coinc_lmfit module

This is a set of convenience functions written by Aaron Mueninghoff in July 2022 as part of his work on the Quantum Astrometry project, part of the Quantum Information Science research group at Brookhaven National Labs, working under Andrei Nomerotsky and Paul Stankus. Questions and comments can be sent to aaron.mueninghoff@stonybrook.edu.

These functions can be used to take a numpy array of channel pairs in the format produced by C++ code written by Denis Dolzhenko and extract information about rate of coincidence pairs.

```
coinc_lmfit.make_nice(arr)
```

Makes the array loaded in easier to work with.

This function takes channel pair data *arr* in the format provided by Denis, changes the dt (second) column to be in nanoseconds, changes the t (first) column to be in minutes, and changes the t column to start at zero.

Parameters: arr: ndarray

A $n \times 2$ numpy array in the the format of Denis' output, where the first column is

the t value and the second column is dt value.

Returns: arr: ndarray

A $n \times 2$ numpy array of channel pairs in an easier to use format.

Notes

It is important that the array elements not be integers for the division.

Passing an array into make_nice a second time will make it not nice.

coinc_lmfit.create_barr(arrdt, Lower=-20, upper=20, dtwidth=0.05, nbins=None)
Bins an array of dt values.

Given an array of dt values, this function separates the dt values in bins ranging from *lower* to *upper* of width *dtwidth* or across *nbins* number of bins. It is assumed that the dt quantities are in nanoseconds.

Parameters: arrdt: array-like

1d array-like object containing dt values to be binned.

Returns: binmids: ndarray

1d array of the dt values of the centers of the bins.

barrdt : ndarray

1d array of the bin values.

Other Parameters: lower: float or int, optional

The lower bound on the values to be binned (default is -20).

upper : float or int, optional

The upper bound on the values to be binned (default is 20).

dtwidth : float, optional

The width of the bins (default is 0.05). Ignored if *nbins* is given.

nbins: int, optional

The number of bins to bin *arrdt* into. An alternative to selecting the bin width using *dtwidth*.

coinc_lmfit.time_bin(arr, twidth)

Bins an array of t and dt values by the t values.

Takes a $n \times 2$ array and separates it into sections that are *twidth* wide (in seconds). Returns a 1d array of the center of the bins and a list of 1d arrays of dt values.

Parameters: arr: ndarray

A $n \times 2$ array with t values in the first column and dt values in the second column.

twidth: int or float

The width of the bins (in seconds).

Returns: midlist: ndarray

1d array containing the center values of the bins (in minutes).

arrdt_tlist : list of ndarrays

A list of arrays of the dt values in each bin.

coinc_lmfit.get_amp_unc(result, explicit_unc, i, maxiter, ci_verbose)

Tries to get the uncertainty on the amplitude of a peak fit.

Takes a ModelResult and tries to calculate or estimate the uncertainty on the best-fit value of the amplitude using <code>lmfit.conf_interval</code> or the covariance matrix, respectively.

Parameters: result: LMFIT ModelResult

A ModelResult from a fit to a peak-like model using the LMFIT package.

explicit_unc : bool

Whether to calculate the uncertainty explicitly (True, default) using result.conf_interval() or to estimate the uncertainty using the covariance matrix (False).

i : int

The index of the tbin that coinc_sum is at. Used for warning when explicit uncertainty calculation has reached *maxiter*.

maxiter: int

To be passed to result.conf_interval. The maximum number of iterations to do when calculating the uncertainty.

Returns: amp_unc : float

The uncertainty on the amplitude or np.nan if the function failed (see Notes).

Other Parameters: ci_verbose : bool, optional

Passed to lmfit.conf_interval. If *ci_verbose* is True (default is False) then lmfit.conf_interval will print out the result of every iteration in calculating the uncertainties.

Raises: UserWarning

Raises a UserWarning if the *maxiter* is reached. Does not raise UserWarning's from result.conf_interval about convergence, because

Notes

Calculating the uncertainties explicitly takes much longer than estimating them.

