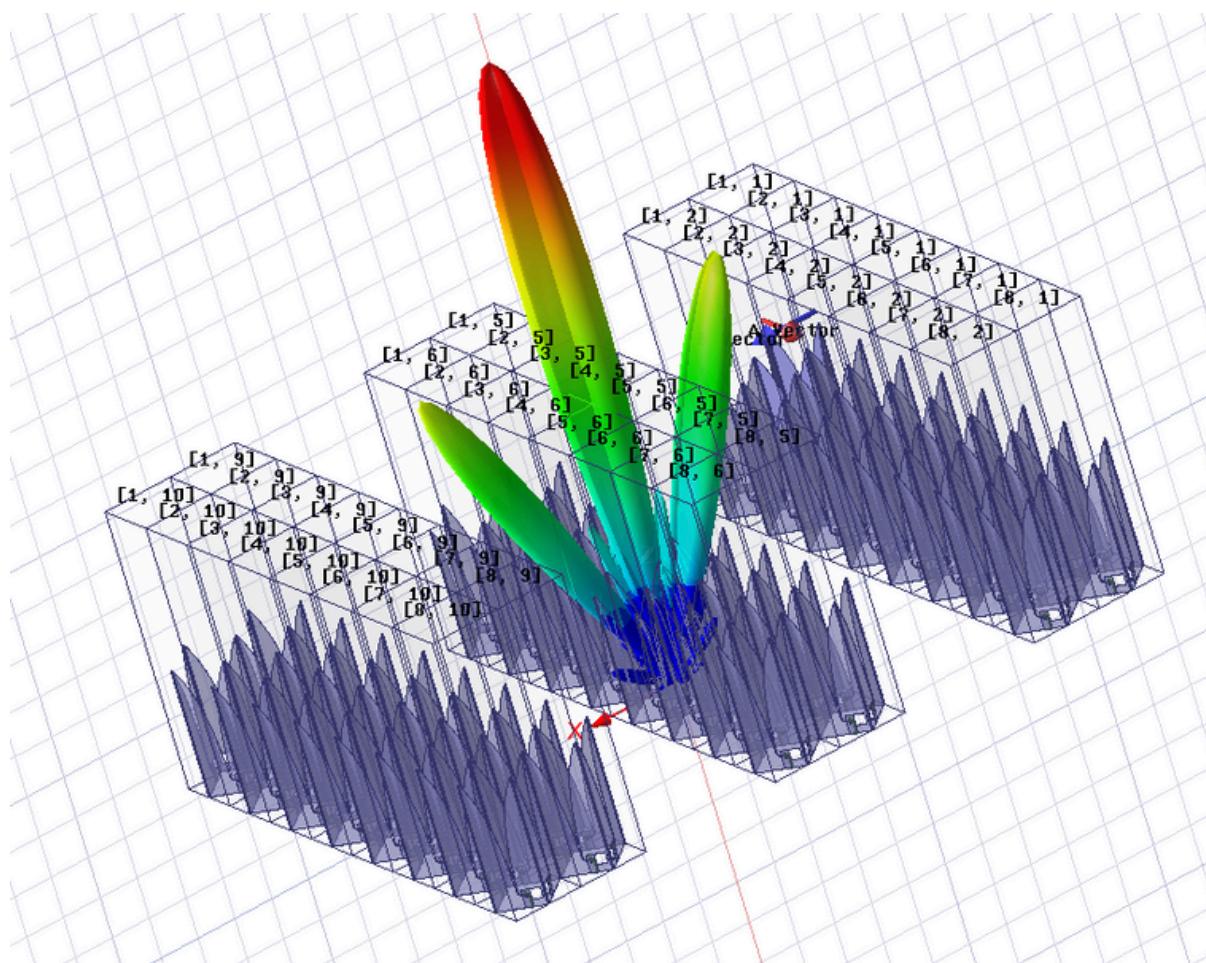
# Effective area

$$A_{\text{eff}} = \frac{\lambda^2}{4\pi} G$$

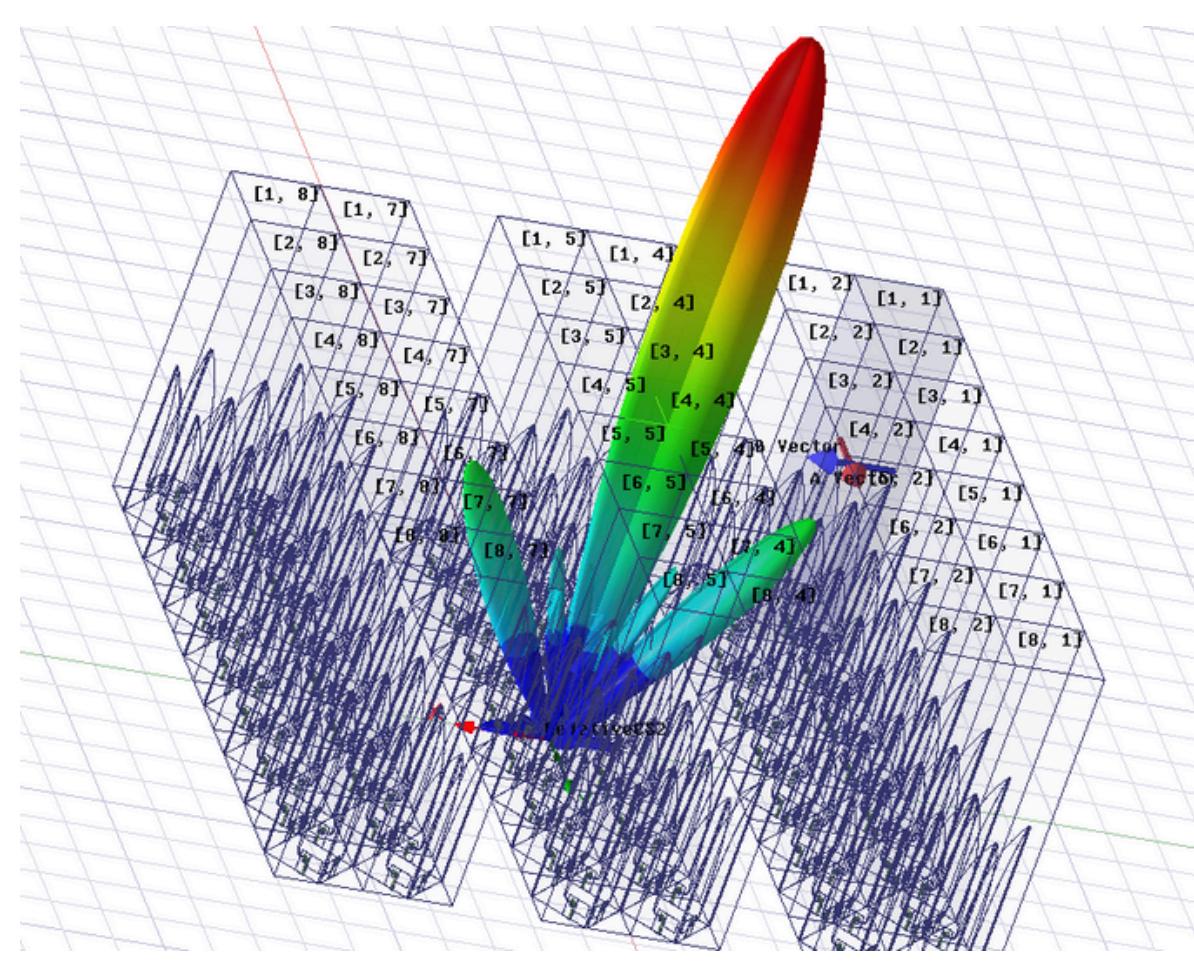


# **Array of 3 tiles**

