Cooling Improves Cosmic Microwave Background Map-Making When Low-Frequency Noise is Large

Bai-Qiang Qiang¹ and Kevin M. Huffenberger D¹

¹Department of Physics, Florida State University, Tallahassee, Florida 32306

ABSTRACT

In the context of Cosmic Microwave Background data analysis, we study the solution to the equation that transforms scanning data into a map. As originally suggested in "messenger" methods for solving linear systems, we split the noise covariance into uniform and non-uniform parts and adjust their relative weights during the iterative solution. With simulations, we study mock instrumental data with different noise properties, and find that this "cooling" or perturbative approach is particularly effective when there is significant low-frequency noise in the timestream. In such cases, a conjugate gradient algorithm applied to this modified system converges faster and to a higher fidelity solution than the standard conjugate gradient approach. We give an analytic estimate for the parameter that controls how gradually the linear system should change during the course of the solution.

Keywords: Computational methods — Cosmic microwave background radiation — Astronomy data reduction

1. INTRODUCTION

10

11

12

13

14

15

In observations of the Cosmic Microwave Background (CMB), map-making is an intermediate step between the collection of raw scanning data and the scientific analyses, such as the estimation of power spectra and cosmological parameters. Next generation CMB observations will generate much more data than those today, and so it is worth exploring efficient ways to process the data even though, on paper, the map-making problem has long been solved.

The time-ordered scanning data is summarized by

$$\mathbf{d} = P\mathbf{m} + \mathbf{n} \tag{1}$$

where **d**, **m**, and **n** are the vectors of time-ordered data (TOD), the CMB sky-map signal, and measurement noise. *P* is a sparse matrix in the time-by-pixel domain that encodes the telescope's pointing. Of several map-making methods (Tegmark 1997), one of the most common is the method introduced for the Cosmic Back-33 ground Explorer (COBE, Janssen & Gulkis 1992). This optimal, linear solution is

$$(P^{\dagger}N^{-1}P)\hat{\mathbf{m}} = P^{\dagger}N^{-1}\mathbf{d} \tag{2}$$

 $_{37}$ where $\hat{\mathbf{m}}$ provides the standard generalized least squares $_{38}$ minimization of the χ^2 statistic,

$$\chi^{2}(\mathbf{m}) \equiv (\mathbf{d} - P\mathbf{m})^{\dagger} N^{-1} (\mathbf{d} - P\mathbf{m}). \tag{3}$$

⁴¹ Here we assume that the noise has zero mean $\langle \mathbf{n} \rangle = \mathbf{0}$, ⁴² and the noise covariance matrix $N = \langle \mathbf{n} \mathbf{n}^{\dagger} \rangle$ is diago- ⁴³ nal in the frequency space. In the case where the noise ⁴⁴ is Gaussian, the COBE solution is also the maximum ⁴⁵ likelihood solution.

With current computational power, we cannot solve for $\hat{\mathbf{m}}$ by calculating $\left(P^{\dagger}N^{-1}P\right)^{-1}P^{\dagger}N^{-1}\mathbf{d}$ directly. The noise covariance matrix N is often sparse in the frequency domain and the pointing matrix P is sparse in the time-by-pixel domain. In experiments currently under design, there may be $\sim 10^{16}$ time samples and $\sim 10^9$ pixels, so these matrix inversions are intractable unless the covariance is uniform (proportional to the identity matrix I). We can use iterative methods, such as conjugate gradient descent, to avoid the matrix inversions, and execute each matrix multiplication in a basis where the matrix is sparse, using a fast Fourier transform to go between the frequency and time domain.

As an alternative to conjugate gradient descent, Huffor fenberger & Næss (2018) showed that the "messenger"
iterative method could be adapted to solve the linear
map-making system, based on the approach from Elsner & Wandelt (2013) to solve the linear Wiener filter. This technique splits the noise covariance into a
uniform part and the remainder, and introduces an additional vector that represent the signal plus uniform
noise. This messenger field acts as an intermediary between the signal and the data and has a covariance that
signal sconveniently sparse in every basis. Elsner & Wandelt

70 (2013) also introduced a cooling scheme that takes ad71 vantage of the split covariance: over the course of the
72 iterative solution, we adjust the relative weight of those
73 two parts. Starting with the uniform covariance, the
74 modified linear system gradually transforms to the final
75 system, under the control of a cooling parameter. In nu76 merical experiments, Huffenberger & Næss (2018) found
77 that a map produced by the cooled messenger method
78 converged significantly faster than for standard conju79 gate gradient methods, and to higher fidelity, especially
80 on large scales.

Papež et al. (2018) showed that the messenger field approach is equivalent to a fixed point iteration scheme, and studied its convergence properties in detail. Furthermore, they showed that the split covariance and the modified system that incorporates the cooling can be solved by other means, including a conjugate gradient technique, which should generally show better conversence properties than the fixed-point scheme. However in numerical tests, Papež et al. (2018) did not find benefits to the cooling modification of the map-making system, in contrast to the findings of Huffenberger & Næss (2018).

In our numerical experiments, we confirm the conclusion from Papež et al. (2018) that the conjugate gradient approach is converging faster than the fixed point iterations suggested by the messenger map-making method in Huffenberger & Næss (2018).

