

Introduction

- Energetic electron precipitation couples magnetosphere and ionosphere
- DMSP provides low-energy measurements (30 keV)
- ELFIN extends to higher energies (50 keV to 6 MeV)

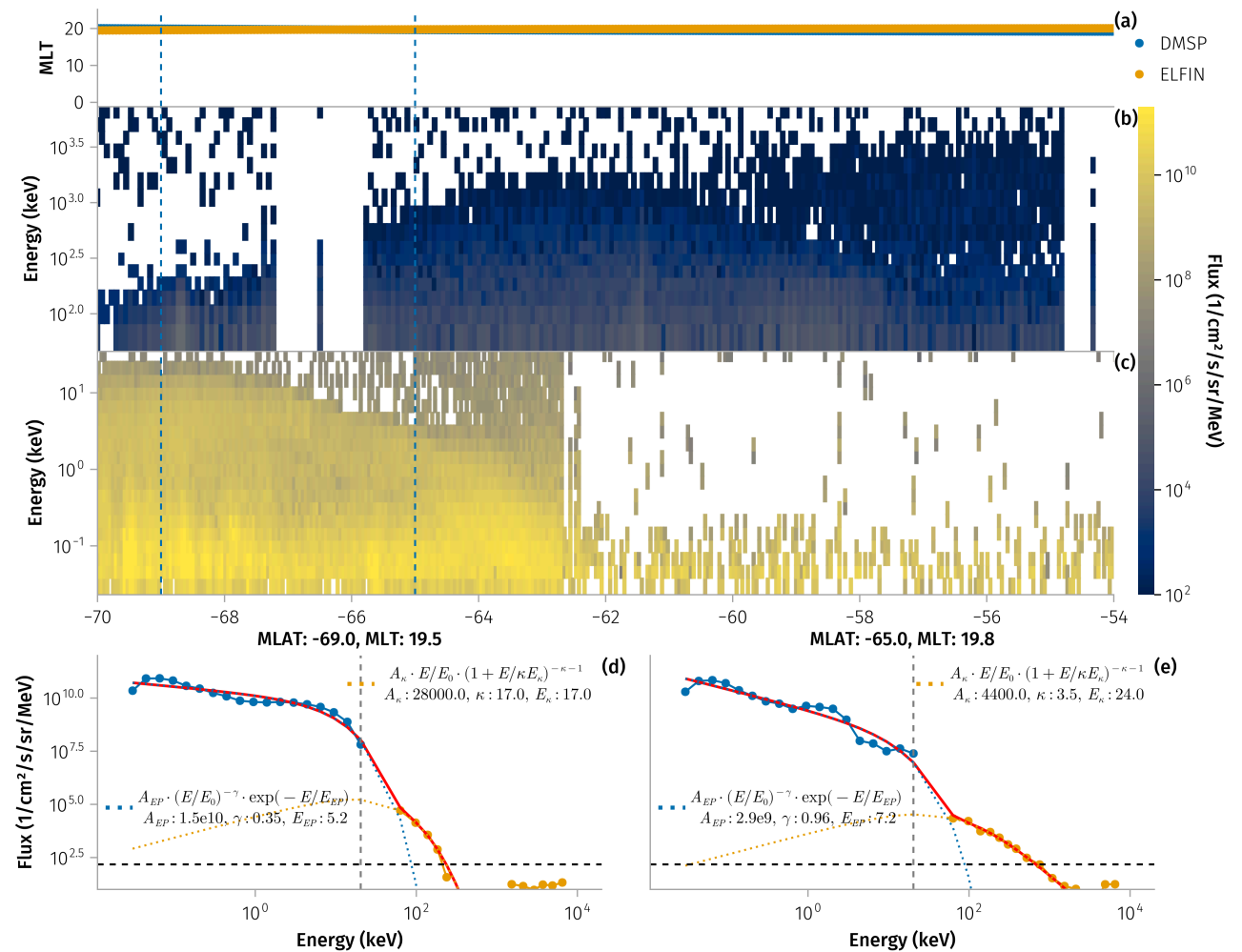


Figure 1: Example of a satellite conjunction event between 2021-12-01T22:19 and 2021-12-01T22:28. (a) ELFIN precipitating electron flux, (b) DMSP electron flux, (c) MLT, (d-e) averaged precipitating flux spectra.

