Cosmic Microwave Background map-making solutions improve with cooling

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ABSTRACT

In the context of Cosmic Microwave Background data analysis, we study the solution to the equation that transforms scanning data into a map. As originally suggested in "messenger" methods for solving linear systems, we split the noise covariance into uniform and non-uniform parts and adjusting their relative weight during the iterative solution. This "cooling" or perturbative approach is particularly effective when there is significant low-frequency noise in the timestream. A conjugate gradient algorithm applied to this modified system converges faster and to a higher fidelity solution than the standard conjugate gradient approach, for the same computational cost per iteration. We conclude that cooling is helpful separate from its appearance in the messenger methods. We give an analytical expression for the parameter that controls how gradually should change during the course of the solution.

Keywords: Computational methods — Cosmic microwave background radiation — Astronomy data reduction

1. INTRODUCTION

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In observations of the Cosmic Microwave Background (CMB), map-making is an intermediate step between the collection of raw scanning data and the scientific analyses, such as the estimation of power spectra and cosmological parameters. Next generation CMB observations will generate much more data that today, and so it is worth exploring efficient ways to process the data, even though, on paper, the map-making problem has long been solved.

The time-ordered scanning data is summarized by

$$\mathbf{d} = P\mathbf{m} + \mathbf{n} \tag{1}$$

where **d**, **m**, and **n** are the vectors of time-ordered data (TOD), the CMB sky-map signal, and measurement noise, and P is the sparse matrix that encodes the telescope's pointing. Of several mapmaking methods (Tegmark 1997a), one of the most common is the method introduced for the Cosmic Background Explorer (COBE, Janssen & Gulkis 1992). This optimal, linear solution is

$$(P^{\dagger}N^{-1}P)\hat{\mathbf{m}} = P^{\dagger}N^{-1}\mathbf{d}$$
 (2)

 $_{\rm 37}$ where ${\bf \hat{m}}$ provides the generalized least squares mini- $_{\rm 38}$ mization of the χ^2 statistic

$$\chi^{2}(\mathbf{m}) \equiv (\mathbf{d} - P\mathbf{m})^{\dagger} N^{-1} (\mathbf{d} - P\mathbf{m}). \tag{3}$$

⁴¹ Here we assume that the noise has zero mean $\langle \mathbf{n} \rangle = \mathbf{0}$, ⁴² and noise covariance matrix could be written as N = ⁴³ $\langle \mathbf{n} \mathbf{n}^{\dagger} \rangle$. We cast mapmaking as a standard linear regres- ⁴⁴ sion problem. In case the noise is Gaussian, the COBE ⁴⁵ solution is also the maximum likelihood solution.

With current computation power, we cannot solve for 47 $\hat{\mathbf{m}}$ by calculating $(P^{\dagger}N^{-1}P)^{-1}P^{\dagger}N^{-1}\mathbf{d}$ directly, since 48 the $(P^{\dagger}N^{-1}P)$ matrix is too large to invert. The noise 49 covariance matrix N is sparse in frequency domain and 50 the pointing matrix P is sparse in the time-by-pixel do-51 main, and their product is dense. In experiments cur- $_{52}$ rently under design, there may be $\sim 10^{16}$ time sam-₅₃ ples and $\sim 10^9$ pixels, so these matrix inversions are 54 intractable. Therefore we use iterative methods, such 55 as conjugate gradient descent, to avoid the matrix in-56 versions, while executing each matrix multiplication in 57 a basis where the matrix is sparse, using a fast Fourier 58 transform to go between the frequency and time domain. As an alternative technique, Huffenberger & Næss 60 (2018) showed that the "messenger method" could be 61 adapted to solve the linear mapmaking system, based 62 on the approach from Elsner & Wandelt (2013) to solve 63 the linear Wiener filter. This technique splits the noise 64 covariance into a uniform part and the remainder, and, 65 over the course of the iterative solution, it adjusts the 66 relative weight of those two parts. Starting with the 67 uniform covariance, the modified linear system gradu-68 ally transforms to the final system via a cooling param-69 eter. The cooling idea again comes from Elsner & Wan-

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70 delt (2013). In numerical experiments, Huffenberger & 71 Næss (2018) found that the large scales of map produced 72 by the cooled messenger method converged significantly ₇₃ faster than for standard methods, and to higher fidelity. Papež et al. (2018) showed that the splitting of the co-75 variance in the messenger field approach is equivalent to 76 a fixed point iteration scheme, and studied its conver-77 gence properties in detail. Furthermore, they showed 78 that the modified system that incorporates the cool-79 ing scheme can be solved by other means, including 80 a conjugate gradient technique, which should generally 81 show better convergence properties than the fixed-point 82 scheme. However in numerical tests, Papež et al. (2018) 83 did not find benefits to the cooling modification of the 84 mapmaking system, in contrast to findings of Huffen-85 berger & Næss (2018).

