Cosmic Microwave Background map-making solutions improve with cooling

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ABSTRACT

In the context of Cosmic Microwave Background data analysis, we study the solution to the equation that transforms scanning data into a map. As originally suggested in "messenger" methods for solving linear systems, we split the noise covariance into uniform and non-uniform parts and adjusting their relative weight during the iterative solution. This "cooling" or perturbative approach is particularly effective when there is significant low-frequency noise in the timestream. A conjugate gradient algorithm applied to this modified system converges faster and to a higher fidelity solution than the standard conjugate gradient approach, for the same computational cost per iteration. We conclude that cooling is helpful separate from its appearance in the messenger methods. We give an analytical expression for the parameter that controls how gradually should change during the course of the solution.

Keywords: Computational methods — Cosmic microwave background radiation — Astronomy data reduction

1. INTRODUCTION

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In observations of the Cosmic Microwave Background (CMB), map-making is an intermediate step between the collection of raw scanning data and the scientific analyses, such as the estimation of power spectra and cosmological parameters. Next generation CMB observations will generate much more data that today, and so it is worth exploring efficient ways to process the data, even though, on paper, the map-making problem has long been solved.

The time-ordered scanning data is summarized by

$$\mathbf{d} = P\mathbf{m} + \mathbf{n} \tag{1}$$

where **d**, **m**, and **n** are the vectors of time-ordered data (TOD), the CMB sky-map signal, and measurement noise, and P is the sparse matrix that encodes the telescope's pointing. Of several mapmaking methods (Tegmark 1997a), one of the most common is the method introduced for the Cosmic Background Explorer (COBE, Janssen & Gulkis 1992). This optimal, linear solution is

$$(P^{\dagger}N^{-1}P)\hat{\mathbf{m}} = P^{\dagger}N^{-1}\mathbf{d}$$
 (2)

 $_{\rm 37}$ where ${\bf \hat{m}}$ provides the generalized least squares mini- $_{\rm 38}$ mization of the χ^2 statistic

$$\chi^{2}(\mathbf{m}) \equiv (\mathbf{d} - P\mathbf{m})^{\dagger} N^{-1} (\mathbf{d} - P\mathbf{m}). \tag{3}$$

⁴¹ Here we assume that the noise has zero mean $\langle \mathbf{n} \rangle = \mathbf{0}$, ⁴² and noise covariance matrix could be written as N = ⁴³ $\langle \mathbf{n} \mathbf{n}^{\dagger} \rangle$. We cast mapmaking as a standard linear regres- ⁴⁴ sion problem. In case the noise is Gaussian, the COBE ⁴⁵ solution is also the maximum likelihood solution.

With current computation power, we cannot solve for 47 $\hat{\mathbf{m}}$ by calculating $(P^{\dagger}N^{-1}P)^{-1}P^{\dagger}N^{-1}\mathbf{d}$ directly, since 48 the $(P^{\dagger}N^{-1}P)$ matrix is too large to invert. The noise 49 covariance matrix N is sparse in frequency domain and 50 the pointing matrix P is sparse in the time-by-pixel do-51 main, and their product is dense. In experiments cur- $_{52}$ rently under design, there may be $\sim 10^{16}$ time sam-₅₃ ples and $\sim 10^9$ pixels, so these matrix inversions are 54 intractable. Therefore we use iterative methods, such 55 as conjugate gradient descent, to avoid the matrix in-56 versions, while executing each matrix multiplication in 57 a basis where the matrix is sparse, using a fast Fourier 58 transform to go between the frequency and time domain. As an alternative technique, Huffenberger & Næss 60 (2018) showed that the "messenger method" could be 61 adapted to solve the linear mapmaking system, based 62 on the approach from Elsner & Wandelt (2013) to solve 63 the linear Wiener filter. This technique splits the noise 64 covariance into a uniform part and the remainder, and, 65 over the course of the iterative solution, it adjusts the 66 relative weight of those two parts. Starting with the 67 uniform covariance, the modified linear system gradu-68 ally transforms to the final system via a cooling param-69 eter. The cooling idea again comes from Elsner & Wan-

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70 delt (2013). In numerical experiments, Huffenberger & 71 Næss (2018) found that the large scales of map produced 72 by the cooled messenger method converged significantly ₇₃ faster than for standard methods, and to higher fidelity. Papež et al. (2018) showed that the iterations in the 75 messenger field approach is equivalent to a fixed point 76 iteration scheme, and studied its convergence proper-77 ties in detail. Furthermore, they showed that the split 78 covariance and the modified system that incorporates 79 the cooling can be solved by other means, including 80 a conjugate gradient technique, which should generally 81 show better convergence properties than the fixed-point 82 scheme. However in numerical tests, Papež et al. (2018) 83 did not find benefits to the cooling modification of the 84 mapmaking system, in contrast to findings of Huffen-85 berger & Næss (2018).