Sometimes the fit cannot calculate the covariance matrix, and sometimes *conf_interval* cannot calculate the uncertainty. These usually happen when the amplitude is very small. If one of these methods fails, the function will try to get the uncertainty using the other method. If both fail, then *get_amp_unc* returns np.nan so that it can be fixed later by fix amp_unc.

This function uses a context manager to temporarily change the way Warnings are filtered. This is so that warnings from conf_interval about convergence are ignored but warnings about hitting the maximum iteration are raised as exceptions. The maxiter warning is raised as an exception so it can be caught with a try except clause, modified, and then raised as a warning.

coinc lmfit.fix amp unc(coinc arr, unc arr)

Estimates amplitude uncertainty where get_amp_unc couldn't.

Sometimes, get_amp_unc will fail, outputing np.nan. This function estimates the uncertainties where get_amp_unc could not, by taking the average of the uncertainties of the closest 10% of amplitudes by best-fit value.

Parameters: coinc_arr : ndarray

1d array of the best-fit amplitudes.

unc_arr : ndarray

1d array of the uncertainties on the best-fit ampltides, where some values may be

np.nan.

Returns: unc_arr: ndarray

Id array of the uncertainties on the best-fit amplitudes containing no np.nan's. coinc_lmfit.coinc_sum(arr, sigma_width=1.5, tbinwidth=10, dtbinwidth_full=0.005, dtbinwidth=0.05, peakmodel=None, explicit_unc=True, maxiter=2000, ci_verbose=False) Calculates the number of coincident pairs as function of time.

Takes an array in the format that make_nice outputs. Fits a single gaussian to all of the dt values, then separates the array into bins by time and fits a gaussian with the same width and center as the gaussian fit to the full array. Then calculates the area under each gaussian (without the background) and the uncertainty on the area.

The gaussian fitting function is $F(x; A, \mu, \sigma, c) = \frac{A}{\sigma\sqrt{2\pi}}e^{-(x-\mu)^2/2\sigma^2} + c$ where A, μ, σ, c are the amplitude, mean, sigma, and constant background, respectively.

Parameters: arr: ndarray

 $n \times 2$ array with t values (in seconds) in the first column and dt values (in nanoseconds) in the second column.

sigma width : float, optional

The distance in sigma units from the center of the gaussian to integrate over. That is $\int_{\mu-w_\sigma\sigma}^{\mu+w_\sigma\sigma}G(x)\,dx$ where μ,σ,w_σ,G are the center, the sigma, the $sigma_width$, and the gaussian function, respectively.

tbinwidth : *float*, *optional*

The width in time (seconds) in which to bin *arr*. A gaussian will be fit to the dt values in each of these bins.

Returns: tbinmids: ndarray

1d array of the center values of the bins (in minutes).

coinc_arr : ndarry

1d array of the areas under the gaussian within $sigma_width$ in each time bin.

unc_arr : ndarray

1d array of the uncertainties on the areas in *coinc_arr*.

Other Parameters: dtbinwidth_full : float, optional

The width (in nanoseconds) of the bins for the dt values of the entirety of *arr*, to be used for the initial gaussian fit. Default is 0.005.

dtbinwidth : float, optional

The width (in nanoseconds) of the bins for the dt values of each time bin. This should be larger than *dtbinwidth_full* because there will be fewer values to bin. Default is 0.05.

peakmodel: LMFIT peak-like built-in model, optional

The peak-like model to use to fit the HBT peaks in the data (default is GaussianModel). Must be a model that has a guess function (all built-in models have this).

explicit_unc : bool, optional

Whether or not to calculate the uncertainies on the amplitude explicitly. Default is True.

maxiter: int, optional

Passed to get_amp_unc which passes it to lmfit.conf_interval. The maximum number of iterations (default is 2000) for lmfit.conf_interval to use when calculating explicitly. If a UserWarning is raised about reaching the maxiter, then the maxiter should be raised. Setting maxiter to np.inf is probably a bad idea.

ci_verbose : bool, optional

Passed to get_amp_enc, which passes it to lmfit.conf_interval. If ci_verbose is True, then lmfit.conf_interval will print out the result of every iteration in calculating the uncertainties.

