# MATLAB GMMNLSE Solver User Guide

Zachary Ziegler<sup>1</sup>, Logan Wright<sup>1</sup> and Frank Wise<sup>1</sup>

<sup>1</sup>School of Applied and Engineering Physics, Cornell University, Ithaca, New York 14853, USA

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# I don't want to read this, how do I run a simulation?

Fair enough. Run the script Examples/MM Propagation/GMMNLSE\_driver\_OM 4\_lowpower\_cpu.m or Examples/MM Propagation/GMMNLSE\_driver\_OM4\_lowpower\_gpu.m. All you need to run this is MATLAB, but if you want to use the GPU version (which is highly recommended), see the Setup section below.

### 1 Overview

This package contains all the code needed to simulate multimode pulse propagation in multimode optical fiber, based around the Massively Parallel Algorithm implemented efficiently on the GPU. The core of the package steps a multimode field through a number of small steps  $\Delta z$  in parallel; the rest of the code provides user-friendly functions to propagate over an arbitrary number of steps, to set up the simulations, and to model out-of-cavity auxiliary components such as saturable absorbers or filters. While a modest speed-up is achieved through the use of the MPA algorithm, the use of the GPU alone gives a significant speed-up compared to the same calculation on the CPU. Using a GPU is not required, therefore, but it is strongly suggested if possible. Run on a modern GPU, this code enables realistic multimode laser cavity simulations and multimode propagation simulations that are faster and include more modes than was previously possible.

Specifically, the code solves the Generalized Multimode Nonlinear Schrödinger equation (GMMNLSE) [1]:

$$\frac{\partial A_{p}}{\partial z} = i(\beta_{0}^{(p)} - \Re[\beta_{0}^{(0)}]) A_{p} - (\beta_{1}^{(p)} - \Re[\beta_{1}^{(0)}]) \frac{\partial A_{p}}{\partial t} + i \sum_{n \geq 2} \frac{\beta_{n}^{(p)}}{n!} \left( i \frac{\partial}{\partial t} \right)^{n} A_{p} 
+ i \frac{n_{2} \omega_{0}}{c} \left( 1 + \frac{i}{\omega_{0}} \frac{\partial}{\partial t} \right) \sum_{l,m,n} \left\{ (1 - f_{R}) S_{plmn}^{K} A_{l} A_{m} A_{n}^{*} + f_{R} S_{plmn}^{R} A_{l} [h * (A_{m} A_{n}^{*})] \right\}$$
(1)

Where  $A_p$  is the field envelope of the pth mode,  $\beta_n^{(p)}$  is the nth order dispersion coefficient for the pth mode,  $n_2$  is the nonlinear index coefficient,  $\omega_0$  is the center angular frequency,  $f_R$  is the fractional Raman contribution to the nonlinearity,  $S_{plmn}^K$  and  $S_{plmn}^R$  are the modal overlap factors for the Kerr term and the Raman term, respectively, h is the Raman response function, t is time in a reference frame moving with the fundamental mode at the central frequency, and z is the distance along the fiber.

## 2 Core files

The following is a list of the files included in the solver package, along with a short description of what they do. The various step methods are described first, followed by the propagation method, followed by the auxiliary functions. More details about the exact input and output can be found in comments at the top of each function.

### 2.1 Step methods

Depending on the options (whether or not to use MPA, and which gain model to use) one of the following functions will be called to perform a single step along the fiber:

### GMMNLSE\_ss\_step\_nogain

Performs one small step  $\Delta z$ , using the split-step method. The dispersion is applied during the first half step exactly, then the nonlinearity is applied over the full time step using a 4th order Runge-Kutta method, and finally the dispersion is applied again over the last half step. No gain is accounted for.

### GMMNLSE\_ss\_step\_SMgain

The same as GMMNLSE\_ss\_step\_nogain except gain is applied in a single-mode fashion during the linear half steps. A single-mode fashion means that gain saturation only takes into account the total energy, i.e. it is not mode-resolved.

### GMMNLSE\_ss\_step\_taylorgain

The same as GMMNLSE\_ss\_step\_nogain except gain is applied in the nonlinear step, using a Taylor series expansion of the gain term. After Taylor expanding the spatially dependant gain saturation and decomposing the equation into a modal basis, the first term in the Taylor series can be calculated with the same complexity as the nonlinear term. See the slides on the treatment of gain for more details.

### GMMNLSE\_ss\_step\_newgain

The same as GMMNLSE\_ss\_step\_nogain except gain is applied once per step, after the final half-step of dispersion, taking into account the full spatially dependant nature of the gain saturation using an implicit method. See the slides on the treatment of gain for more details.

### GMMNLSE\_MPA\_step\_nogain

Performs M small steps of size  $\Delta z$  in parallel, using the Massively Parallel Algorithm. The dispersion is again accounted for exactly, but now the nonlinear phase accumulation is calculated for each small step in parallel and then integrated to calculate the full nonlinear phase accumulation. This process is iterated until a desired tolerance in the change of the solution is met. See the description by Pavel Lushnikov.

