# On Computing Gauss-Kronrod Quadrature Formulae\*

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Abstract. We discuss the use of Newton's method for computing Gauss-Kronrod quadrature formulae from modified moments. The underlying nonlinear maps are analyzed from the point of view of numerical condition. A method is indicated of computing the polynomial whose zeros are the Kronrod nodes. Examples include Gauss-Kronrod formulae for integrals with a logarithmic and algebraic singularity at one endpoint. Pertinent numerical results are tabulated in the supplements section at the end of this issue.

1. Introduction. Given a positive measure  $d\sigma(t)$  on the real line, whose support contains infinitely many points and all of whose moments exist, we call

(1.1) 
$$\int_{\mathbf{R}} f(t) d\sigma(t) = \sum_{\nu=1}^{n} \sigma_{\nu} f(\tau_{\nu}) + \sum_{\mu=1}^{n+1} \sigma_{\mu}^{*} f(\tau_{\mu}^{*}) + R_{n}(f)$$

a Gauss-Kronrod quadrature formula if  $\tau_{\nu} = \tau_{\nu}^{(n)}$  are the Gaussian nodes associated with  $d\sigma$  and the nodes  $\tau_{\mu}^{*}$  and weights  $\sigma_{\nu}$ ,  $\sigma_{\mu}^{*}$  are chosen so as to maximize the degree of exactness of (1.1). Since there are 3n+2 unknowns, we can achieve degree of exactness d=3n+1, i.e.,  $R_n(f)=0$  whenever  $f=\mathbf{P}_{3n+1}$ . It is well known, in fact, that the nodes  $\tau_{\mu}^{*}$  must be the zeros of  $\pi_{n+1}^{*}$ , the (monic) polynomial of degree n+1 satisfying the orthogonality property

(1.2) 
$$\int_{\mathbb{R}} \pi_{n+1}^*(t) \pi_n(t) t^i d\sigma(t) = 0, \qquad i = 0, 1, \dots, n,$$

where  $\pi_n(\cdot) = \pi_n(\cdot; d\sigma)$  is the *n*th degree orthogonal polynomial belonging to the measure  $d\sigma$ . Note that the measure  $d\sigma^*(t) = \pi_n(t; d\sigma)d\sigma(t)$ , with respect to which  $\pi_{n+1}^*$  is orthogonal, changes sign. We cannot expect, therefore, that  $\pi_{n+1}^*$  has all its zeros necessarily real.

We are interested here only in Gauss-Kronrod formulae (1.1) with *real* nodes  $\tau_{\mu}^{*}$ , all contained in [a, b]—the smallest interval containing the support of  $d\sigma$ —and with *positive* weights  $\sigma_{\mu}^{*}$ . Then the interlacing property holds (Monegato [17, Theorem 1]),

$$(1.3) a \leq \tau_{n+1}^* < \tau_n < \tau_n^* < \tau_{n-1} < \cdots < \tau_1 < \tau_1^* \leq b.$$

Our concern is with the actual computation of these nodes and the corresponding weights (provided they exist), given the integer  $n \ge 1$  and the positive measure  $d\sigma$ .

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Constructive methods for this problem have been considered before. For the Legendre measure  $d\sigma(t) = dt$  on [-1, 1], Kronrod [16] in his original work obtains  $\pi_{n+1}^*$  in power form by solving (1.2). Patterson [22] expands  $\pi_{n+1}^*$  in Legendre polynomials, while Piessens and Branders [23] use Chebyshev polynomials, the nodes  $\tau_{\mu}^*$  each time being computed by an appropriate rootfinding procedure. The weights can be obtained by interpolation. Monegato [17], [19], Baratella [1], and Dagnino and Fiorentino [3] use similar procedures to compute  $\pi_{n+1}^*$  for Gegenbauer measures and other classical measures. Kahaner et al. [13], in the case of the Laguerre measure, and assuming  $k \ge n + 1$  Kronrod nodes  $\tau_u^*$  in (1.1), expand  $L_n \pi_k^*$  in Laguerre polynomials  $L_i$ . All these methods require three distinct phases to obtain (1.1): The computation of the appropriate polynomial, say  $\pi_{n+1}^*$ , finding the zeros of  $\pi_{n+1}^*$ , and computing the weights  $\sigma_{\nu}$ ,  $\sigma_{\mu}^*$ . An entirely different approach is taken by Kautsky and Elhay [15] and Elhay and Kautsky [5], who compute the nodes  $\tau_u^*$  as eigenvalues of a certain matrix derived by matrix decomposition methods—an approach which extends the well-known method of Golub and Welsch [12] for ordinary Gaussian quadratures and its extension by Golub and Kautsky [11]. The weights are then obtained using general methods for constructing interpolatory quadrature rules [14], [6].

Here we propose to compute (1.1) directly by solving a set of nonlinear equations that express the exactness of (1.1) for a given set of 3n + 2 polynomials. These polynomials are chosen so as to generate a well-conditioned problem. The respective system of nonlinear equations is solved by Newton's method, careful attention being given to the choice of initial approximations and to monitoring the progress of the iteration. We make no claims for our method to be superior in any way to other methods, but merely demonstrate its feasibility and stability. It would indeed be interesting to undertake a detailed comparative study of the various methods now in use.

While aiming directly at the unknown quantities is certainly a virtue if our method is successful, it is less than satisfactory otherwise, since no information about  $\pi_{n+1}^*$  is generated when the method fails. Nevertheless, if one wishes to examine the polynomial  $\pi_{n+1}^*$ , one can express it in terms of the orthogonal polynomials  $\{\pi_k(\cdot;d\sigma)\}$  and obtain the coefficients by solving a triangular system of linear equations.

The implementation of Newton's method for computing the Gauss-Kronrod rule (1.1) is discussed in Section 2. In Section 3 we study the condition of the underlying problem. Section 4 deals with the computation of the polynomial  $\pi_{n+1}^*$ . Examples will be given in Section 5, and numerical results for  $d\sigma(t) = t^{\alpha} \ln(1/t) dt$  on [0, 1],  $\alpha = 0, \pm \frac{1}{2}$ , are tabulated in the supplements section at the end of this issue.

For additional questions related to Gauss-Kronrod quadrature we refer the reader to surveys by Monegato [20], [21].

2. The Computation of Gauss-Kronrod Rules by Newton's Method. The Gaussian nodes  $\tau_{\nu}$  for the measure  $d\sigma$  can be computed by well-known methods; see, e.g., [8, p. 290]. We assume therefore that they are available. For a given system of monic polynomials  $\{p_k(\cdot)\}$ , with deg  $p_k = k$ ,  $k = 0, 1, 2, \ldots$ , we further assume that the

first 3n + 2 modified moments of  $d\sigma$ ,

(2.1) 
$$m_k = \int_{\mathbf{R}} p_k(t) \, d\sigma(t), \qquad k = 0, 1, \dots, 3n + 1,$$

are known. The system of nonlinear equations defining the Gauss-Kronrod formula (1.1) then is

(2.2) 
$$\sum_{\nu=1}^{n} \sigma_{\nu} p_{k}(\tau_{\nu}) + \sum_{\mu=1}^{n+1} \sigma_{\mu}^{*} p_{k}(\tau_{\mu}^{*}) = m_{k}, \qquad k = 0, 1, \dots, 3n+1.$$

We propose to solve this system for the 3n + 2 unknowns  $\tau_{\mu}^*$ ,  $\sigma_{\nu}$ ,  $\sigma_{\mu}^*$  by Newton's method. A number of practical issues need to be addressed.

First the choice of polynomials  $p_k$ . We assume that they satisfy a three-term recurrence relation,

(2.3) 
$$p_{-1}(t) = 0, p_0(t) = 1, p_{k+1}(t) = (t - a_k)p_k(t) - b_k p_{k-1}(t), k = 0, 1, 2, \dots,$$

with known coefficients  $a_k$ ,  $b_k$ . This makes the evaluation of these polynomials, and of their derivatives, which is required in the computation of the left-hand side of the system (2.2) and its Jacobian, easy and straightforward. If, in addition, the polynomials  $\{p_k\}$  are orthogonal with respect to some given measure ds(t), then this computation is also numerically stable. If  $d\sigma$  is one of the classical measures, the simplest choice is  $p_k(\cdot) = \pi_k(\cdot; d\sigma)$ , the corresponding recursion coefficients  $a_k = \alpha_k(d\sigma)$ ,  $b_k = \beta_k(d\sigma)$  then being explicitly known; the modified moments become

(2.4) 
$$m_k = \int_{\mathbf{R}} \pi_k(t; d\sigma) d\sigma(t) = \beta_0 \delta_{0,k}, \qquad k = 0, 1, 2, ...,$$

where  $\beta_0 = \int_{\mathbf{R}} d\sigma(t)$  and  $\delta_{0,k}$  is the Kronecker delta,  $\delta_{0,0} = 1$ ,  $\delta_{0,k} = 0$  if k > 0. For nonclassical measures  $d\sigma$ , the choice of  $p_k$  is usually dictated by the necessity of being able to compute the modified moments (2.1).