UserWarning

Calls get_amp_unc which may raise a UserWarning about lmfit.conf_interval reaching maxiter. If this happens, *maxiter* should be raised.

Notes

Raises:

The expression for the area is derived here. Let w_{dt} be the width of the dt bins, and let w_{σ} be the width of the integration window in terms of sigma. Then the fitting equation is written as $f(x;A,\mu,\sigma,c)=\frac{A}{\sigma\sqrt{2\pi}}e^{-(x-\mu)^2/2\sigma^2}c$, where the parameters are defined as above.

$$egin{align*} Area &= \int_{x=\mu-w_{\sigma}\sigma}^{x=\mu+w_{\sigma}\sigma} rac{A}{\sigma\sqrt{2\pi}} e^{-(x-\mu)^2/2\sigma^2} \, dx \ u &= x-\mu, \quad \Longrightarrow \quad du = dx \quad \Longrightarrow \quad dx = du \ &= rac{2A}{\sigma\sqrt{2\pi}} \int_{u=0}^{u=w_{\sigma}\sigma} e^{-u^2/2\sigma^2} \, du \ v^2 &= rac{u^2}{2\sigma^2} \quad \Longrightarrow \quad v = rac{u}{\sqrt{2}\,\sigma} \quad \Longrightarrow \quad dv = rac{du}{\sqrt{2}\,\sigma} \quad \Longrightarrow \quad du = \sqrt{2}\,\sigma dv \ &= rac{2A}{\sqrt{\pi}} \int_{v=0}^{v=w_{\sigma}/\sqrt{2}} e^{-v^2} \, dv \ &= A \cdot rac{2}{\sqrt{\pi}} \int_{0}^{w_{\sigma}/\sqrt{2}} e^{-v^2} \, dv \ &= A \cdot \mathrm{erf}\left(rac{w_{\sigma}}{\sqrt{2}}
ight) \end{aligned}$$

This value is then divided by w_{dt} to get the total number of counts under the gaussian (unitless). coinc_lmfit.make_slices(bounds_list, tbinmids)

Creates a list of slices using the centers of time bins.

Parameters: bounds_list: array-like

array or list of 2-tuples, where first entry in each tuple is the beginning of the region (in minutes) and the second entry is the end (in minutes).

tbinmids: ndarray

1d array of the center values of time bins (in minutes).

Returns: slice_list: *list*

List of slices in the same order as bounds_list.

coinc_lmfit.fit_sin(tbinmids, coinc_arr, unc_arr=None, init_vals=None, maxiter=200, ci_verbose=False, suppress_warnings=False)

Fits a sine function to the result of coinc sum.

Takes the results of coinc_sum and fits it to $F(x;A,f,\phi,c)=A\sin(fx+\phi)+c$, where A,ϕ,f,c are the amplitude, frequency, phase shift, and background constant, respectively. Can explicitly calculate the uncertainties on the best-fit values.

Parameters: tbinmids: ndarray

1d array of the center values of time bins (in minutes).

coinc_arr : ndarray

1d array of the area under the gaussians fitted in each bin by coinc_sum.

unc_arr : ndarray, optional

1d array of the uncertainties on *coinc_arr*; to be used for weighting in the

fit.

