# Band Stacking, RGB & False Color Images Using Functions

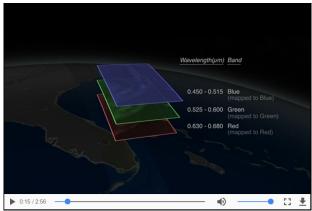
#### **Objectives**

In the first lesson, we learned how to read in hdf5 data using h5py, apply the no-data value and scale factor, and plot a single band of a reflectance data tile. We can do these tasks more efficiently using functions. In this lesson, we will introduce functions aop\_h5ref12array to . We will also plot different combinations of bands, and demonstrate how to create IPython widgets for more interactive data visualization.

#### **Background on RGB & False-Color Images**

We can combine any three bands from the NEON reflectance data to make an **RGB** image that allows us to visualize spectral data collected at the Earth's surface. A **natural-color** image, made with bands from the red, green, and blue wavelengths looks close to what we would see with the naked eye, or what we would see in a typical photograph. We can also choose band combinations from other wavelengths, and map them to the red, blue, and green colors to highlight different spectral features. A **false-color** image is made with one or more bands from a non-visible portion of the electromagnetic spectrum that are mapped to red, green, and blue colors. These images can display other information about the landscape that is not easily seen with a natural-color image.

The NASA Goddard Media Studio video *Peeling Back Landsat's Layers of Data* gives a good overview of natural and false color band combinations. Note that Landsat collects information from 11 wavelength bands, while NEON AOP hyperspectral data collects information from 426 bands!



(https://www.youtube.com/watch?v=YP0et8l\_bvY)

https://www.youtube.com/watch?v=YP0et8l\_bvY (https://www.youtube.com/watch?v=YP0et8l\_bvY) https://svs.gsfc.nasa.gov//vis/a010000/a011400/a011491/index.html (https://svs.gsfc.nasa.gov//vis/a010000/a011400/a011491/index.html)

Further Reading: Check out the NASA Earth Observatory Article on "How to Interpret a False-Color Satellite Image"

https://earthobservatory.nasa.gov/Features/FalseColor/ (https://earthobservatory.nasa.gov/Features/FalseColor/)

Before we get started, let's import our standard packages, and set up our plot and warning preferences:

```
In [1]:
    import matplotlib.pyplot as plt
    import numpy as np
    import h5py, os, osr, copy
    %matplotlib inline
    import warnings
    warnings.filterwarnings('ignore')
```

The first function we will use is aop\_h5ref12array . This function is loaded into the cell below, we encourage you to look through the code to understand what it is doing -- most of these steps should look familiar to you from the first lesson. This function can be thought of as a wrapper to automate the steps required to read AOP hdf5 reflectance tiles into a Python format. This function also cleans the data: it sets any no data values within the reflectance tile to nan (not a number) and applies the reflectance scale factor so the final array that is returned represents unitless scaled reflectance, with values ranging between 0 and 1 (0-100%).

```
def aop h5refl2array(refl filename):
In [2]:
            """aop h5refl2array reads in a NEON AOP reflectance hdf5 file and returns
                   1. reflectance array (with the no data value and reflectance scale factor applied)
                   2. dictionary of metadata including spatial information, and wavelengths of the bands
            Parameters
                refl_filename -- full or relative path and name of reflectance hdf5 file
            Returns
            reflArray:
                array of reflectance values
            metadata:
                dictionary containing the following metadata:
                    bad band window1 (tuple)
                    bad band window2 (tuple)
                    bands: # of bands (float)
                    data ignore value: value corresponding to no data (float)
                    epsg: coordinate system code (float)
                    map info: coordinate system, datum & ellipsoid, pixel dimensions, and origin coordinates (string)
                    reflectance scale factor: factor by which reflectance is scaled (float)
                    wavelength: wavelength values (float)
                    wavelength unit: 'm' (string)
            NOTE: This function applies to the NEON hdf5 format implemented in 2016, and should be used for
            data acquired 2016 and after. Data in earlier NEON hdf5 format (collected prior to 2016) is
            expected to be re-processed after the 2018 flight season.
            Example Execution:
            sercRef1, sercRef1 metadata = h5ref12array('NEON D02 SERC DP3 368000 4306000 reflectance.h5') """
            import h5py
            #Read in reflectance hdf5 file
            hdf5_file = h5py.File(refl_filename, 'r')
            #Get the site name
            file_attrs_string = str(list(hdf5_file.items()))
            file_attrs_string_split = file_attrs_string.split("'")
            sitename = file_attrs_string_split[1]
            #Extract the reflectance & wavelength datasets
            refl = hdf5_file[sitename]['Reflectance']
            reflData = refl['Reflectance Data']
            reflRaw = refl['Reflectance_Data'].value
            #Create dictionary containing relevant metadata information
            metadata = {}
            metadata['map info'] = ref1['Metadata']['Coordinate_System']['Map_Info'].value
            metadata['wavelength'] = refl['Metadata']['Spectral_Data']['Wavelength'].value
            #Extract no data value & scale factor
            metadata['data ignore value'] = float(reflData.attrs['Data_Ignore_Value'])
            metadata['reflectance scale factor'] = float(reflData.attrs['Scale_Factor'])
            #metadata['interleave'] = reflData.attrs['Interleave']
            #Apply no data value
            reflClean = reflRaw.astype(float)
            arr size = reflClean.shape
            if metadata['data ignore value'] in reflRaw:
                print('% No Data: ',np.round(np.count_nonzero(reflClean==metadata['data ignore value'])*100/(arr_size[
                nodata_ind = np.where(reflClean==metadata['data ignore value'])
                reflClean[nodata_ind]=np.nan
            #Apply scale factor
            reflArray = reflClean/metadata['reflectance scale factor']
            #Extract spatial extent from attributes
            metadata['spatial extent'] = reflData.attrs['Spatial Extent meters']
            #Extract bad band windows
            metadata['bad band window1'] = (refl.attrs['Band Window 1 Nanometers'])
            metadata['bad band window2'] = (refl.attrs['Band_Window_2_Nanometers'])
```

```
#Extract projection information
#metadata['projection'] = refl['Metadata']['Coordinate_System']['Proj4'].value
metadata['epsg'] = int(refl['Metadata']['Coordinate_System']['EPSG Code'].value)

#Extract map information: spatial extent & resolution (pixel size)
mapInfo = refl['Metadata']['Coordinate_System']['Map_Info'].value

hdf5_file.close
return reflArray, metadata
```