In this paper, we show that the difference arose because the numerical tests in Papež et al. (2018) used
much less low-frequency (or 1/f) noise than Huffenberger & Næss (2018), and show that the cooling
technique improves map-making performance especially
when the low-frequency noise is large. This performance
boost depends on a proper choice for the pace of cooling. Kodi Ramanah et al. (2017) showed that for Wiener
filter the cooling parameter should be chosen as a geometric series. In this work, we give an alternative interpretation of the parameterizing process and show that
for map-making the optimal choice (unsurprisingly) is
also a geometric series.

In Section 2 we describe our methods for treating the map-making equation and our numerical experiments. In Section 3 we present our results. In Section 4, we list our conclusions. In Appendix A we derive the prescription for our cooling schedule.

2. METHODS

116

117

2.1. Parameterized Conjugate Gradient Method

The messenger field approach introduced an extra cooling parameter λ to the map-making equation, and solved the linear system with an alternative parame-

121 terized covariance $N(\lambda) = \lambda \tau I + \bar{N}$. The paramele 122 ter $\tau = \min(\operatorname{diag}(N))$ represents the uniform level of 123 (white) noise in the original covariance. The remain- 124 der $\bar{N} \equiv N - \tau I$ is the non-uniform part of the original 125 noise covariance. (Here N without any arguments dele 126 notes the original noise covariance matrix $N = \langle \mathbf{n} \mathbf{n}^{\dagger} \rangle$.) 127 In this work we find it more convenient to work with 128 the reciprocal of cooling parameter $\eta = \lambda^{-1}$ which represents the degree of heteroscedasticity (non-uniformity) 130 in the parameterized covariance

$$N(\eta) = \tau I + \eta \bar{N},\tag{4}$$

132 -a choice that leads to the same system of map-making equation. (This is because $N(\eta) = \lambda^{-1}N(\lambda)$ and the map-making equation (2) is insensitive to scalar multiples of the covariance.) When $\eta = 1$ this parameter-136 ized covariance $N(\eta)$ equals N.

Papež et al. (2018) showed that the conjugate gradient method can be easily applied to the cooled map-making problem. In our notation, this is equivalent to iterating on the parameterized map-making equation

$$(P^{\dagger}N(\eta_i)^{-1}P)\,\hat{\mathbf{m}}(\eta_i) = P^{\dagger}N(\eta_i)^{-1}\mathbf{d},\tag{5}$$

141

143 as we adjust the parameter through a set of levels $\{\eta_i\}$.
144 (We use $\hat{\mathbf{m}}$ with no η argument to mean the estimated 145 $\hat{\mathbf{m}}$ in Eq. 2, independent of η .) This equation leads to 146 the same system as the parameterized equation in the 147 messenger field method, because $N(\eta) = \lambda^{-1}N(\lambda)$ and 148 the condition number does not change upon scalar mul-149 tiples of both sides of the equation. For concreteness we 150 fix the preconditioner to $M = P^{\dagger}P$ for all calculations.

In our numerical experiments, we confirm the conclusion from Papež et al. (2018) that the conjugate gradient approach is converging faster than the fixed point iterations suggested by the messenger map making method in Huffenberger & Næss (2018).

When $\eta = 0$, the noise covariance matrix N(0) is homoscedastic (uniform), and the solution is given by the simple binned map $\hat{\mathbf{m}}(0) = (P^{\dagger}P)^{-1}P^{\dagger}\mathbf{d}$, which can be solved directly.

Since the non-white part \bar{N} is the troublesome portion of the covariance, we can think of the η parameter as increasing the heteroscedasticity of the system, adding a perturbation to the solution achieved at a particular stage, building ultimately upon the initial uniform covariance model. Therefore, this quasi-static process requires η increase as $0 = \eta_0 \leq \eta_1 \leq \cdots \leq \eta_{\text{final}} = 1$, at which point we arrive at the desired map-making equation, and the solution $\hat{\mathbf{m}}(1) = \hat{\mathbf{m}}$.

We may iterate more than once at each intermediate η_i : we solve equation (5) with conjugate gradient iterations using the result from the previous calculation

 $\hat{\mathbf{m}}(\eta_{i-1})$ as the initial value. We move to next parameter η_{i+1} when the norm of residual vector

$$\|\mathbf{r}(\mathbf{m}, \eta_i)\| \equiv \|P^{\dagger} N(\eta_i)^{-1} P \mathbf{m} - P^{\dagger} N(\eta_i)^{-1} \mathbf{d}\|$$
 (6)

176 is an order of magnitude smaller than the norm of the 177 right hand side of Eq. 5.

$$\|\mathbf{r}(\mathbf{m}, \eta_i)\| < 0.1 \|P^{\dagger} N(\eta_i)^{-1} \mathbf{d}\|$$
 (7)

¹⁸⁰ This is not stringent enough to completely converge at ¹⁸¹ this η -level, but we find that it causes the system to ¹⁸² converge sufficiently to allow us to move on to the next ¹⁸³ η .

2.2. Analytical expression for $\{\eta_i\}$ series

184

197

198

210

The next question is how to appropriately choose these monotonically increasing parameters η . We also want to determine $\eta_1, \cdots, \eta_{n-1}$ before starting conjugate gradient entiterations, because the time ordered data \mathbf{d} is very large, and we do not want to keep it in the system memory during calculation or repeatedly read it in from disk. If we determine $\eta_1, \cdots, \eta_{n-1}$ before the iterations, then we can precompute the right-hand side of Eq. 5 for each η_i and keep these map-sized objects, instead of the entire time-ordered data.