In this paper, we show that the difference arose begraph cause the tests in Papež et al. (2018) used much less low-frequency (1/f) noise, and show that the cooling graph technique improves mapmaking performance especially when the low frequency noise is large. This performance boost depends on a proper choice for the pace of coolgraphical ing. Kodi Ramanah et al. (2017) showed that for Wiener filter the cooling parameter should be chosen as a geometric series. In this work, we give an alternative interpretation of the parameterizing process and show that for map-making the optimal choice (unsurprisingly) is also a geometric series.

In Section 2 we describe our methods for treating the mapmaking equation and our numerical experiments. In Section 3 we present our results. In Section 4 we interpret the mapmaking approach and its computational cost. In Section 5 we conclude. In appendices we derives how we set our cooling schedule.

2. METHODS

2.1. Parameterized Conjugate Gradient Method

The messenger field approach introduced an extra cooling parameter λ to the map-making equation, and solved the linear system with the alternative covariance $N(\lambda) = \lambda \tau I + \bar{N}$. The parameter τ represents the unition form level of (white) noise in the covariance, \bar{N} is the balance of the noise, and the parameterized covariance equals the original covariance when the cooling parameter $\lambda = 1$. In this work we find it more convenient to work with the inverse cooling parameter $\eta = \lambda^{-1}$ and define the covariance as

$$N(\eta) = \tau I + \eta \bar{N} \tag{4}$$

which leads to the same system of mapmaking equations. (This is because $N(\eta) = \lambda^{-1} N(\lambda)$ and the mapmaking

equation is insensitive to to scalar multiple of the covariance since is appears on both sides.)

Papež et al. (2018) showed that the conjugate gradi-122 ent method can be easily applied to parameterized map-123 making equation by iterating on

$$P^{\dagger}N(\eta)^{-1}P \hat{\mathbf{m}} = P^{\dagger}N(\eta)^{-1}\mathbf{d}$$
 (5)

126 as the cooling is adjusted. In our numerical experiments, 127 we confirm that the conjugate gradient approach is con128 verging faster than the fixed point iterations suggested 129 by the messenger mapmaking method in Huffenberger 130 & Næss (2018). For simplicity we fix the preconditioner 131 to $M = P^{\dagger}P$ for all of calculations. For some inter132 mediate η_i , we use the conjugate gradient method to 133 solve equation $\left(P^{\dagger}N(\eta_i)^{-1}P\right)\hat{\mathbf{m}}(\eta_i) = P^{\dagger}N(\eta_i)^{-1}\mathbf{d}$, us134 ing $\hat{\mathbf{m}}(\eta_{i-1})$ as the initial value. KMH: In this description, it is not totally clear whether you intend to update 136 the eta after every iteration.

When $\eta=0$, the noise covariance matrix N(0) is proportional to identity matrix I, and solution is given by simple binned map $\mathbf{m}_0 = (P^\dagger P)^{-1} P^\dagger \mathbf{d}$, which can be solved directly. From this starting point, the cooling scheme requires the inverse cooling parameter η increase as $0=\eta_0 \leq \eta_1 \leq \cdots \leq \eta_{\text{final}}=1$, at which point we arrive at the desired mapmaking equation.

The non-white part N is the troublesome portion of the covariance, and we can think of the η parameter as turning it on slowly, adding a perturbation to the solution achieved at a particular stage, building ultimately upon the initial uniform covariance model.

2.2. Choice of inverse cooling parameters η

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The next question is how we choose these monotonisi ically increasing parameters η . If we choose them inappropriately, the solution converge slowly, because we
waste effort converging on the wrong system. We also
want to determine $\eta_1, \cdots, \eta_{n-1}$ before starting conjuising gate gradient iterations. The time ordered data \mathbf{d} is very
large, and we do not want to keep it in the system memiron ory during calculation. If we determine $\eta_1, \cdots, \eta_{n-1}$ before the iterations, then we can precompute the righthand side $P^{\dagger}N(\eta)^{-1}\mathbf{d}$ for each η_i and keep these mapising sized objects in memory, instead of the entire timeordered data.

In the appendix, we show that a generic good choice for the η parameters are the geometric series

$$\eta_i = \min\left\{ \left(2^i - 1\right) \frac{\tau}{\max(\bar{N}_f)}, \ 1 \right\},\tag{6}$$

where \bar{N}_f is the frequency representation of the non-167 uniform part of the covariance. This is the main result. It tells us not only how to choose parameters η_i , but also when we should stop the perturbation, and set $\eta=170$ 1. For example, if noise covariance matrix N is almost white noise, then $\bar{N}=N-\tau I\approx 0$, and we would have $\tau=170$ $\tau=$

2.3. Numerical Simulations

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To compare these algorithms, we need to do some simple simulation of scanning processes, and generate time ordered data from random sky signal. Our sky is a small rectangular area, with two orthogonal directions x and y, both with range from -1° to $+1^{\circ}$. The signal has first three stokes parameters (I, Q, U).

