

# Notes on Quantum Mechanics

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# Chapter 1

## A Historical Overview

Rishi's article + JP sir's slides

### 1.0.1 Blackbody Radiation

### 1.0.2 The de Broglie Hypothesis

In 1924, the French physicist de Broglie proposed that this wave like structure applies to electrons too and follows the equation:

$$p = \frac{h}{\lambda} = \frac{2\pi\hbar}{\lambda} \quad (1.1)$$





# Chapter 2

## Mathematical Preliminaries

This chapter is a discussion of all the mathematical tools and tricks one would require to master Quantum mechanics. We assume that the reader has a lucid understanding of matrices and vector calculus. If not the reader may refer to:

•

to refresh themselves or learn those concepts before

### 2.1 Matrix Inversion

### 2.2 Complex Numbers

A complex number is an order pair  $\in \mathbb{C}$  where  $a, b \in \mathbb{R}$  where we can denote it as  $z = a + ib$  where  $i = \sqrt{-1}$

#### 2.2.1 Addition

$$z_1 = a_1 + ib_1, \quad z_2 = a_2 + ib_2$$

$$z_1 + z_2 = (a_1 + a_2) + i(b_1 + b_2)$$

#### 2.2.2 Multiplication

$$z_1 = a_1 + ib_1, \quad z_2 = a_2 + ib_2$$

$$z_1 z_2 = (a_1 + ib_1)(a_2 + ib_2) = (a_1 a_2 - b_1 b_2) + i(a_1 b_2 + a_2 b_1)$$

### 2.2.3 Properties

Where,  $\mathcal{W}, \mathcal{Z}, \lambda \in \mathbb{C}$

#### Commutativity

$$\mathcal{W} + \mathcal{Z} = \mathcal{Z} + \mathcal{W}$$

$$\mathcal{W}\mathcal{Z} = \mathcal{Z}\mathcal{W}$$

#### Associativity

$$(\mathcal{Z}_1 + \mathcal{Z}_2) + \mathcal{Z}_3 = \mathcal{Z}_1 + (\mathcal{Z}_2 + \mathcal{Z}_3)$$

$$(\mathcal{Z}_1\mathcal{Z}_2)\mathcal{Z}_3 = \mathcal{Z}_1(\mathcal{Z}_2\mathcal{Z}_3)$$

#### Identities

$$\mathcal{Z} + 0 = \mathcal{Z}$$

$$\mathcal{Z}1 = \mathcal{Z}$$

#### Additive Inverse

$$\forall \mathcal{Z} \exists \mathcal{Z}^{-1} \mid \mathcal{Z} + \mathcal{Z}^{-1} = 0$$

#### Multiplicative Inverse

$$\forall \mathcal{Z} \neq 0 \exists \mathcal{W} \mid \mathcal{Z}\mathcal{W} = 1$$

#### Distributive Property

$$\lambda(\mathcal{W} + \mathcal{Z}) = \lambda\mathcal{W} + \lambda\mathcal{Z}$$

### 2.2.4 Notation

*n-tuple* refers to an ordered set of  $n$  numbers over a field  $\mathcal{F}$ .<sup>1</sup>

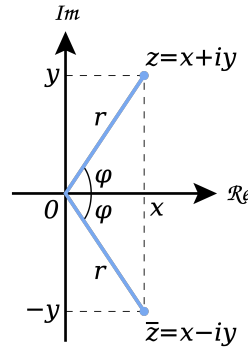


Figure 2.1: Wessel Plane Plot: (Complex conjugate picture.svg from Wikimedia Commons)

### 2.2.5 Wessel Plane

## 2.3 Linear Vector Spaces

A linear vector space or simply a vector space  $\mathbb{V}$  is a set along with the regular multiplication and addition operations over a field  $\mathcal{F}$ , such that the following axioms hold: <sup>2</sup>