In Appendix A, we show that a generic good choice for the η parameters is given by this geometric series

$$\eta_i = \min\left\{ \left(2^i - 1\right) \frac{\tau}{\max(\bar{N}_f)}, \ 1 \right\},\tag{8}$$

where \bar{N}_f are the eigenvalues of \bar{N} under frequency representation. This is one of our main results. It not only tells us how to choose parameters η_i , but also when we should stop the perturbation, and set $\eta=1$. For example, if the noise covariance matrix N is almost uniform, then $\bar{N}=N-\tau I\approx 0$, and we would have $\tau/\max(\bar{N}_f)>1$. This tell us that we don't need to use the parameterized method at all, because $\eta_0=0$ and $\eta_1=\eta_2=\cdots=1$. This corresponds to the standard conjugate gradient method with simple binned map as the initial guess (as recommended by Papež et al. 2018).

2.3. Intuitive Interpretation of η

Here is a way to interpret the role of η that is less technical than Appendix A. Our ultimate goal is to find $\hat{\mathbf{m}}(1)$ which minimizes $\chi^2(\mathbf{m})$ in Eq. 3. Since N is diagonal in the frequency space, χ^2 could be written as a sum of all frequency modes $|(\mathbf{d}-P\mathbf{m})_f|^2$ with weight N_f^{-1} , such as $\chi^2(\mathbf{m}) = \sum_f |(\mathbf{d}-P\mathbf{m})_f|^2 N_f^{-1}$. The weight is large for low-noise frequency modes (small N_f), and small for high-noise modes. Which means $\chi^2(\mathbf{m})$ would

²¹⁹ favor the low-noise modes, and therefore the conjugate ²²⁰ gradient map-making focuses on minimizing the error ²²¹ $\varepsilon \equiv \mathbf{d} - P\mathbf{m}$ in the low-noise part.

After introducing η , we minimize $\chi^2(\mathbf{m}, \eta)$ in Eq. A1 223 instead. For $\eta = 0$, $N^{-1}(0) \propto I$ the system is homoscedastic and the estimated map $\hat{\mathbf{m}}(0)$ does not pri-225 oritize any frequency modes. As we slowly increase η , 226 we decrease the weight for the high-noise modes, and 227 focusing minimizing error for the low-noise part. If we 228 start with $\eta_1=1$ directly, which corresponds to the 229 vanilla conjugate gradient method, then it will focus 230 most on minimizing the low-noise part, such that χ^2 231 would converge very fast on the low-noise modes (typi-232 cally high temporal frequencies and small spatial scales), 233 but slowly on the high-noise part (low frequencies and 234 large scales). The solver may get stuck at some local 235 minimum point in χ^2 and have trouble getting to the global minimum. (Question: if χ^2 is a quadratic form, 237 are there any local minima other than the global one?) 238 (Good point, I think this part is an error, and I'm not 239 sure why vanilla CG would be slow,.) However by in- $_{240}$ troducing the η parameter, we let the solver first treat every frequency equally, then as η slowly increases, it 242 gradually gives more focus to the lowest noise part.

2.4. Computational Cost

243

To properly compare the performance cost of this 244 245 method with respect to the vanilla conjugate gradi-246 ent method with the simple preconditioner, we need 247 to compare their computational cost at each itera-We could define $A(\eta) \equiv P^{\dagger}N(\eta)^{-1}P$ and $\mathbf{b}(\eta) \equiv P^{\dagger} N(\eta)^{-1} \mathbf{d}$, and equation 5 could be written 250 as $A(\eta_i)\hat{\mathbf{m}}(\eta_i) = \mathbf{b}(\eta_i)$. The right-hand side $\mathbf{b}(\eta_i)$ could 251 be computed before iterating, since we have determined $\{\eta_i\}$ in advance, so it will not introduce extra computa-253 tional cost. The most demanding part of conjugate gradient method is calculating its left hand side $A(\eta_i)\mathbf{m}$, 255 because it contains a Fourier transform of Pm from the 256 time domain to frequency domain and an inverse Fourier 257 transform of $N(\eta_i)^{-1}P\mathbf{m}$ from the frequency domain 258 back to time domain, which is order $\mathcal{O}(n \log n)$ with n 259 being the length of time ordered data. Compared to the 260 traditional conjugate gradient method, we swap N^{-1} with $N(\eta)^{-1}$, and the cost is the same for one step, 262 since both methods need a fast Fourier transform and 263 inverse fast Fourier transform at one iteration.