Results

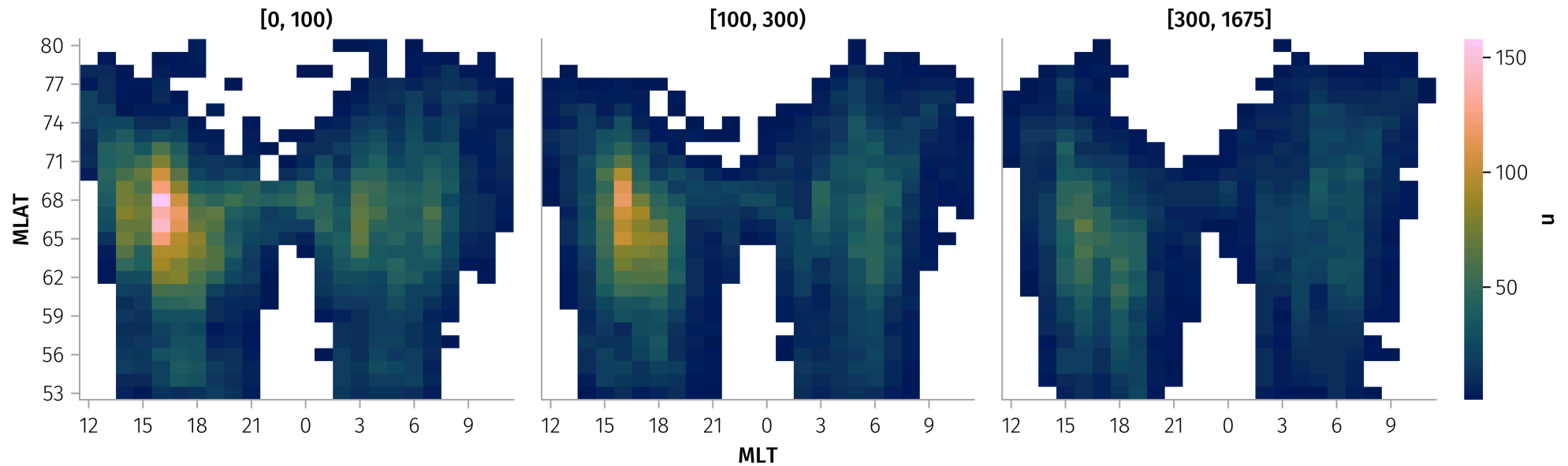


Figure 2: The total number of MLAT-averaged spectral data points across different MLT and MLAT regions, separated by varying levels of the AE index.