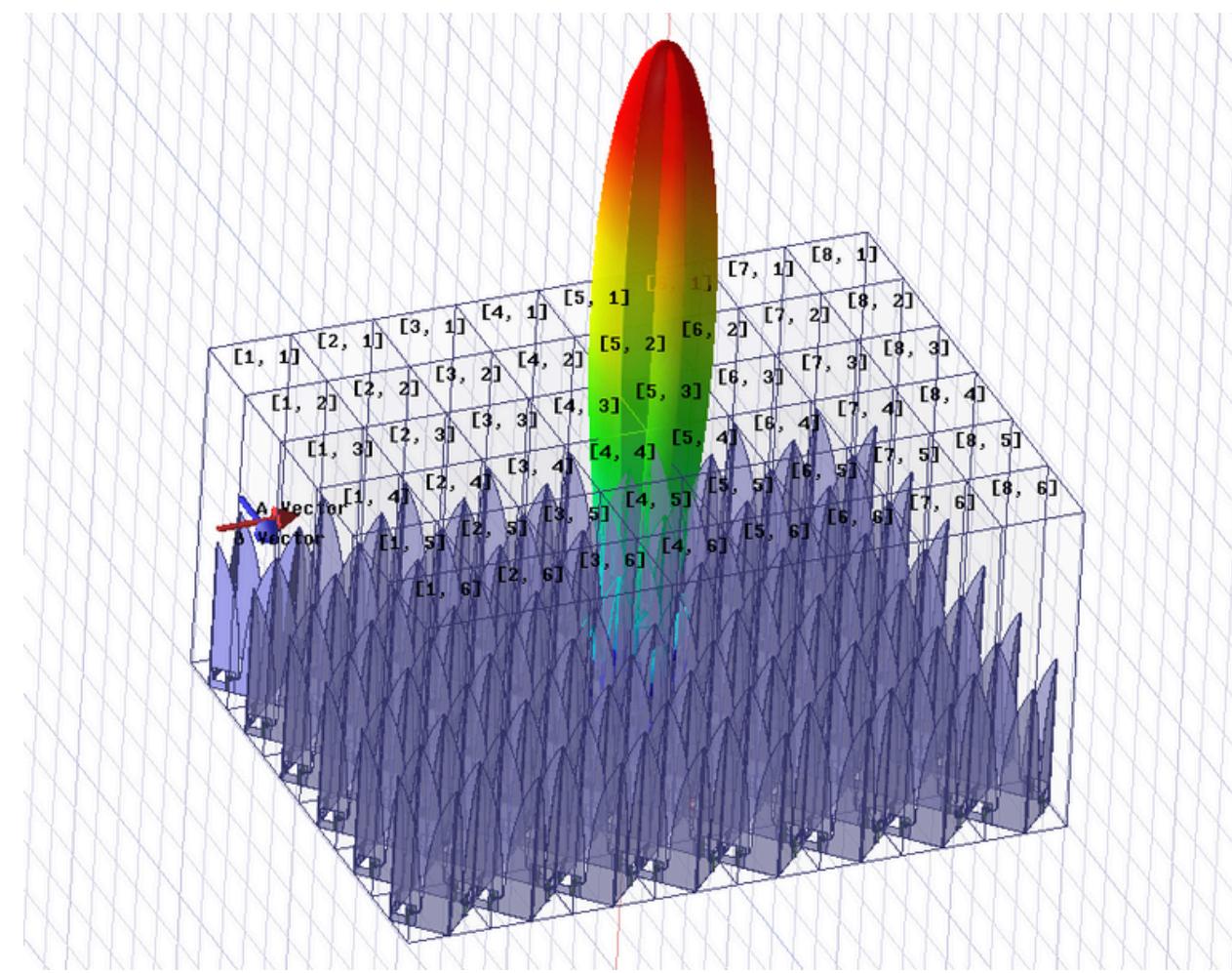
# Array of 3 tiles



Model 1