### 2.3.1 Commutativity

$$\mathcal{U} + \mathcal{V} = \mathcal{V} + \mathcal{U}$$

### 2.3.2 Associativity

$$\begin{aligned} (\mathcal{U} + \mathcal{V}) + \mathcal{W} &= \mathcal{V} + (\mathcal{U} + \mathcal{W}) \\ (\alpha\beta)\mathcal{V} &= \alpha(\beta\mathcal{V}) \end{aligned}$$

### 2.3.3 Additive Identity

$$\exists 0 \in \mathbb{V} \mid \mathcal{V} + 0 = 0 + \mathcal{V} = \mathcal{V}$$

### 2.3.4 Additive Inverse

$$\forall \mathcal{V} \exists \mathcal{V}^{-1} \mid \mathcal{V} + \mathcal{V}^{-1} = 0$$

---

<sup>1</sup>For our case  $\mathcal{F}$  simply refers to  $\mathbb{C}$

<sup>2</sup>Here,  $\alpha, \beta \in \mathcal{F}$  and  $\mathcal{U}, \mathcal{V}$  and  $\mathcal{W} \in \mathbb{V}$

### 2.3.5 Multiplicative identity

$$\exists 1 \in \mathbb{V} \mid 1\mathcal{V} = \mathcal{V}$$

### 2.3.6 Distributive properties

$$\alpha(\mathcal{U} + \mathcal{V}) = \alpha\mathcal{U} + \alpha\mathcal{V}$$

$$(\alpha + \beta)\mathcal{U} = \alpha\mathcal{U} + \beta\mathcal{U}$$

## 2.4 Inner Product Spaces

An inner product is simply an operation that takes a Dual  $|\psi\rangle$  and it's corresponding vector  $\langle\psi|$  and maps them to  $\mathbb{R}$ :

$$\langle expression1 | expression2 \rangle$$

## 2.5 Dual Spaces

## 2.6 Dirac Notation

Operators are represented with respect to a particular basis (in this case  $\{e_m, e_n\}$ ) by their matrix elements

$$\langle e_m | \hat{O} | e_n \rangle = \hat{O}_{mn} \quad (2.1)$$

## 2.7 Subspaces

Given a vector space  $\mathbb{V}$ , a subset of its elements that form a vector space among themselves is called a subspace. We will denote a particular subspace  $i$  of dimensionality  $n_i$  by  $\mathbb{V}_i^{n_i}$ .

Given two subspaces, and , we define their sum  $\mathbb{V}_i^{n_i} \oplus \mathbb{V}_i^{m_i} = \mathbb{V}_i^{l_i}$ <sup>3</sup> as the set containing:

1. All the elements of  $\mathbb{V}_i^{n_i}$
2. All the elements of  $\mathbb{V}_j^{m_j}$
3. And all possible linear combinations of the above

However for the elements of (3), closure is lost. The dimensionality of such a subspace is  $n + m$ .

---

<sup>3</sup>Here  $\oplus$  is the direct sum defined as:

## 2.8 Hilbert Spaces

A Hilbert space  $H$  is simply a normed vector space (a Banach space), whose norm is defined as:

$$\|V\| := \sqrt{\langle V|V \rangle} \quad (2.2)$$

This is an axiomatic definition of a Hilbert space, but we are more concerned with the corollaries of it. All the Cauchy sequences<sup>4</sup> of functions in a Hilbert space always converge to a function that is also a member of the space i.e. it is said to be **complete** which implies that the integral of the absolute square of a function must converge<sup>5</sup>

$$\int_a^b |f(x)|^2 dx < \infty \quad (2.3)$$

Moreover this means that, any function in Hilbert space can be expressed as a linear combination of other functions i.e. it is closed/complete

$$f(x) = \sum_{n=1}^{\infty} c_n f_n(x) \quad (2.4)$$

Where,  $c_n \in \mathbb{C}$

## 2.9 Linear Operators

## 2.10 Eigenvalue Problem

## 2.11 Eigenfunctions of a Hermitian Operator

## 2.12 Transformations

## 2.13 Active Transformation

In a loose sense this can be thought of as,

---

<sup>4</sup>Definition

<sup>5</sup>we simply state this but a proof can be found in

## 2.14 Passive Transformation

From our discussion before it is also clear that the same transformation can be implemented as,

$$\hat{O} \rightarrow U^\dagger \hat{O} U \quad (2.5)$$

This is a very different viewpoint, we can understand this by visualizing it to be a