For the scanning process, our single telescope contains nine detectors, each has different sensitivity to polarization Q and U. It scans the sky with a raster scanning pattern and scanning frequency $f_{\rm scan}=0.1$ Hz sampling frequency $f_{\rm sample}=100$ Hz. The telescope scans the sky horizontally and then vertically, and then digitizes the position (x,y) into 512×512 pixel. This gives noiseless signal s.

The noise power spectrum is given by

$$P(f) = \sigma^2 \left(1 + \frac{f_{\text{knee}}^{\alpha} + f_{\text{apo}}^{\alpha}}{f^{\alpha} + f_{\text{apo}}^{\alpha}} \right)$$
 (7)

¹⁹⁵ Here we fixed $\sigma^2=10~\mu\text{K}^2$, $\alpha=2$ and $f_{\text{knee}}=10$ Hz, and change f_{apo} to compare the performance under ¹⁹⁷ different noise models. Note that as $f_{\text{apo}} \to 0$, $P(f) \to$ ¹⁹⁸ $\sigma^2 \left(1 + (f/f_{\text{knee}})^{-1}\right)$, it becomes a 1/f noise model. The ¹⁹⁹ noise covariance matrix

$$N_{ff'} = P(f) \frac{\delta_{ff'}}{\Delta_f} \tag{8}$$

 $_{201}$ is a diagonal matrix in frequency space, where Δ_f is $_{202}$ equal to reciprocal of total scanning time T. In our $_{203}$ calculations we choose the $f_{\rm apo}$ such that the condition numbers κ are 10^2 , 10^6 , and 10^{12} . The corresponding power spectrum are shown in Figure(1).

Finally, we get the simulated time ordered data $\mathbf{d} = \mathbf{s} + \mathbf{n}$ by adding up signal and noise.

KMH: Compare to the noise power spectrum of Papez. Remark how little 1/f is in their test. What is the effect of changing the noise slope?

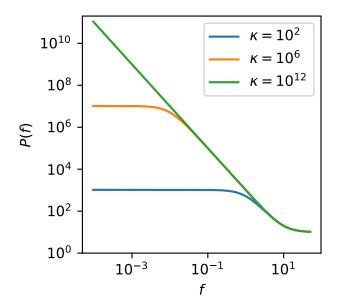


Figure 1. The noise power spectum based on Eq. (7) with $\sigma^2 = 10 \ \mu\text{K}^2$, $\alpha = 2$ and $f_{\text{knee}} = 10 \ \text{Hz}$. And fixing the condition number κ of noise covariance matrix Eq. (8) by choosing f_{apo} . KMH: show the scanning frequency with a vertical (dashed?) line. Can use axyline().

3. RESULTS

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First let's compare the results with vanilla conjugate gradient method with simple preconditioner $P^{\dagger}P$. The results are showed in Figure (2) for different kinds of noise power spectra. Here note that χ^2 in all figures are calculated based on Eq. (3) not $\chi^2(\mathbf{m},\eta)$ in Eq. (??). The χ^2_{\min} is calculated from perturbative conjugate gradient method with more intermediate η values, and more iterations after $\eta=1$.

As we can see in the left graph in Figure(2), when the condition number of noise covariance matrix $\kappa(N)$ is small, the performance between different these two methods are small. The vanilla conjugate gradient method converge faster, because its perturbation parameter goes to 1 at the first iteration, however for the perturbation method its η value will slowly reach 1 in about ten iterations.

Notice that as we increase $\kappa(N)$, or equivalently decrease $f_{\rm apo}$, the perturbation parameter η starts showing its benefits, as showed in the second and third graph in Figure(2). It outperforms the vanilla conjugate gradient method when $f_{\rm apo}\approx 0$ and the noise power spectrum becomes the 1/f noise model, which usually is the intrinsic noise of instruments (Tegmark (1997b)).

Now let us compare the performance difference between choosing η parameters based on Eq. (A6) and manually fixing number of η parameters n_{η} manually. We manually choose the η_i values us-

¹ The source code and other information are available at https://github.com/Bai-Qiang/map_making_perturbative_approach

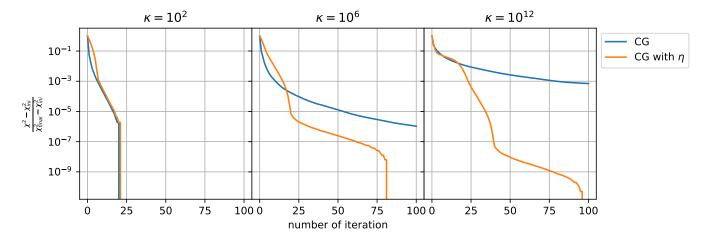


Figure 2. These three figures show the $\frac{\chi^2(\mathbf{m}) - \chi^2_{\text{ini}}}{\chi^2_{\text{min}} - \chi^2_{\text{ini}}}$ changes for each iteration under different noise covariance matrix with condition number being 10^2 , 10^6 , and 10^{12} .

when $\kappa(N)$ is small, and Eq. (A6) tells us that only a few η parameters are good enough, see the orange line in the first Figure(3), where we have $\sim 10~\eta$ levels. If unfortunately we choose n_{η} being large value, like 15 or 30, then it will ends up converge slowly, because it needs at least 15 or 30 iterations to reach $\eta=1$, at least one iteration per η level.