The next important issue is the computation of the initial approximations for the Kronrod nodes  $\tau_{\mu}^{*}$  and for the weights  $\sigma_{\nu}$ ,  $\sigma_{\mu}^{*}$ . Since we are interested only in Gauss-Kronrod formulae with nodes in [a, b] and with positive weights  $\sigma_{\mu}^{*}$ , we can assume the interlacing property (1.3). This suggests as initial approximations  $\hat{\tau}_{\mu}^{*}$ ,  $2 \le \mu \le n$ , the midpoints

(2.5) 
$$\mathring{\tau}_{\mu}^* = \frac{1}{2} (\tau_{\mu-1} + \tau_{\mu}), \qquad \mu = 2, 3, \dots, n,$$

between the Gaussian nodes  $\tau_{\mu-1}$  and  $\tau_{\mu}$ . The choice of the initial approximations  $\mathring{\tau}_1^*$  and  $\mathring{\tau}_{n+1}^*$  is less obvious. If the endpoint b is finite, we select a number of trial values  $\mathring{\tau}_1^*$  equally spaced between  $\tau_1$  and b. If a is also finite, we do the same for  $\mathring{\tau}_{n+1}^*$  in the interval  $(a,\tau_n)$  and let both  $\mathring{\tau}_{n+1}^*$  and  $\mathring{\tau}_1^*$  move inward simultaneously, or, if necessary, let them move independently from one another. If an endpoint, say b, is infinite, we select  $\mathring{\tau}_1^* = \tau_1 + (\tau_1 - \tau_2)$ , or try a number of values  $\mathring{\tau}_1^*$  equally spaced between  $\tau_1$  and, say,  $\tau_1 + 2(\tau_1 - \tau_2)$ . For each set of initial approximations  $\mathring{\tau}_\mu^*$  we compute corresponding approximations  $\mathring{\sigma}_\nu$ ,  $\mathring{\sigma}_\mu^*$  to the weights by solving the first 2n+1 equations in (2.2), where  $\tau_\mu^*$  is replaced by  $\mathring{\tau}_\mu^*$ . Since the matrix of this system forms part of the Jacobian matrix used in the first Newton step, the only overhead of this computation is the solution of the linear system.

Each Newton step is monitored as to the location of the iterates of  $\tau_{\mu}^{*}$ . If at any stage one of these iterates falls outside the interval [a, b] (if either a or b is finite), the iteration is terminated and restarted with a new set of initial approximations. The same action is taken if the number of Newton iterates exceeds a preset limit. If none of the initial approximations leads to a convergent process, the attempt of computing (1.1) is declared a failure. This is usually an indication (not a proof!) that the desired Gauss-Kronrod rule does not exist.

If  $d\sigma(t) = d\sigma(-t)$  is an even measure and the support of  $d\sigma$  is symmetric with respect to the origin, then both the Gaussian nodes  $\tau_{\nu}$  and the Kronrod nodes  $\tau_{\mu}^*$  are located symmetrically with respect to the origin, and weights corresponding to symmetric nodes are equal. As a result, (2.2) is trivially true if  $p_k(t)$  is an odd polynomial. If we choose for  $\{p_k\}$  a system of polynomials satisfying

$$(2.6) p_{\nu}(-t) = (-1)^{k} p_{\nu}(t), k = 0, 1, 2, \dots,$$

the system (2.2), therefore, is equivalent to the system

(2.7e) 
$$\sum_{\nu=1}^{n/2} \sigma_{\nu} p_{2k}(\tau_{\nu}) + \sum_{\mu=1}^{n/2} \sigma_{\mu}^{*} p_{2k}(\tau_{\mu}^{*}) + \frac{1}{2} \sigma_{(n/2)+1}^{*} p_{2k}(0) = \frac{1}{2} m_{2k}, \\ k = 0, 1, \dots, 3n/2,$$

if n is even, and to the system

(2.70) 
$$\sum_{\nu=1}^{\lfloor n/2 \rfloor} \sigma_{\nu} p_{2k}(\tau_{\nu}) + \frac{1}{2} \sigma_{\lfloor n/2 \rfloor+1} p_{2k}(0) + \sum_{\mu=1}^{\lfloor n/2 \rfloor+1} \sigma_{\mu}^{*} p_{2k}(\tau_{\mu}^{*}) = \frac{1}{2} m_{2k}, \\ k = 0, 1, \dots, (3n+1)/2,$$

if n is odd. This in effect reduces the size of the problem by a factor of 2.

3. The Condition of the Underlying Problem. Let  $m^T = [m_0, m_1, \ldots, m_{3n+1}]$  be the vector of modified moments, and  $\gamma^T = [\sigma_1, \ldots, \sigma_n, \sigma_1^*, \ldots, \sigma_{n+1}^*, \tau_1^*, \ldots, \tau_{n+1}^*]$  the vector of the weights and Kronrod nodes of the Gauss-Kronrod formula (1.1). The procedure of Section 2 is an attempt of carrying out the nonlinear map

$$G_n: \mathbf{R}^{3n+2} \to \mathbf{R}^{3n+2} \quad m \to \gamma,$$

where the Gauss-Kronrod formula is assumed to have real nodes. We now wish to examine the sensitivity of the map  $G_n$  to perturbations in the modified moments. The development parallels the treatments given for Gaussian formulae in Gautschi [8], [9]; see also Gautschi [10, Section 5].

We assume that the polynomials  $\{p_k\}$  defining the modified moments are orthogonal on the real line with respect to some measure ds. The support of this measure normally coincides with the support of  $d\sigma$ , but does not have to. We define normalized modified moments by

(3.1) 
$$\tilde{m}_k = d_k^{-1} m_k, \qquad d_k^2 = \int_{\mathbb{R}} p_k^2(t) \, ds(t),$$

and consider, in place of  $G_n$ , the map

(3.2) 
$$\tilde{G}_n \colon \mathbf{R}^{3n+2} \to \mathbf{R}^{3n+2} \quad \tilde{m} \to \gamma,$$

where  $\tilde{m}^T = [\tilde{m}_0, \tilde{m}_1, \dots, \tilde{m}_{3n+1}]$ . We analyze the sensitivity of  $\tilde{G}_n$  by computing the Frobenius norm of the Jacobian matrix,  $J_{\tilde{G}_n}$ , of the map  $\tilde{G}_n$ .

The basic equations can be written in the form

$$\Phi(\gamma) = \tilde{m}.$$

where

(3.4) 
$$\Phi_{k}(\gamma) = d_{k}^{-1} \left\{ \sum_{\nu=1}^{n} \sigma_{\nu} p_{k}(\tau_{\nu}) + \sum_{\mu=1}^{n+1} \sigma_{\mu}^{*} p_{k}(\tau_{\mu}^{*}) \right\},$$

$$k = 0, 1, 2, \dots, 3n+1.$$

Since the map  $\tilde{G}_n$  amounts to solving (3.3) for  $\gamma$ , the Jacobian of  $\tilde{G}_n$  is the inverse of the Jacobian  $\partial \Phi / \partial \gamma$  of  $\Phi$ ,

$$J_{\tilde{G}_n} = (\partial \Phi / \partial \gamma)^{-1}.$$

An elementary computation shows that

$$\partial \Phi / \partial \gamma = D^{-1} P \Sigma^*,$$

where  $D = \text{diag}(d_0, d_1, \dots, d_{3n+1}), \ \Sigma^* = \text{diag}(1, \dots, 1, 1, \dots, 1, \sigma_1^*, \dots, \sigma_{n+1}^*),$  and

$$P = \begin{bmatrix} p_0(\tau_1) & \cdots & p_0(\tau_n) & p_0(\tau_1^*) & \cdots & p_0(\tau_{n+1}^*) & p_0'(\tau_1^*) & \cdots & p_0'(\tau_{n+1}^*) \\ p_1(\tau_1) & \cdots & p_1(\tau_n) & p_1(\tau_1^*) & \cdots & p_1(\tau_{n+1}^*) & p_1'(\tau_1^*) & \cdots & p_1'(\tau_{n+1}^*) \\ \cdots & \cdots & \cdots & \cdots & \cdots & \cdots & \cdots \\ p_{3n+1}(\tau_1) & \cdots & p_{3n+1}(\tau_n) & p_{3n+1}(\tau_1^*) & \cdots & p_{3n+1}(\tau_{n+1}^*) & p_{3n+1}'(\tau_1^*) & \cdots & p_{3n+1}'(\tau_{n+1}^*) \end{bmatrix}.$$