If you forget what this function does, or don't want to scroll up to read the docstrings, remember you can use help or ? to display the associated docstrings.

```
In [3]:
        help(aop_h5refl2array)
        aop_h5refl2array?
        Help on function aop_h5refl2array in module __main__:
        aop h5refl2array(refl filename)
            aop h5refl2array reads in a NEON AOP reflectance hdf5 file and returns
                   1. reflectance array (with the no data value and reflectance scale factor applied)
                   2. dictionary of metadata including spatial information, and wavelengths of the bands
            Parameters
                refl_filename -- full or relative path and name of reflectance hdf5 file
            Returns
            reflArray:
                array of reflectance values
            metadata:
                dictionary containing the following metadata:
                    bad band window1 (tuple)
                    bad_band_window2 (tuple)
                    bands: # of bands (float)
                    data ignore value: value corresponding to no data (float)
                    epsg: coordinate system code (float)
                    map info: coordinate system, datum & ellipsoid, pixel dimensions, and origin coordinates (strin
        g)
                    reflectance scale factor: factor by which reflectance is scaled (float)
                    wavelength: wavelength values (float)
                    wavelength unit: 'm' (string)
            NOTE: This function applies to the NEON hdf5 format implemented in 2016, and should be used for
            data acquired 2016 and after. Data in earlier NEON hdf5 format (collected prior to 2016) is
            expected to be re-processed after the 2018 flight season.
            Example Execution:
            sercRef1, sercRef1 metadata = h5ref12array('NEON D02 SERC DP3 368000 4306000 reflectance.h5')
```

Now that we have an idea of how this function works, let's try it out. First, define the path where the reflectance data is stored and use os.path.join to create the full path to the data file. Note that if you want to run this notebook later on a different reflectance tile, you just have to change this variable.

```
In [4]: serc_h5_tile = ('../data/NEON_D02_SERC_DP3_368000_4306000_reflectance.h5')
```

Now that we've specified our reflectance tile, we can call aop\_h5ref12array to read in the reflectance tile as a python array called sercRef1, and the associated metadata into a dictionary sercMetadata

```
In [5]: sercRef1,sercMetadata = aop_h5ref12array(serc_h5_tile)
```

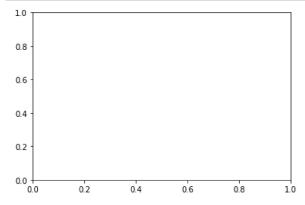
We can use the shape method to see the dimensions of the array we read in. NEON tiles are (1000 x 1000 x # of bands), the number of bands may vary depending on the hyperspectral sensor used, but should be in the vicinity of 426.

```
In [6]: sercRefl.shape
Out[6]: (1000, 1000, 426)
```