On the other hand if $\kappa(N)$ is very large and the power spectrum is 1/f noise, we need more η parameters. If n_{η} is too small, for example $n_{\eta}=5$ the green line in last Figure(3), it may be better than the vanilla conjugate gradient method, but it is still far from optimal.

Since the η values determined from Eq. (A6)

$$\eta_i = \min\left\{1, \ \frac{\tau}{\max(\bar{N}_f)} \left(2^i - 1\right)\right\} \tag{A6}$$

²⁵⁷ are not dependent on any scanning information, it only depends on noise power spectrum P(f), or noise covariance matrix N. Figure (4) and Figure (5) show two examples with same parameters as in Figure (3) except for the scanning frequency $f_{\rm scan}$ (also we need to change $f_{\rm apo}$ to fix condition number), KMH: I'm confused because scanning frequency does not affect the condition number of $f_{\rm apo}$ in Figure (4) it scans very slow and in Figure (5) it's very fast. In these two cases under $f_{\rm apo}$ to hoise model, our $f_{\rm apo}$ values based on Eq. (A6) are better than manually selected values. Based on these two results we know, the $f_{\rm apo}$ values should somehow depends on scanning scheme.

4. DISCUSSION

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4.1. Intuitive Interpretation of η

KMH: most of this is pretty similar to discussion in Huffenberger and Naess. The last paragraph is new.

In this section, let me introduce another way to understand the role of η . Our ultimate goal is to find $\hat{\mathbf{m}}(\eta=1)$ which minimizes $\chi^2(\mathbf{m})=(\mathbf{d}-P\mathbf{m})^\dagger N^{-1}(\mathbf{d}-P\mathbf{m})$. Since N is diagonal in frequency space, χ^2 could be written as a sum of all frequency mode $|(\mathbf{d}-P\mathbf{m})_f|^2$ with weight N_f^{-1} , such as $\chi^2(\mathbf{m})=\sum_f|(\mathbf{d}-P\mathbf{m})_f|^2N_f^{-1}$. Now the same when there is little noise at that frequency, and vice versa. Which means $\chi^2(\mathbf{m})$ would favor the low noise frequency mode over high noise ones. In other words the optimal map $\hat{\mathbf{m}}$ focusing on minimize the eresult of \mathbf{r} and \mathbf{r} in the low-noise part.

After introducing η , we minimize $\chi^2(\mathbf{m},\eta)=(\mathbf{d}-P\mathbf{m})^\dagger N_\eta^{-1}(\mathbf{d}-P\mathbf{m})$. For $\eta=0$, $N_{\eta=0}^{-1}\propto I$ and the estimated map $\hat{\mathbf{m}}(\eta=0)$ does not prioritize any frequency mode. As we slowly increase η , we decrease the weight for the frequency modes which have large noise, and focusing minimizing error for low noise part. If we start with $\eta_1=1$ directly, which corresponds to the vanilla conjugate gradient method, then the entire conjugate gradient solver will focus most on minimizing the low noise part, such that χ^2 would converge very fast at low noise region, but slowly on high noise part. However by introducing η parameter, we let the solver first treat every frequency equally. Then as η slowly increases, it gradually shifts focus from the highest noise to the lowest noise part. KMH: I feel what this is missing is why the high-noise modes get stuck though.

If we write the difference between final and initial χ^2 value as $\chi^2(\hat{\mathbf{m}}(1),1) - \chi^2(\hat{\mathbf{m}}(0),0) = \frac{1}{2}$ $\int_0^1 \mathrm{d}\eta \, \frac{\mathrm{d}}{\mathrm{d}\eta} \chi^2(\hat{\mathbf{m}}(\eta),\eta)$, and use Eq. (B7). We note that when η is very small, the $\frac{\mathrm{d}}{\mathrm{d}\eta} \chi^2(\hat{\mathbf{m}}(\eta),\eta)$ would have relatively large contribution from medium to large noise region, comparing to large η . So introducing η might

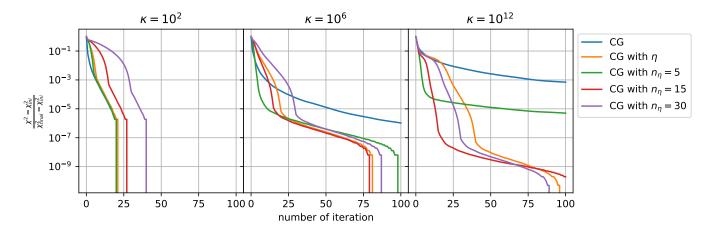


Figure 3. The blue line and the orange line are the same as Figure (2). For three extra lines, we fix the number of η parameter n_{η} manully. Instead of using Eq. (A6), we use numpy.logspace(start=ln(η_1), stop=0, num= n_{η} , base=e). to get all η parameters. KMH: In these figures you use χ^2_{final} as the label, but I think "min" is better than "final". You used "min" in Figure (2).