Results

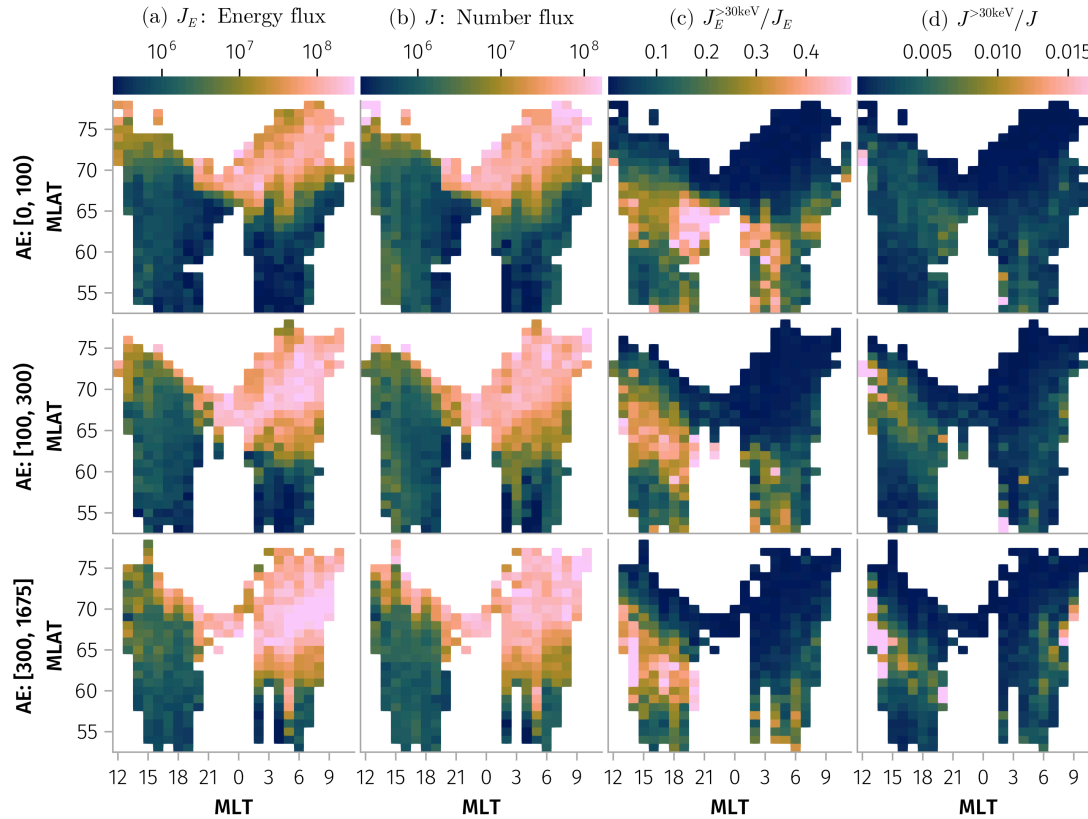


Figure 3: Median distributions of (a) total energy flux [$\text{keV cm}^{-2} \text{s}^{-1} \text{sr}^{-1}$], (b) total number flux [$\text{cm}^{-2} \text{s}^{-1} \text{sr}^{-1}$], and (c-d) fractional contribution of energetic particles ($>30 \text{ keV}$) to total energy and number flux.

