\$RTA\$ Documentation Using the RTA library to compute on-line sound descriptors

Jean-Philippe Lambert

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1 Doxygen: Code self-contained documentation

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The RTA library is documented within the code, using the Doxygen format. This is the reference documentation. To generate HTML and LaTeX output, use the command doxygen Doxyfile within the rta directory. The HTML documentation will then be accessible from rta/doc/html/index.html. To compile a PDF document, use the command make within the rta/doc/latex directory, which is generated by the previous command.

2 Generalities

The RTA library is frame-based, which means that for any vectors of samples, a set of descriptors can be computed without adding any delay. A noticeable exception to this is the *delta* (and *delta-delta*) computation, as it is based on more than on frame.

There are two variants of each function, one being _stride post-fixed. It allows to directly access interleaved data without copying them (like a stereophonic samples vector). However, using a big *stride* value may lead to a bad usage of the memory cache.

Any index starts at 0.

Some functions require an initialisation before any processing, in order to pre-calculate what will not depend on the incoming frame. Every allocation must be done before anything else, outside of the functions themselves except mentioned otherwise.

Some descriptors can be computed by several functions, and the results may slightly differ for several reasons: the functions does not rely on the same algorithms and the signal used for the computation may differ (due to windowing, filtering, etc.). The auto-correlation from yin and from the LPC are not the same and there is a lot of ways to get the energy: from yin, as the sum of the squares of the samples, from LPC, or as the first MFCC coefficient.

3 Library configuration

A file named $rta_configuration.h$ must be present within your sources in order to use the RTA library. An empty file means that the defaults settings are used for the compilation.

Instead of using the malloc, realloc and free functions from the stdlib.h, one can respectively define rta_malloc, rta_realloc and rta_free.

The floating-point precision can be simple, double or long double, according to the definition of RTA_REAL_TYPE to respectively RTA_FLOAT_TYPE, RTA_DOUBLE_TYPE or RTA_LONG_DOUBLE_TYPE. The constants in rta_float.h are then redefined according to the proper type from float.h. The same applies to the functions in rta_math.h from math.h.

Note that the long double precision is not supported when using the Apple's

VecLib by setting RTA_USE_VECLIB to 1.

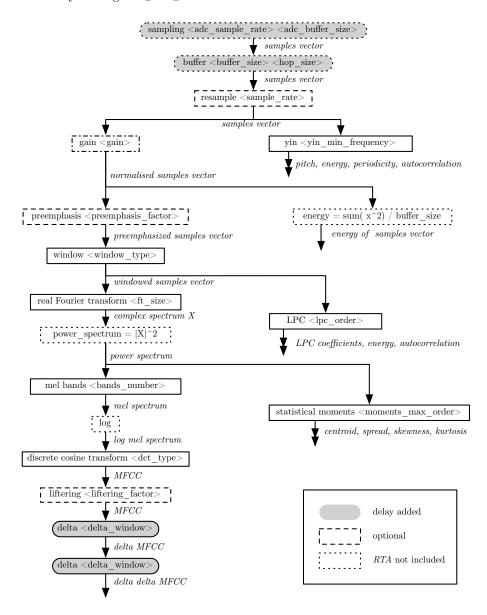


Figure 1: Sound descriptors data-flow

4 Descriptors based on a vector of samples

The vector of samples is characterised by its sample-rate and its size. Moreover, to use sizes larger than those provided by the sound card (e.g. for Fourier transforms), the hop-size gives the number of samples between two consecutive vectors that overlap if the hop-size is smaller than the vector-size.

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4.1 Re-sampling

Down-sampling a signal is interesting to lower the further computations, especially for the computational-intensive algorithms, like *yin*. It can helps to concentrate on a range of the spectrum where the information is pertinent for the analysis: the *LPC* (which is linear, as its name suggests it) finds poles and zeros to fit the whole spectrum; to keep the results under 5 kHz, we simply use a sample-rate of 11 kHz. If the original signal sample-rate is 44 kHz, we can use the function <code>rta_downsample_int_mean</code> with a factor of 4. This function implies a low-pass filtering. The function <code>rta_downsample_int_remove</code> can produce aliasing if the original signal contains information above the resulting sample-rate. Note that the resulting samples vector-size is smaller, according to the given factor.

4.2 Yin

The *yin* algorithm computes the periodicity of a samples vector, finding its most probable *lag*. To get the period in Hz, on simply multiply this lag by the input vector sample-rate.