At each η_i level, we use the residual to determine whether to switch to the next level (η_{i+1}) , as is Equation (7). Calculation of the residual vector $\mathbf{r}(\mathbf{m}, \eta_i)$ is part of the conjugate gradient algorithm, so this will not add extra cost either. Therefore, overall introduc-

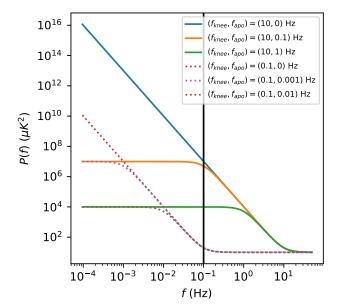


Figure 1. Noise power spectra that we use in our mapmaking simulations. These show a variety of low-frequency behavior, parameterized by Eq. 9, with white noise at high frequency and a low-frequency power-law slope $\alpha=3$. Here we show two knee frequencies, $f_{\rm knee}=10~{\rm Hz}$ (solid lines) and $f_{\rm knee}=0.1~{\rm Hz}$ (dashed lines). For each knee frequency, we have shown an unflattened spectrum ($f_{\rm apo}=0~{\rm Hz}$), and two flattened ones ($f_{\rm apo}=0.1f_{\rm knee}$ and $0.01f_{\rm knee}$). The vertical line shows our scanning frequency.

 $_{269}$ ing the η will not have extra computational cost within $_{270}$ the conjugate gradient iterations.

However, we start a new conjugate gradient algorithm whenever η_i updates to η_{i+1} . Thus we must re-initialize the conjugate gradient algorithm, re-calculating the residual $\mathbf{r}(\mathbf{m}, \eta_{i+1})$ based on new η_{i+1} . This residual calculation contains an extra $A(\eta_i)\mathbf{m}$ operation. Therefore, if we have a series $\eta_1, \eta_2, \eta_3, \cdots, \eta_{n_\eta}$, there will have $\eta_i = 1$ extra $\eta_i = 1$ extra $\eta_i = 1$ extra $\eta_i = 1$ extra to the traditional conjugate gradient method. If the total number of iterations is much larger than $\eta_i = 1$, then this extra cost is negligible. For our simulation, this extra step would have rather significant impact on final result. To have a fair comparison between the parameterized and traditional conjugate gradient method, we will present our results with number of $\eta_i = 1$ extra the parameterized and traditional conjugate gradient method, we will present our results with number of $\eta_i = 1$ extra the parameterized and traditional conjugate gradient method, we will present our results with number of $\eta_i = 1$ extra the parameterized and traditional conjugate gradient method, we will present our results with number of $\eta_i = 1$ extra the parameterized and traditional conjugate gradient method, we will present our results with number of $\eta_i = 1$ extra the parameterized and traditional conjugate gradient method, we will present our results with number of $\eta_i = 1$ extra the parameterized and traditional conjugate gradient method the parameterized and

2.5. Numerical Simulations

To compare these algorithms, we need to do some simple simulations of scanning processes, and generate the ²⁸⁹ time ordered data from a random sky signal. Our sky ²⁹⁰ is a small rectangular area, with two orthogonal directions x and y, both with range from -1° to $+1^{\circ}$. The ²⁹² signal has Stokes parameters (I,Q,U) for intensity and ²⁹³ linear polarization.

For the scanning process, our mock telescope contains nine detectors, each with different sensitivity to polarization Q and U. It scans the sky with a raster scanning pattern. Its scanning frequency is $f_{\rm scan}=0.1$ Hz and sampling frequency is $f_{\rm sample}=100$ Hz. The telescope scans the sky horizontally then vertically. This gives the noiseless signal s. The sky signal in the timestream has an rms a root-mean-square (RMS) of 56 μ K. The signal is continuous, so that it has structure on sub-pixel scales, but we find that our main conclusions remain the same when the input signal is pixelized. In map-making, we digitize the position (x,y) into 512×512 pixels.

We model the noise power spectrum with

307

324

330

$$P(f) = \sigma^2 \left(1 + \frac{f_{\text{knee}}^{\alpha} + f_{\text{apo}}^{\alpha}}{f^{\alpha} + f_{\text{apo}}^{\alpha}} \right)$$
(9)

which is white at high frequencies, a power law below the knee frequency, and gives us the option to flatten the low-frequency noise below an apodization frequency (like in Papež et al. 2018). Note that as $f_{\rm apo} \to 0$, $P(f) \to \sigma^2 (1 + (f/f_{\rm knee})^{-\alpha})$, and it becomes a 1/f-type noise model.

Dünner et al. (2013) measured the slopes of the atmospheric noise in the Atacama under different water vapor conditions, finding $\alpha=2.7$ to 2.9. Here we use $\sigma^2=10~\mu\mathrm{K}^2$, $\alpha=3$, and compare the performance under different noise models. In our calculations, we choose different combinations of f_{knee} and f_{apo} as in Figure 1. The noise spectra with the most low frequency noise have high f_{knee} or low cut-off f_{apo} .

The noise covariance matrix

$$N_{ff'} = P(f) \frac{\delta_{ff'}}{\Delta_f} \tag{10}$$

325 is a diagonal matrix in the frequency space, where Δ_f 326 is equal to the reciprocal of total scanning time $T\approx$ 327 1.05×10^4 seconds.

Finally, we get the simulated time ordered data $\mathbf{d} = \mathbf{s} + \mathbf{n}$ by adding up the signal and noise.

