1. Prob 12.9

(a) The magnetic induction generated by dipole M is

$$\mathbf{B}(\mathbf{x}) = \frac{1}{r^3} [3(\mathbf{M} \cdot \hat{\mathbf{r}}) \hat{\mathbf{r}} - \mathbf{M}]$$
 (1)

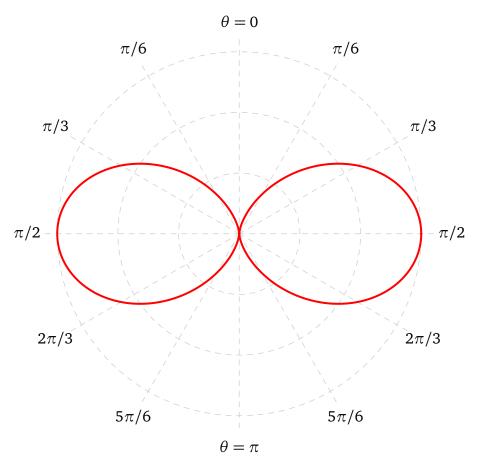
Let **M** be along the z-axis, and with $\hat{\mathbf{z}} = \cos\theta \hat{\mathbf{r}} - \sin\theta \hat{\boldsymbol{\theta}}$, we can write the above as

$$\mathbf{B}(\mathbf{x}) = \frac{1}{r^3} \left[3M \cos \theta \,\hat{\mathbf{r}} - M \left(\cos \theta \,\hat{\mathbf{r}} - \sin \theta \,\hat{\boldsymbol{\theta}} \right) \right] = \frac{1}{r^3} \left(2M \cos \theta \,\hat{\mathbf{r}} + M \sin \theta \,\hat{\boldsymbol{\theta}} \right)$$
(2)

The tangent of the field line must be proportional to the ratio of the corresponding components of B, i.e.

$$\frac{dr}{rd\theta} = \frac{B_r}{B_\theta} = \frac{2\cos\theta}{\sin\theta} \qquad \Longrightarrow \qquad \frac{dr}{r} = 2\frac{d\left(\sin\theta\right)}{\sin\theta} \qquad \Longrightarrow \qquad r \propto \sin^2\theta \qquad (3)$$

Below is a plot of a field line.



(b) At $\theta = \pi/2$, the radius of curvature is

$$R_C = \frac{\left[r^2 + \left(\frac{dr}{d\theta}\right)^2\right]^{3/2}}{r^2 + 2\left(\frac{dr}{d\theta}\right)^2 - r\frac{d^2r}{d\theta^2}}\bigg|_{r=R} = \frac{R}{3}$$

$$(4)$$

The gradient drift velocity can be obtained via (12.61) with $v_{\parallel} = 0$,

$$\mathbf{v}_C = \frac{1}{\omega_B R_C} \cdot \frac{v_\perp^2}{2} \left(\frac{\mathbf{R}_C \times \mathbf{B}_0}{R_C B_0} \right) = \frac{3}{2} \frac{v_\perp^2}{\omega_B R} = \frac{3}{2} \frac{\omega_B a^2}{R} \hat{\boldsymbol{\phi}}$$
 (5)

Note that at $\theta = \pi/2$, **B** is along the direction $\hat{\theta}$, and \mathbf{R}_C is pointing from the center of curvature to the charge, so \mathbf{v}_C is in the direction of $\hat{\phi}$. Since **M** points to the south, ϕ points to the west. The drift angular frequency is given by \mathbf{v}_C/R , or in terms of time dependency of east longitude, we have

$$\phi(t) = \phi_0 - \frac{3}{2} \left(\frac{a}{R} \right)^2 \omega_B(t - t_0) \tag{6}$$

(c) Let $\xi = \pi/2 - \theta$ be the (south) latitude angle (recall that **M** points to the south). Then (2) can be rewritten in terms of ξ

$$\mathbf{B}(\mathbf{x}) = \frac{M}{r^3} \left(2\sin \xi \,\hat{\mathbf{r}} + \cos \xi \,\hat{\boldsymbol{\theta}} \right) \tag{7}$$

For the force line at r = R, we can substitude $r = R \sin^2 \theta = R \cos^2 \xi$ into (7) and get the magnitude of **B** as

$$B = \frac{M}{R^3} \frac{\sqrt{\cos^2 \xi + 4\sin^2 \xi}}{\cos^6 \xi} \approx B_0 \left(1 + 3\xi^2\right)^{1/2} \left(1 - \frac{\xi^2}{2}\right)^{-6} \approx B_0 \left(1 + \frac{9\xi^2}{2}\right)$$
(8)

We can apply the adiabatic invariance discussed in section 12.5 to the particle and invoke equation (12.72)

$$v_{\parallel}^{2}(\xi) = v_{0}^{2} - v_{\perp 0}^{2} \frac{B(\xi)}{B_{0}} \approx v_{0}^{2} - v_{\perp 0}^{2} \left(1 + \frac{9\xi^{2}}{2}\right) = v_{\parallel}^{2}(0) - \frac{9}{2}(\omega_{B}a)^{2} \xi^{2}$$
(9)