In this paper, we show that the difference arose be- 87 cause the numerical tests in Papež et al. (2018) used 88 much less low-frequency (1/f) noise than Huffenberger 89 & Næss (2018), and show that the cooling technique 90 improves mapmaking performance especially when the 91 low frequency noise is large. This performance boost de- 92 pends on a proper choice for the pace of cooling. Kodi 93 Ramanah et al. (2017) showed that for Wiener filter the 94 cooling parameter should be chosen as a geometric se- 95 ries. In this work, we give an alternative interpretation 96 of the parameterizing process and show that for map- 97 making the optimal choice (unsurprisingly) is also a ge- 98 ometric series.

In Section 2 we describe our methods for treating the mapmaking equation and our numerical experiments. In Section 3 we present our results. In Section 4 we interpret the mapmaking approach and its computational cost. In Section 5 we conclude. In appendices we derives how we set our cooling schedule.

2. METHODS

2.1. Parameterized Conjugate Gradient Method

The messenger field approach introduced an extra cooling parameter λ to the map-making equation, and solved the linear system with the alternative covariance $N(\lambda) = \lambda \tau I + \bar{N}$. The parameter τ represents the uniform level of (white) noise in the covariance, \bar{N} is the balance of the noise, and the parameterized covariance equals the original covariance when the cooling parameter $\lambda = 1$. In this work we find it more convenient to work with the inverse cooling parameter $\eta = \lambda^{-1}$ and define the covariance as

$$N(\eta) = \tau I + \eta \bar{N} \tag{4}$$

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which leads to the same system of mapmaking equations. (This is because $N(\eta) = \lambda^{-1} N(\lambda)$ and the mapmaking

¹²⁰ equation is insensitive to to scalar multiple of the co-¹²¹ variance since is appears on both sides.)

Papež et al. (2018) showed that the conjugate gradi-123 ent method can be easily applied to parameterized map-124 making equation by iterating on

$$P^{\dagger}N(\eta)^{-1}P \hat{\mathbf{m}} = P^{\dagger}N(\eta)^{-1}\mathbf{d}$$
 (5)

127 as the cooling is adjusted. In our numerical experiments, 128 we confirm that the conjugate gradient approach is con129 verging faster than the fixed point iterations suggested 130 by the messenger mapmaking method in Huffenberger 131 & Næss (2018). For simplicity we fix the preconditioner 132 to $M = P^{\dagger}P$ for all of calculations.

When $\eta=0$, the noise covariance matrix N(0) is proportional to identity matrix I, and solution is given by simple binned map $\mathbf{m}_0=\left(P^\dagger P\right)^{-1}P^\dagger\mathbf{d}$, which can be solved directly. From this starting point, the cooling scheme requires the inverse cooling parameter η increase as $0=\eta_0\leq\eta_1\leq\cdots\leq\eta_{\mathrm{final}}=1$, at which point we arrive at the desired mapmaking equation. For each intermediate η_i , we treat it as a seperate conjugate gradient method to solve equation $(P^\dagger N(\eta_i)^{-1}P)\hat{\mathbf{m}}(\eta_i)=P^\dagger N(\eta_i)^{-1}\mathbf{d}$, using the result from previous calculation $\hat{\mathbf{m}}(\eta_{i-1})$ as the initial value, and move to next parameter η_{i+1} when $(P^\dagger N(\eta_i)^{-1}P)\hat{\mathbf{m}}(\eta_i)-P^\dagger N(\eta_i)^{-1}\mathbf{d}\simeq 0$, the KMH: In this description, it is not totally clear whether you intend to update the eta after every iteration.

The non-white part \bar{N} is the troublesome portion of the the covariance, and we can think of the η parameter as turning it on slowly, adding a perturbation to the solution achieved at a particular stage, building ultimately upon the initial uniform covariance model.

2.2. Choice of inverse cooling parameters η

The next question is how we choose these monotonising ically increasing parameters η . If we choose them inappropriately, the solution converge slowly, because we
solve waste effort converging on the wrong system. We also
want to determine $\eta_1, \dots, \eta_{n-1}$ before starting conjugate gradient iterations. The time ordered data \mathbf{d} is very
large, and we do not want to keep it in the system memory during calculation. If we determine $\eta_1, \dots, \eta_{n-1}$ before the iterations, then we can precompute the righthand side $P^{\dagger}N(\eta)^{-1}\mathbf{d}$ for each η_i and keep these mapisized objects in memory, instead of the entire timeordered data.

In the appendix, we show that a generic good choice for the η parameters are the geometric series

$$\eta_i = \min\left\{ \left(2^i - 1\right) \frac{\tau}{\max(\bar{N}_f)}, \ 1 \right\},\tag{6}$$

where \bar{N}_f is the frequency representation of the non-uniform part of the covariance. This is the main result.