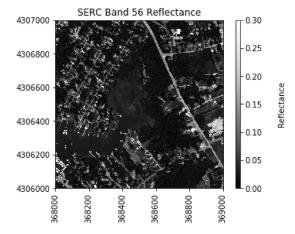
#### Use plot\_aop\_refl to plot a single band of reflectance data

Next we'll use the function plot\_aop\_refl to plot a single band of reflectance data. Read the Parameters section of the docstring to understand the required inputs & data type for each of these; only the band and spatial extent are required inputs, the rest are optional inputs that, if specified, allow you to set the range color values, specify the axis, add a title, colorbar, colorbar title, and change the colormap (default is to plot in greyscale).

```
In [7]: def plot_aop_refl(band_array,refl_extent,colorlimit=(0,1),ax=plt.gca(),title='',cbar ='on',cmap_title='',color
             '''plot_refl_data reads in and plots a single band or 3 stacked bands of a reflectance array
            Parameters
                band array: array of reflectance values, created from aop h5refl2array
                refl extent: extent of reflectance data to be plotted (xMin, xMax, yMin, yMax)
                             use metadata['spatial extent'] from aop_h5refl2array function
                colorlimit: optional, range of values to plot (min, max).
                            - helpful to look at the histogram of reflectance values before plotting to determine colo
                ax: optional, default = current axis
                title: optional; plot title (string)
                cmap_title: optional; colorbar title
                colormap: optional (string, see https://matplotlib.org/examples/color/colormaps_reference.html) for li
            Returns
                plots flightline array of single band of reflectance data
            Examples:
            plot_aop_refl(sercb56,
                      sercMetadata['spatial extent'],
                      colorlimit=(0,0.3),
                      title='SERC Band 56 Reflectance',
                      cmap_title='Reflectance',
                      colormap='Greys_r') '''
            import matplotlib.pyplot as plt
            plot = plt.imshow(band_array,extent=refl_extent,clim=colorlimit);
                cbar = plt.colorbar(plot,aspect=40); plt.set_cmap(colormap);
                cbar.set label(cmap title,rotation=90,labelpad=20)
            plt.title(title); ax = plt.gca();
            ax.ticklabel_format(useOffset=False, style='plain'); #do not use scientific notation for ticklabels
            rotatexlabels = plt.setp(ax.get_xticklabels(),rotation=90); #rotate x tick labels 90 degrees
```



Now that we have loaded this function, let's extract a single band from the SERC reflectance array and plot it:



#### **RGB Plots - Band Stacking**

It is often useful to look at several bands together. We can extract and stack three reflectance bands in the red, green, and blue (RGB) spectrums to produce a color image that looks like what we see with our eyes; this is your typical camera image. In the next part of this tutorial, we will learn to stack multiple bands and make a geotif raster from the compilation of these bands. We can see that different combinations of bands allow for different visualizations of the remotely-sensed objects and also conveys useful information about the chemical makeup of the Earth's surface.

For this exercise, we'll first use the neon\_aop\_module function  $stack_rgb$  to extract the bands we want to stack. This function uses splicing to extract the nth band from the reflectance array, and then uses the numpy function stack to create a new 3D array (1000 x 1000 x 3) consisting of only the three bands we want.

```
In [10]: def stack_rgb(reflArray,bands):
    import numpy as np

    red = reflArray[:,:,bands[0]-1]
    green = reflArray[:,:,bands[1]-1]
    blue = reflArray[:,:,bands[2]-1]

    stackedRGB = np.stack((red,green,blue),axis=2)

    return stackedRGB
```

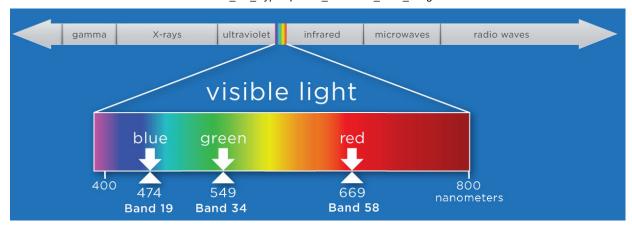
First we will look at red, green, and blue bands, whos indices are defined below. To confirm that these band indices correspond to wavelengths in the expected portion of the spectrum, we can print out the wavelength values stored in <code>metadata['wavelength']</code>:

```
In [11]: rgb_bands = (58,34,19)

print('Band 58 Center Wavelength = %.2f' %(sercMetadata['wavelength'][57]),'nm')
print('Band 33 Center Wavelength = %.2f' %(sercMetadata['wavelength'][33]),'nm')
print('Band 19 Center Wavelength = %.2f' %(sercMetadata['wavelength'][18]),'nm')

Band 58 Center Wavelength = 668.98 nm
Band 33 Center Wavelength = 548.80 nm
Band 19 Center Wavelength = 473.68 nm
```