Therefore,

(3.8) 
$$J_{\tilde{G}_n} = (\Sigma^*)^{-1} P^{-1} D.$$

For the inversion of P, define  $g_{\nu}$ ,  $h_{\mu}$ ,  $k_{\mu}$  to be the elementary Hermite interpolation polynomials of degree 3n + 1, belonging to the nodes  $\tau_{\nu}$  and  $\tau_{\mu}^{*}$ , defined by

$$g_{\nu}(\tau_{\lambda}) = \delta_{\nu\lambda}, \quad \lambda = 1, 2, \dots, n$$

$$g_{\nu}(\tau_{\mu}^{*}) = 0, \quad g_{\nu}'(\tau_{\mu}^{*}) = 0, \quad \mu = 1, 2, \dots, n + 1$$

$$p = 1, 2, \dots, n + 1$$

$$h_{\mu}(\tau_{\lambda}) = 0, \quad \lambda = 1, 2, \dots, n$$

$$h_{\mu}(\tau_{\nu}^{*}) = \delta_{\mu\nu}, \quad h_{\mu}'(\tau_{\nu}^{*}) = 0, \quad \nu = 1, 2, \dots, n + 1$$

$$k_{\mu}(\tau_{\lambda}) = 0, \quad \lambda = 1, 2, \dots, n$$

$$k_{\mu}(\tau_{\nu}^{*}) = 0, \quad k_{\mu}'(\tau_{\nu}^{*}) = \delta_{\mu\nu}, \quad \nu = 1, 2, \dots, n + 1$$

$$\mu = 1, 2, \dots, n + 1.$$

Writing

$$g_{\nu}(t) = \sum_{\rho=1}^{3n+2} a_{\nu\rho} p_{\rho-1}(t), \quad h_{\mu}(t) = \sum_{\rho=1}^{3n+2} b_{\mu\rho} p_{\rho-1}(t), \quad k_{\mu}(t) = \sum_{\rho=1}^{3n+2} c_{\mu\rho} p_{\rho-1}(t),$$

it is easily seen that

(3.10) 
$$P^{-1} = \begin{bmatrix} A \\ B \\ C \end{bmatrix}, \quad A = [a_{\nu\rho}], \quad B = [b_{\mu\rho}], \quad C = [c_{\mu\rho}].$$

By a computation similar to the one in [8, pp. 303-304] one finds from (3.8) and (3.10) that the Frobenius norm of  $J_{\tilde{G}_n}$  is given by

(3.11) 
$$\|J_{\tilde{G}_n}\|_F = \left\langle \int_{\mathbf{R}} \left[ \sum_{\nu=1}^n g_{\nu}^2(t) + \sum_{\mu=1}^{n+1} \left( h_{\mu}^2(t) + \frac{1}{\sigma_{\mu}^{*2}} k_{\mu}^2(t) \right) \right] ds(t) \right\rangle^{1/2}.$$

Its magnitude, therefore, is critically influenced by the magnitude of the polynomial

(3.12) 
$$f_n(t) = \sum_{\nu=1}^n g_{\nu}^2(t) + \sum_{\mu=1}^{n+1} \left( h_{\mu}^2(t) + \frac{1}{\sigma_{\mu}^{*2}} k_{\mu}^2(t) \right)$$

on the support of ds. The degree of  $f_n$  is 6n + 2. The integral in (3.11) can therefore be computed exactly (up to rounding errors) by a (3n + 2)-point Gaussian quadrature rule belonging to the measure ds.

Explicit forms of the polynomials  $g_{\nu}$ ,  $h_{\mu}$ ,  $k_{\mu}$  can easily be given in terms of the fundamental Lagrange polynomials

$$l_{\nu}(t) = \prod_{\substack{\lambda=1\\\lambda\neq\nu}}^{n} \frac{t-\tau_{\lambda}}{\tau_{\nu}-\tau_{\lambda}}, \qquad l_{\mu}^{*}(t) = \prod_{\substack{\lambda=1\\\lambda\neq\mu}}^{n+1} \frac{t-\tau_{\lambda}^{*}}{\tau_{\mu}^{*}-\tau_{\lambda}^{*}}$$

belonging to the nodes  $\tau_{\nu}$  and  $\tau_{\mu}^*$ , respectively. One obtains

$$g_{\nu}(t) = l_{\nu}(t) \left[ \frac{\pi_{n+1}^{*}(t)}{\pi_{n+1}^{*}(\tau_{\nu})} \right]^{2}, \qquad \nu = 1, 2, \dots, n,$$

(3.13) 
$$h_{\mu}(t) = \frac{\pi_{n}(t)}{\pi_{n}(\tau_{\mu}^{*})} \left[ l_{\mu}^{*}(t) \right]^{2} \left\{ 1 - \left( t - \tau_{\mu}^{*} \right) \left[ \frac{\pi'_{n}(\tau_{\mu}^{*})}{\pi_{n}(\tau_{\mu}^{*})} + 2 l_{\mu}^{*'}(\tau_{\mu}^{*}) \right] \right\},$$

$$\mu = 1, 2, \dots, n+1,$$

$$k_{\mu}(t) = \frac{\pi_{n}(t)}{\pi_{n}(\tau_{\mu}^{*})} [l_{\mu}^{*}(t)]^{2} (t - \tau_{\mu}^{*}), \qquad \mu = 1, 2, \dots, n+1,$$

where  $\pi_n(\cdot) = \pi_n(\cdot; d\sigma)$  and  $\pi_{n+1}^*(t) = \prod_{\lambda=1}^{n+1} (t - \tau_{\lambda}^*)$ . We also note the following properties of  $f_n$ , which follow directly from (3.9) and (3.12),

(3.14) 
$$f_{n}(t) > 0 \quad \text{for all } t \in \mathbb{R},$$

$$f_{n}(\tau_{\nu}) = 1, \quad f'_{n}(\tau_{\nu}) = 2g'_{\nu}(\tau_{\nu}), \qquad \nu = 1, 2, \dots, n,$$

$$f_{n}(\tau_{\mu}^{*}) = 1, \quad f'_{n}(\tau_{\mu}^{*}) = 0, \qquad \mu = 1, 2, \dots, n + 1.$$

These conditions, of course, are not sufficient to determine the polynomial  $f_n$ . There are 2n + 1 degrees of freedom left, which allow  $f_n$  considerable room for movement.

**4. Computation of the Polynomial**  $\pi_{n+1}^*$ . Expressing the polynomial  $\pi_{n+1}^*$  in terms of the orthogonal polynomials  $\pi_k(\cdot) = \pi_k(\cdot; d\sigma)$ ,

(4.1) 
$$\pi_{n+1}^*(t) = \pi_{n+1}(t) + c_0 \pi_n(t) + c_1 \pi_{n-1}(t) + \cdots + c_n \pi_0(t),$$

and replacing the powers  $t^i$  in (1.2) by the polynomials  $\pi_i(t)$ , one obtains the conditions

$$\int_{\mathbf{R}} \left[ \pi_{n+1}(t) + \sum_{k=0}^{n} c_k \pi_{n-k}(t) \right] \pi_n(t) \pi_i(t) \, d\sigma(t) = 0, \qquad i = 0, 1, \dots, n,$$

hence the linear system

(4.2) 
$$\sum_{k=0}^{n} a_{ik} c_k = b_i, \qquad i = 0, 1, \dots, n,$$

for the coefficients  $c_k$ , where

(4.3) 
$$a_{ik} = \int_{\mathbf{R}} \pi_i(t) \pi_{n-k}(t) \pi_n(t) d\sigma(t), \qquad i, k = 0, 1, \dots, n,$$
$$b_i = -\int_{\mathbf{R}} \pi_i(t) \pi_{n+1}(t) \pi_n(t) d\sigma(t), \qquad i = 0, 1, \dots, n.$$

By orthogonality,  $a_{ik} = 0$  if i < k, so that the matrix  $A = [a_{ik}]$  is lower triangular, and  $a_{ii} = \int_{\mathbb{R}} \pi_n^2(t) d\sigma(t) > 0$ , so that A is nonsingular. The solution of (4.2), therefore, can be effected by forward substitution.

The coefficients  $a_{ik}$  and  $b_i = -a_{i,-1}$  satisfy a two-dimensional recursion relation, which could be used for their computation (see [2]). Noting, however, that the integrands in (4.3) are polynomials of degree at most equal to 3n + 1, we can also use m-point Gauss-Christoffel quadrature relative to the measure  $d\sigma$ , with m = [(3n + 3)/2], to compute  $a_{ik}$  and  $b_i$ . This might be preferable, since it requires nothing beyond standard software.