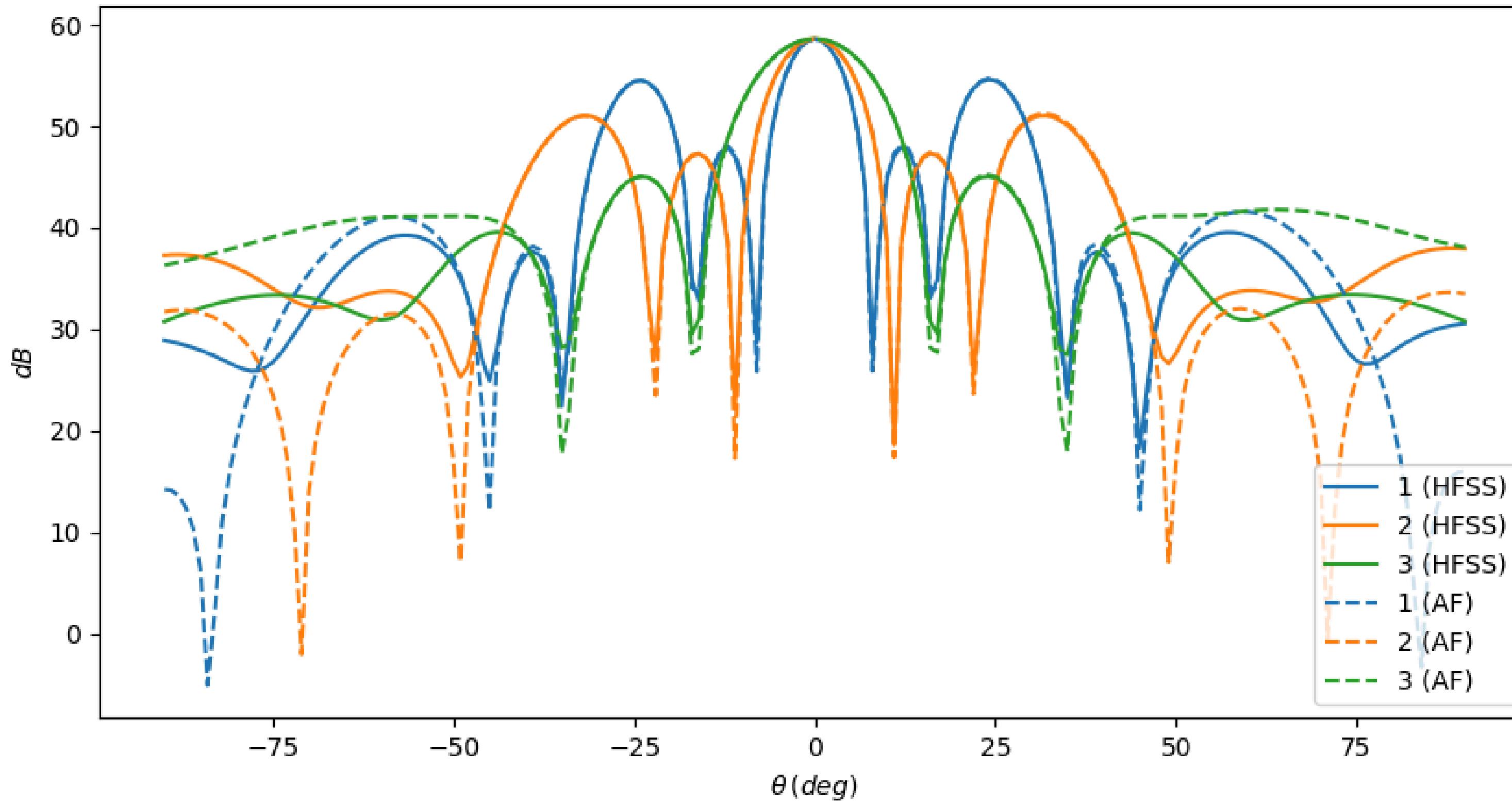
Model 2



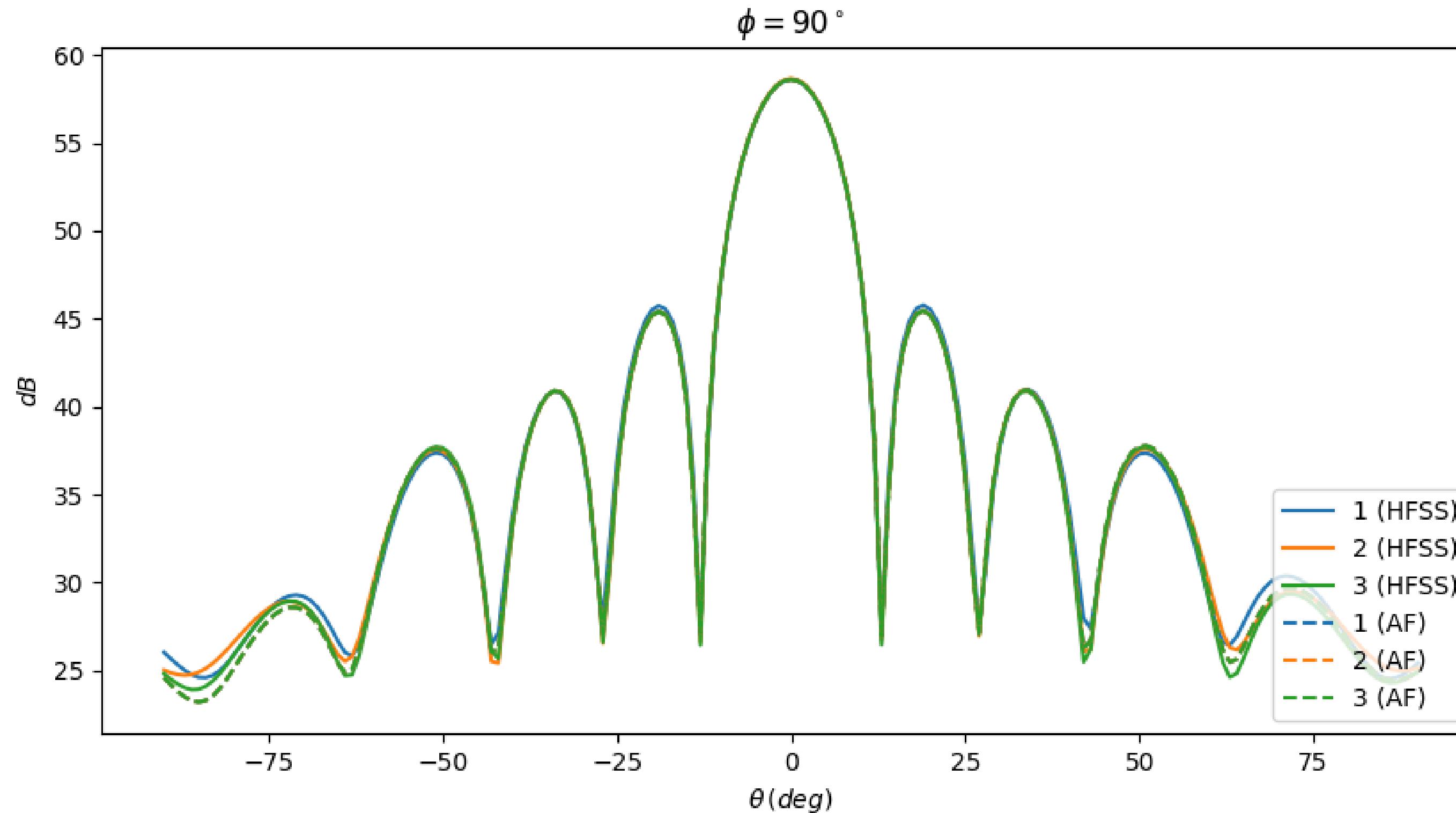
Model 3