Results

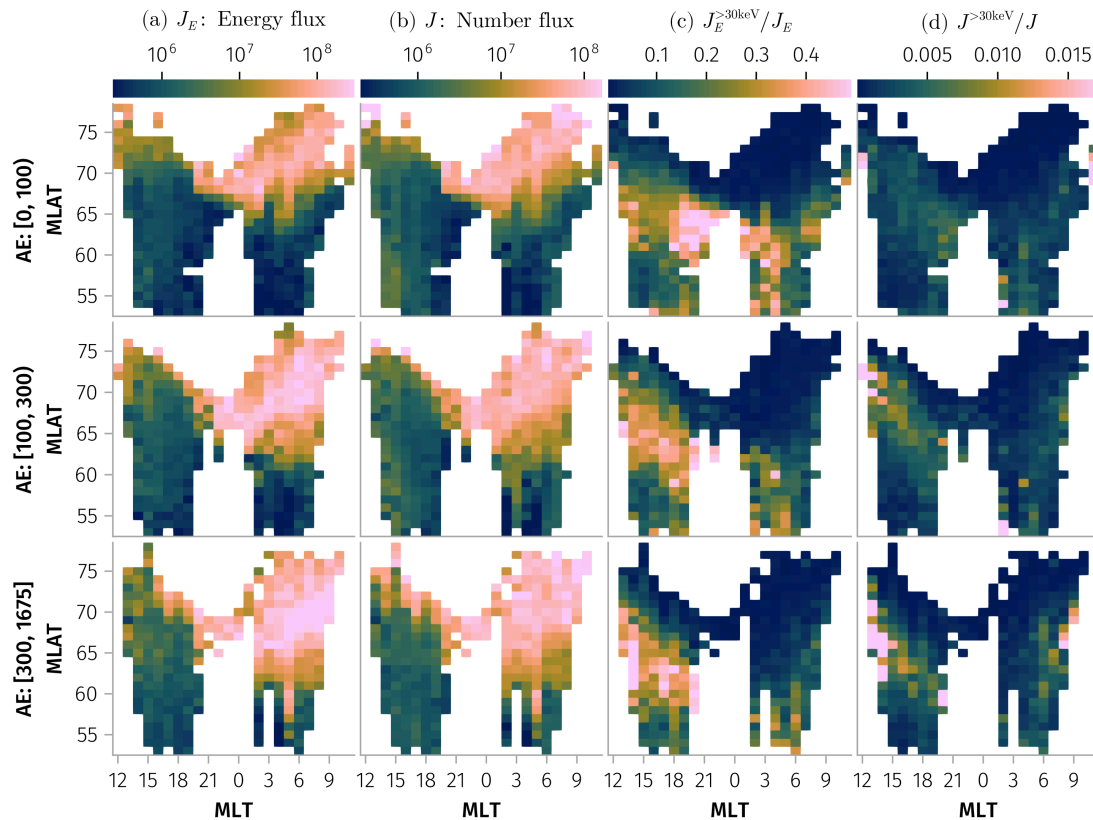


Figure 3: Median distributions of (a) total energy flux [$\text{keV cm}^{-2} \text{s}^{-1} \text{sr}^{-1}$], (b) total number flux [$\text{cm}^{-2} \text{s}^{-1} \text{sr}^{-1}$], and (c-d) fractional contribution of energetic particles ($>30 \text{ keV}$) to total energy and number flux.

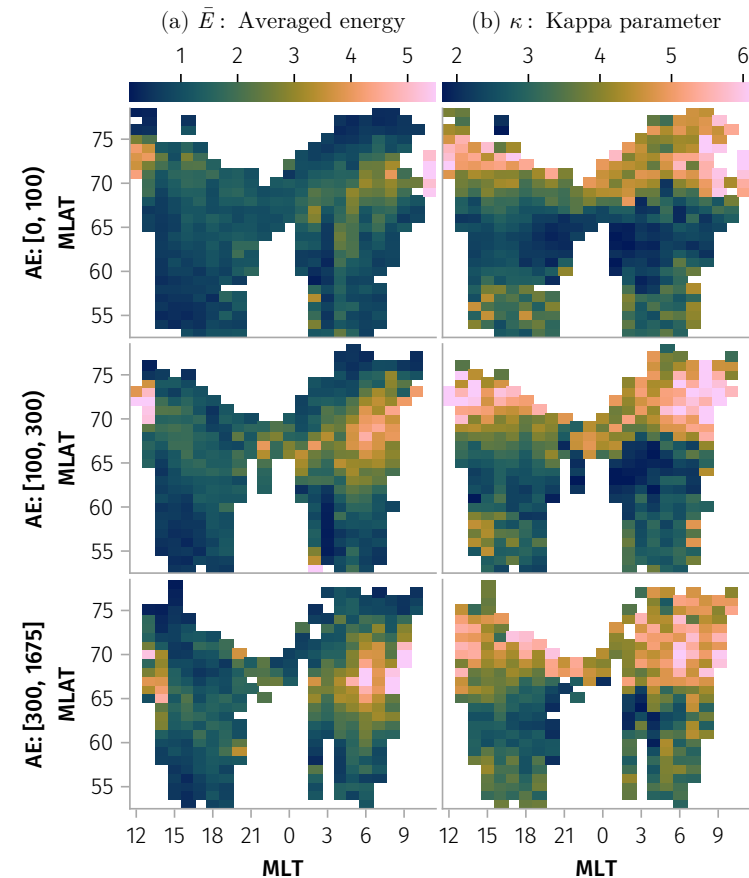


Figure 4: Median distributions of (a) averaged energy (\bar{E}) [keV] and (b) kappa parameter (κ) as functions of MLT and MLAT, sorted by AE index levels.