We selected these bands so that they fall within the visible range of the electromagnetic spectrum (400-700 nm):



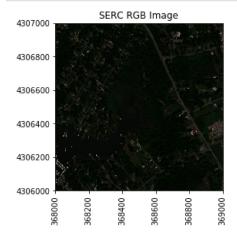
Below we use stack\_rgb to create an RGB array. Check that the dimensions of this array are as expected.

• TIP: checking the shape of arrays with .shape is a good habit to get into when creating your own workflows, and can be a handy tool for troubleshooting.

```
In [12]: SERCrgb = stack_rgb(sercRefl,rgb_bands)
SERCrgb.shape
Out[12]: (1000, 1000, 3)
```

### Use plot\_aop\_refl to plot an RGB band combination

Next, we can use the function <code>plot\_aop\_ref1</code> , even though we have more than one band. This function only works for a single or 3-band array, so ensure the array you use has the proper dimensions before using. You do not need to specify the colorlimits as the <code>matplotlib.pyplot</code> automatically scales 3-band arrays to 8-bit color (256).

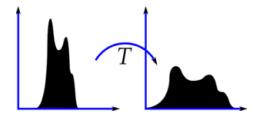


You'll notice that this image is very dark; it is possible to make out some of the features (roads, buildings), but it is not ideal. Since colorlimits don't apply to 3-band images, we have to use some other image processing tools to enhance the visibility of this image.

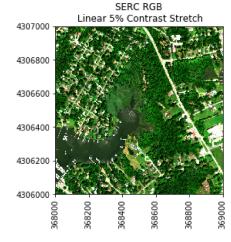
#### Image Processing -- Contrast Stretch & Histogram Equalization

We can also try out some image processing routines to better visualize the reflectance data using the ski-image package.

Histogram equalization is a method in image processing of contrast adjustment using the image's histogram. Stretching the histogram can improve the contrast of a displayed image by eliminating very high or low reflectance values that skew the display of the image.



Let's see what the image looks like using a 5% linear contrast stretch using the skiimage module's exposure function.



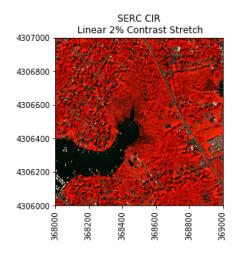
# False Color Image - Color Infrared (CIR)

We can also create an image from bands outside of the visible spectrum. An image containing one or more bands outside of the visible range is called a **false-color image**. Here we'll use the green and blue bands as before, but we replace the red band with a near-infrared (NIR) band.

For more information about non-visible wavelengths, false color images, and some frequently used false-color band combinations, refer to NASA's Earth Observatory page:

https://earthobservatory.nasa.gov/Features/FalseColor/ (https://earthobservatory.nasa.gov/Features/FalseColor/)

```
Band 90 Center Wavelength = 829.24 nm
Band 34 Center Wavelength = 548.80 nm
Band 19 Center Wavelength = 473.68 nm
```



# **Demo: Exploring Band Combinations Interactively**

Now that we have made a couple different band combinations, we can demo a Python widget to explore different combinations of bands in the visible and non-visible portions of the spectrum.

```
In [18]: from IPython.html.widgets import *
```

```
In [19]:
           array = copy.copy(sercRefl)
           metadata = copy.copy(sercMetadata)
           def RGBplot_widget(R,G,B):
               #Pre-allocate array size
               rgbArray = np.zeros((array.shape[0],array.shape[1],3), 'uint8')
               Rband = array[:,:,R-1].astype(np.float)
               #Rband clean = clean band(Rband, Refl md)
               Gband = array[:,:,G-1].astype(np.float)
               #Gband clean = clean band(Gband, Refl md)
               Bband = array[:,:,B-1].astype(np.float)
               #Bband clean = clean band(Bband, Refl md)
               rgbArray[..., 0] = Rband*256
               rgbArray[..., 1] = Gband*256
               rgbArray[..., 2] = Bband*256
               # Apply Adaptive Histogram Equalization to Improve Contrast:
               img nonan = np.ma.masked invalid(rgbArray) #first mask the image
               img_adapteq = exposure.equalize_adapthist(img_nonan, clip_limit=0.10)
               plot = plt.imshow(img_adapteq,extent=metadata['spatial extent']);
               plt.title('Bands: \nR:' + str(R) + ' (' + str(round(metadata['wavelength'][R-1])) + 'nm)' + '\n G:' + str(G) + ' (' + str(round(metadata['wavelength'][G-1])) + 'nm)' + '\n B:' + str(B) + ' (' + str(round(metadata['wavelength'][B-1])) + 'nm)');
               ax = plt.gca(); ax.ticklabel format(useOffset=False, style='plain')
               rotatexlabels = plt.setp(ax.get_xticklabels(),rotation=90)
           interact(RGBplot widget, R=(1,426,1), G=(1,426,1), B=(1,426,1))
```