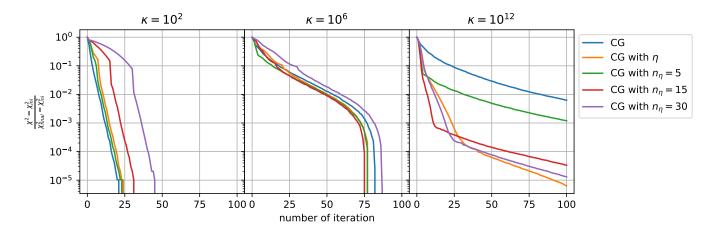


Figure 4. In this case all parameters are the same as Figure (3) except $f_{\text{scan}} = 0.001$, and corresponding f_{apo} to fix the condition number. KMH: I'm mixed up about what is changing here. If you change the scan frequency only, I don't see why the apodization should have to change. Since this has the scan slower than the knee, this seems like a lot like fig 3, although the convergence is slower it seems.

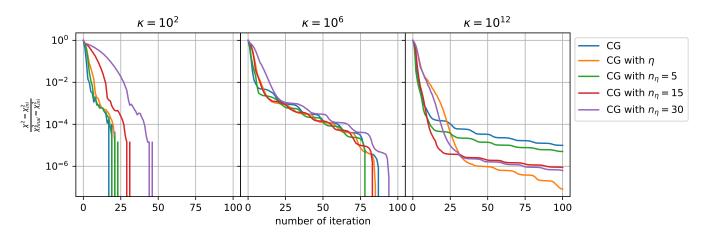


Figure 5. In this case all parameters are the same as Figure (3) except $f_{\text{scan}} = 10$, and corresponding f_{apo} to fix the condition number. KMH: I wonder what a scan faster than the knee would make.

 $_{307}$ improve the convergence of χ^2 at these regions, because $_{308}$ the vanilla conjugate gradient method only focuses on

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309 the low noise part and it may have difficulty at these 310 regions.

4.2. Computational Cost To properly compare the performance cost of this

method with respect to vanilla conjugate gradient method with simple preconditioner, we need to com-

315 pare their computational cost at each iteration. The 316 right hand side of parameterized map-making equation Eq. (5) could be computed before iterations, so it won't 318 introduce extra computational cost. The most demand-319 ing part of conjugate gradient method is calculating $P^{\dagger}N^{-1}P\hat{\mathbf{m}}$, because it contains a Fourier transform of $P\hat{\mathbf{m}}$ from time domain to frequency domain and an in-322 verse Fourier transform of $N^{-1}P\hat{\mathbf{n}}$ from frequency domain back to time domain, which is order $\mathcal{O}(n \log n)$ n with n being the length of time ordered data. If we s25 change N^{-1} to $N(\eta)^{-1}$, it won't add extra cost, since 326 both matrices are diagonal in frequency domain. There-327 fore the computational cost it the same for one step. However our previous analysis is based $\chi^2(\hat{\mathbf{m}}(\eta_i), \eta_i)$ which is evaluated at $\hat{\mathbf{m}}(\eta_i)$ the esti-330 mated map at η_i . So We should update η_i to η_{i+1} 331 when $\mathbf{m} \approx \hat{\mathbf{m}}(\eta_i)$. How do we know this condition is 332 satisfied? Since for each new η_i value, we are solving 333 a new set of linear equations $A(\eta_i)\hat{\mathbf{m}} = \mathbf{b}(\eta_i)$ with 334 $A(\eta_i) = P^{\dagger} N(\eta_i)^{-1} P$ and $\mathbf{b}(\eta_i) = P^{\dagger} N(\eta_i)^{-1} \mathbf{d}$, and we 335 could stop calculation and moving to next value η_{i+1} when the norm of residual $||\mathbf{r}(\eta_i)|| = ||\mathbf{b}(\eta_i) - A(\eta_i)\mathbf{m}||$ 337 smaller than some small value. Calculate $||\mathbf{r}(\eta_i)||$ is 338 part of conjugate gradient algorithm, so this won't 339 add extra cost compare to vanilla conjugate gradient 340 method. Therefore, overall introducing η won't have 341 extra computational cost.