While this procedure of generating  $\pi_{n+1}^*$  is similar to Kronrod's original method, it is significantly more stable, since the use of powers as a polynomial basis is completely avoided.

**5. Examples.** It is known that for  $d\sigma(t) = (1-t^2)^{\lambda-1/2} dt$  on [-1,1] all Gauss-Kronrod formulae exist if  $0 \le \lambda \le 2$ . Furthermore, the interlacing property (1.3) holds and  $\sigma_{\mu}^* > 0$  for  $\mu = 1, 2, \ldots, n+1$  (Monegato [17]). If  $0 \le \lambda \le 1$ , one has in addition  $\sigma_{\nu} > 0$  for  $\nu = 1, 2, \ldots, n$  (Monegato [18]). Our first example deals with the case  $\lambda = \frac{1}{2}$ , i.e., with the Legendre measure. All computations reported were done on the CDC 6500 computer in single precision (machine precision  $\approx 3.55 \times 10^{-15}$ ), unless noted otherwise.

Example 5.1. 
$$d\sigma(t) = dt$$
 on  $[-1, 1]$ .

Taking advantage of symmetry, we apply Newton's method to the system (2.7), using for  $p_{2k}$  the (monic) Legendre polynomials of degree 2k. The required modified moments are then given by (2.4). We had no difficulty with convergence. By symmetry, only one "trial" initial approximation,  $\mathring{\tau}_1^*$ , is needed, which was programmed to move in nine equal steps of length  $h = (1 - \tau_1)/10$  from 1 - h to  $\tau_1 + h$ . Convergence was invariably achieved for the first choice of  $\mathring{\tau}_1^*$ . Moreover, the problem, suitably scaled, turns out to be extremely stable. In the first four columns of Table 5.1 we report on the number of iterations required for 12 decimal place accuracy, the maximum of  $f_n(t)$  (cf. Eq. (3.12)) on [-1,1] and the value of  $\|J_{\tilde{G}_n}\|_F$  (cf. Eq. (3.11)) for n=5, 10, 20, 40, 80. The maximum of  $f_n(t)$ —an even function of t—is typically assumed between  $\tau_2$  and  $\tau_2^*$ , if n is even, there being a couple of smaller maxima on either side of it. If n is odd, the maximum seems to occur at t=1. Through most of the interval (-1,1), however,  $f_n$  remains  $\leq 1$ .

Table 5.1
Performance and stability characteristics of Newton's method for generating the
$(2n + 1)$ -point Gauss-Kronrod formula with $d\sigma(t) = dt$ on $[-1, 1]$ .

n	#iter.	$  f_n  _{\infty}$	$\ J_{\check{G}_n}\ _F$	$cond_1$	$cond_2$	cond <sub>1</sub> (scaled)	cond <sub>2</sub> (scaled)
5	6	1.126	1.293	1.4(3)	3.3(4)	1.2(1)	1.4(1)
10	6	2.456	1.397	1.5(6)	5.2(8)	1.8(1)	2.3(1)
20	6	2.520	1.333	1.5(12)	5.3(17)	4.1(1)	4.8(1)
40	6	2.535	1.310	1.6(24)	5.8(35)	6.6(1)	7.7(1)
80	6	2.540	1.302			1.3(2)	1.7(2)

The linear systems of equations for determining the initial approximations  $\mathring{\sigma}_{\nu}$ ,  $\mathring{\sigma}_{\mu}^*$  for the weights, as well as the Jacobian matrices in Newton's method, appear to become rapidly ill-conditioned as n increases. Typical condition number estimates (furnished by the LINPACK routine SGECO; cf. [4, Chapter 1]) for the former are shown in the fifth column of Table 5.1, while those for the latter are shown in the sixth column. (Numbers in parentheses indicate decimal exponents.) In spite of the large condition numbers, numerical difficulties were not observed, except in the case n=80, when the computation was aborted due to an arithmetic error. We believe that the apparent ill-conditioning is caused by the use of monic polynomials  $p_{2k}$  in (2.7); their  $L_2$ -norm goes to zero rather quickly,

$$\|p_{2k}\|_2 = \frac{2^{2k}(2k)!^2}{(4k)!} \sqrt{\frac{2}{4k+1}} \sim \sqrt{\pi} 2^{-2k} \text{ as } k \to \infty,$$

thereby introducing a systematic diminution of the rows down the matrices. If the row involving  $p_{2k}$  is scaled by dividing it by  $2^{2k}(2k)!^2/(4k)!$ , the condition numbers indeed become much more reasonable (the solutions remaining the same); they are shown in the last two columns of Table 5.1.

Example 5.2. 
$$d\sigma(t) = \ln(1/t) dt$$
 on [0, 1].

It appears that this measure also admits Gauss-Kronrod formulae for all n, satisfying the interlacing property (1.3) and having all weights positive. A summary of our numerical experience with Newton's method is given in Table 5.2; it contains information analogous to the one given in Table 5.1 for Example 5.1.

TABLE 5.2

Newton's method for Gauss-Kronrod formulae with  $d\sigma(t) = \ln(1/t) dt$  on [0,1].

n	#iter.	$  f_n  _{\infty}$	$\ J_{\tilde{G}_n}\ _F$	$cond_1$	cond <sub>2</sub>	cond <sub>1</sub> (scaled)	cond <sub>2</sub> (scaled)
5	6	2.90(4)	12.14	2.0(6)	6.9(9)	5.8(1)	2.7(3)
10	6	4.09(5)	24.35	2.2(12)	7.5(18)	2.8(1)	1.6(4)
20	6	5.77(6)	47.43	2.4(24)	8.2(36)	9.8(2)	1.0(5)
40	6	8.86(7)	94.59	<u>.</u> .	<u>.</u>	3.8(3)	6.4(5)

We have used modified moments with respect to the (monic) shifted Legendre polynomials,  $p_k(t) = [k!^2/(2k)!]P_k^*(t)$ ; they are known to be (cf., e.g., Gautschi [7])

$$m_0 = 1,$$
  $m_k = \int_0^1 p_k(t) \ln(1/t) dt = \frac{k!^2}{(2k)!} \frac{(-1)^k}{k(k+1)},$   $k = 1, 2, ....$ 

Row scaling of the matrices was performed through division by  $k!^2/(2k)!$  (of the row numbered k+1).

The initial approximations  $\mathring{\tau}_1^*$  and  $\mathring{\tau}_{n+1}^*$  were programmed to first move inward symmetrically in nine steps of length  $h=(1-\tau_1)/10$  and  $h'=\tau_n/10$ , respectively, and if this did not work,  $\mathring{\tau}_1^*$  was varied independently over the same set of points for each fixed  $\mathring{\tau}_{n+1}^*$ . Convergence, in general, was achieved for the very first choice of  $\mathring{\tau}_1^*$  and  $\mathring{\tau}_{n+1}^*$ , i.e., for  $\mathring{\tau}_1^*=1-h$  and  $\mathring{\tau}_{n+1}^*=h'$ , except when n is small, for example, n=1 and n=3, in which cases convergence was realized when  $\mathring{\tau}_{n+1}^*=h'$  and  $\mathring{\tau}_1^*=1-4h$  (for n=1),  $\mathring{\tau}_1^*=1-2h$  (for n=3).

The polynomial  $f_n(t)$  invariably assumes its global maximum at t = 1, has a few much smaller relative maxima between the first few neighboring nodes  $\tau_{\mu}^*$  and  $\tau_{\nu}$ , and then settles down to magnitudes around 1 for the remaining portion of the interval [0, 1]. The condition of the problem, though slightly worse than in Example 5.1, is still remarkably good.

Numerical results for the nodes and weights of the (2n + 1)-point Gauss-Kronrod formula for n = 5(5)25 can be found in Table S.1 of the supplements section at the end of this issue. They have been computed in double precision to an accuracy of 25 decimal places after the decimal point. As the results are displayed in D-format, some of the end figures may not be reliable in those numbers that are much smaller than 1.