3. RESULTS

We compare the standard conjugate gradient method versus the conjugate gradient with our perturbed linear

¹ The source code and other information are available at https://github.com/Bai-Qiang/map_making_perturbative_approach

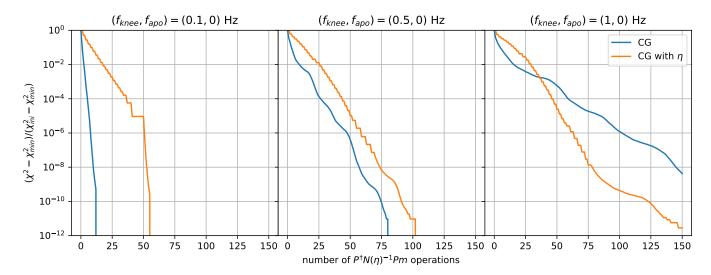


Figure 2. Convergence properties depend on the amount of low-frequency noise, which increases from the left panel to the right panel with increasing knee frequency. The map-making equation 2 minimize the $\chi^2(\mathbf{m})$, so the curve which falls fastest versus the number of operations is the preferred method. We compare the traditional conjugate gradient method ("CG," blue line) with the parameterized conjugate gradient method ("CG with η ," orange line) under different 1/f noise models (fixed $f_{\rm apo} = 0$ Hz but different $f_{\rm knee}$ in Eq. 9). When $f_{\rm knee} \gtrsim 10\,f_{\rm scan} = 1$ Hz, the significant amount of low-frequency noise causes the parameterized conjugate gradient method to start showing its advantage. The vertical axis is rescaled such that all curves start from 1.

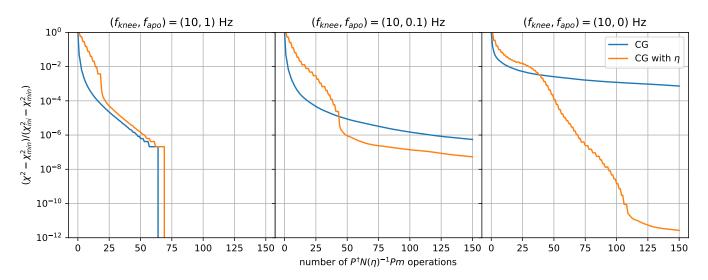


Figure 3. Like Figure 2, low-frequency noise increases from left to right, but by flattening the low-frequency noise at an apodization frequency. Low-frequency noise increases with decreasing apodization frequency (compare Figure 1). We again compare the traditional conjugate gradient method ("CG," blue line) with the parameterized conjugate gradient method ("CG with η ," orange line). When f_{apo} is much smaller than f_{knee} , there is a lot of low-frequency noise and the parameterized conjugate gradient method is better (ultimately falls faster) than the traditional one.

333 system. Both methods use the simple preconditioner $P^{\dagger}P$. Figure 2 shows the χ^2 results for 1/f noise mod335 els $(f_{\rm apo}=0)$ with different knee frequencies. Note that 336 the χ^2 values in all figures are calculated based on the 337 standard $\chi^2(\mathbf{m})$ in Eq. 3, not the η -dependent $\chi^2(\mathbf{m}, \eta)$ 338 of the modified system (Eq. A1). The minimum $\chi^2_{\rm min}$ 339 that we use for comparison is calculated from a delib-

340 erately slowed and well-converged parameterized conju-341 gate gradient method: one with 100 η values and that 342 halts when the final norm of the residual $\|\mathbf{r}(\mathbf{m}, 1)\|$ is 343 smaller than $10^{-5} \times \|P^{\dagger}N^{-1}\mathbf{d}\|$, or 100 iterations after 344 $\eta = 1$. From Figure 2, we can see for the 1/f noise 345 model, when $f_{\text{knee}} \gtrsim 10 f_{\text{scan}}$ the parameterized method ³⁴⁶ starts showing an advantage over the vanilla conjugate ³⁴⁷ gradient method.

In Figure 3, we fixed $f_{\text{knee}} = 10 \text{ Hz}$, and change f_{apo} .

As we decrease f_{apo} relative to f_{knee} , increasing the amount of low-frequency noise, the parameterized conjugate gradient method performs better.

Looking at the power spectrum in Figure 1, when $f_{\rm knee}$ is small, or $f_{\rm apo}$ is large, there is not much low-frequency noise. These situations corrrespond to the left-side plots in Figure 2 and Figure 3. The right-side graphs have significant amount of low frequency noise. We conclude that the introduction of the slowly-varying η parameter improves performance most when there are large low-frequency noise contributions.

We also tried different 1/f noise slopes α . For $\alpha = 2$, 360 the conclusion is the same as $\alpha = 3$. When $\alpha = 1$, the 362 low-frequency noise is reduced compared to the cases 363 with steeper slopes, and the vanilla conjugate gradient method is preferred, except some cases with very large 365 knee frequency like $f_{\rm knee}=100~{\rm Hz}$ and $f_{\rm apo}=0$ which favors the parameterized method. In Papež et al. (2018), $_{367}$ the slope α = 1 and the noise power spectrum is flat- $_{368}$ tened at $f_{
m apo} \approx 0.1 f_{
m knee}$. Their knee frequency is the 369 same as their scanning frequency, so is most like our 370 case when $f_{\rm knee} = f_{\rm scan} = 0.1$ Hz. Their case had little 371 low-frequency noise, and we confirm their specific result 372 that the standard conjugate gradient method converges 373 faster in that case. In general, however, we find cases 374 with significantly more low-frequency noise benefit from 375 the cooling/parameterized approach.

