1. The effective transitional magnetization is

$$\mathcal{M}(\mathbf{x},t) = \frac{\mathbf{x} \times \mathbf{J}(\mathbf{x},t)}{2} = -\frac{i\nu_0}{4} \cdot \frac{a_0}{2} \mathbf{x} \times \hat{\mathbf{z}} \rho (\mathbf{x},t)$$

$$= -\frac{i\nu_0 a_0}{4} \left( \frac{r \sin \theta \sin \phi \hat{\mathbf{x}} - r \sin \theta \cos \phi \hat{\mathbf{y}}}{r \cos \theta} \right) \rho (\mathbf{x},t)$$

$$= -\frac{i\alpha c a_0}{4} \tan \theta (\hat{\mathbf{x}} \sin \phi - \hat{\mathbf{y}} \cos \phi) \rho (\mathbf{x},t)$$

$$= \frac{i\alpha c a_0}{4} \tan \theta \rho (\mathbf{x},t) \hat{\boldsymbol{\phi}}$$
(1)

Since  $\mathcal{M}$  has only  $\hat{\phi}$  component and that component is independent of  $\phi$ , we know

$$\nabla \cdot \mathcal{M}(\mathbf{x}, t) = \frac{1}{r \sin \theta} \frac{\partial \mathcal{M}_{\phi}(\mathbf{x}, t)}{\partial \phi} = 0$$
 (2)

For long-wavelength, (9.172) applies, so all the magnetic multipole moments vanish. For electric multipole moments, by definition

$$Q_{lm} = \int Y_{lm}^{*}(\theta, \phi) r^{l} \rho(\mathbf{x}) d^{3}x = \int Y_{lm}^{*}(\theta, \phi) r^{l} \frac{2q_{e}}{\sqrt{6}a_{0}^{4}} r e^{-3r/2a_{0}} Y_{00}(\theta, \phi) Y_{10}(\theta, \phi) d^{3}x$$

$$= \delta_{l1} \delta_{m0} \cdot \frac{2q_{e}}{\sqrt{6}a_{0}^{4}} \sqrt{\frac{1}{4\pi}} \int_{0}^{\infty} r^{4} e^{-3r/2a_{0}} dr \qquad \text{recall } \int_{0}^{\infty} x^{n} e^{-ax} = \frac{n!}{a^{n+1}}$$

$$= \delta_{l1} \delta_{m0} \frac{2q_{e}}{\sqrt{6}a_{0}^{4}} \sqrt{\frac{1}{4\pi}} \frac{4!}{(3/2a_{0})^{5}} = \delta_{l1} \delta_{m0} \frac{256}{81\sqrt{6\pi}} q_{e} a_{0} \qquad (3)$$

We see the only multipole moment that enters is  $Q_{10}$ , which corresponds to the harmonically oscillating dipole along the  $\hat{\mathbf{z}}$  direction.

With the presence of **J**, or equivalently the effective magnetization  $\mathcal{M}$ , there is also a contribution to the electric multipole from magnetization source (9.170),

$$Q'_{lm} = -\frac{ik}{(l+1)c} \int r^{l} Y_{lm}^{*}(\theta, \phi) \overbrace{\nabla \cdot (\mathbf{x} \times \mathcal{M})}^{-\mathbf{x} \cdot (\nabla \times \mathcal{M})} d^{3}x$$
 (4)

Note

$$\mathbf{x} \cdot (\nabla \times \mathcal{M}) = \mathbf{x} \cdot \left[ \nabla \times \left( \frac{i\alpha c a_0}{4} \tan \theta \frac{2q_e}{\sqrt{6}a_0^4} r e^{-3r/2a_0} \sqrt{\frac{1}{4\pi}} \sqrt{\frac{3}{4\pi}} \cos \theta \hat{\boldsymbol{\phi}} \right) \right]$$

$$= \frac{i\alpha c a_0}{4} \cdot \frac{2q_e}{\sqrt{6}a_0^4} \sqrt{\frac{1}{4\pi}} \sqrt{\frac{3}{4\pi}} r^2 e^{-3r/2a_0} \frac{1}{r \sin \theta} \frac{d \sin^2 \theta}{d\theta}$$

$$= \frac{i\alpha c a_0}{4} \cdot 2\rho \left( \mathbf{x} \right)$$
(5)

Thus

$$Q'_{lm} = -\frac{k\alpha a_0}{2(l+1)}Q_{lm} = -\frac{k\alpha a_0}{4}\delta_{l1}\delta_{m0}Q_{10}$$
(6)

At typical wavelengths,  $Q'_{10} \ll Q_{10}$ .

2. Ignoring  $Q'_{10}$ , we can use (4.5) to find **p** from  $Q_{10}$ ,

$$\mathbf{p} = \sqrt{\frac{4\pi}{3}} Q_{10} \hat{\mathbf{z}} \tag{7}$$

The total power is given by (9.24)

$$P = \frac{c^2 Z_0 k^4}{12\pi} |\mathbf{p}|^2 = \frac{c^2 Z_0 k^4}{9} Q_{10}^2 = \left(\frac{2}{3}\right)^8 \cdot \hbar \omega_0 \frac{\alpha^4 c}{a_0} \qquad \text{where } \omega_0 = \frac{3q_e^2}{32\pi \epsilon_0 \hbar a_0}$$
 (8)

3. Numerically, the reciprocal time is

$$\left(\frac{2}{3}\right)^8 \frac{\alpha^4 c}{a_0} \approx 6.27 \times 10^8 \text{ s}^{-1} \tag{9}$$

4. Under this model, the dipole will be

$$\mathbf{p}(t) = 2q_e a_0 (\cos \omega t, \sin \omega t, 0) = 2q_e a_0 \operatorname{Re} \left[ e^{-i\omega t} (1, i, 0) \right]$$
(10)

The power derived from this dipole is

$$P = \frac{c^2 Z_0 k^4}{12\pi} \cdot 8q_e^2 a_0^2 \tag{11}$$

The ratio to the power from part b is

$$\frac{P_{(d)}}{P_{(b)}} = \frac{8}{\left(\frac{256}{81\sqrt{6\pi}}\sqrt{\frac{4\pi}{3}}\right)^2} \approx 3.60$$
(12)