We used Example 5.2 to further experiment with alternative choices of initial approximations. In particular, we examined how inaccuracies in individual initial approximations affect the speed of convergence. To obtain a basis for meaningful comparison, we first obtained "reference" values for the number of iterations required when all initial approximations are at the same level of accuracy. This was achieved by imposing on the "exact" results for  $\sigma_{\nu}$ ,  $\sigma_{\mu}^{*}$ ,  $\tau_{\mu}^{*}$  (computed to 12 decimal digits) a random perturbation at level  $\varepsilon$ , i.e., by taking  $\mathring{\sigma}_{\nu} = \sigma_{\nu}(1 + r_{\nu}\varepsilon)$ ,  $\mathring{\sigma}_{\mu}^{*} = \sigma_{\mu}^{*}(1 + r_{\mu}\varepsilon)$ ,  $\mathring{\sigma}_{\mu}^{*} = \tau_{\mu}^{*}(1 + s_{\mu}^{*}\varepsilon)$ , where  $r_{\nu}$ ,  $r_{\mu}^{*}$ ,  $s_{\mu}^{*}$  are random numbers from [-1, 1]. The results for  $\varepsilon = 10^{-2}$ ,  $10^{-5}$ ,  $10^{-8}$ , and  $10^{-11}$  are shown in Table 5.3. (For  $\varepsilon = 10^{-2}$  and n = 40, Newton's method did not converge within 20 iterations.)

Table 5.3

The number of iterations required for initial approximations at accuracy level  $\epsilon$ .

ε	n = 5	n = 10	n = 20	n = 40
10 <sup>-2</sup> 10 <sup>-5</sup> 10 <sup>-8</sup>	6	6	5	
$10^{-5}$	3	3	3	4
10-8	2	2	2	2
10-11	2	2	2	2

Table 5.4
The number of iterations required when two pairs of initial approximations
(with indices $\mu$ and $n + 2 - \mu$ ) are inaccurate.

ε	l	μ	= 1			μ≈	n/4	μ =	$\mu = n/2$					
	n=5	n = 10	n = 20	n = 40	n = 5	n = 10	n = 20	n = 40	n = 5	n = 10	n = 20	n = 40		
10-2	5	4	6	6	5	4	6	5	4	4	5	6		
10-5	3	3	3	4	3	3	3	3	3	3	3	3		
$10^{-8}$	2	2	2	2	2	2	2	2	2	2	2	2		
$10^{-11}$	2	2	2	2	2	2	2	2	2	2	2	2		
					l				1					

We now contrast this with the case in which all initial approximations are at the accuracy level  $\frac{1}{2}10^{-12}$  of the initially computed results, except for two pairs of Kronrod nodes and weights (situated symmetrically with respect to the midpoint of the interval [0, 1]), which are randomly perturbed at level  $\varepsilon$ . Choosing the inaccurate pairs  $\mathring{\tau}_{\mu}^{*}$ ,  $\mathring{\sigma}_{\mu}^{*}$  to be those corresponding to  $\mu = 1$ ,  $\mu \approx n/4$ , and  $\mu \approx n/2$  (and to the symmetric indices  $n + 2 - \mu$ ), we observed the results shown in Table 5.4.

Comparing Tables 5.3 and 5.4, we note some improvement, particularly for  $\mu \approx n/2$ , in the case  $\varepsilon = 10^{-2}$ , when only two pairs of initial approximations are inaccurate (though for n = 20 there are two instances of deterioration), but in all other cases the performance of Newton's iteration is practically the same. This seems to suggest that it is the maximum relative error in the initial approximations (usually associated with a Kronrod or Gauss node near the end of the interval) which determines the speed of convergence.

The choice of initial approximations proposed in Section 2 leads to initial (relative) errors that are reasonably small for "interior" nodes, but comparatively larger near the "boundary". Specifically, if we regard the three Kronrod nodes nearest to each of the endpoints of [0,1], and the two Gauss nodes between them, as belonging to the "boundary", and all others to the "interior", then the relative errors of the initial approximations in our scheme range for n = 10, 20, 40 respectively from 2.0(-3) to 9.5(-2), 2.4(-5) to 5.5(-2), and 4.3(-6) to 2.9(-2) in the interior, and from 4.3(-3) to 5.4(-1), 1.3(-3) to 5.6(-1), and 3.3(-4) to 5.7(-1), respectively, at the boundary. The relatively large number of 6 iterations reported in Table 5.2 appears to be due to the large maximum error of the initial approximations in the boundary zones.

Attempts to improve the initial approximations for  $\sigma_{\nu}$  by using, for example,  $\sigma_{\nu} = \frac{1}{2}\sigma_{\nu}^{(n)}$ ,\*\*\* where  $\sigma_{\nu}^{(n)}$  are the Christoffel numbers for  $d\sigma$ , and by obtaining the remaining initial approximations  $\sigma_{\mu}^{*}$  from a reduced system (2.2) of n+1 linear equations, do not speed up Newton's iteration (in fact, require 7 iterations, instead of 6, when n=5 and n=10), precisely because of the initial approximations in the boundary zones remaining at the same low accuracy level.

Example 5.3. 
$$d\sigma(t) = t^{\alpha} \ln(1/t) dt$$
 on [0, 1],  $\alpha = \pm \frac{1}{2}$ .

Modified moments with respect to the shifted Legendre polynomials are again available (Gautschi [7]) and suggest the same scaling as in Example 5.2. Results for

<sup>\*\*</sup> This approximation was proposed to us by the referee.

n	#iter.	$  f_n  _{\infty}$	$\ J_{\tilde{G}_n}\ _F$	cond <sub>1</sub> (scaled)	cond <sub>2</sub> (scaled)
5	6	1.03(4)	7.542	3.8(3)	7.8(4)
10	6	1.35(5)	14.64	7.3(4)	1.0(6)
20	6	2.06(6)	29.60	5.2(5)	1.4(7)
<b>4</b> 0	6	3.40(7)	61.04	1.4(7)	3.0(8)

TABLE 5.5

Newton's method for Gauss-Kronrod formulae with  $d\sigma(t) = t^{1/2} \ln(1/t) dt$  on [0,1].

Table 5.6

Newton's method for Gauss-Kronrod formulae with  $d\sigma(t) = t^{-1/2} \ln(1/t) dt$  on [0,1].

n	#iter.	$  f_n  _{\infty}$	$\ J_{\tilde{G}_n}\ _F$	cond <sub>1</sub> (scaled)	cond <sub>2</sub> (scaled)			
4	7	9.78(6)	272.21	1.3(2)	9.0(5)			
8	9	3.66(8)	889.01	2.0(3)	1.3(7)			
16	9	1.52(10)	2961.5	6.3(3)	2.0(8)			
32		6.56(11)	9907.2	8.8(3)	3.1(9)			

 $\alpha=\frac{1}{2}$  are summarized in Table 5.5. (The computation was done in double precision, the number of iterations referring again to 12 decimal place accuracy.) The performance of Newton's method is similar as in Example 5.2, except for the (scaled) matrices now being more ill-conditioned. The ill-conditioning of the Jacobian matrices is still worse in the case  $\alpha=-\frac{1}{2}$ , as can be seen from Table 5.6. For this value of  $\alpha$  the independent variation of the starting approximations  $\hat{\tau}_{n+1}^*$ ,  $\hat{\tau}_{n+1}^*$  proved to be rather essential, since convergence was *never* achieved for the first choices of these starting values. This is in contrast to the case  $\alpha=\frac{1}{2}$ , where the first choice of  $\hat{\tau}_{n+1}^*$  always worked.

When  $\alpha = -\frac{1}{2}$ , Newton's method could not be made to converge for n = 32, using the implementation described in Section 2. The difficulty, we believe, is caused by the smallest Kronrod node being almost equal to zero. Computing (in double precision)  $\pi_{33}^*$  by the procedure of Section 4, and applying Newton's method to  $\pi_{33}^*$ , we find  $\tau_{33}^* = 3.05867... \times 10^{-9}$ . Using all zeros of  $\pi_{33}^*$  computed in this way as initial approximations to Newton's method, and lowering the accuracy requirement to 20 decimal places, indeed restores convergence and yields the data for n = 32 in Table 5.6 after 1 iteration.

Nonconvergence (in the case  $\alpha = -\frac{1}{2}$ ) was also observed for odd values of n, this time because of the presence of negative Kronrod nodes. When n = 1, for example, one computes directly  $\pi_2^*(t) = t^2 - (198/343)t - (3671/117649)$ , which has the zeros  $\tau_1^* = .627023...$  and  $\tau_2^* = -.0497636...$  We verified that for all odd n < 32 the polynomial  $\pi_{n+1}^*$  has exactly one negative zero, while all other zeros are between 0 and 1.

Numerical results for  $\alpha = \frac{1}{2}$ , n = 5(5)25, computed in double precision, are given in Table S.2 and for  $\alpha = -\frac{1}{2}$ , n = 4(4)24, in Table S.3 of the supplements section at the end of this issue.