Returns: res: LMFIT ModelResult

A lmfit.ModelResult object for the fit

Other Parameters: init_vals: array-like, optional

An iterable with 4 float elements. Used as the initial values in the fit for the parameters in this order: amplitude, frequency, phase shift, constant. Note that the <code>lmfit.guess</code> function for <code>lmfit.SineModel</code> seems to work very well. You probably dont need to specify the initial values.

maxiter: int, optional

To be passed to lmfit.conf_interval; the maximum number of iterations to be used. Default is 200.

ci_verbose : bool, optional

Passed to lmfit.conf_interval. If *ci_verbose* is True (default is False), then lmfit.conf_interval will print out the result of every iteration in calculating the uncertainties.

suppress_warnings : bool, optional

Whether to suppress warnings from conf_interval. Default is False.

Raises: UserWarning

lmfit.conf_interval may raise UserWarnings about bad convergence or reaching maxiter.

Notes

It is often easiest to give the first three arguments by unpacking (an array in the shape of) the output of coinc_sum.

coinc_lmfit.make_label(result, percent=True)

Makes a legend label given the ModelResult of a SineModel fit.

Parameters: result : LMFIT ModelResult

The lmfit.ModelResult object, usually from fit sin.

percent: bool

If True (default), include the percent uncertainty on the best fit values.

Returns: str

The legend label as a single string.

Notes

The fitting function is not in the preferred form.

$$F_{ ext{fit}}(x;A,f,\phi,c) = A\sin(fx+\phi) + c, \quad F_{ ext{plot}}(t;A,T,\delta,D) = A\cos((2\pi t/T) - \delta) + D$$

This function does the conversion of both the values and uncertanties, and converts the time-units to seconds.

This function uses sigfig.round to handle significant figures and uncertanties properly.

coinc_lmfit.plot_chpairs(coinc_lists, res_lists=None, slice_lists=None,
color_list=None, sin_pointnum=2000, ylabel_binwidth=None, figure_kwargs=None,
errorbar_kwargs=None, sin_kwargs=None, legend_kwargs=None)

Plots the results of coinc sum with or without fit sin.

Takes one or multiple results of coinc_sum and plots it. Can also plot the results of one or multiple sin fits on top. Written with customization in mind.

Parameters: coinc_lists : array-like

A list of 3-tuples, where the elements of the 3-tuples are the results from a single call of coinc sum (tbinmids, coinc_arr, unc_arr).

res_lists: array-like containing LMFIT ModelResults, optional

A list of lists of ModelResults from sin_fit. The length of res_lists should match the length of coinc_list, and the length of each list in res_lists should match the length of the corresponding list in slice_lists.

slice_lists: array-like containing slice, optional

A list of lists of slices (easily made with make_slices). The length of slice_lists should match the length of coinc_lists, and the length of each list in slice_lists should match the length of the corresponding list in res_lists.

Returns: fig: Matplotlib figure object

axs: ndarray of Matplotlib AxesSubplot ndrray of axes that the errorbars (and sin fits) are plotted on.

extra: Matplotlib AxesSubplot

The axis used for having common axis labels.

Other Parameters: color_list: *list of strings, optional*

A list of the colors the sine plot lines should be. Passed to ax.plot, look at MatPlotLib documentation for acceptable color strings. Default is the MatPlotLib base colors (except black and white).

sin_pointnum : int, optional

The number of points used to plot the sine fits.

ylabel_binwidth : int or str, optional

If specified, then the common ylabel is automatically created.

figure kwargs: dict, optional

A dict of keyword arguments to pass to plt.figure. Only overrides default arguments that are specified.

errorbar_kwargs : dict, optional

A dict of keyword arguments to pass to ax.errorbar. Only overrides default arguments that are specified.

sin_kwargs : dict, optional

A dict of keyword arguments to pass to ax.plot for plotting the sin fits. The color of the line should be specified with *color_list*, not in this dict. Only overrides default arguments that are specified.

legend_kwargs : dict, optional

A dict of keyword arguments to pass to ax.legend. Only overrides default arguments that are specified.

Notes

Look at the example file to see how to properly enter the required argument(s). Since this function returns the figure and the axes, additional formatting can be done on the returned figure (such as changing axis limits or labels, and saving the figure).