### 2.14.1 Equivalence of Transformation types

It's pretty simple to see that both types of transformation constitute the same physical picture. Thus, we can take both viewpoints to mean the same physical transformation in each case, and later on we will see how this leads us two different pictures of Quantum Mechanics and how they are related.

## 2.15 Functions of Operators

## 2.16 Generalization to Infinite Dimensions

## 2.17 Probability

### 2.17.1 Discrete Variables

Suppose we have a frequency distribution

$$N = \sum_{j=0}^{\infty} N(j) \quad (2.6)$$

The probability of  $N_j$  is defined as,

$$P(j) = \frac{N(j)}{N} \quad (2.7)$$

### 2.17.2 Continuous Variables

We now move to a continuous probability distribution, we'll create continuous analogs of all the quantities we just introduced. Let's start with probability, the probability of that  $x$  lies between  $a$  and  $b$

$$P_{ab} = \int_a^b \rho(x) dx \quad (2.8)$$

where  $\rho(x)$  is called the probability density i.e. the probability of getting  $x$ , or more concretely,

$\rho(x)dx$  = Probability that an individual is chosen at random lies between  $x$  and  $x+dx$

Now supposing the rules we held for discrete variables hold, the continuous analogs look like this:

$$1 = \int_{-\infty}^{\infty} \rho(x) dx \quad (2.9)$$

$$\langle x \rangle = \int_{-\infty}^{\infty} x \rho(x) dx \quad (2.10)$$

$$\langle f(x) \rangle = \int_{-\infty}^{\infty} f(x) \rho(x) dx \quad (2.11)$$

$$\sigma^2 := \langle (\Delta x)^2 \rangle = \langle x^2 \rangle - \langle x \rangle^2 \quad (2.12)$$

## 2.18 Expectation Values

In this section we'll explore how we express the expectation values of a few operators. Let's start with the position operator in the position representation (i.e. position basis):

$$\langle x \rangle = \int_{-\infty}^{\infty} x |\psi(\vec{x}, t)|^2 dx \quad (2.13)$$

We can differentiate 2.13 with respect to time to find the expectation value for "velocity":

$$\frac{d \langle x \rangle}{dt} =$$

Throwing away

$$\langle v \rangle = \frac{d \langle x \rangle}{dt} = -\frac{i\hbar}{m} \int \psi^* \frac{\partial \psi}{\partial x} dx \quad (2.14)$$

Therefore we can write the expectation value of momentum as,

$$\langle p \rangle = m \frac{d \langle x \rangle}{dt} = -i\hbar \int \left( \psi^* \frac{\partial \psi}{\partial x} \right) dx \quad (2.15)$$

In general, every observable is a function of position and momentum, thus for an observable  $\hat{O}(x, p)$ , the expectation value is given by,

$$\langle \hat{O}(x, p) \rangle = \int \psi^* \hat{O}(x, -i\hbar \nabla) \psi dx \quad (2.16)$$

For example, the expectation value of kinetic energy is,

$$\langle T \rangle = -\frac{\hbar^2}{2m} \int \psi^* \frac{\partial^2 \psi}{\partial x^2} dx \quad (2.17)$$