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- 1. P.BARATELLA, "Un' estensione ottimale della formula di quadratura di Radau," Rend. Sem. Mat. Univ. Politec. Torino, v. 37, 1979, pp. 147-158.
- 2. F. CALIO, E. MARCHETTI & G. PIZZI, "Valutazione numerica di alcuni integrali con singolarità di tipo logaritmico," Rend. Sem. Fac. Sci. Univ. Cagliari, v. 54, 1984, pp. 31-40.
- 3. C. DAGNINO & C. FIORENTINO, "Computation of nodes and weights of extended Gaussian rules," Computing, v. 32, 1984, pp. 271-278.
  - 4. J. J. Dongarra et al., LINPACK Users' Guide, SIAM, Philadelphia, Pa., 1979.
- 5. S. ELHAY & J. KAUTSKY, "A method for computing quadratures of the Kronrod Patterson type," Proc. 7th Australian Computer Science Conf., Austral. Comput. Sci. Comm., v. 6, no. 1, February 1984, pp. 15.1–15.9. Department of Computer Science, University of Adelaide, Adelaide, South Australia.
- 6. S. ELHAY & J. KAUTSKY, IQPACK: Fortran Subroutines for the Weights of Interpolatory Quadratures, School of Mathematical Sciences, The Flinders University of South Australia, April 1985.
- 7. W. GAUTSCHI, "On the preceding paper 'A Legendre polynomial integral' by James L. Blue," *Math. Comp.*, v. 33, 1979, pp. 742–743.
- 8. W. GAUTSCHI, "On generating orthogonal polynomials," SIAM J. Sci. Statist. Comput., v. 3, 1982, pp. 289-317.
- 9. W. GAUTSCHI, "On the sensitivity of orthogonal polynomials to perturbations in the moments," *Numer. Math.*, v. 48, 1986, pp. 369-382.
- 10. W. GAUTSCHI, "Questions of numerical condition related to polynomials," in MAA Studies in Numerical Analysis (G. H. Golub, ed.), Math. Assoc. America, Washington, D. C., 1984, pp. 140-177.
- 11. G. H. GOLUB & J.KAUTSKY, "Calculation of Gauss quadratures with multiple free and fixed knots," *Numer. Math.*, v. 41, 1983, pp. 147-163.
- 12. G. H. GOLUB & J. H. WELSCH, "Calculation of Gauss quadrature rules," *Math. Comp.*, v. 23, 1969, pp. 221–230.
- 13. D. K. KAHANER, J. WALDVOGEL & L. W. FULLERTON, "Addition of points to Gauss-Laguerre quadrature formulas," SIAM J. Sci. Statist. Comput., v. 5, 1984, pp. 42-55.
- 14. J. KAUTSKY & S. ELHAY, "Calculation of the weights of interpolatory quadratures," *Numer. Math.*, v. 40, 1982, pp. 407-422.
- 15. J. KAUTSKY & S. ELHAY, "Gauss quadratures and Jacobi matrices for weight functions not of one sign," *Math. Comp.*, v. 43, 1984, pp. 543-550.
- 16. A. S. Kronrod, Nodes and Weights for Quadrature Formulae. Sixteen-place Tables, Izdat. "Nauka", Moscow, 1964. (In Russian.) [English Translation: Consultants Bureau, New York, 1965.]
- 17. G. Monegato, "A note on extended Gaussian quadrature rules," Math. Comp., v. 30, 1976, pp. 812-817.
- 18. G. Monegato, "Positivity of the weights of extended Gauss-Legendre quadrature rules," *Math. Comp.*, v. 32, 1978, pp. 243–245.
- 19. G. MONEGATO, "Some remarks on the construction of extended Gaussian quadrature rules," *Math. Comp.*, v. 32, 1978, pp. 247–252.
- 20. G. MONEGATO, "An overview of results and questions related to Kronrod schemes," in *Numerische Integration* (G. Hämmerlin, ed.), Internat. Ser. Numer. Math., vol. 45, Birkhäuser, Basel, 1979, pp. 231–240.
- 21. G. MONEGATO, "Stieltjes polynomials and related quadrature rules," SIAM Rev., v. 24, 1982, pp. 137-158.
- 22. T. N. L. PATTERSON, "The optimum addition of points to quadrature formulae," *Math. Comp.*, v. 22, 1968, pp. 847-856. Loose microfiche suppl. cl-cl1. [Errata: *ibid.*, v. 23, 1969, p. 892.]
- 23. R. PIESSENS & M. BRANDERS, "A note on the optimal addition of abscissas to quadrature formulas of Gauss and Lobatto type," *Math. Comp.*, v. 28, 1974, pp. 135-139. Suppl., *ibid.*, pp. 344-347.

# **Supplement to On Computing Gauss-Kronrod Quadrature Formulae**

## By Franca Caliò, Walter Gautschi, and Elena Marchetti

Table S.1. Nodes and weights for (2n+1)-point Gauss-Kronrod quadrature with respect to the weight function  $\ln(1/t)$  on [0,1], n=5(5)25.

#### n=5

nodes	we1ghts
9.670317011313184162899526d-01	1.817674998509004844923280a-03
8.947713610310082836388862g-01	9.820514710459464849401820g-03
7.946457674432365150600413a-01	2.528143279850861173750120d-02
6.773141745828203807018027a-01	4.8359486624192653084139940-02
5.474241572272107646167938a-01	8.111987879942632270559929a-02
4.117025202849020431749319a-01	1.1928870631016403584921680-01
2.833909712982217042042441d-01	1.515254284520667107440880a-01
1.739772133208976287011397a-01	1.715203967327575628055522a-01
8.793120246479704019290778a-02	1.785532970704892670281547d-01
2.913447215197205330372676a-02	1.518624375163058221223388a-01
3.055453450374047608522276a-03	6.085074598712054422908396d-02

#### n=10

n=	10
nodes	weights
nodes  9.904014566346610399228104a-01 9.688479887186335393915045a-01 9.375650909249013043766844a-01 8.981610912190035381669532a-01 8.503062699837251446095142a-01 7.941904160119662173585093a-01 7.319872866081715824749702a-01 6.657752055164245972223813a-01 5.961171169960002092317760a-01 5.961171169960002092317760a-01 5.961171169960002092317760a-01 3.802125396093323339723413a-01 3.802125396093323339723413a-01 3.119327193419556033237433a-01 2.4705241628715982422225454a-01 1.875727130810663812513927a-01 1.353118246392507748702321a-01 9.068916719091722376416700a-02 5.397126622250062950420123a-02 2.6405646249468567738494414-02	weights  1.533569184207643623003683d-048.487068240706659310736070d-042.284556832745782346462822d-034.671211724894484356713219d-038.467273369906779154553426d-031.373507422471821554874215d-022.009874297264113701298160d-022.764325282485899850930559d-023.686595352281781058596218d-024.718382784133641587909658d-025.734613169174315560725749d-026.733491954479115487650944d-027.784758916696079582379503d-028.74825885263594360502868d-029.381826910705523922356189d-029.485544515771952988656812d-029.794446461628536838979348d-029.456636039965041400178269d-028.19090316926699549075249d-0
9.042630962199650636946628d-03 1.283562335805386114259342d-03	5.774125174710110186821323a-02 2.520201991637924753934347a-02

2,079484873867849035127780d-02 8,577830778258255066645536d-03

3.694444650593027343659999d-04

2.588327955921955428332736a-03

S58

n=20

9T=U

S59

S	
7=	
5	

nodes

weights

78-01 4.7299637 28-01 2.6373405 28-01 7.1990643	30d-01 1.502106716669550701254	22d-01 4.63833809168626112491517	U9G-UI /.U3UUI338338330933908430 16G-OI I.OO863947279785714046350	27d-01 1.40618220614814007433955	95d-01 1.89243793774408211184387 79d-01 2.44966079249722285983532	d=01 3.096482034768542390	53d-01 4.74554501031732551998737	83d-01 5.69430742107474118661543	266-01 6./40099545/22/8642551415 11d-01 7.92873215863290721095995	80d-01 9.2246335902533143174432	55d=01 1.20006991870964886407029	514-01 1.35758785715296426600449	49d-01 1.5238663968764787810283	#3.d=01	21g-01 2.04721400000328576720000	39d-01 2.23659715009142268608638	2.d=01	264-01 2.79360274924039059480461	97a-01 2.98382265843419425766551	339-01 3.15774716260463519241169	69d-01 3.32362393927983382386530 64d-01 3.49378728282568001197096	400-01 3.65391630872385777054323	51d-01 3.78540645047158269635431	20d-01 3.8988677808580781845668 22d-01 4.0096933709528131684847	97d-01 4.09960512381625022319677	11d-01 4.14544998865505177609133	4.163137141338749900085971	560-02 4.12990577768560317561726	05d-02 4.03243774586159772416916	18a-02 3.8862477801763191811488 93a-02 3.7096914523491159772729	-02 3.46990264199000761995395	754-03 2.71425184860231261656868	01 d-03 2.20311417496101007946778	d-03 I.50026/29/6II8019021456
6278789961661 7646927524797 8544905070619	816252725992549675149 725798827984935169482	616242619673757147608	.489694797742318211236 .347912842966051819430	189964680454576632653	.015261575671396041060 .825755888546463808008	623241609572156561755	177687343472579641766	.936380012583529233399	. 685 <i>2213</i> 43520897819970 . 424142881943629420746	153184705742639531113	.8/43232558648904442/1	298760168905099694178	.002446668992273705187	40083612983388338833883	.097739109651696368322	. 793555262000715691454	.4902168/2094010/99525 	. 892032480187998615929	.597894506450982806557	309020855010095615519	.027179664539716304239 752681778745490731973	485857025771854935199	.228388124017029203218	.981872403062508468259	.522591278588089225362	.311523567594944725896	. 114683666474060405652 321492439096544167459	640165348265870132682	.114754710816483918926	.755526608752405384819 .561500389827781933660	. 532505988961362761827	. 994559916853792794865	.969956845265878995266	. 71 000 7201 966834333767