# Simulation of the 3 tiles array

$\phi = 0^\circ$

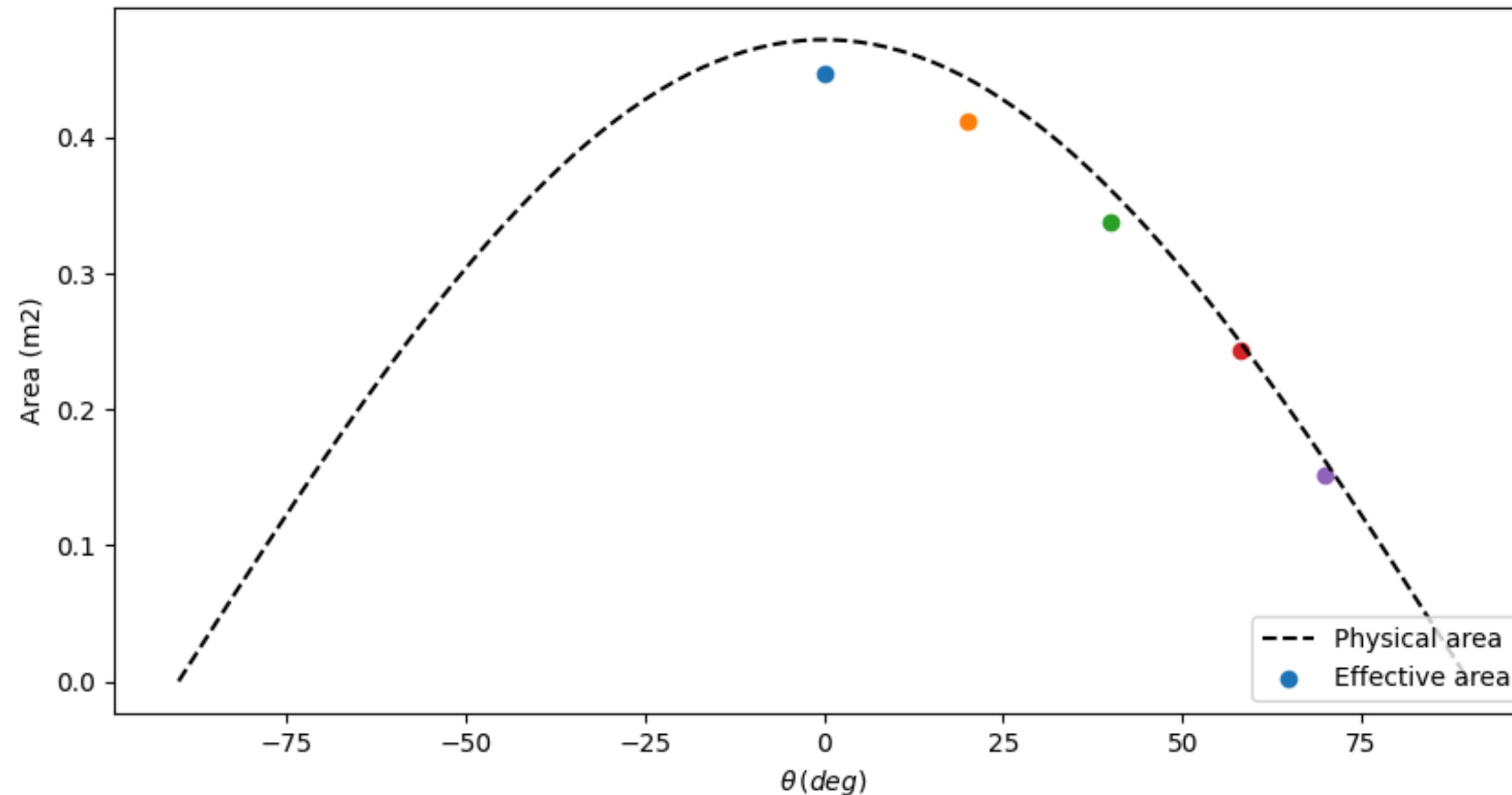