As is interpreted in the text, this is equivalent to a harmonic potential

$$V(z) = \frac{1}{2}m \cdot \frac{9}{2} \left(\frac{\omega_B a}{R}\right)^2 z^2 \tag{10}$$

giving rise to an angular frequency of

$$\Omega = \frac{3}{\sqrt{2}} \omega_B \left(\frac{a}{R}\right) \tag{11}$$

The change in longitude per cycle of oscillation in latitude is then

$$\Delta \phi = \frac{2\pi}{\Omega} \cdot \frac{3}{2} \left(\frac{a}{R}\right)^2 \omega_B = \sqrt{2}\pi \left(\frac{a}{R}\right) \tag{12}$$

(d) For $M = 8.1 \times 10^{25} \text{ gauss-cm}^3$, at $R = 3 \times 10^9 \text{ cm}$, we have

$$B_0 = \frac{M}{R^3} = 3 \times 10^{-7} \text{Tesla} \tag{13}$$

For an electron with kinetic energy of 10MeV,

$$(\gamma - 1)mc^2 = 10 \text{MeV}$$
 \Longrightarrow $\gamma = 1 + \frac{10 \text{MeV}}{0.511 \text{MeV}} \approx 20.6$ (14)

Then (in SI units)

$$\omega_B = \frac{eB_0}{\gamma m} = 2.57 \times 10^3 \text{rad/s} \tag{15}$$

giving the gyration radius

$$a \approx \frac{c}{\omega_B} \approx 1.17 \times 10^5 \text{m}$$
 (16)

Furthermore, from (11) and (5), we have

$$T_{\theta} = \frac{2\pi}{\Omega} \approx 0.3$$
s $T_{\phi} = \frac{2\pi}{\frac{3}{2} \left(\frac{a}{R}\right)^2 \omega_B} \approx 107$ s (17)

Repeating the calculation for kinetic energy of 10KeV or $\gamma \approx$ 1.02, we have

$$\omega_B \approx 5.19 \times 10^4 \text{rad/s}$$
 $a \approx 1.13 \times 10^3 \text{m}$ $T_\theta \approx 1.52 \text{s}$ $T_\phi \approx 5.7 \times 10^4 \text{s}$ (18)

2. Prob 12.10

The adiabatic invariance derivation allows us to plug the exact form of (8) into (9), giving

$$v_{\parallel}^{2}(\xi) = v_{0}^{2} - v_{\perp 0}^{2} \frac{\sqrt{1 + 3\sin^{2}\xi}}{\cos^{6}\xi} \qquad \Longrightarrow$$

$$v_{\parallel}^{2}(\xi) = v_{\parallel}^{2}(0) - v_{\perp 0}^{2} \left(\frac{\sqrt{1 + 3\sin^{2}\xi}}{\cos^{6}\xi} - 1\right) \qquad \Longrightarrow$$

$$\frac{v_{\parallel}^{2}(\xi)}{v_{\perp 0}^{2}} = \tan^{2}\alpha - \left(\frac{\sqrt{1 + 3\sin^{2}\xi}}{\cos^{6}\xi} - 1\right)$$
(19)

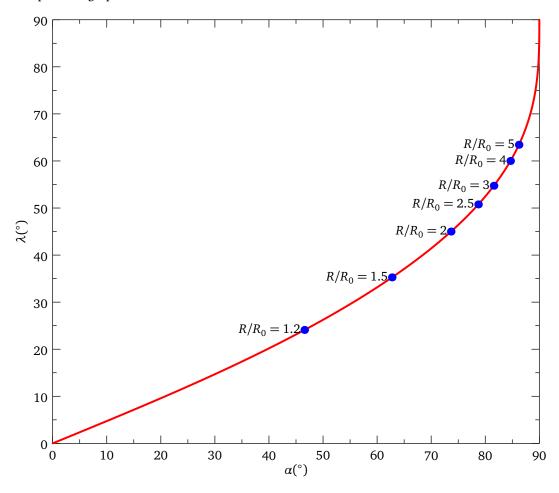
The maximum ξ is achieved at $\xi = \lambda$ where the LHS is zero, or

$$\frac{\sqrt{1+3\sin^2\lambda}}{\cos^6\lambda} = \sec^2\alpha \tag{20}$$

For the plot, it is easier to calculate the inverse function

$$\alpha = \cos^{-1} \left[\frac{\cos^3 \lambda}{\left(1 + 3\sin^2 \lambda \right)^{1/4}} \right] \tag{21}$$

and then transpose the graph.



From the field line equation in the previous problem, for the particle to hit earth, we must have

$$\cos^2 \lambda = \frac{R_0}{R} \qquad \Longrightarrow \qquad \lambda = \cos^{-1} \sqrt{\frac{R_0}{R}} \tag{22}$$

The numerical values are recorded in the table below and marked in the graph above.

R/R_0	1.2	1.5	2	2.5	3	4	5
α	46.58°	62.76°	73.67°	78.72°	81.59°	84.66°	86.22°