5. CONCLUSIONS

KMH: some of this future prospects should move to discussion As you may have noticed in the second and third Figure(3), the perturbation parameter based on Eq. (A6) is more than needed, especially for 1/f noise are case. For the case $\kappa=10^{12}$, we notice that based on Eq. (A6) it gives us $n_{\eta}\approx 40$, however from χ^2 result in the last Figure(3) $n_{\eta}\approx 30$ or even $n_{\eta}\approx 15$ is good enough. Also, for the nearly-white-noise case, we could certainly choose $n_{\eta}=1$ such that $\eta_1=1$ which corresponds to vanilla conjugate gradient method, based on χ^2 result in first Figure(3). However Eq. (A6) gives us $\eta_{\eta}\approx 6$, even though it does not make the final χ^2 result much different at the end.

Is it possible to further improve the analysis, such that it produces smaller n_{η} ? Let's examine how we get η_i series. Remember that we determine $\delta \eta$ value based

on the upper bound of $-\delta \chi^2(\hat{\mathbf{m}}(\eta), \eta)/\chi^2(\hat{\mathbf{m}}(\eta), \eta)$, in Eq. (A2). For $\eta \neq 0$, the upper bound is

$$\delta \eta \frac{\hat{\mathbf{r}}_{\eta}^{\dagger} N(\eta)^{-1} \bar{N} N(\eta)^{-1} \hat{\mathbf{r}}_{\eta}}{\hat{\mathbf{r}}_{\eta}^{\dagger} N(\eta)^{-1} \hat{\mathbf{r}}_{\eta}} \le \frac{\delta \eta}{\eta + \frac{\tau}{\max(N_f) - \tau}}$$
(9)

363 with $\mathbf{r}_{\eta} = \left[1 - P(P^{\dagger}N(\eta)^{-1}P)^{-1}P^{\dagger}N(\eta)^{-1}\right]\mathbf{d} \equiv \mathcal{P}_{\eta}\mathbf{d}.$ To get the upper bound we treated $\mathbf{d} - P\hat{\mathbf{m}}(\eta)$ as an ar-365 bitrary vector in frequency domain, since we don't know 366 how to calculate \mathcal{P}_{η} for $\eta \neq 0$, and it's hard to analyze 367 the projection matrix \mathcal{P}_{η} in frequency space, as it con-368 tains $(P^{\dagger}N(\eta)^{-1}P)^{-1}$. Note that we have to determine η all of η value before calculation, because we don't want 370 to keep the time ordered data in system RAM, so we 371 need to somehow analytically analyze \mathcal{P}_{η} , and its be- $_{372}$ havior in frequency space. Unless \mathbf{r}_{η} almost only has 373 large noise modes, $\left|\frac{\mathrm{d}}{\mathrm{d}\eta}\chi^2(\hat{\mathbf{m}}(\eta),\eta)/\chi^2(\hat{\mathbf{m}}(\eta),\eta)\right|$ won't 374 get close to the upper bound $1/\left(\eta + \frac{\tau}{\max(N_f) - \tau}\right)$. Based 375 on the analysis in Section(4.1), for small η the esti-376 mated map $\hat{\mathbf{m}}(\eta)$ does not only focusing on minimiz- \mathbf{r}_n ing error \mathbf{r}_n at low noise region. So we would expect 378 that there would be a fair amount of low noise modes $_{379}$ contribution in \mathbf{r}_{η} especially for the first few η values. 380 Which means if we could somehow know the frequency \mathbf{r}_{η} distribution of \mathbf{r}_{η} , we could tighten the boundary of $\frac{1}{282} \left| \frac{\mathrm{d}}{\mathrm{d}\eta} \chi^2(\hat{\mathbf{m}}(\eta), \eta) / \chi^2(\hat{\mathbf{m}}(\eta), \eta) \right|$, and get larger $\delta \eta$ value. This should make η goes to 1 faster, and yields the fewer 384 η parameters we need.

Also notice that the η values determined from Eq. (A6) are not dependent on any scanning informaion, it only depends on noise power spectrum P(f), or noise covariance matrix N. In Appendix $\ref{Appendix}$ we would show two examples with same parameters as in Figure(3) except scanning frequency f_{scan} . It turns out the η values should somehow depends on scanning scheme. Again that's because when we determine the upper boundwe treated \mathbf{r}_{η} as an arbitrary vector, such that we lose all information related to scanning scheme in the pointing matrix P.

KMH: We need some discussion of the things that haven't yet been demonstrated with the PCG, like mulsiple messenger fields. Has the Kodi-Ramanah dual messenger field scheme been demonstrated in a PCG scheme by Papez?

Even though the perturbation parameter η get from Eq. (A6) are not the most optimal, it still performs much better than traditional conjugate gradient method under 1/f noise scenario without adding extra computational cost. The only extra free parameter added is to determine whether the error at current step $\mathbf{r}(\eta_i)$ =

 $_{407}$ $||\mathbf{b}(\eta_i) - A(\eta_i)\mathbf{m}||$ is small enough such that we advance to next value η_{i+1} .