Table S.2. Nodes and weights for (2n+1)-point Gauss-Kronrod quadrature with respect to the weight function  $t^2 \ln(1/t)$  on [0,1], n=5(5)25.

n=5

# nodes

weights

1.122054492767190248450108a-0	1.084416999529989053225187a-02
3.592188934246314021399989a-0	4.947701935575340163420805a-02
5.955012620582518424233754a-0	1.1738777405388591583647459-01
7.333914141939224778997501d-0	2.096522296417729123298881a-01
7.702597204357045876511760d-0	3.210905075966306604889925a-01
6.997805838224885451283229a-0	4.4671237319486446718151890-01
5.363653907515083214871332a-0	5.769087058303224900777386a-01
3.492536726049289162501093a-0	7.001427291343135134908556d-01
1.940997445389493706984277a-0	8.1011072880352710908493300-01
7.918209630488930680875090a-0	9.0293549763026936974403260-01
1.518621703245064911238920g-0	9.6961660531944313860345696-01

222222222222

01=u

1.394496655684032669349753a-04 7.635011344553433632030752a-04 2.024857904781603108605194a-03 4.063789872820607031523504a-03 3.153900421376590286697619d-02 2.465406934320781026244711d-02 1.6555919485540565843122956-02 8.6750973580069683587651396-03 2.6140189692613024973947386-03 8.566056607706878019699342d-01 8.026806142819284908704270d-01 7.427730638296861355174625d-01 6.787593289760173472075718d-01 6.11733227010825674487500d-01 5.408058388275156878481092d-01 6.781818479343935804646819d-02 4.697468016339130402377083d-01 3.999403178414447653361729d-01 3,322045316253965362431151d-01 2,674429071474993800705367d-01 2,074532308974391232607778d-01 1,538036758794403479214318d-C1 1.070797600006763569921467d-01 3.6217228727290359001821150-03 9.702208699408183350814147a-01 9.402965535372893169161932d-01 9.025384582552350528421821d-01 9.908249255019305165744127a-01

5.920141048904754145795907d-05

.974660223723762994377495a-01 .917367397325117087867489a-01

n=20

n=15

	60 40	0.4	40.5	50	93	03	.03	0.5	.02	0.2	.02	02	.02
weights	3.148188081324152319829025d-05 1.741679371363330396600107d-04	4.697566331459988000048351d-04	9.6475338512854318452499180-04	2,8520234446234207624047789-03	4.200501745800172624049450d-03	5.831239523911026933767281d-03	7.833185179012937434611582d-03	1.010663620433614594284828d-02	1.247194640152809995224219a-02	1.4950422792039293308263490-02	1.761778314420711474668436d-02	2.0264687214481146401068270-02	2.262739672072958670783827d-02
nodes	9.956437735044021502692128d-01 9.858130450027474222285917d-01	9.7140453364862332661706704-01	9. 529754524595691561897181d-01 9. 301746463403134780234265d-01	9,0285971914411330285375589-01	8.717287103016314373415490d-01	8.374001820934952597481019d-01	7.998370565553648198159345d-01	7.5911687122882514602257540-01	7.159542588193601129720480d-01	6.7101966847577117552032030-01	6.244592145904029741841424d-01	5.764907527804063730798203d-01	5.278310621284453211150285a-01

2.823131549448067237614801d-02 2.907873800974671420627266d-02 2.934519342726095829740092d-02 2.9149553735702209530041343d-02 2.8265542202557352716d-02 2.656207179145095975286140d-02 2,4237625708991149767246230-02 2,1456576652266407665616060-02 1,8156283859254539403084340-02 6.862793823771532409339749a-03 9.978271892698620745027816d-04 2.669993597133315119047438d-02 1.442921155204044134329441d-02 1.058972012609887498501608d-02 3.490958422386958576167944a-03 2.4783951836090780748137788-01 2.0708274468941015184153340-01 1.691465202192176190569540-01 1.349132095515663067227940-01 1.0330020031698188614595460-01 7.578106237672045127378528d-02 5.228726581769844567883073d-02 1.756217551177943375333811a-03 4.306893358185325262578644d-01 3.305091043962676544548240d-02 1.814357408264750996250969a-02 7.692542996206947006087645a-03 4.7915205738849495467483020-01 3.8272307600237649234837629-01 3.359167188949541422464380a-01 2.908744520453002250635029d-01

1. 6067825755251176238986766-04 3.22771984364878292174156-04 6. 110975811069729472080563-0-04 1. 0064394798441642010122056-03 1. 1504918484278653417341734-03 2. 1252899032655020297453496-03 3. 484265233527448527022456-03 4. 8684147756134357668388660-03 7. 3086413218954751482305656-03 8. 7118976043952316331404370-03 1.6111476589067155560134734-02 1.7458669731715111210280644-02 2.208044127947931080226739d-02 2.152795766765041018721722d-02 2.067237610829580113347144d-02 1.943536719178504826376461d-02 1.792793585481223265169154d-02 1.468334711839712012158403d-02 1.878321992914134523720709a-02 1.996928465589146806659658d-02 2.0889301903007879501843150-02 2.217218609842731863766559d-02 1.012886689369375208144031a-02 1.158667272800433606202130d-02 1.313613533970197667538385d-02 2.160011623307030328874774d-02 2.248684146211996345886165a-02 1.6253523176176361570210259-02 1.435607658523903943762713d-02 1.221693853158328285572143d-02 9.976362115056468821983689a-03 7.759749319371855842282736d-03 5.582040833518154685688934a-03 1.744365871092624877536576d-03 4.9873579485676292537143690-04 9.833107160803341873960998G-01 9.724798714438548129297555G-01 9.589965474977613593095175G-01 9.4272271006527521409947080G-01 9.240037561077215848918337G-01 9.031336968543576052358897G-01 8.800127347822681227226298G-01 8.545990943043285247267384G-01 8.27226578422210802066822-01 7.982045619607285834734980G-01 6.315214643529458967934773d-01 5.951685585094725659131014d-01 5.58026547608479528741499d-01 5.21243889165180447657242d-01 7.675233805702575521431915a-01 7.352107744146125210893367a-01 7.016043222156841471592541d-01 6.670248909610046887894631d-01 4.8407505732144461406871559-01 4.468960895737404100777051g-01 4.100287465227397538441176d-01 3.737770493538123076514239d-01 3.382349802169044855150777d-01 3.035072060934090080163193d-01 2.698849891448891834762372d-01 2.376385873018692908670225d-01 8.019783064321909343718488d-02 4.487866857518176891107446a-02 4.557108230851008498202103d-03 1.057243174302654236678512d-03 2.068530977173010872503285a-01 1.776190751740269135090476a-01 1.501809117049637920211031a-01 1.247586250473093000845315d-01 1.014110585574666679706981 0-01 6.1301511880103641527396940-02 3.0946875024473267822970270-02 1.069009822521478628577196a-02