Or to sum it up in Dirac notation,

$$\langle \hat{O} \rangle = \langle \psi | \hat{O} | \psi \rangle \quad (2.18)$$

## 2.19 Fourier Analysis

### 2.19.1 Dirichelet's Theorem

### 2.19.2 Fourier Transform

## 2.20 Delta Function

### 2.20.1 The Divergence of $\frac{\hat{r}}{r^2}$

We can see why the divergence is,

$$\nabla \cdot \frac{\hat{r}}{r^2} = 0 \quad (2.19)$$

But if we calculate this using the Divergence theorem, we find that ,

$$\oint v \cdot da = \int \left( \frac{\hat{r}}{r^2} \right) \cdot (r^2 \sin(\theta) d\theta d\phi \hat{r}) = \left( \int_0^\pi \sin(\theta) d\theta \right) \left( \int_0^{2\pi} d\phi \right) = 4\pi \quad (2.20)$$

This is paradoxical. The issue is that it blows up at  $r = 0$  but is negligible everywhere else. How do we fix this? The Dirac Delta functional!

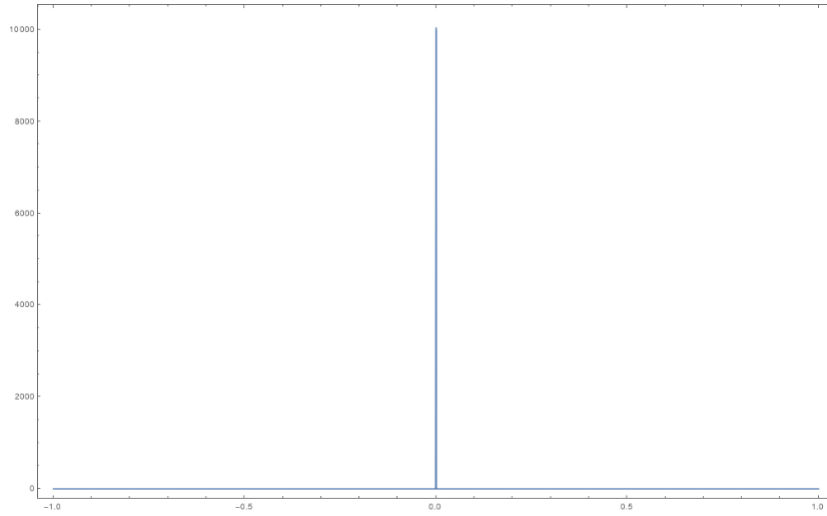
### 2.20.2 The One-Dimensional Dirac Delta Functional

The Dirac Delta is a functional <sup>6</sup> which we define as,

$$\delta(x - a) = \begin{cases} 0, & \text{if } x \neq a \\ \infty, & \text{if } x = a \end{cases} \quad (2.21)$$

$$\int_{-\infty}^{+\infty} \delta(x - a) dx = 1 \quad (2.22)$$



Figure 2.2: A Plot of  $\delta(x)$ 

$\forall a \in \mathbb{R}$  We can visualize it as a sharp peak at  $a$ , We can interpret 2.22 as saying "the area of the delta distribution is always 1".

$$f(x)\delta(x-a) = f(a) \quad (2.23)$$

We can combine these to get,

$$\int_{-\infty}^{+\infty} \delta(x-a)f(x)dx = f(a) \quad (2.24)$$

### A few interesting properties

$$\delta(kx) = \frac{1}{|k|}\delta(x) \quad (2.25)$$

$$\frac{d}{dx}(\delta(x)) = -\delta(x) \quad (2.26)$$

where  $k$  is a constant

$$\frac{d\theta}{dx} = \delta(x) \quad (2.27)$$

Where  $\theta$  is the step function defined as,

$$\theta(x) = \begin{cases} 1, & \text{if } x > 0 \\ 0, & \text{if } x \leq 0 \end{cases} \quad (2.28)$$

---

<sup>6</sup>An object that is a map between functions

### 2.20.3 The Three-Dimensional Dirac Delta Function

We generalize (2.21) to three dimensions,

$$\delta^3(\vec{r} - \vec{a}) = \delta(x - a_x)\delta(y - a_y)\delta(z - a_z) \quad (2.29)$$

$$\int_{-\infty}^{+\infty} \delta^3(\vec{r} - \vec{a}) dV = 1 \quad (2.30)$$

We can also define the three-dimensional delta function as

$$\delta^3(\mathbf{z}) = \frac{1}{4\pi} \left[ \nabla \cdot \left( \frac{\hat{\mathbf{z}}}{z^2} \right) \right] \quad (2.31)$$