It tells us not only how to choose parameters η_i , but also when we should stop the perturbation, and set $\eta=173$ 1. For example, if noise covariance matrix N is almost white noise, then $\bar{N}=N-\tau I\approx 0$, and we would have $\tau=175$ $\tau/\max(\bar{N}_f)\gg 1$. This tell us that we don't need to use parameterized method at all, because $\tau=175$ due to $\tau=175$ $\tau=175$ $\tau=175$ This corresponds to the standard $\tau=175$ $\tau=17$

2.3. Numerical Simulations

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To compare these algorithms, we need to do some simple simulation of scanning processes, and generate time ordered data from random sky signal. Our sky is a small rectangular area, with two orthogonal directions x and y, both with range from -1° to $+1^{\circ}$. The signal has first three stokes parameters (I,Q,U).

For the scanning process, our single telescope contains nine detectors, each has different sensitivity to polarization Q and U. It scans the sky with a raster scanning pattern and scanning frequency $f_{\rm scan}=0.1$ Hz sampling frequency $f_{\rm sample}=100$ Hz. The telescope scans the sky horizontally and then vertically, and then digitizes the position (x,y) into 512×512 pixel. This gives noiseless signal s.

The noise power spectrum is given by

$$P(f) = \sigma^2 \left(1 + \frac{f_{\text{knee}}^{\alpha} + f_{\text{apo}}^{\alpha}}{f^{\alpha} + f_{\text{apo}}^{\alpha}} \right) \tag{7}$$

Here we fixed $\sigma^2=10~\mu\text{K}^2,~\alpha=2$ and $f_{\rm knee}=10$ Hz, and change $f_{\rm apo}$ to compare the performance under different noise models. Note that as $f_{\rm apo}\to 0,~P(f)\to \sigma^2 \left(1+(f/f_{\rm knee})^{-1}\right)$, it becomes a 1/f noise model. The noise covariance matrix

$$N_{ff'} = P(f) \frac{\delta_{ff'}}{\Delta_f} \tag{8}$$

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 $_{204}$ is a diagonal matrix in frequency space, where Δ_f is equal to reciprocal of total scanning time T. In our calculations we choose the $f_{\rm apo}$ such that the condition numbers κ are 10^2 , 10^6 , and 10^{12} . The corresponding power spectrum are shown in Figure(1).

Finally, we get the simulated time ordered data $\mathbf{d} = \mathbf{s} + \mathbf{n}$ by adding up signal and noise.

KMH: Compare to the noise power spectrum of Papez. Remark how little 1/f is in their test. What is the effect of changing the noise slope? In Papež et al. 2018, the noise power spectrum is apodized at

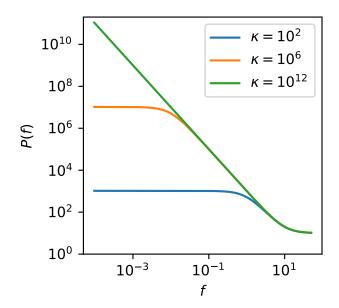


Figure 1. The noise power spectum based on Eq. (7) with $\sigma^2 = 10 \ \mu\text{K}^2$, $\alpha = 2$ and $f_{\text{knee}} = 10 \ \text{Hz}$. And fixing the condition number κ of noise covariance matrix Eq. (8) by choosing f_{apo} . KMH: show the scanning frequency with a vertical (dashed?) line. Can use axyline().

 $_{216}$ $0.1f_{\mathrm{knee}}$, which corresponds to $f_{\mathrm{apo}} \approx 0.1f_{\mathrm{knee}}$ For the $_{217}$ case $f_{\mathrm{knee}} = 10$ and $\kappa = 100$, the apodization frequency $_{218}$ $f_{\mathrm{apo}} \approx 0.99 \approx 0.1f_{\mathrm{knee}}$. Therefore the blue line in Figure $_{219}$ (1), is close to the power spectrum used in Papež et al. $_{220}$ 2018.

3. RESULTS

First let's compare the results with vanilla conjugate gradient method with simple preconditioner $P^{\dagger}P$. The results are showed in Figure (2) for different kinds of noise power spectra. Here note that χ^2 in all figures are calculated based on Eq. (3) not $\chi^2(\mathbf{m},\eta)$ in Eq. (??). The χ^2_{\min} is calculated from perturbative conjugate gradient method with more intermediate η values, and more iterations after $\eta=1$.

As we can see in the left graph in Figure (2), when the condition number of noise covariance matrix $\kappa(N)$ is small, the performance between different these two methods are small. The vanilla conjugate gradient method converge faster, because its perturbation parameter goes to 1 at the first iteration, however for the perturbation method its η value will slowly reach 1 in about ten iterations.