1.013039813045428716972524a-03 2.836424696312020050370199a-04

6.974541561033979183012355a-04

3.0154513622380437757206720-03

n=25

4.9036157695121270676228534-03 5.7167101338983494623563884-03 1.791629909977009118883295d-02 1.817158755078314799294321d-02 1.810015620682106641723902d-02 1.782264331346983867708252d-02 1.7388894363846906650904680-02 1.684391045606071694613058d-02 1, 61 3941 3682 941 550 7091 9989 d-02 1.524216079727533777099470d-02 L. 420833635665940046221989a-02 1.308229828524520774480091a-02 1.1836616074468040414114190-02 1.046761988804355593972789a-02 0355176418241060156636500-03 7.581591102495519880601995a-03 6.109978931326805650411763a-03 4.658158223804104374174715d-03 2.055179152044591605108541a-03 1.4327135156692751318411650-04 4.3796389222857492515301660-04 9.397398328365766209330049a-04 732274251192089226999094d-03 2.219824396755677886209283d-03 2.775402719408686294340720a-03 3.424178441403852272969322d-03 4.1451053003421653808792280-03 613159901002410624804741d-03 7.556262004409136993274396a-03 8.496697801706256450662582a-03 9.4559111327904303637121420-03 1.046615023627082994080665a-02 1.147684760209533122451069a-02 1.242743134523052335999358a-02 .. 334504039292318992048795a-02 1.426558257800085265027897d-02 1.512874018959156682410576d-02 1.586843916783805040193235d-02 1.651883331210355790789400<del>d-</del>02 1.711962256422544022588934d-02 1.7606789359321333281098630-02 1,809020802232089038186257d-02 4.548106526616023854472569d-06 53091 81 41 93 55841 4379201 50-05 6.889794071646865465493743a-05 2.643807922079948138965084a-04 6.596250171105473811207737a-04 1.299316577295864014318008a-03 9,62346064736065172654908d-01 9,4992606688781005999522506-01 9,3600628034301262200411956-01 9,033481655244656053843246-01 8,84723663517615234455287769d-01 8,6482090427128934396211336-01 8,43590369114769034300488d-01 8,210147033904835055394196-01 7,725552865337124659484674-01 7,4684985633772862560410-01 983449410853950788744021a-01 9,8197324917850745085444420-01 9,7309695193248214813911430-01 6.066424203986215816011536d-01 5.770098863968475208391041d-01 5. 471 73353448994964291 40230-01 5. 1 716945788848778316863080-01 975522118378792382158838d-01 3.395279537225274298795493a-01 3.113998220526236718264130d-01 2.8395956754232910noa3aen 2.572600834918173348706592d-01 3144797924049295965125830-01 2.066577385857402195544410d-01 1.829320973674932278918869d-01 1.6031891819536440801600880-01 1.389405582680830798830988d-01 1.1890474911946654183876940-01 002451306311717159206499d-01 3000982204046772686556990-02 6.726422568402615513922525d-02 .310827484735858775824096d-02 4.055443951133245943529628d-02 963276095443142866082606a-02 039755904888424004579264a-02 . 2875577678850893126099699-02 9.945995753894035464780109d-01 9.890822920253952821670840d-01 7.2016218027469421836567690-01 6.6458656538807066217599690-01 6.358957861268925981530241d-01 4.870443791409003470537304a-01 4.569749619692095667873420a-01 4.2712912711689154866024650-01 3.682969082751076033697547a-01 6.926820939724622897627391a-01

Table S.3. Nodes and weights for (2n+1)-point Gauss-Kronrod quadrature with respect to the weight function  $t^{-\frac{1}{2}}\ln(1/t)$  on [0,1], n=4(4)24.

n=4

les

we 19 hts

9.467823860260810350329786G-01 4.8914400551791895723794436-03 8.32437860057592332291369G-01 2.77089860735405590525-02 8.3243786057592332291369G-01 2.750595082719505454237052G-02 5.188100896811898410745488G-01 2.87477610097394017935257416-01 3.4518109261778984107787834974579496417778146151778246-01 7.86018080745033205600824G-02 8.34339417778146131782896-01 7.86018080745033205600824G-02 1.099943109487781826857564G-01 1.60183873842733176459018G-03 1.0999431094847781826858582670

n=β

6.001432418657058561816630d-02 8.624072799621082058522394d-02 3.9269153464451853106852510-04 2.202117153247343158681574d-03 6.023994172220388039372040a-03 2.3387947899915241719106640-02 1,221846359360077873613050d-01 1,688882563842434531022851d-01 2, 223967752688921495142328d-01 2, 842232970531557362894586d-01 3.725619240143217463016735a-01 5.016165133560980438739997a-01 1.254391445564177294731867a-02 3.939444491974970289051482d-02 6.640377748617388785722559a-01 8.060646307386342503543205a-01 6.278260300687075811410164a-01 we 1ghts 7.686135746463162861002726d-01 6.859668444430305674743988d-01 5.972846551081826271071169d-01 5.069890430489052040480914d-01 4.166481922447431769794810G-01 3.2783132043305244454565G-01 2.44975673169443416G-01 1.722653496947714628389G-01 1.106116624165237457561707d-01 6.026637319421194494424290d-02 5.079445421534194356518628d-03 9.0149766550940624642145160-01 8.4079089641561199312475850-01 2.4204305482699448380273870-02 1.428455519688558285056329a-04 9.8471005992113268139204720-01 9.505517502619251862996247d-01

9T=u

2T=u

weights	-01 2.796696253016289349963365d-05 1-01 1.563618047866580836374612d-04		1-01 8.953/18434/2169651259/104d-04 1-01 1.667823441448298692849240d-03						•			1-01 3.112637634051024021346720d-02		1-01 4.559175304495993122835237d-02	F-01 5.476881407984725842184861d-02	1-01 6.526768640264185711280638d-02		1-01 8.876066795067189137014152d-02	⊢01 1.035400689846584558875641¤-01	1-01 1.204351082860888294292007d-01	H01 1.379688245775762562052125a-01	1-01 1.570768111358994027147340d-01						1-02 3.835631688506201888693871d-01	•	i-03 5.424729284517174404220322d-01	-05 3.784969915426641107134499a-01
nodes	9.959030014166557799095396d-01 9.866515060015691419583797d-01	9.730831485036508224494794d=01	9.55/24685551666/8/22498296U 9.342322202953951064273636GU	9.084410487719504138811946a-01	8.790043208903614626232110a-01	8.465172197510391627033991d-01	8.109094289793264441170613a-01	7.721954238899551020288427a-01	7.310578764357724062707320a-01	6.8816363429182762322359434-01	6.435911779555280266552427d-01	5.974569405535026844411241d-01	5.504729697239363220234514d-01	5.033457796482281675897882d-01	4.562129900337846201935433d-01	4.092153034701697007042927a-01	3.6305749398671497283735850-01	3.184402442336167895674019d-01	2.754787120787447394771186a-01	2.342596648033869713090313d-01	1.954353091919325313914467d-01	1.596557467773047315218671d-01	1.269382086713585331460270d-01	9.724006597479101015701976d-02	7.112145494114210204176551d-02	4.914898627259817786969019a-02	3.1171227583366603977937520-02	1.694404804397794929883851d-02	6.907658421540979336488189a-03	1.485765788500239757310222d-03	1,760376132934953188281986d-05
weights	8.458735678948550037906433d-05 4.733122955476183293624621d-04	1.295334109572110901102865d=03	2./06/3168312/645324403101d=03 5.043077178842719760579588d=03	8.464378140053237244756604a-03	1.290184488023427186184698d-02	1.860574263854033135894625d-02	2.620807444877358112292989d-02	3.574276734938528850187790d-02	4.6745789989785164315648440-02	5.967444746985659301284190a-02	7.589043686446729665694696d-02	9.526118334566874264250961a-02	1.163886052639084540070262d-01	1.401920773080115885220792d-01	1.703148279001835272137846d-01	2.067657388086946130994607a-01	2.453472778680508216014776d-01	2.8781074625727006707829380-01	3.488252075194488774741921d-O1	4.371382853483828410959049a-01	5.443017169369761164032346d-01	6.448506073230140879888332a-01	4.689672017154149189815816a-01								
nodes	9.928826120909483027505245a-01 9.768566552363760599297629a-01	9.534839312427130712288346d-01	9.238428533052712944323204d-01 8.875285378330856871129993a-01	8.4447571195006357721362300-01	7.961003718805113291006520d-01	7.437592675800603295615879d-01	6.876293672741867836731351d-01	6.2802556103145050483551044-01	5.664847326900428451563138a-01	5.045314832870891371017139d-01	4.425648565847031835767200d-01	3.809913524608552951366326d-01	3.213287813866706015400765d-01	2,6507802002278668627000904-01	2.125435271249162729620080a-01	1.6392301681011559572951880-01	1.205253501937517224566087d-01	8.364282470170227024739391d-02	5.3199562625101674056399639-02	2,890555423183584924268035d-02	1.170669272088357088680876d-02	2.4937276776315936766831120-03	4.489625718700550057552605a-05								

3.611073729279555201724547a-01 4.179383868248953910138205a-01

2.773785441746910131887578<del>a-</del>01

3.139932348481914050544306a-06

1.730385187601515161300166a-01 1.924729940396529788410872a-01 2.117532723618073902167795a-01 2.324824382014282543255341 a-01 2.6357282116813677352865229-013.0853973706667850533420519-01

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