Since,

$$\nabla \left( \frac{1}{z} \right) = -\frac{\hat{\mathbf{z}}}{z^2}$$

We can rewrite as,

$$\delta^3(\mathbf{z}) = -\frac{1}{4\pi} \left[ \nabla^2 \left( \frac{1}{z} \right) \right] \quad (2.32)$$

## 2.21 Gaussian Integrals

## 2.22 The $i\epsilon$ Prescription

We will now derive and interpret the formula:

$$\frac{1}{x \mp i\epsilon} = \mathcal{P} \frac{1}{x} \pm \pi \delta(x) \quad (2.33)$$

where  $\epsilon \rightarrow 0$  is a positive infinitesimally small quantity. Now we'll consider the integral

$$content... \quad (2.34)$$

$$a \quad (2.35)$$

$$asdfkj h \quad (2.36)$$

## 2.23 Permutation Functions

### 2.23.1 Kronecker delta

It simply has the ‘function’ of ‘renaming’ an index:

$$\delta_\nu^\mu x^\nu = x^\mu$$

it is in a sense simply the identity matrix. Or it is sometimes defined as:

$$\delta_{ij} = \begin{cases} 1 & \text{if } i = j \\ 0 & \text{if } i \neq j \end{cases} \quad (2.37)$$

### 2.23.2 Levi-Civita Pseudotensor

The Levi-Civita Pseudotensor i.e. Tensor density is a completely anti-symmetric i.e.  $\epsilon_{ijk} = -\epsilon_{jik} = -\epsilon_{ikj} = -\epsilon_{kji}$ , we define it as:

$$\epsilon_{ijk} = \begin{cases} 1 & \text{if } ijk \text{ is an even permutation of } 123 \\ -1 & \text{if } ijk \text{ is an odd permutation of } 123 \\ 0 & \text{if two indices are equal} \end{cases} \quad (2.38)$$

#### Identities

$$\epsilon_{\alpha\beta\nu}\epsilon_{\alpha\beta\sigma} = \delta_{\mu\rho}\delta_{\nu\sigma} - \delta_{\mu\sigma}\delta_{\nu\rho} \quad (2.39)$$

From this it follows that,

$$\epsilon_{\alpha\beta\nu}\epsilon_{\alpha\beta\sigma} = 2\delta_{\nu\sigma} \quad (2.40)$$

and

$$\epsilon_{\alpha\beta\gamma}\epsilon_{\alpha\beta\gamma} = 6 \quad (2.41)$$



# Chapter 3

## Formalism

In Quantum Mechanics, we start with an object called the state vector  $|\psi\rangle$ . All the information about the system is contained in it. The position basis representation of the state vector is called the wavefunction  $\psi(\vec{x}, t)$ . If we wish to know about a particular physical measurable such as an object's position or momentum, we can extract this information from the State vector by means of acting on it with an Operator that corresponds to the measurable quantity.

To get down to even more specifics if I consider an observable  $\hat{O}$ , then in general I have the form:

$$\hat{O} |\psi\rangle = o |\psi\rangle \quad (3.1)$$

Where,  $o$  is an Eigenvalue. The only types of operators that are constrained in such a fashion are "Hermitian Operators", they are identified with the condition:

$$\text{content} \dots \quad (3.2)$$

Where

If we consider the Schrodinger picture i.e. the State vector evolves with time whereas the Observables are in a loose sense eternal. The time evolution of the state vector is given by the Schrodinger equation:

$$i\hbar \frac{\partial |\psi\rangle}{\partial t} = \hat{H} |\psi\rangle \quad (3.3)$$