Note that the *yin* algorithm operates on a non-windowed samples vector. As it can be computational-intensive, a down-sampling of the incoming vector is often performed: to track a pitch below 1 kHz, one can use a sampling-rate of 11 kHz, or even 5.5 kHz, depending on the results quality request. One can then check the *absolute minimum* found, which gives the *periodicity* as long as the absolute minimum is positive:

$$periodicity = 1 - \sqrt{absolute_minimum}$$

Before computing anything, a new *yin* structure must be allocated and filled with the rta_yin_setup_new function. It will be released by the function rta_-yin_setup_delete. (The auto-correlation result vector must also be allocated beforehand, like any results vector.)

4.3 Gain and integer-float conversion

The samples are generally coded by floating-points number over 32 bits, within the range [-1.0, 1.0]. To convert them into 16 bits signed integers (which is the format used by some systems, like HTK), one can apply a simple gain of 2^{15} by multiplying every sample.

4.4 Pre-emphasis

The pre-emphasis is a simple first-order difference between a current sample and the previous one (weighted by a factor):

$$s(n) = s(n) - f \times s(n-1)$$

It is often used for voice analysis with a factor of 0.97 as it reduces the low frequencies while raising the high frequencies, thus amplifying the contrast.

4.5 Windowing

If the samples vector-size is known, it is possible to pre-calculate the weights that will be used to apply a given function. The function rta_window_hamming_-weights computes a *Hamming* window while the function rta_window_hamn_-weights computes a *von Hann* window. These (or any weights vector) can be applied with the rta_window_apply function. The _in_place post-fixed functions change the input samples vector values directly.

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If the samples vector-size is not known in advance, one can still apply the window using the rta_window_rounded_apply function. There is no interpolation, then. The weights vector indexes are simply scaled and rounded: this is efficient but the rounding error may be unacceptable if the size of the weights vector is too small comparing with the samples vector size. It is also possible to compute and apply a window on the fly, with the functions rta_window_-hann_apply and rta_window_hamming_apply.

4.6 Linear predictive coefficients (LPC)

The rta_lpc function calculates the linear predictive coefficients (LPC) for a samples vector, using an auto-correlation and a Levinson-Durbin decomposition. Note that the LPC order is one value less than the LPC size.

The first LPC coefficient is always 1 and is often replaced (e.g. in HTK) by the prediction error, which gives the energy of the samples vector. If the LPC is computed on overlapping samples vectors, they are often windowed in order for the coefficients to evolve smoothly from frame to frame, and the energy is then reduced (by a constant factor depending on the window).

5 Descriptors based on the power spectrum

Some descriptors are based on the power spectrum of a samples vector. The samples vector is first windowed. Then a real *Fourier* transform is applied, giving a complex spectrum. The power spectrum is the square of the magnitude of the complex spectrum.

5.1 Real Fourier transform

Before computing a real Fourier transform, a new real Fourier transform setup must be allocated and filled with the rta_fft_real_setup_new function, with the type real_to_complex_1d. It will be released by the function rta_fft_setup_delete. The function rta_fft_execute applies the Fourier transform to a samples vector.

The transform size must be a power of 2. If the transform size is bigger than the actual samples vector, it is then padded with zeros.

By convention (e.g. HTK), no scale is applied to this direct transform. The inverse of the transform size can later be applied to the inverse transform in order to obtain the identity transform.

5.2 Complex spectrum to power spectrum

The power spectrum is the square of the magnitude the complex spectrum. Its size is half the size of the *Fourier* transform plus one (the last element corresponds to the *Nyquist* frequency). To get a correspondence between the power spectrum index and the corresponding frequency, one can apply a simple ratio between the maximum frequency (which is half of the sample-rate) and the maximum index (which is half of the *Fourier* transform size, as all the indexes start at 0):

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$$frequency = index \frac{sample_rate}{transform_size}$$

5.3 Statistical moments

The statistical moments can be computed from any samples vector, as long as the weights are positive, as they represent the probability of the random variables to appear. The power spectrum conforms to this, as any value is positive, but not the amplitude spectrum (if not translated above 0). The same applies for the moments of the mel bands.

The moments are calculated over the indexes and weighted by the input values. They will be normalised by the sum of the input values in order to get the indexes probability. Note that all moments (but the first) are centred. Any moment above the second can be standardised.

The moments described hereafter describe the power spectrum moments.