4. CONCLUSIONS

We analyzed the parameterized conjugate gradient map-making method that is inspired by the messenger-field idea of separating the white noise out of the noise covariance matrix. Then we gave an analytical expression for the series of η parameters that govern how quickly the modified covariance adjusts to the correct covariance, and showed that this method adds only the

extra computational cost of re-initializing the conjugate gradient process based on a new η parameter.

We tested this method for different noise power spectra, both flattened and non-flattened at low frequency. The results showed that the parameterized method is faster than the traditional conjugate gradient method when there is a significant amount of low-frequency noise. It could be further improved if we could get a more accurate estimation for the change in χ^2 as a function of the η parameter, either before iteration or without using time ordered data during iteration.

Also note that we fixed the preconditioner as $M=P^{\dagger}P$ during our calculation, this parameterizing process could be applied to any preconditioner and possibly improve performance when there is significant amount of low-frequency noise.

This type of analysis for the cooling parameter may also be apply to other areas, like the Wiener filter. Papež et al. (2018) showed that the messenger field method of Elsner & Wandelt (2013) for solving Wiener filter problem could also be written as a parameterized conjugate gradient algorithm. It stands to reason that such a system may also benefit from splitting and parameterizing its noise covariance, depending on the noise properties. (In the Wiener filter, Kodi Ramanah et al. (2017) additionally suggests the splitting the signal covariance and combining the uniform parts of the signal and noise.)

The benefits to map-making from a cooled messenger method seem to come from the cooling and not actually from the messenger field that inspired it. However, the messenger field approach may still have a role in the production and analysis of CMB maps. In particular, the close connection between the messenger method and Gibbs sampling may allow us to cheaply generate noise realizations of a converged map by generating samples from the map posterior distribution, something that we will continue to explore in future work.

⁴²¹ For this work, BQQ QBQ (just feel BQQ sounds like ⁴²² BBQ) and KMH are supported by NSF award 1815887.

APPENDIX

A. THE DERIVATION OF η PARAMETER SERIES

We know that the initial degree of heteroscedasticity $\eta_0=0$, which means the system is homoscedastic (uniform noise) to start. What would be a good value for the next parameter η_1 ? To simplify notation, we use N_η to denote the parameterized covariance matrix $N(\eta)=\tau I+\eta\bar{N}$. For some specific η value, the estimated map $\hat{\mathbf{n}}(\eta)=\left(P^\dagger N_n^{-1}P\right)^{-1}P^\dagger N_n^{-1}\mathbf{d}$ minimizes

$$\chi^{2}(\mathbf{m}, \eta) = (\mathbf{d} - P\mathbf{m})^{\dagger} N_{\eta}^{-1} (\mathbf{d} - P\mathbf{m}). \tag{A1}$$

376

423

424

with η being fixed. We restrict to the case that the noise covariance matrix N is diagonal in the frequency domain, and represent the frequency-domain eigenvalues as N_f .

The perturbative scheme works like this. We start with $\chi^2(\hat{\mathbf{m}}(\eta_0), \eta_0)$ with $\hat{\mathbf{m}}(\eta_0) = (P^{\dagger}P)^{-1}P^{\dagger}\mathbf{d}$ which could be solved directly. Then we use conjugate gradient method to find $\hat{\mathbf{m}}(\eta_1)$ and its corresponding $\chi^2(\hat{\mathbf{m}}(\eta_1), \eta_1)$. So let us consider $\eta_1 = \eta_0 + \delta \eta = \delta \eta$ such that $\eta_1 = \delta \eta$ is very small quantity, $\delta \eta \ll 1$. (Remember $\eta_0 = 0$.) Since $\hat{\mathbf{m}}(\eta)$ minimizes $\chi^2(\mathbf{m}, \eta)$ with η being fixed, we have $\frac{\partial}{\partial \hat{\mathbf{m}}} \chi^2(\hat{\mathbf{m}}(\eta), \eta) = 0$, and using the chain rule

$$\frac{\mathrm{d}}{\mathrm{d}\eta}\chi^{2}(\hat{\mathbf{m}}(\eta),\eta) = \frac{\partial}{\partial\eta}\chi^{2}(\hat{\mathbf{m}}(\eta),\eta) = -(\mathbf{d} - P\hat{\mathbf{m}}(\eta))^{\dagger}N_{\eta}^{-1}\bar{N}N_{\eta}^{-1}(\mathbf{d} - P\hat{\mathbf{m}}(\eta))$$
(A2)

Then the fractional decrease of $\chi^2(\mathbf{\hat{m}}(\eta_0),\eta_0)$ from η_0 to $\eta_1=\delta\eta$ is

437

451

457

462

$$-\frac{\delta \chi^{2}(\hat{\mathbf{m}}(\eta_{0}), \eta_{0})}{\chi^{2}(\hat{\mathbf{m}}(\eta_{0}), \eta_{0})} = -\delta \eta \frac{\frac{\mathrm{d}}{\mathrm{d}\eta} \chi^{2}(\hat{\mathbf{m}}(\eta_{0}), \eta_{0})}{\chi^{2}(\hat{\mathbf{m}}(\eta_{0}), \eta_{0})} = \delta \eta \frac{1}{\tau} \frac{(\mathbf{d} - P\hat{\mathbf{m}}(\eta_{0}))^{\dagger} \bar{N}(\mathbf{d} - P\hat{\mathbf{m}}(\eta_{0}))}{(\mathbf{d} - P\hat{\mathbf{m}}(\eta_{0}))^{\dagger}(\mathbf{d} - P\hat{\mathbf{m}}(\eta_{0}))}$$
(A3)