Notice that as we increase $\kappa(N)$, or equivalently decrease $f_{\rm apo}$, the perturbation parameter η starts showing its benefits, as showed in the second and third graph in Figure(2). It outperforms the vanilla conjugate gradient method when $f_{\rm apo} \approx 0$ and the noise power spectrum be-

¹ The source code and other information are available at https://github.com/Bai-Qiang/map_making_perturbative_approach

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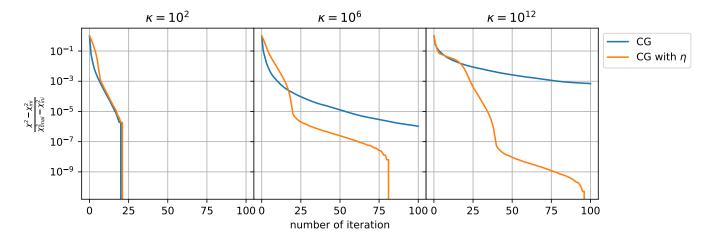


Figure 2. These three figures show the $\frac{\chi^2(\mathbf{m}) - \chi^2_{\text{ini}}}{\chi^2_{\text{min}} - \chi^2_{\text{ini}}}$ changes for each iteration under different noise covariance matrix with condition number being 10^2 , 10^6 , and 10^{12} .

comes the 1/f noise model, which usually is the intrinsic noise of instruments (Tegmark (1997b)).

Now let us compare the performance difference between choosing η parameters based on Eq. (6) 247 and manually fixing number of η parameters n_n 248 manually. We manually choose the η_i values us- $_{ ext{249}}$ ing function numpy.logspace(start= $\ln(\eta_1)$, stop=0, 250 num= n_n , base=e). The results are showed in Figure(3). When $\kappa(N)$ is small, and Eq. (6) tells us that only a 251 252 few η parameters are good enough, see the orange line 253 in the first Figure(3), where we have $\sim 10~\eta$ levels. If unfortunately we choose n_{η} being large value, like 15 or 255 30, then it will ends up converge slowly, because it needs 256 at least 15 or 30 iterations to reach $\eta = 1$, at least one 257 iteration per η level.

On the other hand if $\kappa(N)$ is very large and the power spectrum is 1/f noise, we need more η parameters. If n_{η} is too small, for example $n_{\eta} = 5$ the green line in last 261 Figure(3), it may be better than the vanilla conjugate 262 gradient method, but it is still far from optimal.

Since the η values determined from Eq. (6)

$$\eta_i = \min\left\{1, \ \frac{\tau}{\max(\bar{N}_f)} (2^i - 1)\right\} \tag{6}$$

²⁶⁶ are not dependent on any scanning information, it only depends on noise power spectrum P(f), or noise covariance matrix N. Figure (??) and Figure (??) show two 269 examples with same parameters as in Figure (3) except 270 for the scanning frequency f_{scan} in Figure (??) it scans 271 very slow and in Figure (??) it's very fast. In these 272 two cases under 1/f noise model, our η values based on 273 Eq. (6) are better than manually selected values. Based 274 on these two results we know, the η values should some-275 how depends on scanning scheme.

4. DISCUSSION

4.1. Intuitive Interpretation of n

KMH: most of this is pretty similar to discussion in 279 Huffenberger and Naess. The last paragraph is new.

In this section, let me introduce another way to understand the role of η . Our ultimate goal is to find $\hat{\mathbf{m}}(\eta=1)$ which minimizes $\chi^2(\mathbf{m}) = (\mathbf{d} - P\mathbf{m})^{\dagger} N^{-1} (\mathbf{d} - P\mathbf{m}).$ Since N is diagonal in frequency space, χ^2 could be writ-²⁸⁴ ten as a sum of all frequency mode $|(\mathbf{d} - P\mathbf{m})_f|^2$ with ²⁸⁵ weight N_f^{-1} , such as $\chi^2(\mathbf{m}) = \sum_f |(\mathbf{d} - P\mathbf{m})_f|^2 N_f^{-1}$. 286 N_f^{-1} is large when there is little noise at that frequency, 287 and vice versa. Which means $\chi^2(\mathbf{m})$ would favor the 288 low noise frequency mode over high noise ones. In other words the optimal map $\hat{\mathbf{m}}$ focusing on minimize the er-290 ror $\mathbf{r} \equiv \mathbf{d} - P\mathbf{m}$ in the low-noise part.

After introducing η , we minimize $\chi^2(\mathbf{m}, \eta) = (\mathbf{d} - 292 \ P\mathbf{m})^{\dagger} N_{\eta}^{-1} (\mathbf{d} - P\mathbf{m})$. For $\eta = 0, N_{\eta=0}^{-1} \propto I$ and the esti-293 mated map $\hat{\mathbf{m}}(\eta=0)$ does not prioritize any frequency 294 mode. As we slowly increase η , we decrease the weight 295 for the frequency modes which have large noise, and fo-296 cusing minimizing error for low noise part. If we start with $\eta_1 = 1$ directly, which corresponds to the vanilla 298 conjugate gradient method, then the entire conjugate 299 gradient solver will focus most on minimizing the low 300 noise part, such that χ^2 would converge very fast at low 301 noise region, but slowly on high noise part. Since it fo-302 cus on low noise part only, it may be stuck at some local minimum point. To get to the global minimum, it need 304 to adjust the low noise part, that would be difficult if it's 305 stuck at an local minimum. However by introducing η 306 parameter, we let the solver first treat every frequency $_{307}$ equally. Then as η slowly increases, it gradually shifts 308 focus from the highest noise to the lowest noise part. 309 KMH: I feel what this is missing is why the high-noise 310 modes get stuck though.