Also this analysis of η value also explains why cooling parameters $\lambda=1/\eta$ in messenger field are chosen to be geometric series or logspace used in Huffenberger & Næss (2018).

All of the calculation are using simple preconditioner $P^{\dagger}P$, but the entire analysis is independent of precon-

ditioner. Better preconditioners would also lead to improvements.

⁴¹⁷ BQ and KH are supported by NSF award 1815887.

APPENDIX

A. THE SEQUENCE OF INVERSE COOLING PARAMETERS

We know that the initial inverse cooling parameter $\eta_0 = 0$. What would be good value for the next parameter η_1 ?

To simplify notation, we use N_{η} to denote $N(\eta) = \tau I + \eta \bar{N}$. For some specific η value, the minimum χ^2 value is given by the optimized map $\hat{\mathbf{m}}(\eta)$,

$$\chi^{2}(\hat{\mathbf{m}}(\eta), \eta) = (\mathbf{d} - P\hat{\mathbf{m}}(\eta))^{\dagger} N_{\eta}^{-1} (\mathbf{d} - P\hat{\mathbf{m}}(\eta))$$
(A1)

We restrict to the case that the noise covariance matrix N is diagonal in the frequency domain.

Let's first consider $\eta_1 = \eta_0 + \delta \eta = \delta \eta$ such that $\eta_1 = \delta \eta$ is very small quantity. Then the relative decrease of $\chi^2(\hat{\mathbf{m}}(0), 0)$ from $\eta_0 = 0$ to $\eta_1 = \delta \eta$ is

$$-\frac{\delta \chi^2(\hat{\mathbf{m}}(0), 0)}{\chi^2(\hat{\mathbf{m}}(0), 0)} = \delta \eta \frac{1}{\tau} \frac{(\mathbf{d} - P\hat{\mathbf{m}}(0))^{\dagger} \bar{N} (\mathbf{d} - P\hat{\mathbf{m}}(0))}{(\mathbf{d} - P\hat{\mathbf{m}}(0))^{\dagger} (\mathbf{d} - P\hat{\mathbf{m}}(0))}$$
(A2)

Here we put a minus sign in front of this expression such that it's non-negative. KMH: This statement is kind of non trivial... I think you need to explain more the step $(\tau I + \delta \eta \bar{N})^{-1} \approx \tau I - \delta \eta \bar{N}$. (I think that's right, even if you don't restrict to the frequency diagonal case.)

Ideally, we want $\delta\chi^2(\hat{\mathbf{m}}(0),0) = \chi^2(\hat{\mathbf{m}}(1),1) - \chi^2(\hat{\mathbf{m}}(0),0)$, such that it would get close to the final χ^2 at next iteration. Here if we assume that initial χ^2 value $\chi^2(\hat{\mathbf{m}}(0),0)$ is much larger than final value $\chi^2(\hat{\mathbf{m}}(1),1)$, then we would expect $|\delta\chi^2(\hat{\mathbf{m}}(0),0)/\chi^2(\hat{\mathbf{m}}(0),0)| \approx 1^-$, strictly smaller than 1. To make sure it will not start too fast, we could set its upper bound equal to 1, $\delta\eta \max(\bar{N}_f)/\tau = 1$ KMH: explain this reasoning a bit better. This gives

$$\eta_1 = \frac{\tau}{\max(\bar{N}_f)} = \frac{\min(N_f)}{\max(N_f) - \min(N_f)}$$
(A3)

Here N_f and \bar{N}_f are the eigenvalues of N and \bar{N} in the frequency domain. If the condition number of noise covariance matrix $\kappa(N) = \max(N_f)/\min(N_f) \gg 1$, then $\eta_1 \approx \kappa^{-1}(N)$.

What about the other parameters η_m with m > 1? We could use a similar analysis, let $\eta_{m+1} = \eta_m + \delta \eta_m$ with a small $\delta \eta_m$, and set the upper bound of relative decrease equal to 1. See Appendix B for detailed derivation. We would get

$$\delta \eta_m = \min\left(\frac{\tau + \eta_m \bar{N}_f}{\bar{N}_f}\right) = \eta_m + \frac{\tau}{\max(\bar{N}_f)}.$$
 (A4)

445 Therefore

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$$\eta_{m+1} = \eta_m + \delta \eta_m = 2\eta_m + \frac{\tau}{\max(\bar{N}_f)}$$
(A5)

448 As we can see, η_1, \dots, η_n increase like a geometric series.

$$\eta_i = \min\left\{1, \ \frac{\tau}{\max(\bar{N}_f)} (2^i - 1)\right\} \tag{A6}$$

⁴⁵¹ Here we need to truncate the series when $\eta_i > 1$.