### GMMNLSE\_MPA\_step\_SMgain

The same as GMMNLSE\_MPA\_step\_nogain except gain is applied in a single-mode fashion, i.e. with the saturation only taking into account the total energy. Here the gain is applied as a part of the linear operator, because the term is only nonlinear with respect to the total energy, which can be calculated exactly for each of the M small steps.

### GMMNLSE\_MPA\_step\_newgain

The same as GMMNLSE\_MPA\_step\_nogain except gain is applied using the new implicit method which takes into account the full spatially dependant nature of the gain saturation. Because the calculation involves almost exactly the same computations as the nonlinear term, the transfer matrix is computed at the same time. This is therefore not a true split step approach and some extra terms are necessarily missing, but if the step size is small enough these extra terms should be small.

## 2.2 Propagation (multiple steps)

### GMMNLSE\_propagate

This function takes in parameters of the fiber, parameters of the simulation, and an initial condition, and propagates the field through a number of steps to the end of the fiber. It first performs some checks to ensure the input is valid, formats the S^K and S^R tensors to maximize efficiency and precomputes as much as possible (the dispersion term, part of the gain term, and the Raman response). Then it enters the main loop, calling the appropriate step function each iteration.

## 2.3 Pre-propagation functions

### solve\_for\_modes

Given an index profile function and the parameters of the profile, this function solves Maxwell's equations to find the modes of the fiber and their propagation constants. Specifically, it calculates the modes and propagation constants over a range of frequency, which is required to calculate the dispersion of each mode.

### calc\_dispersion

After calculating the propagation constants for each mode with solve\_for\_modes, this function imports the results, approximates the dispersion with a polynomial, and saves the dispersion coefficients.

### calc\_SRSK\_tensors

After calculating the modes with solve\_for\_modes, this function calculates the overlap tensors as in [1]. The polarization of the modes is assumed to be either linear or circular, in which case the SK and SR tensors are constant multiples of one another. In the more general case the code would have to be slightly modified for other polarization states. The modes used in this simulation package are approximated as scalar modes.

### 2.4 Cavity simulation auxiliary functions

### calc\_filter\_matrix

Calculates the coefficients of the matrix that should be applied to a multimode field to get the multimode field after a gaussian spatial filter.

### spatial\_filter\_moderesolved

Using the transfer matrix calculated with calc\_filter\_matrix, apply the filter to a given input field.

### calc\_mode\_coupling\_matrix

Calculates the coefficients of the matrix that projects from a basis of one fiber's guided modes to a basis of a different fiber's guided modes. To apply the projection, multiply the field by the transfer matrix as with the spatial filter.

### gaussian\_spectral\_filter

Apply a Gaussian spectral filter with a given FWHM to each mode of an input field.

### recompose\_into\_space

Given the spatial mode profiles and an input multimode field, produce the full 3D spatiotemporal field.

### saturable\_absorber\_action\_simple

Apply an ideal temporal saturable absorber with a given saturation power and modulation depth to each mode in the field.

### saturable\_absorber\_action\_3d

Apply an ideal spatio-temporal saturable absorber with a given saturation power and modulation depth to the full field. The multimode field is first composed into the full 2D spatial field at each time grid point, the saturable absorber is applied to the spatial field, and then the spatial field is recomposed back into the modal basis.

# 3 Setup

Run on the CPU, the GMMNLSE solver only requires the base version of MATLAB without any extra packages. So far, the code has been tested on MATLAB 2016b.

This GMMNLSE solver has also been optimized to run efficiently on Nvidia GPUs (due to CUDA support). The code has been tested on 800-series, 900-series, and 1000-series GPUs, although it should work for older GPUs as well. In general the newer GPUs will give better performance, in addition to having more memory. The GPU memory primarily determines the number of modes; for a sense of scale, on the Titan X with M=10 and  $N=2^{13}$ , the GPU can support 125 modes. So almost all modern GPUs will have enough memory to simulate a reasonable number of modes .

Using a GPU with MATLAB requires the MATLAB Parallel Computing Toolbox. Furthermore, because the solver uses a small section of custom CUDA code to greatly accelerate the simulations, the Nvidia CUDA Toolkit is also required. To enable GPU simulations, the following steps should be taken:

- 1. Install the Nvidia drivers. If the GPU has been used before these will probably already be installed. If on Windows, install Nvidia's GeForce Experience software and follow the steps to download the drivers. If on Linux, you will get the best performance with Nvidia's proprietary drivers which can be installed either via most package managers (recommended) or directly from Nvidia's website.
- 2. Install a C++ compiler, if you do not have one already. The CUDA compiler driver nvcc relies on a pre-existing C++ compiler, and will not function without one. If on Windows, the free Visual Studio Community is recommended. If on Linux, gcc is recommended and should be installed already.
- 3. Install the Nvidia CUDA Toolkit. You can find the download page here, and the installation instruction guides here for Linux and here for Windows. This will install the Nvidia CUDA compiler driver nvcc.
- 4. Add the C++ compiler and nvcc to the PATH. On Windows nvcc.exe is located by default in C:\Program Files\NVIDIA GPU Computing Toolkit\CUDA\v8.0\bin or similar. If using Visual Studio the C++ compiler cl.exe may be located in C:\Program Files (x86)\Microsoft Visual Studio 14.0\VC\bin, although the path may be different for different versions. These can be added to the PATH by adding them to the list located at Control Panel → System and Security → System → Advanced system settings → Environment Variables in Windows 10 and Windows 8. In Windows 7, right click on 'My Computer' then choose Properties → Advanced system settings → Environment Variables

## 4 Use

Some examples of how to use the tools included in this package to simulate pulses according to the GMMNLSE can be found in the Examples folder. Generally, the workflow is as follows:

- 1. Choose the type and parameters of fiber, and solve for the modes with <code>solve\_for\_modes.m</code>. The type of fiber is selected through the function used the generate the refractive index profile. Functions to build common fiber types such as step index fiber or GRIN fiber are included, and the specific fiber parameters can be modified. The modes can be solved for first at a single wavelength, to get an idea of the number of guided modes and shape, and then again over a range of wavelengths which is needed to calculate the dispersion.
- 2. Calculate the dispersion coefficients with calc\_dispersion.m. This script takes propagation constants from the modes that have been previously calculated, fits them to a polynomial, and extracts the dispersion coefficients.
- 3. Calculate the mode overlap tensors with calc\_SRSK\_tensors.m. This script takes the spatial fields of the modes at a single wavelength and computes the modes overlap tensors according to [1].
- 4. Build the system to simulate in a file similar to any of the GMMNLSE\_driver.m files. This is a good place to set the simulation parameters and then run the GMMNLSE\_propagate function to do the actual propagation. The other auxiliary function such as filters can be called in this script as well.

### 5 Fast Fourier Transform convention

As the code included here uses a spectral method to solve the GMMNLSE, the FFT plays an important role in the simulation code. The optics community tends to use the convention where the Fourier transform pair is defined as

$$\mathcal{F}_{optics}\{f(t)\} = \int f(t)e^{i\omega t}dt \tag{2}$$

$$\mathcal{F}_{optics}^{-1}\{f(\omega)\} = \frac{1}{2\pi} \int f(\omega)e^{-i\omega t}d\omega$$
 (3)

whereas MATLAB follows the more traditional definition where the negative sign in the exponent is included in the forward transform. Ultimately as long as the convention is consistent it does not matter physically which convention is used, however for easy comparison with the rest of the optics community in the frequency domain the optics convention is used in this simulation package. To accomplish this in MATLAB, the forward transform is performed with ifft and the backward transform is performed with fft. This does not affect the results of the simulation, but when visualizing the output it is suggested to retain the convention as well.

# 6 Massively Parallel Algorithm (MPA)

The core simulation code in this package uses the MPA algorithm to take a number of small steps together in parallel. The GMMNLSE MPA algorithm was developed by Pavel Lush-

nikov, Department of Mathematics and Statistics, University of New Mexico, as an extension of a similar algorithm for the single mode NLSE. In short, the algorithm treats the linear dispersion term exactly in the Fourier domain and solves for the nonlinear phase accumulation over a number of small steps  $\Delta z$  in parallel, initially assuming the nonlinear phase accumulated over the whole step is small. This process is then iterated until a self-consistent solution with a desired level of accuracy is obtained. See the notes in the Documentation folder for more details.

# 7 Treatment of spatial gain saturation

This solver supports three ways of implementing gain saturation, in addition to a stepper without gain at all: single-mode-like gain saturation, Taylor expansion-based gain saturation, and a new implicit gain saturation method. All methods model spatially dependent gain saturation the same way, but they differ in how they map the spatial saturation onto the modal basis. The single-mode gain saturation does not take into account the spatial dependence of gain saturation, and instead saturates the gain based on the total energy in all modes. The Taylor expansion gain saturation Taylor expands the saturation term and keeps only up to the first order. Finally, the new implicit gain method solves the gain term alone in a modal basis, using an implicit method, and then assumes that this solution is valid even in the presence of the other terms of the GMMNLSE. For more details, see the notes in the Documentation folder.

# 8 Acknowledgements

The varycolor and symodes functions are used from the MATLAB file exchange.

# 9 Contact

Technical questions may be directed to Zachary Ziegler: zmz4@cornell.edu

## References

[1] Peter Horak and Francesco Poletti. Multimode Nonlinear Fibre Optics: Theory and Applications. In Moh Yasin, editor, *Recent Prog. Opt. Fiber Res.* InTech, 2012.