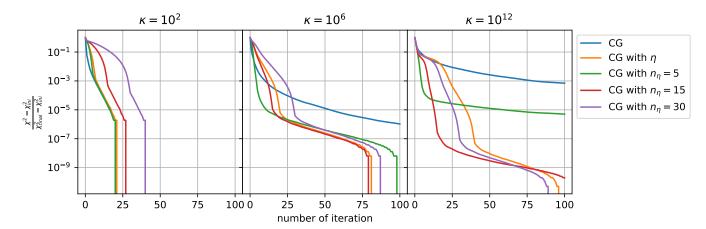


Figure 3. The blue line and the orange line are the same as Figure (2). For three extra lines, we fix the number of η parameter n_{η} manully. Instead of using Eq. (6), we use numpy.logspace(start=ln(η_1), stop=0, num= n_{η} , base=e). to get all η parameters.

If we write the difference between final and initial χ^2 value as $\chi^2(\hat{\mathbf{n}}(1),1)-\chi^2(\hat{\mathbf{n}}(0),0)=$ $_{313}\int_0^1\mathrm{d}\eta\,\frac{\mathrm{d}}{\mathrm{d}\eta}\chi^2(\hat{\mathbf{n}}(\eta),\eta),$ and use Eq. (A2). We note that when η is very small, the $\frac{\mathrm{d}}{\mathrm{d}\eta}\chi^2(\hat{\mathbf{n}}(\eta),\eta)$ would have relatively large contribution from medium to large noise region, comparing to large η . So introducing η might improve the convergence of χ^2 at these regions, because the vanilla conjugate gradient method only focuses on the low noise part and it may have difficulty at these regions.

$4.2. \ \ Computational \ Cost$ To properly compare the performance cost of this

323 method with respect to vanilla conjugate gradient

method with simple preconditioner, we need to com-

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325 pare their computational cost at each iteration. The 326 right hand side of parameterized map-making equation Eq. (5) could be computed before iterations, so it won't 328 introduce extra computational cost. The most demand-329 ing part of conjugate gradient method is calculating 330 $P^{\dagger}N^{-1}P\hat{\mathbf{m}}$, because it contains a Fourier transform of $P\hat{\mathbf{m}}$ from time domain to frequency domain and an in-³³² verse Fourier transform of $N^{-1}P\hat{\mathbf{n}}$ from frequency domain back to time domain, which is order $\mathcal{O}(n \log n)$ n with n being the length of time ordered data. If we sas change N^{-1} to $N(\eta)^{-1}$, it won't add extra cost, since 336 both matrices are diagonal in frequency domain. There-337 fore the computational cost it the same for one step. However our previous analysis is based 339 $\chi^2(\hat{\mathbf{m}}(\eta_i), \eta_i)$ which is evaluated at $\hat{\mathbf{m}}(\eta_i)$ the esti-340 mated map at η_i . So We should update η_i to η_{i+1} when $\mathbf{m} \approx \hat{\mathbf{m}}(\eta_i)$. How do we know this condition is satisfied? Since for each new η_i value, we are solving 343 a new set of linear equations $A(\eta_i)\hat{\mathbf{m}} = \mathbf{b}(\eta_i)$ with $A(\eta_i) = P^{\dagger} N(\eta_i)^{-1} P$ and $\mathbf{b}(\eta_i) = P^{\dagger} N(\eta_i)^{-1} \mathbf{d}$, and we

345 could stop calculation and moving to next value η_{i+1} 346 when the norm of residual $||\mathbf{r}(\eta_i)|| = ||\mathbf{b}(\eta_i) - A(\eta_i)\mathbf{m}||$ 347 smaller than some small value. Calculate $||\mathbf{r}(\eta_i)||$ is 348 part of conjugate gradient algorithm, so this won't 349 add extra cost compare to vanilla conjugate gradient 350 method. Therefore, overall introducing η won't have 351 extra computational cost.

4.3. Future Prospects

KMH: some of this future prospects should move to discussion As you may have noticed in the second and third Figure(3), the perturbation parameter based on Eq. (6) is more than needed, especially for 1/f noise ase. For the case $\kappa=10^{12}$, we notice that based on Eq. (6) it gives us $n_{\eta}\approx 40$, however from χ^2 result in the last Figure(3) $n_{\eta}\approx 30$ or even $n_{\eta}\approx 15$ is good enough. Also, for the nearly-white-noise case, we could certainly choose $n_{\eta}=1$ such that $\eta_1=1$ which corresponds to vanilla conjugate gradient method, based on χ^2 result in first Figure(3). However Eq. (6) gives us $n_{\eta}\approx 6$, even though it does not make the final χ^2 result much different at the end.