Here we put a minus sign in front of this expression such that it is non-negative, and use $N_{\eta=0}=\tau I$ at the second equality. We want $\left|\delta\chi^2(\hat{\mathbf{m}}(\eta_0),\eta_0)\right|=\chi^2(\hat{\mathbf{m}}(\eta_0),\eta_0))-\chi^2(\hat{\mathbf{m}}(\eta_1),\eta_1)$ to be large to encourage fast convergence. Therefore $\chi^2(\hat{\mathbf{m}}(\eta_1),\eta_1)$ is much smaller than $\chi^2(\hat{\mathbf{m}}(\eta_0),\eta_0)$, or $\chi^2(\hat{\mathbf{m}}(\eta_1),\eta_1)\ll\chi^2(\hat{\mathbf{m}}(\eta_0),\eta_0)$. Then we would expect

$$-\frac{\delta \chi^2(\hat{\mathbf{m}}(0), 0)}{\chi^2(\hat{\mathbf{m}}(0), 0)} = 1 - \frac{\chi^2(\hat{\mathbf{m}}(\eta_1), \eta_1)}{\chi^2(\hat{\mathbf{m}}(0), 0)} \approx 1^-$$
(A4)

Here we use the notation 1⁻ means the upper bound is close to but strictly smaller than 1. Now we could use Eq. A3 and let it equal to 1, then $\delta \eta = -\chi^2(\mathbf{\hat{m}}(\eta_0), \eta_0) / \frac{\mathrm{d}}{\mathrm{d}\eta} \chi^2(\mathbf{\hat{m}}(\eta_0), \eta_0)$.

Applying this same idea to $\eta_{m+1} = \eta_m + \delta \eta_m$ with $m \ge 1$, we would get

$$\delta \eta_m = -\chi^2(\hat{\mathbf{m}}(\eta_m), \eta_m) / \frac{\mathrm{d}}{\mathrm{d}\eta} \chi^2(\hat{\mathbf{m}}(\eta_m), \eta_m). \tag{A5}$$

⁴⁵³ As mentioned in the main text, we need to determine the entire series $\{\eta_i\}$ before conjugate gradient iterations. We do not have the $\hat{\mathbf{m}}(\eta_m)$ with which to calculate them and need to find another approach.

Let us go back to Eq. A3. Since we can't calculate $\mathbf{d} - P\hat{\mathbf{m}}(\eta_m)$, we treat it as an arbitrary vector, then the least upper bound of Eq. A3 is given by

$$-\frac{\delta \chi^2(\hat{\mathbf{m}}(\eta_0), \eta_0)}{\chi^2(\hat{\mathbf{m}}(\eta_0), \eta_0)} \le \frac{\delta \eta}{\tau} \max(\bar{N}_f)$$
(A6)

where $\max(\bar{N}_f)$ is the maximum eigenvalue of \bar{N} . We want $-\frac{\delta\chi^2(\hat{\mathbf{m}}(\eta_0),\eta_0)}{\chi^2(\hat{\mathbf{m}}(\eta_0),\eta_0)}$ to be as large as possible, but it won't exceed 1. If we combine Eq. A4 and Eq. A6, and choose $\delta\eta$ such that the least upper bound is equal to 1, to make sure the process would not going too fast. Thus we have

$$\eta_1 = \delta \eta = \frac{\tau}{\max(\bar{N}_f)} = \frac{\min(N_f)}{\max(N_f) - \min(N_f)}.$$
(A7)

Here N_f and \bar{N}_f are the eigenvalues of N and \bar{N} in the frequency domain. If the condition number of noise covariance matrix $\kappa(N) = \max(N_f)/\min(N_f) \gg 1$, then $\eta_1 \approx \kappa^{-1}(N)$.

What about the other parameters η_m with m > 1? We use a similar analysis, letting $\eta_{m+1} = \eta_m + \delta \eta_m$ with a small $\delta \eta_m \ll 1$. First, let us find the least upper bound

$$-\frac{\delta \chi^{2}(\hat{\mathbf{m}}(\eta_{m}), \eta_{m})}{\chi^{2}(\hat{\mathbf{m}}(\eta_{m}), \eta_{m})} = \delta \eta_{m} \frac{(\mathbf{d} - P\hat{\mathbf{m}}(\eta_{m}))^{\dagger} N_{\eta_{m}}^{-1} \bar{N} N_{\eta_{m}}^{-1} (\mathbf{d} - P\hat{\mathbf{m}}(\eta_{m}))}{(\mathbf{d} - P\hat{\mathbf{m}}(\eta_{m}))^{\dagger} N_{\eta_{m}}^{-1} (\mathbf{d} - P\hat{\mathbf{m}}(\eta_{m}))}$$
(A8)

$$\leq \delta \eta_m \, \max \left(\frac{\bar{N}_f}{\tau + \eta_m \bar{N}_f} \right) \tag{A9}$$