Failed to display Jupyter Widget of type interactive.

If you're reading this message in the Jupyter Notebook or JupyterLab Notebook, it may mean that the widgets JavaScript is still loading. If this message persists, it likely means that the widgets JavaScript library is either not installed or not enabled. See the Jupyter Widgets Documentation (https://ipywidgets.readthedocs.io/en/stable/user\_install.html) for setup instructions.

If you're reading this message in another frontend (for example, a static rendering on GitHub or <u>NBViewer (https://nbviewer.jupyter.org/)</u>), it may mean that your frontend doesn't currently support widgets.

```
Out[19]: <function __main__.RGBplot_widget>
```

#### **Demo: Interactive Linear Stretch & Equalization**

Here is another widget to play around with, demonstrating how to interactively visualize linear contrast stretches with a variable percent.

```
In [20]: rgbArray = copy.copy(SERCrgb)

def linearStretch(percent):
    pLow, pHigh = np.percentile(rgbArray[~np.isnan(rgbArray)], (percent,100-percent))
    img_rescale = exposure.rescale_intensity(rgbArray, in_range=(pLow,pHigh))
    plt.imshow(img_rescale,extent=sercMetadata['spatial extent'])
    plt.title('SERC RGB \n Linear ' + str(percent) + '% Contrast Stretch');
    ax = plt.gca()
    ax.ticklabel_format(useOffset=False, style='plain')
    rotatexlabels = plt.setp(ax.get_xticklabels(),rotation=90)

interact(linearStretch,percent=(0,20,1))
```

Failed to display Jupyter Widget of type interactive.

If you're reading this message in the Jupyter Notebook or JupyterLab Notebook, it may mean that the widgets JavaScript is still loading. If this message persists, it likely means that the widgets JavaScript library is either not installed or not enabled. See the <u>Jupyter Widgets Documentation (https://ipywidgets.readthedocs.io/en/stable/user\_install.html)</u> for setup instructions.

If you're reading this message in another frontend (for example, a static rendering on GitHub or <u>NBViewer (https://nbviewer.jupyter.org/)</u>), it may mean that your frontend doesn't currently support widgets.

```
Out[20]: <function __main__.linearStretch>
```

#### References & Resources

Kekesi, Alex et al. "NASA | Peeling Back Landsat's Layers of Data". https://svs.gsfc.nasa.gov/vis/a010000/a011400/a011491/ (https://svs.gsfc.nasa.gov/vis/a010000/a011400/a011491/). Published on Feb 24, 2014.

Riebeek, Holli. "Why is that Forest Red and that Cloud Blue? How to Interpret a False-Color Satellite Image" <a href="https://earthobservatory.nasa.gov/Features/FalseColor/">https://earthobservatory.nasa.gov/Features/FalseColor/</a> (https://earthobservatory.nasa.gov/Features/FalseColor/)

Histogram Equalization: http://opencv-python-

tutroals.readthedocs.io/en/latest/py\_tutorials/py\_imgproc/py\_histograms/py\_histogram\_equalization/py\_histogram\_equalization.html (http://opencv-python-

tutroals.readthedocs.io/en/latest/py\_tutorials/py\_imgproc/py\_histograms/py\_histogram\_equalization/py\_histogram\_equalization.html)

skikit-image tutorial: <a href="http://scikit-image.org/docs/stable/auto\_examples/color\_exposure/plot\_equalize.html#sphx-glr-auto-examples-color-exposure-plot-equalize.html#sphx-glr-auto-examples/color\_exposure-plot\_equalize.html#sphx-glr-auto-examples-color-exposure-plot-equalize.py">http://scikit-image.org/docs/stable/auto\_examples/color\_exposure/plot\_equalize.html#sphx-glr-auto-examples-color-exposure-plot-equalize-py</a>)