Is it possible to further improve the analysis, such that it produces smaller n_{η} ? Let's examine how we get η_i series. Remember that we determine $\delta \eta$ value based on the upper bound of $-\delta \chi^2(\hat{\mathbf{m}}(\eta), \eta)/\chi^2(\hat{\mathbf{m}}(\eta), \eta)$, in Eq. (??). For $\eta \neq 0$, the upper bound is

$$\delta \eta \frac{\hat{\mathbf{r}}_{\eta}^{\dagger} N(\eta)^{-1} \bar{N} N(\eta)^{-1} \hat{\mathbf{r}}_{\eta}}{\hat{\mathbf{r}}_{\eta}^{\dagger} N(\eta)^{-1} \hat{\mathbf{r}}_{\eta}} \le \frac{\delta \eta}{\eta + \frac{\tau}{\max(N_f) - \tau}}$$
(9)

with $\mathbf{r}_{\eta} = \left[1 - P(P^{\dagger}N(\eta)^{-1}P)^{-1}P^{\dagger}N(\eta)^{-1}\right]\mathbf{d} \equiv \mathcal{P}_{\eta}\mathbf{d}$.

To get the upper bound we treated $\mathbf{d} - P\hat{\mathbf{m}}(\eta)$ as an arbitrary vector in frequency domain, since we don't know how to calculate \mathcal{P}_{η} for $\eta \neq 0$, and it's hard to analyze

377 the projection matrix \mathcal{P}_{η} in frequency space, as it contains $(P^{\dagger}N(\eta)^{-1}P)^{-1}$. Note that we have to determine 379 all of η value before calculation, because we don't want 380 to keep the time ordered data in system RAM, so we need to somehow analytically analyze \mathcal{P}_{η} , and its be- $_{382}$ havior in frequency space. Unless \mathbf{r}_{η} almost only has 383 large noise modes, $\left| \frac{d}{d\eta} \chi^2(\hat{\mathbf{m}}(\eta), \eta) / \chi^2(\hat{\mathbf{m}}(\eta), \eta) \right|$ won't 384 get close to the upper bound $1/\left(\eta + \frac{\tau}{\max(N_f) - \tau}\right)$. Based on the analysis in Section(4.1), for small η the estimated map $\hat{\mathbf{m}}(\eta)$ does not only focusing on minimiz- $_{387}$ ing error \mathbf{r}_{η} at low noise region. So we would expect 388 that there would be a fair amount of low noise modes contribution in \mathbf{r}_{η} especially for the first few η values. 390 Which means if we could somehow know the frequency 391 distribution of \mathbf{r}_{η} , we could tighten the boundary of $\frac{\mathrm{d}}{\mathrm{d}\eta}\chi^2(\hat{\mathbf{m}}(\eta),\eta)/\chi^2(\hat{\mathbf{m}}(\eta),\eta)$, and get larger $\delta\eta$ value. This should make η goes to 1 faster, and yields the fewer η parameters we need.

Also notice that the η values determined from Eq. (6) are not dependent on any scanning information, it only depends on noise power spectrum P(f), or noise covariance matrix N. In Appendix $\ref{eq:spectrum}$ we would show two samples with same parameters as in Figure (3) except scanning frequency f_{scan} . It turns out the η values should somehow depends on scanning scheme. Again that's because when we determine the upper boundwe treated \mathbf{r}_{η}

 $_{403}$ as an arbitrary vector, such that we lose all information $_{404}$ related to scanning scheme in the pointing matrix P.

5. CONCLUSIONS

KMH: We need some discussion of the things that haven't yet been demonstrated with the PCG, like multiple messenger fields. Has the Kodi-Ramanah dual messenger field scheme been demonstrated in a PCG scheme by Papez?

Even though the perturbation parameter η get from Eq. (6) are not the most optimal, it still performs much better than traditional conjugate gradient method under 1/f noise scenario without adding extra computational cost. The only extra free parameter added is to determine whether the error at current step $\mathbf{r}(\eta_i) = |\mathbf{b}(\eta_i) - A(\eta_i)\mathbf{m}|$ is small enough such that we advance to next value η_{i+1} .

Also this analysis of η value also explains why cooling parameters $\lambda=1/\eta$ in messenger field are chosen to be geometric series or logspace used in Huffenberger & Næss (2018).

All of the calculation are using simple preconditioner $P^{\dagger}P$, but the entire analysis is independent of preconditioner. Better preconditioners would also lead to improvements.

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428 APPENDIX

A. THE SEQUENCE OF INVERSE COOLING PARAMETERS

We know that the initial inverse cooling parameter $\eta_0 = 0$. What would be good value for the next parameter η_1 ?