5.3.1 Centroid

The spectral centroid is the first moment over the indexes weighted by the vector of power spectrum values. It is computed by rta_weighted_moment_1_indexes. The result unit is *index* (of the power spectrum, starting at 0). This function returns also the input sum as it can be used in further calculations.

$$m_1 = centroid = \frac{\sum_i i \times input(i)}{\sum_i input(i)}$$

5.3.2 Spread and deviation

The spectral spread is the second central moment over the indexes weighted by the vector of power spectrum values. It is computed by $rta_weighted_-moment_2_indexes$. The result unit is $index^2$ (of the power spectrum, starting at 0).

$$m_2 = spread = \frac{\sum_i (i - centroid)^2 \times input(i)}{\sum_i input(i)}$$

The standard deviation is $std = \sqrt{spread}$.

5.3.3 Skewness

The spectral skewness is the third standard central moment over the indexes weighted by the vector of power spectrum values. It is computed by rta_std_-weighted_moment_3_indexes. The result is without unit.

$$m_{3std} = skewness = \frac{\sum_{i}(i-centroid)^{3} \times input(i)}{std^{3} \sum_{i} input(i)}$$

5.3.4 Kurtosis

The spectral kurtosis is the fourth standard central moment over the indexes weighted by the vector of power spectrum values. It is computed by rta_std_-weighted_moment_4_indexes. The result is without unit.

$$m_{4std} = kurtosis = \frac{\sum_{i}(i-centroid)^{4} \times input(i)}{std^{4}\sum_{i}input(i)}$$

Note that the kurtosis is often defined as the fourth cumulant divided by the square root of the variance, which gives $kurtosis = \frac{m_4}{std^4} - 3$. This function does not include the "-3" term.

6 Descriptors based on the mel bands

The mel scale can be derived from the frequencies in hertz. The conversion functions are in rta_mel.h. They are based on two slightly different formulas, according to HTK or the $Auditory\ Toolbox$ with the respective suffix _htk or _slaney.

The power spectrum is integrated into several bands, according again to HTK or the $Auditory\ Toolbox$. The integration window peak is 1 for HTK, while the sum of any channel is 1 for the $Auditory\ Toolbox$.

In order to reproduce the results of one of these tools, one must obviously choose the desired variant among the whole computation process.

6.1 Mel bands

The mel bands integration is done in the magnitude (abs) or the power (abs^2) domain, using respectively the function rta_spectrum_to_bands_abs or rta_spectrum_to_bands_square_abs. This respectively gives

$$mel_bands = weights_matrix \times spectrum$$

or

$$mel_bands = (weights_matrix \times \sqrt{spectrum})^2$$

The latter is the default (for HTK and $Auditory\ Toolbox$) but it involves more computation.

The matrix to multiply the power spectrum vector with, in order to obtain the mel bands, must be computed beforehand, using the rta_spectrum_to_mel_bands_weights function.

6.2 Mel-frequency cepstral coefficients (MFCC)

First, the logarithm of the mel bands values is taken. In order to avoid log(0), one can add a very small value to the mel bands values before taking the log.

Then a cepstrum is computed for the log of the mel spectrum, using a discrete cosine transform (DCT) of type II. HTK uses an orthogonal but not unitary transform, while the Auditory Toolbox uses an orthogonal and unitary transform. First, a weights matrix is constructed with the function rta_dct_weights and it is then applied to the mel bands vector using the function rta_dct. The coefficients obtained are the MFCC.

The *MFCC* as any *DCT* coefficients, are ordered by the order of importance to model the spectrum. However, one can need to modify them (for visualisation or further processing) using the liftering functions in rta_lifter.h. These functions are provided for the *HTK* or *Auditory Toolbox* compatibility.

6.3 Delta and delta-delta MFCC

The function rta_delta computes a simple linear slope on a sequence of fixedrate sampled data (the frames). The *delta* values correspond to the mid-point frame (which is not used, by the way). *It means that the delta values are not* those of the last frame. Considering the filter-size, which is the size of the sequence taken into account, the delay (in frames) introduced is half of the filter-size (rounded down as it is always odd).

Note that the HTK DELTAWINDOW variable is not the same as the filter-size (the same applies for the ACCWINDOW variable):

$$filter_size = 1 + 2 \times DELTAWINDOW$$

Beforehand, a matrix of weights to multiply the sequence vector with is constructed by the function rta_delta_weights. A normalisation factor, computed by the function rta_delta_normalization_factor gives delta values, which are independent of the filter-size. The normalisation factor can be applied directly to the weights matrix but the rounding errors may be unacceptable when using the simple floating-point precision.

Applying the delta computation again gives the delta-delta values, adding a new delay of half of the delta-delta filter-size.