

# Introduction to Cosmology

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# Dark Matter

## Dark Matter in Clusters

The acceleration of the  $i^{th}$  galaxy in the coma cluster is given by ,

$$\ddot{\vec{x}}_i = G \sum_{j \neq i} m_j \frac{\vec{x}_j - \vec{x}_i}{(\vec{x}_j - \vec{x}_i)^3} \quad (1)$$

The gravitational potential energy of the system of N galaxies is

$$W = -\frac{G}{2} \sum_{j \neq i} \frac{m_i m_j}{|\vec{x}_j - \vec{x}_i|} \quad (2)$$

The potential energy of the cluster can also be written as,

$$W = -\alpha \frac{GM^2}{r_h} \quad (3)$$

# Dark Matter

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The kinetic energy associated with the relative motion of the galaxies in the cluster is,

$$K = \frac{1}{2} \sum_i m_i |\dot{\vec{x}}_i|^2 \quad (4)$$

The kinetic energy  $K$  can also be written in the form

$$K = \frac{1}{2} M \langle v^2 \rangle \quad (5)$$

# Dark Matter

## Dark Matter in Clusters

The moment of inertia is also an important factor to understand when we try to understand the radial velocities of galaxies. Hence the moment of inertia can be written as,

$$I = \sum_i m_i |\dot{\vec{x}}_i|^2 \quad (6)$$

The moment of inertia in terms of the potential and kinetic energy can be written as,

$$\ddot{I} = 2 \sum_i m_i (\vec{x}_i \times \ddot{\vec{x}}_i + \dot{\vec{x}}_i \times \dot{\vec{x}}_i) \quad (7)$$

(7) can be written as,

$$\ddot{I} = 2 \sum_i m_i (\vec{x}_i \times \ddot{\vec{x}}_i) + 4K \quad (8)$$

# Dark Matter

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From equation (8) we can formulate the expression for potential energy.

$$\sum_i m_i (\mathbf{x}_i \times \ddot{\mathbf{x}}_i) = G \sum_i m_i m_j \frac{\dot{\mathbf{x}}_i \times (\mathbf{x}_j - \mathbf{x}_i)}{|\mathbf{x}_i - \mathbf{x}_j|^3} \quad (9)$$

$$\sum_i m_i (\mathbf{x}_i \times \ddot{\mathbf{x}}_i) = \sum_j m_j (\mathbf{x}_j \times \ddot{\mathbf{x}}_j) \quad (10)$$

Using equations (9) and (10),

$$\sum_i m_i (\mathbf{x}_i \times \ddot{\mathbf{x}}_i) = \frac{1}{2} \left[ \sum_i m_i (\mathbf{x}_i \times \ddot{\mathbf{x}}_i) + \sum_j m_j (\mathbf{x}_j \times \ddot{\mathbf{x}}_j) \right] \quad (11)$$

$$\sum_i m_i (\mathbf{x}_i \times \ddot{\mathbf{x}}_i) = -\frac{G}{2} \sum_{ij} \frac{m_i m_j}{|\mathbf{x}_i - \mathbf{x}_j|} = W \quad (12)$$

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We can substitute this relation in equation (8),

$$\ddot{I} = 2W + 4K \quad (13)$$

This relation is known as the virial theorem

If  $I = \text{constant}$ , then the steady-state virial theorem is,

$$0 = W + 2K \quad (14)$$

$$K = -\frac{W}{2} \quad (15)$$

Using (3) and (5) in (15),

$$\frac{1}{2}M\langle v^2 \rangle = \frac{\alpha GM^2}{2r_h} \quad (16)$$

# Dark Matter

## Dark Matter in Clusters

Using this we can calculate the mass of the cluster with the help of virial theorem,

$$M = \frac{\langle v^2 \rangle r_h}{\alpha G} \quad (17)$$

## Dark Matter in Clusters

About the coma Cluster.....

$$z = 0.0232 \quad (18)$$

$$d_{coma} = \left(\frac{C}{H_0}\right) \langle z \rangle = 99 Mpc \quad (19)$$

# Dark Matter

## Dark Matter in Clusters

$$v_r = cz = 6960 \text{ km s}^{-1} \quad (20)$$

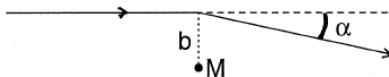
$$r_h = 1.5 \text{ Mpc} = 4.6 \times 10^{22} \quad (21)$$

$$M_{coma} = \frac{\langle v^2 \rangle r_h}{\alpha G} \approx 2 \times 10^{15} M_0 \quad (22)$$



# Dark Matter

## Gravitational Lensing



From Einstein's theory of General Relativity, If a photon passes such a compact massive object at an impact parameter  $b$ , the local curvature of space-time will cause the photon to be deflected by an angle

$$\alpha = \frac{4GM}{c^2 b} \quad (23)$$

If the MACHO is exactly along the line of sight between the observer and the lensed star, the angular radius of the image produced is given by,

$$\theta_E = \left( \frac{4GM}{c^2 b} \frac{1-x}{x} \right)^{\frac{1}{2}} \quad (24)$$

# Dark Matter

Dark Matter - Probably the matter which matters the most.

What is it? Why do we care so much about it? What is its significance in our universe?