Or,

$$i\hbar \frac{\partial \psi}{\partial t} = \hat{H} \psi \quad (3.4)$$

in terms of the Wavefunction. Where,  $\hat{H}$  is the Hamiltonian operator, which

can be expressed as:

$$\hat{H} = -\frac{\hbar^2 \nabla^2}{2m} + V(\vec{x}) \quad (3.5)$$

for a free particle. According to Born's rule

$$\int_a^b |\psi(\vec{x}, t)|^2 dx = \text{Probability of finding the particle at a time } t \text{ between positions } a \text{ and } b \quad (3.6)$$

Thus, . Physically speaking this lends a kind of indeterminacy to the wavefunction. We can only speak of probabilities. Therefore, we can only , this brings to the measurement hypothesis, that is the State vector evolves to the state corresponding to the measurement being made. And unlike the Schrodinger equation, this evolution is non-deterministic. This tension is often called the "measurement problem", i.e. why is the measurement of an observable a special process distinct from others? Several theories and models claim to have resolved this, but we shall save that discussion for another time. We will fully focus on understanding the theory of Quantum Mechanics in a pragmatic lens before we question its foundations (although the converse isn't necessarily a bad thing, it isn't the purpose of this manuscript).

### 3.1 Normalization

Normalization is a process through which we ensure that,

$$\int_{-\infty}^{\infty} |\psi(\vec{x}, t)|^2 dx = 1 \quad (3.7)$$

This is a natural consequence of Born's rule, we simply want all the probabilities to add up to 1. Thus, to rule out any other absurd scenarios, we make a ruling that non-Normalizable and non-square integrable Wavefunctions are unphysical.

We can also prove that once normalized, the wavefunction always remains normalized, we start by differentiating 3.7 with respect to time

$$\frac{d}{dt} \int_{-\infty}^{\infty} |\psi(\vec{x}, t)|^2 dx = \frac{\partial}{\partial t} \int_{-\infty}^{\infty} |\psi(\vec{x}, t)|^2 dx$$

Dealing with the term inside the integral,

$$|\psi(\vec{x}, t)|^2 =$$

## 3.2 Summary of Postulates

## 3.3 Generalized Uncertainty Principle

Suppose we have a ket  $|\psi\rangle$  and two operators  $\hat{A}$  and  $\hat{B}$ , we define two new vectors

,

,

We use the Cauchy-Schwarz inequality,

$$2|X||Y| \geq |\langle X|Y\rangle + \langle Y|X\rangle|$$

Substituting in the left-hand side,  $2\sqrt{\langle X|X\rangle\langle Y|Y\rangle} \geq |\langle X|Y\rangle + \langle Y|X\rangle|$   
 Plugging in Eqs. (4) and (5),  $2\sqrt{\langle\psi|A^2|\psi\rangle\langle\psi|-B^2|\psi\rangle} \geq |\langle X|Y\rangle + \langle Y|X\rangle|$   
 Taking the  $-1$  outside,  $2i\sqrt{\langle\psi|A^2|\psi\rangle\langle\psi|B^2|\psi\rangle} \geq |\langle X|Y\rangle + \langle Y|X\rangle|$  We now  
 substitute in the right hand of the equation  $2i\sqrt{\langle\psi|A^2|\psi\rangle\langle\psi|B^2|\psi\rangle} \geq |\langle\psi|\hat{A}\hat{B}|\psi\rangle -$   
 $|\langle\psi|\hat{B}\hat{A}|\psi\rangle|$  The negative sign is due to the  $i$ , this also seems to represent  
 the commutator, so we substitute  $2i\sqrt{\langle\psi|A^2|\psi\rangle\langle\psi|B^2|\psi\rangle} \geq |\langle\psi|[\hat{A}, \hat{B}]|\psi\rangle|$   
 Again, the right hand side looks like the expectation value of a quantity,  
 so  $2i\sqrt{\langle A^2\rangle\langle B^2\rangle} \geq |\langle[\hat{A}, \hat{B}]\rangle|$   $\sqrt{\langle A^2\rangle\langle B^2\rangle} \geq \frac{1}{2i}|\langle[\hat{A}, \hat{B}]\rangle|$  We use Eq. (2),