Results

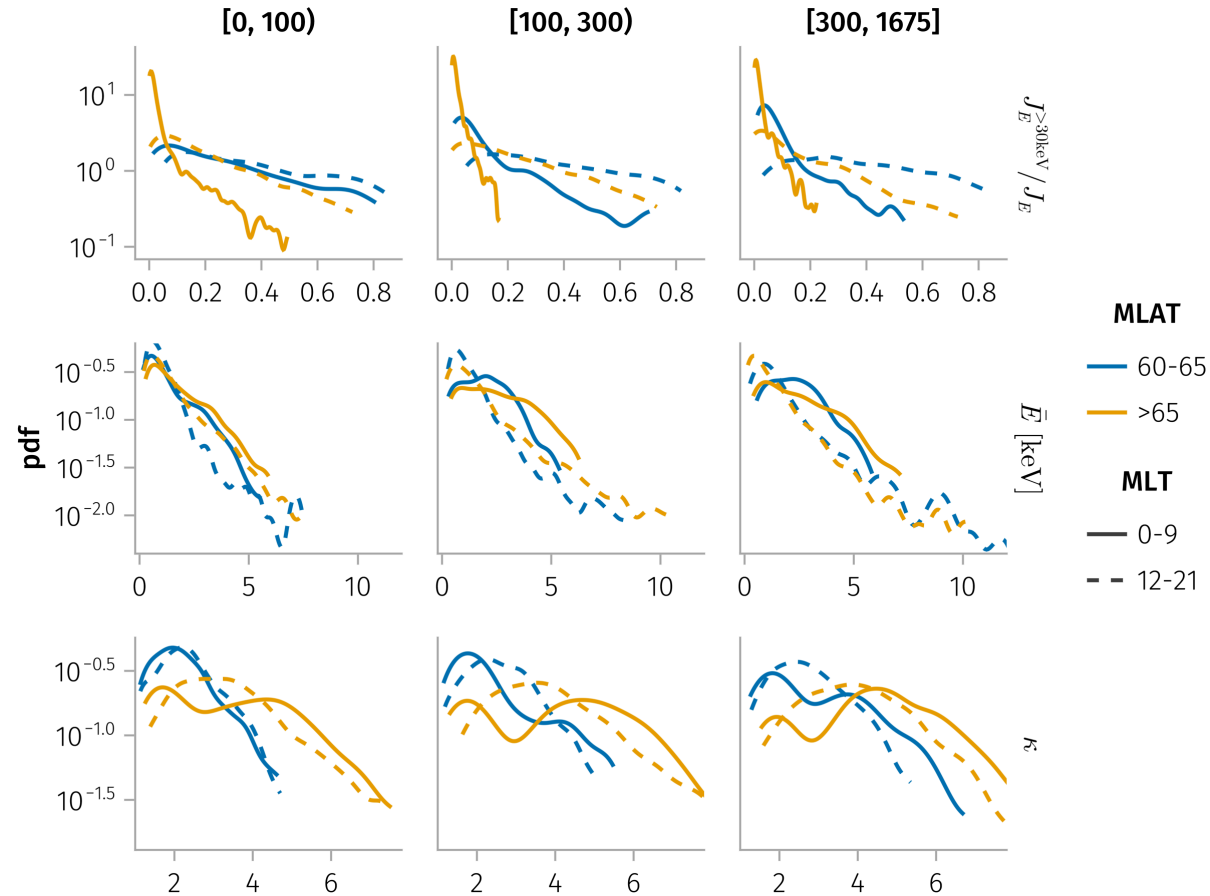


Figure 5: Probability density functions for (a) energy flux ratio $\frac{J_E^{>30\text{ keV}}}{J_E}$, (b) kappa parameter (κ), and (c) mean energy \bar{E} . Data organized by AE ranges, MLT sectors, and MLAT bands.