



1. For the radiation zone $kr \gg 1$, we can use (9.8) to find the vector potential

$$\mathbf{A}(\mathbf{x}) = \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int \mathbf{J}(\mathbf{x}') e^{-ik\mathbf{n} \cdot \mathbf{x}'} d^3x' \quad (1)$$

We choose the coordinate frame so that the observation point \mathbf{x} is at $(r \sin \theta, 0, r \cos \theta)$, hence $\mathbf{n} = \sin \theta \hat{\mathbf{x}} + \cos \theta \hat{\mathbf{z}}$ and a point on the loop has the form $\mathbf{x}' = a \cos \phi' \hat{\mathbf{x}} + a \sin \phi' \hat{\mathbf{y}}$.

The integral in (1) becomes

$$\begin{aligned} \mathbf{A}(\mathbf{x}) &= \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int_0^{2\pi} d\phi' \hat{\boldsymbol{\phi}} I_0 e^{-ika \sin \theta \cos \phi'} \\ &= \frac{\mu_0 I_0}{4\pi} \frac{e^{ikr}}{r} \int_0^{2\pi} e^{-ika \sin \theta \cos \phi'} (-\sin \phi' \hat{\mathbf{x}} + \cos \phi' \hat{\mathbf{y}}) d\phi' \end{aligned} \quad (2)$$

The $\hat{\mathbf{x}}$ component of the integral is obviously zero, the $\hat{\mathbf{y}}$ component can be obtained by recalling [DLMF 10.9.E2](#)

$$J_n(z) = \frac{i^{-n}}{\pi} \int_0^\pi e^{iz \cos \alpha} \cos(n\alpha) d\alpha \quad (3)$$

and letting $n = 1, z = -ka \sin \theta$, which gives

$$\mathbf{A}(\mathbf{x}) = \hat{\mathbf{y}} \frac{\mu_0 I_0}{4\pi} \frac{e^{ikr}}{r} \frac{2\pi}{i^{-1}} J_1(-ka \sin \theta) = -\hat{\mathbf{y}} \frac{i\mu_0 I_0}{2} \frac{e^{ikr}}{r} J_1(ka \sin \theta) \quad (4)$$

The radiation zone fields can be obtained from the vector potential by

$$\mathbf{H} = \frac{ik}{\mu_0} \mathbf{n} \times \mathbf{A} = \frac{kI_0}{2} \frac{e^{ikr}}{r} J_1(ka \sin \theta) \mathbf{n} \times \hat{\mathbf{y}} = \frac{kI_0}{2} \frac{e^{ikr}}{r} J_1(ka \sin \theta) (\sin \theta \hat{\mathbf{z}} - \cos \theta \hat{\mathbf{x}}) \quad (5)$$

$$\mathbf{E} = Z_0 \mathbf{H} \times \mathbf{n} = \frac{Z_0 k I_0}{2} \frac{e^{ikr}}{r} J_1(ka \sin \theta) (\sin \theta \hat{\mathbf{z}} - \cos \theta \hat{\mathbf{x}}) \times \mathbf{n} = \frac{Z_0 k I_0}{2} \frac{e^{ikr}}{r} J_1(ka \sin \theta) \hat{\mathbf{y}} \quad (6)$$

and the power radiated per unit solid angle is

$$\frac{dP}{d\Omega} = \frac{1}{2} \text{Re} [r^2 \mathbf{n} \cdot (\mathbf{E} \times \mathbf{H}^*)] = \frac{Z_0 k^2 I_0^2}{4} [J_1(ka \sin \theta)]^2 \quad (7)$$

2. By (9.170), (9.172), the only feasible multipole moments are M_{lm}

$$M_{lm} = -\frac{1}{l+1} \int r^l Y_{lm}^*(\theta, \phi) \nabla \cdot (\mathbf{x} \times \mathbf{J}) d^3x \quad (8)$$

where \mathbf{J} is most conveniently expressed in spherical coordinates (see 5.33)

$$\mathbf{J}(r, \theta, \phi) = I_0 \frac{\delta(r-a)}{a} \sin \theta \delta(\cos \theta) \hat{\boldsymbol{\phi}} = I_0 \frac{\delta(r-a)}{a} \delta\left(\theta - \frac{\pi}{2}\right) \hat{\boldsymbol{\phi}} \quad (9)$$

thus

$$\begin{aligned}
\mathbf{x} \times \mathbf{J} &= -r I_0 \frac{\delta(r-a)}{a} \delta\left(\theta - \frac{\pi}{2}\right) \hat{\boldsymbol{\theta}} \\
\nabla \cdot (\mathbf{x} \times \mathbf{J}) &= -r I_0 \frac{\delta(r-a)}{a} \cdot \frac{1}{r \sin \theta} \frac{d}{d\theta} \left[\delta\left(\theta - \frac{\pi}{2}\right) \sin \theta \right] \\
&= -I_0 \frac{\delta(r-a)}{a} \left[\delta'\left(\theta - \frac{\pi}{2}\right) + \delta\left(\theta - \frac{\pi}{2}\right) \cot \theta \right]
\end{aligned} \tag{10}$$

giving

$$\begin{aligned}
M_{lm} &= \frac{I_0}{l+1} \int_0^\infty r^l \frac{\delta(r-a)}{a} r^2 dr \int Y_{lm}^*(\theta, \phi) \left[\delta'\left(\theta - \frac{\pi}{2}\right) + \overbrace{\delta\left(\theta - \frac{\pi}{2}\right) \cot \theta}^{\text{no effect on integral}} \right] d\Omega \\
&= \frac{I_0 a^{l+1}}{l+1} \delta_{m0} \cdot 2\pi \sqrt{\frac{2l+1}{4\pi}} \int_0^\pi P_l(\cos \theta) \delta'\left(\theta - \frac{\pi}{2}\right) \sin \theta d\theta \\
&= \delta_{m0} \frac{2\pi I_0 a^{l+1}}{l+1} \sqrt{\frac{2l+1}{4\pi}} \left\{ -\frac{d[P_l(\cos \theta) \sin \theta]}{d\theta} \right\}_{\theta=\pi/2} \\
&= \delta_{m0} \frac{2\pi I_0 a^{l+1}}{l+1} \sqrt{\frac{2l+1}{4\pi}} P'_l(0)
\end{aligned} \tag{11}$$

The only non-vanishing multipole moments are M_{l0} where l is odd, the lowest of which is the dipole $l = 1$,

$$M_{10} = \sqrt{\frac{3}{4\pi}} \cdot I_0 \pi a^2 \tag{12}$$

where $\mathbf{m} = \hat{\mathbf{z}} I_0 \pi a^2$ is the dipole moment of the current loop in Cartesian tensor form, and the factor $\sqrt{3/4\pi}$ is required by the spherical tensor normalization (see 4.5).