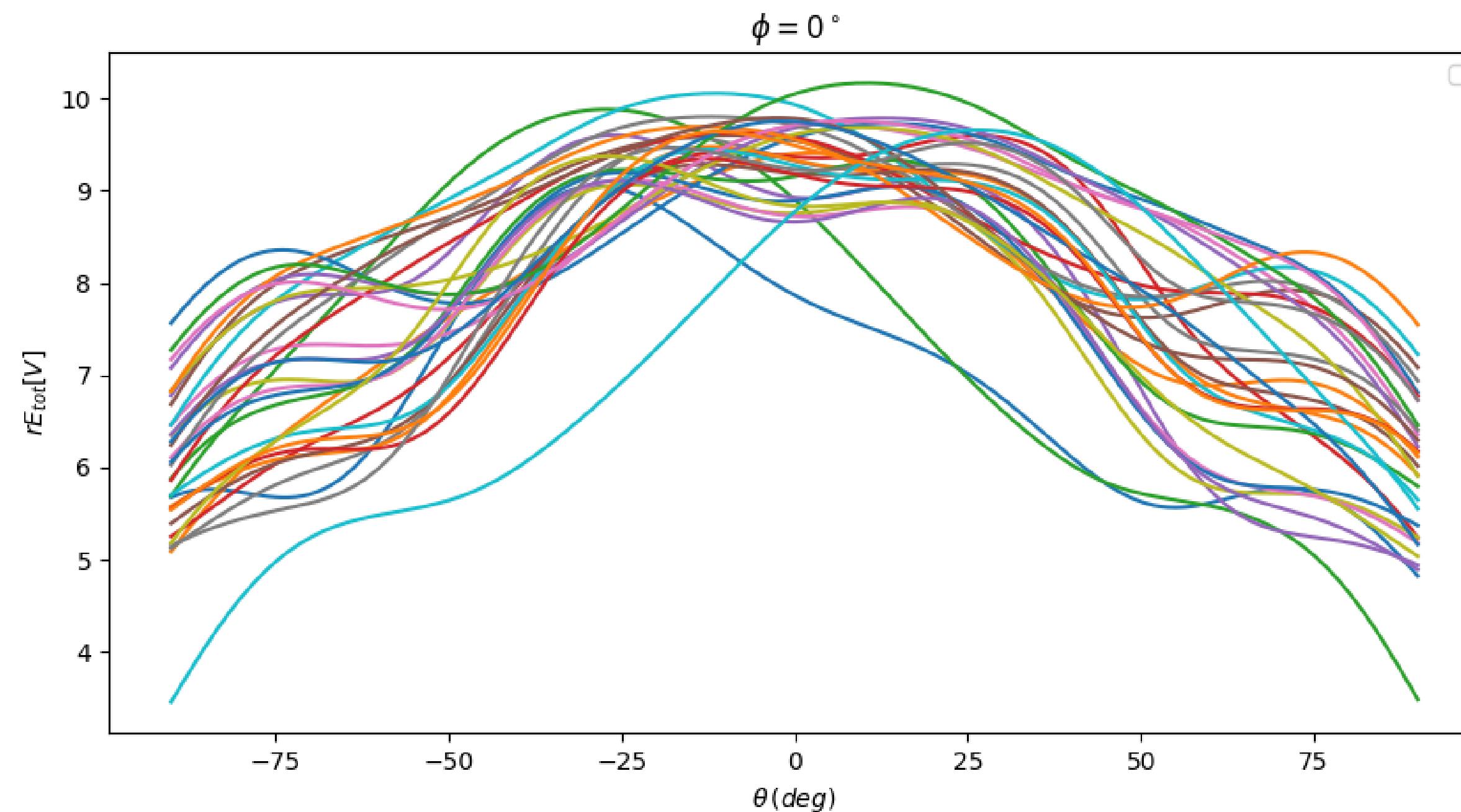
# Simulation of the 3 tiles array