To simplify notation, we use N_{η} to denote $N(\eta) = \tau I + \eta \bar{N}$. For some specific η value, the minimum χ^2 value is given by the optimized map $\hat{\mathbf{n}}(\eta) = \left(P^{\dagger}N_{\eta}^{-1}P\right)^{-1}P^{\dagger}N_{\eta}^{-1}\mathbf{d}$, which minimize

$$\chi^{2}(\hat{\mathbf{m}}(\eta), \eta) = (\mathbf{d} - P\hat{\mathbf{m}}(\eta))^{\dagger} N_{\eta}^{-1} (\mathbf{d} - P\hat{\mathbf{m}}(\eta))$$
(A1)

We restrict to the case that the noise covariance matrix N is diagonal in the frequency domain.

Let's first consider $\eta_1 = \eta_0 + \delta \eta = \delta \eta$ such that $\eta_1 = \delta \eta$ is very small quantity. Since $\hat{\mathbf{m}}(\eta)$ minimize $\chi^2(\hat{\mathbf{m}}(\eta), \eta)$, we have $\frac{\partial}{\partial \hat{\mathbf{m}}} \chi^2(\hat{\mathbf{m}}(\eta), \eta) = 0$, and using chain rule

$$\frac{\mathrm{d}}{\mathrm{d}\eta}\chi^{2}(\hat{\mathbf{m}}(\eta),\eta) = \frac{\partial}{\partial\eta}\chi^{2}(\hat{\mathbf{m}}(\eta),\eta) = -(\mathbf{d} - P\hat{\mathbf{m}}(\eta))^{\dagger}N_{\eta}^{-1}\bar{N}N_{\eta}^{-1}(\mathbf{d} - P\hat{\mathbf{m}}(\eta))$$
(A2)

Then the fractional decrease of $\chi^2(\hat{\mathbf{m}}(0),0)$ from $\eta_0=0$ to $\eta_1=\delta\eta$ is

$$-\frac{\delta \chi^{2}(\hat{\mathbf{m}}(0),0)}{\chi^{2}(\hat{\mathbf{m}}(0),0)} = -\delta \eta \frac{\frac{\mathrm{d}}{\mathrm{d}\eta} \chi^{2}(\hat{\mathbf{m}}(0),0)}{\chi^{2}(\hat{\mathbf{m}}(0),0)} = \delta \eta \frac{1}{\tau} \frac{(\mathbf{d} - P\hat{\mathbf{m}}(0))^{\dagger} \bar{N}(\mathbf{d} - P\hat{\mathbf{m}}(0))}{(\mathbf{d} - P\hat{\mathbf{m}}(0))^{\dagger}(\mathbf{d} - P\hat{\mathbf{m}}(0))}$$
(A3)

Here we put a minus sign in front of this expression such that it's non-negative, and use $N_{\eta=0} = \tau I$ at the second equality. Since it's hard to analyze $\mathbf{d} - P\hat{\mathbf{m}}$ under frequency domain, we treat it as an arbitary vector, then the least

445 upper bound is given by

$$-\frac{\delta \chi^2(\hat{\mathbf{m}}(0), 0)}{\chi^2(\hat{\mathbf{m}}(0), 0)} \le \frac{\delta \eta}{\tau} \max(\bar{N}_f) \tag{A4}$$

where $\max(\bar{N}_f)$ is the maxim eigenvalue of \bar{N} . Here if we assume that initial χ^2 value $\chi^2(\hat{\mathbf{m}}(0),0)$ is much larger than final value $\chi^2(\hat{\mathbf{m}}(1),1), \chi^2(\hat{\mathbf{m}}(0),0) \gg \chi^2(\hat{\mathbf{m}}(1),1)$, then we would expect

$$-\frac{\delta \chi^2(\hat{\mathbf{m}}(0), 0)}{\chi^2(\hat{\mathbf{m}}(0), 0)} = 1 - \frac{\chi^2(\hat{\mathbf{m}}(1), 1)}{\chi^2(\hat{\mathbf{m}}(0), 0)} \approx 1^-$$
(A5)

The upper bound is strictly smaller than 1. Ideally, if $\delta \chi^2(\hat{\mathbf{m}}(0),0) = \chi^2(\hat{\mathbf{m}}(1),1) - \chi^2(\hat{\mathbf{m}}(0),0)$, then it would get dose to the final χ^2 at next iteration, but we don't know final $\chi^2(\hat{\mathbf{m}}(1),1)$. So we want $\left|\frac{\delta \chi^2(\hat{\mathbf{m}}(0),0)}{\chi^2(\hat{\mathbf{m}}(0),0)}\right|$ being as large as possible, so it could converge fast. But there is another constrain that the least upper bound cannot exceed 1. Therefore could choose $\delta \eta$, such that the least upper bound is equal to 1. Then we find

$$\eta_1 = \frac{\tau}{\max(\bar{N}_f)} = \frac{\min(N_f)}{\max(N_f) - \min(N_f)} \tag{A6}$$

Here N_f and \bar{N}_f are the eigenvalues of N and \bar{N} in the frequency domain. If the condition number of noise covariance matrix $\kappa(N) = \max(N_f)/\min(N_f) \gg 1$, then $\eta_1 \approx \kappa^{-1}(N)$.