474

479

490 491

The upper bound in the second line is a little bit tricky. Both matrix \bar{N} and $N_{\eta_m}^{-1}$ can be simultaneously diagonalized in the frequency space. For each eigenvector \mathbf{e}_f , the corresponding eigenvalue of the matrix on the numerator $N_{\eta_m}^{-1}\bar{N}N_{\eta_m}^{-1}$ is $\lambda_f = \bar{N}_f(\tau + \eta_m\bar{N}_f)^{-2}$, and the eigenvalue for the matrix in the denominator $N_{\eta_m}^{-1}$ is $\gamma_f = (\tau + \eta_m\bar{N}_f)^{-1}$. Their eigenvalues are related by $\lambda_f = [\bar{N}_f/(\tau + \eta_m\bar{N}_f)]\gamma_f$. For any vector $\mathbf{v} = \sum_f \alpha_f \mathbf{e}_f$, we have

$$\frac{\mathbf{v}^{\dagger} N_{\eta_m}^{-1} \bar{N} N_{\eta_m}^{-1} \mathbf{v}}{\mathbf{v}^{\dagger} N_{\eta_m}^{-1} \mathbf{v}} = \frac{\sum_f \alpha_f^2 \lambda_f}{\sum_f \alpha_f^2 \gamma_f} = \frac{\sum_f \alpha_f^2 \gamma_f \bar{N}_f / (\tau + \eta_m \bar{N}_f)}{\sum_f \alpha_f^2 \gamma_f} \le \max\left(\frac{\bar{N}_f}{\tau + \eta_m \bar{N}_f}\right). \tag{A10}$$

Again assuming $\chi^2(\hat{\mathbf{m}}(\eta_{m+1}), \eta_{m+1}) \ll \chi^2(\hat{\mathbf{m}}(\eta_m), \eta_m)$, which we expect it to be satisfied for $\eta_m \ll 1$. That is because if $\eta \lesssim 1$, $\chi^2(\hat{\mathbf{m}}(\eta), \eta)$ would close to the minimum χ^2 which means $\chi^2(\hat{\mathbf{m}}(\eta_{m+1}), \eta_{m+1}) \lesssim \chi^2(\hat{\mathbf{m}}(\eta_m), \eta_m)$, which would violate our assumption. Luckily, the final result (Eq. A14) is a geometric series, only the last few η_m values fail to satisfy this condition. Similarly, we could set the least upper bound equal to 1. Then we get

$$\delta \eta_m = \min\left(\frac{\tau + \eta_m \bar{N}_f}{\bar{N}_f}\right) = \eta_m + \frac{\tau}{\max(\bar{N}_f)}.$$
(A11)

481 Therefore

$$\eta_{m+1} = \eta_m + \delta \eta_m = 2\eta_m + \frac{\tau}{\max(\bar{N}_f)}$$
(A12)

484 If written in the form $\eta_{m+1} + \tau/\max(\bar{N}_f) = 2(\eta_m + \tau/\max(\bar{N}_f))$ it's easy to see that for $m \ge 1$, $\eta_m + \tau/\max(\bar{N}_f)$ 485 forms a geometric series

$$\eta_m + \frac{\tau}{\max(\bar{N}_f)} = \left(\eta_1 + \frac{\tau}{\max(\bar{N}_f)}\right) 2^{m-1} = \frac{\tau}{\max(\bar{N}_f)} 2^m \tag{A13}$$

where we used $\eta_1 = \tau/\max(\bar{N}_f)$. Note that m = 0 and $\eta_0 = 0$ also satisfy this expression and we've got final expression and η_m

$$\eta_m = \min\left\{1, \ \frac{\tau}{\max(\bar{N}_f)}(2^m - 1)\right\} \tag{A14}$$

⁴⁹² Here we need to truncate the series when $\eta_m > 1$.

In numerical simulations, we find that if we update the η parameter based on the more precise Eq. A5, there is only marginally improvements over the η series given in Eq. A14 for the 1/f noise model, and slight improvements when there is not much low frequency noise. If we use Eq. A5, it ends up with fewer η parameters in the series, but the interval between η_i and η_{i+1} gets larger. In our simulation this sometimes causes one more iteration at certain η value, so in the end there is only slight improvements. For large data set that need lots of iterations to converge from η_i to η_{i+1} , where several extra iterations may not be significant, using Eq. A5 may provide a larger performance boost.

REFERENCES

Dünner, R., Hasselfield, M., Marriage, T. A., et al. 2013,
ApJ, 762, 10, doi: 10.1088/0004-637X/762/1/10
Elsner, F., & Wandelt, B. D. 2013, A&A, 549, A111,
doi: 10.1051/0004-6361/201220586
Huffenberger, K. M., & Næss, S. K. 2018, The
Astrophysical Journal, 852, 92,
doi: 10.3847/1538-4357/aa9c7d
Janssen, M. A., & Gulkis, S. 1992, in NATO Advanced
Science Institutes (ASI) Series C, ed. M. Signore &

C. Dupraz, Vol. 359 (Springer), 391–408

509 Kodi Ramanah, D., Lavaux, G., & Wandelt, B. D. 2017,

510 MNRAS, 468, 1782, doi: 10.1093/mnras/stx527

511 Papež, J., Grigori, L., & Stompor, R. 2018, A&A, 620, A59,

doi: 10.1051/0004-6361/201832987

513 Tegmark, M. 1997, ApJL, 480, L87, doi: 10.1086/310631