$\sqrt{\sigma_A^2\sigma_B^2} \geq \frac{1}{2i}|\langle[\hat{A}, \hat{B}]\rangle|$  Removing the square root we get the expression:  
 $\sigma_A\sigma_B \geq \frac{1}{2i}|\langle[\hat{A}, \hat{B}]\rangle|$

This is called the generalized uncertainty principle. This basically states  
 that two variables that do not commute cannot be measured with precision  
 simultaneously.

Talking about position and momentum

We know that observable properties can be represented using operators,  
 here we'll

$\hat{x} = x$   $\hat{P} = -i\hbar\frac{\partial}{\partial x}$  So we now try to find the commutator now  $[\hat{x}, \hat{p}] =$   
 $\hat{x}\hat{p} - \hat{p}\hat{x}$   $[\hat{x}, \hat{p}] = -ix\hbar\frac{\partial}{\partial x} + i\hbar\frac{\partial}{\partial x}$  Now let's apply this to state vector to obtain  
 the expectation value  $[\hat{x}, \hat{p}]|\psi\rangle = -ix\hbar\frac{\partial}{\partial x}|\psi\rangle + i\hbar\frac{\partial x|\psi\rangle}{\partial x}$

$$[\hat{x}, \hat{p}]|\psi\rangle = -ix\hbar\frac{\partial}{\partial x}|\psi\rangle + ix\hbar\frac{\partial(|\psi\rangle)}{\partial x} + i\hbar$$

$[\hat{x}, \hat{p}]|\psi\rangle = i\hbar$  Substituting this into Eq.(),  $\sigma_x\sigma_p \geq \frac{1}{2i}i\hbar$   $\sigma_x\sigma_p \geq \frac{\hbar}{2}$   $\sigma_x\sigma_p \geq \frac{\hbar}{4\pi}$

## 3.4 Generalized Statistical Interpretation





# Chapter 4

## Toy Models

- 4.1 Time-Dependent Schrodinger Equation
- 4.2 Time-Independent Schrodinger Equation
- 4.3 Stationary States
- 4.4 The Infinite Square Well
- 4.5 Harmonic Oscillator
- 4.6 Free Particle
- 4.7 Delta-Function Potential
- 4.8 Finite Square Well
- 4.9 Wave-Packets



## Chapter 5

### Systems with $N$ degrees of freedom



# Chapter 6

## Symmetries and their Consequences

Shankar chapters 11,12 13, 14 ,15



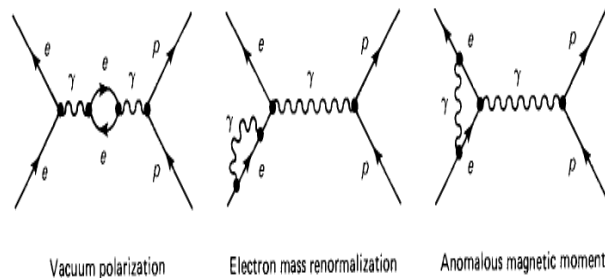
# Chapter 7

## Hydrogen Atom

### 7.1 Fine Structure

#### 7.1.1 Lamb Shift

An interesting feature of the fine structure formula is that it depends only on  $j$  and not  $l$ , moreover in general two different values of  $l$  share the same energy. For example, the  $2S$  and  $2P$  states should remain perfectly degenerate. However in 1947 Lamb and Retherford performed an experiment that displayed something to the contrary. The  $S$  state is slightly higher in energy than the  $p$  state. The explanation of this "Lamb" shift was later explained by Bethe, Feynman, Schwinger and Tomonaga (the founders of QED) as a corollary of the electromagnetic field itself being quantised. Sharply in contrast to the hyperfine structure of Hydrogen, the Lamb shift is a completely novel i.e. non-classical (as the hyperfine structure is explained through Coulomb's law and the Biot-Savart Law) phenomena. It arises from a radiative correction in Quantum Electrodynamics to which classical theories are mute. In Feynman lingo, this arises from loop corrections as potrayed below. Naively,