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B. UPPER BOUND FOR η

We want to find the upper bound for $-\delta\chi^2(\hat{\mathbf{m}}(\eta_m), \eta_m)/\chi^2(\hat{\mathbf{m}}(\eta_m), \eta_m)$. KMH: I think you need more details on the steps of this derivation. First let's calculate $\frac{\mathrm{d}}{\mathrm{d}n}\chi^2(\hat{\mathbf{m}}(\eta), \eta)$

$$\frac{\mathrm{d}}{\mathrm{d}\eta} \chi^{2}(\hat{\mathbf{m}}(\eta), \eta) = \frac{\partial}{\partial \eta} \chi^{2}(\hat{\mathbf{m}}(\eta), \eta)$$

$$= \frac{\partial}{\partial \eta} (\mathbf{d} - P\hat{\mathbf{m}}(\eta))^{\dagger} N(\eta)^{-1} (\mathbf{d} - P\hat{\mathbf{m}}(\eta))$$

$$= -(\mathbf{d} - P\hat{\mathbf{m}}(\eta))^{\dagger} N(\eta)^{-1} \bar{N} N(\eta)^{-1} (\mathbf{d} - P\hat{\mathbf{m}}(\eta))$$

$$= -\mathbf{r}^{\dagger}(\eta) N(\eta)^{-1} \bar{N} N(\eta)^{-1} \mathbf{r}(\eta). \tag{B7}$$

where the first line comes from, $\chi^2(\hat{\mathbf{m}}(\eta), \eta)$ is minimum χ^2 value for certain η , therefore $\frac{\partial}{\partial \mathbf{m}} \chi^2(\mathbf{m}, \eta) \Big|_{\mathbf{m} = \hat{\mathbf{m}}(\eta)} = 0$. So the third line we only take partial derivative with respect to $N(\eta)^{-1}$. The last line we define $\mathbf{r}(\eta) = \mathbf{d} - P\hat{\mathbf{m}}(\eta)$.

The upper bound is given by

 $-\frac{\delta \chi^{2}(\hat{\mathbf{m}}(\eta_{m}), \eta_{m})}{\chi^{2}(\hat{\mathbf{m}}(\eta_{m}), \eta_{m})} = \delta \eta_{m} \frac{\mathbf{r}^{\dagger} N(\eta_{m})^{-1} \bar{N} N(\eta_{m})^{-1} \mathbf{r}}{\mathbf{r}^{\dagger} N(\eta_{m})^{-1} \mathbf{r}}$ $\leq \delta \eta_{m} \max \left(\frac{\bar{N}_{f}}{\tau + \eta_{m} \bar{N}_{f}} \right)$ (B8)

For the last line, both matrix \bar{N} and $N(\eta_m)^{-1}$ can be simultaneously diagonalized in frequency space. For each eigenvector \mathbf{e}_f , the corresponding eigenvalues of the matrix $N(\eta_m)^{-1}\bar{N}N(\eta_m)^{-1}$ are $\lambda_f = \bar{N}_f(\tau + \eta_m\bar{N}_f)^{-2}$, and the eigenvalues for matrix $N(\eta_m)^{-1}$ are $\gamma_f = (\tau + \eta_m\bar{N}_f)^{-1}$. Their eigenvalues are related by $\lambda_f = \frac{\bar{N}_f}{\tau + \eta_m\bar{N}_f}\gamma_f$. For vector $\mathbf{r} = \sum_f \alpha_f \mathbf{e}_f$, we have $\frac{\mathbf{r}^\dagger N(\eta_m)^{-1}\bar{N}N(\eta_m)^{-1}\mathbf{r}}{\mathbf{r}^\dagger N(\eta_m)^{-1}\mathbf{r}} = \frac{\sum_f \alpha_f^2 \lambda_f}{\sum_f \alpha_f^2 \gamma_f} = \frac{\sum_f \alpha_f^2 \gamma_f \bar{N}_f/(\tau + \eta_m\bar{N}_f)}{\sum_f \alpha_f^2 \gamma_f} \leq \max\left(\frac{\bar{N}_f}{\tau + \eta_m\bar{N}_f}\right)$.

If we set the upper bound $\delta\eta_m \max\left(\frac{\bar{N}_f}{\tau + \eta_m\bar{N}_f}\right) = 1$, $\frac{1}{2}$ and then we get

$$\delta \eta_m = \min\left(\frac{\tau + \eta_m \bar{N}_f}{\bar{N}_f}\right) = \eta_m + \frac{\tau}{\max(\bar{N}_f)}.$$
 (B9)

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² Here we also assumed that $\chi^2(\hat{\mathbf{n}}(\eta_m), \eta_m) \gg \chi^2(\hat{\mathbf{n}}(1), 1)$, which we expect it to be satisfied for $0 \simeq \eta_m \ll 1$. Since final result Eq. (A6) is geometric series, only a few η_m values won't satisfy this condition.