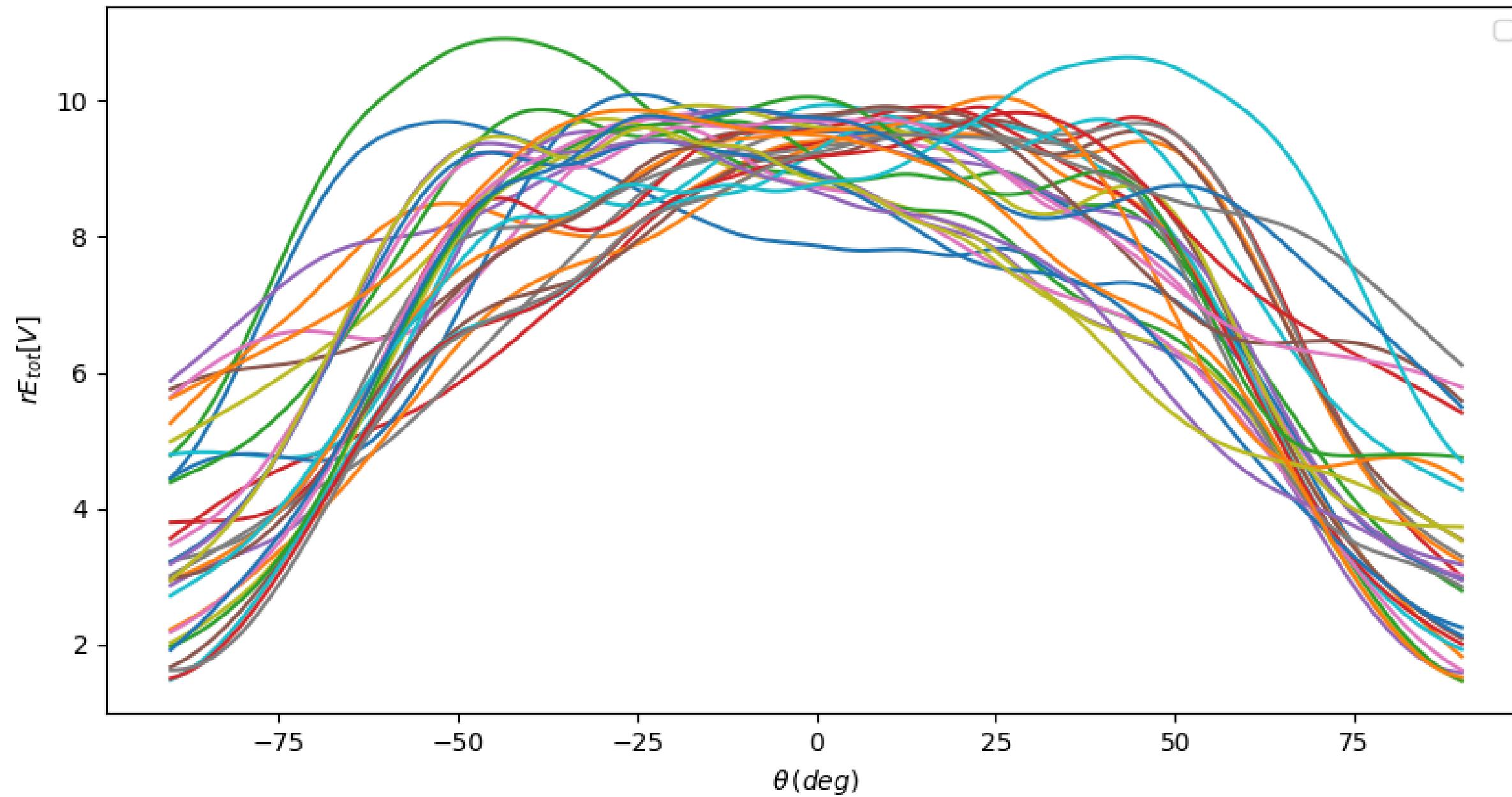
# Effective area

$$A_{\text{eff}} = \frac{\lambda^2}{4\pi} G$$





$$\phi = 90^\circ$$



# Simulated station - 36 vs 64 tiles

$\phi = 0^\circ$