1. the first diagram describes pair-production in the neighborhood of a nucleus, leading to a partial screening effect of the proton's charge;
2. the second diagram reflects the fact that the electromagnetic field has a non-zero ground state
3. the third diagram leads to a tiny modification of the electron's magnetic dipole moment (an addition of  $a + \alpha/2\pi = 1.00116$ )

We shall not discuss the results in depth but rather consider two exemplary cases:

For  $l = 0$ ,

$$\Delta E_{Lamb} = \alpha^5 mc^2 \frac{1}{4n^3} [k(n, 0)] \quad (7.1)$$

Where  $k(n, 0)$  is a numerical factor defined as:

$$k(n, 0) = \begin{cases} 12.7, & \text{if } n = 1 \\ 13.2, & \text{if } n \rightarrow \infty \end{cases}$$

For  $l = 0$  and  $j = l \pm \frac{1}{2}$ ,

$$\Delta E_{Lamb} = \alpha^5 mc^2 \frac{1}{4n^3} \left[ k(n, 0) \pm \frac{1}{\pi(l + \frac{1}{2})(l + \frac{1}{2})} \right] \quad (7.2)$$

Here,  $k(n, l)$  is a very small number ( $< 0.05$ ) which varies a little with its arguments.

The Lamb shift is tiny except for the case  $l = 0$ , for which it amounts to about 10% of the fine structure. However, since it depends on  $l$ , it lifts the degeneracy of the pairs of states with common  $n$  and  $j$  and in particular it splits  $2S_{1/2}$  and  $2P_{1/2}$

## 7.2 The Zeeman Effect

When an atom is placed in a uniform magnetic field  $B_{Ext.}$ , the energy levels are shifted, this is known as the Zeeman effect. For the case of a single electron, the shift is:

$$H'_Z = -(\mu_l + \mu_s) \cdot B_{Ext.} \quad (7.3)$$

Where,

$$\text{content...} \quad (7.4)$$

is the magnetic dipole moment associated with electron spin, and

$$\text{content...} \quad (7.5)$$



is the dipole moment associated with orbital motion. The gyromagnetic ratio in this case is simply classical i.e.  $q/2m$ , it is only for spin that we have an extra factor of 2. We now rewrite ( ) as:

$$asd \quad (7.6)$$

The nature of the Zeeman splitting depends on the strength of the external field vs. the internal one that gives rise to spin-orbit/spin-spin coupling. This table provides a short review of the different cases:

Case	Name	Comments
$B_{Ext.} \gg B_{Int.}$	Strong-Field Zeeman Effect	Zeeman effect dominates; fine structure becomes the perturbation
$B_{Ext.} \ll B_{Int.}$	Weak-Field Zeeman Effect	Fine structure dominates; $H'_z$ can be treated as a small perturbation
$B_{Ext.} = B_{Int.}$	Intermediate Zeeman Effect	Both the fields are equal in strength thus we would need elements of degenerate perturbation theory and will need to diagonalize the necessary portion of the Hamiltonian "by hand"

In the next few sections we'll explore all of them in depth.

#### 7.2.1 Weak-Field Zeeman Effect

#### 7.2.2 Strong-Field Zeeman Effect

#### 7.2.3 Intermediate Zeeman Effect

### 7.3 Hyperfine Splitting in Hydrogen



## Chapter 8

# Approximations



## Chapter 9

# Perturbation Theory



# Chapter 10

## Scattering





# Chapter 11

## Path Integral Formulation



## Chapter 12

### Dirac Equation



## Chapter 13

### The Heisenberg Picture/Theorems that connect to classical mech

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## Chapter 14

## Epilogue: What lies ahead