What about the other parameters η_m with m > 1? We could use a similar analysis, let $\eta_{m+1} = \eta_m + \delta \eta_m$ with a small $\delta \eta_m$, and set the least upper bound of relative decrease equal to 1.

$$-\frac{\delta \chi^{2}(\hat{\mathbf{m}}(\eta_{m}), \eta_{m})}{\chi^{2}(\hat{\mathbf{m}}(\eta_{m}), \eta_{m})} = \delta \eta_{m} \frac{(\mathbf{d} - P\hat{\mathbf{m}}(\eta_{m}))^{\dagger} N_{\eta_{m}}^{-1} \bar{N} N_{\eta_{m}}^{-1} (\mathbf{d} - P\hat{\mathbf{m}}(\eta_{m}))}{(\mathbf{d} - P\hat{\mathbf{m}}(\eta_{m}))^{\dagger} N_{\eta_{m}}^{-1} (\mathbf{d} - P\hat{\mathbf{m}}(\eta_{m}))}$$

$$\leq \delta \eta_{m} \max \left(\frac{\bar{N}_{f}}{\tau + \eta_{m} \bar{N}_{f}}\right)$$
(A7)

The upper bound in the second line is a little bit tricky. Both matrix \bar{N} and $N_{\eta_m}^{-1}$ can be simultaneously diagonalized in frequency space. For each eigenvector \mathbf{e}_f , the corresponding eigenvalue of the matrix on the numerator $N_{\eta_m}^{-1}\bar{N}N_{\eta_m}^{-1}$ is $\gamma_f = (\tau + \eta_m\bar{N}_f)^{-2}$, and the eigenvalue for matrix on the denominator $N_{\eta_m}^{-1}$ is $\gamma_f = (\tau + \eta_m\bar{N}_f)^{-1}$. Their eigenvalues are related by $\lambda_f = \frac{\bar{N}_f}{\tau + \eta_m\bar{N}_f}\gamma_f$. For any vector $\mathbf{v} = \sum_f \alpha_f \mathbf{e}_f$, we have $\frac{\mathbf{v}^\dagger N_{\eta_m}^{-1}\bar{N}N_{\eta_m}^{-1}\mathbf{v}}{\mathbf{v}^\dagger N_{\eta_m}^{-1}\mathbf{v}} = \frac{\sum_f \alpha_f^2\lambda_f}{\sum_f \alpha_f^2\gamma_f} = \frac{\sum_f \alpha_f^2\gamma_f\bar{N}_f/(\tau + \eta_m\bar{N}_f)}{\sum_f \alpha_f^2\gamma_f} \leq \max\left(\frac{\bar{N}_f}{\tau + \eta_m\bar{N}_f}\right)$

Similarly, we could set the least upper bound equal to 1. ² and then we get

$$\delta \eta_m = \min\left(\frac{\tau + \eta_m \bar{N}_f}{\bar{N}_f}\right) = \eta_m + \frac{\tau}{\max(\bar{N}_f)}.$$
 (A8)

471 Therefore

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$$\eta_{m+1} = \eta_m + \delta \eta_m = 2\eta_m + \frac{\tau}{\max(\bar{N}_f)}$$
(A9)

As we can see, η_1, \dots, η_n increase like a geometric series. If written in the form $\eta_{m+1} + \frac{\tau}{\max(\bar{N}_f)} = 2\left(\eta_m + \frac{\tau}{\max(\bar{N}_f)}\right)$ 475 it's easy to see that for $m \ge 1$, $\eta_m + \frac{\tau}{\max(\bar{N}_f)}$ forms a geometric series

$$\eta_m + \frac{\tau}{\max(\bar{N}_f)} = \left(\eta_1 + \frac{\tau}{\max(\bar{N}_f)}\right) 2^{m-1} = \frac{\tau}{\max(\bar{N}_f)} 2^m \tag{A10}$$

² Here we also assumed that $\chi^2(\hat{\mathbf{m}}(\eta_m), \eta_m) \gg \chi^2(\hat{\mathbf{m}}(1), 1)$, which we expect it to be satisfied for $0 \simeq \eta_m \ll 1$. Since final result Eq. (6) is geometric series, only a few η_m values won't satisfy this condition.

where we used $\eta_1 = \frac{\tau}{\max(\tilde{N}_f)}$. Note that m = 0 and $\eta_0 = 0$ also satisfy this expression and we've got final expression and η_i

$$\eta_i = \min\left\{1, \ \frac{\tau}{\max(\bar{N}_f)} (2^i - 1)\right\} \tag{A11}$$

482 Here we need to truncate the series when $\eta_i > 1$.

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