



<https://tinyurl.com/Bennett-Skinner-PLANETSEDGE25>

# Signs of Migrating Icy Worlds in the Radius Valley

Bennett Skinner, Ralph Pudritz, and Ryan Cloutier

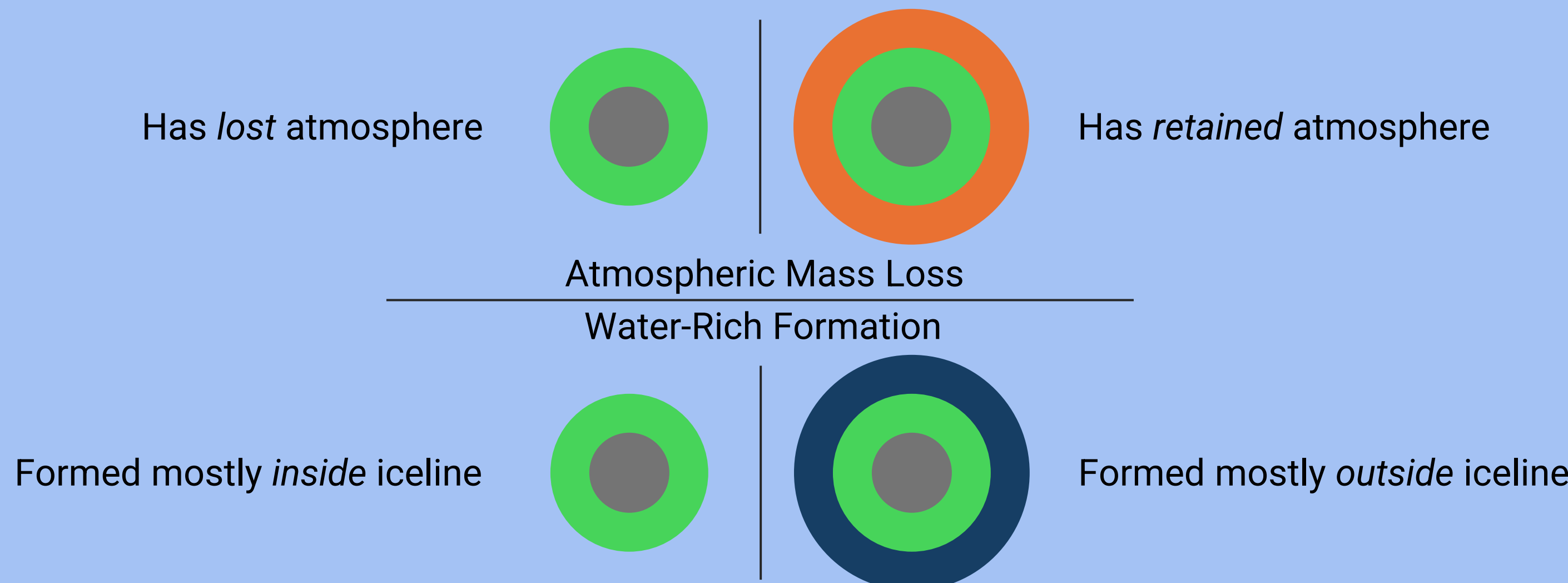
ALL RESULTS NOT FINAL AND WILL CHANGE BEFORE PUBLICATION!



skinnb1@mcmaster.ca

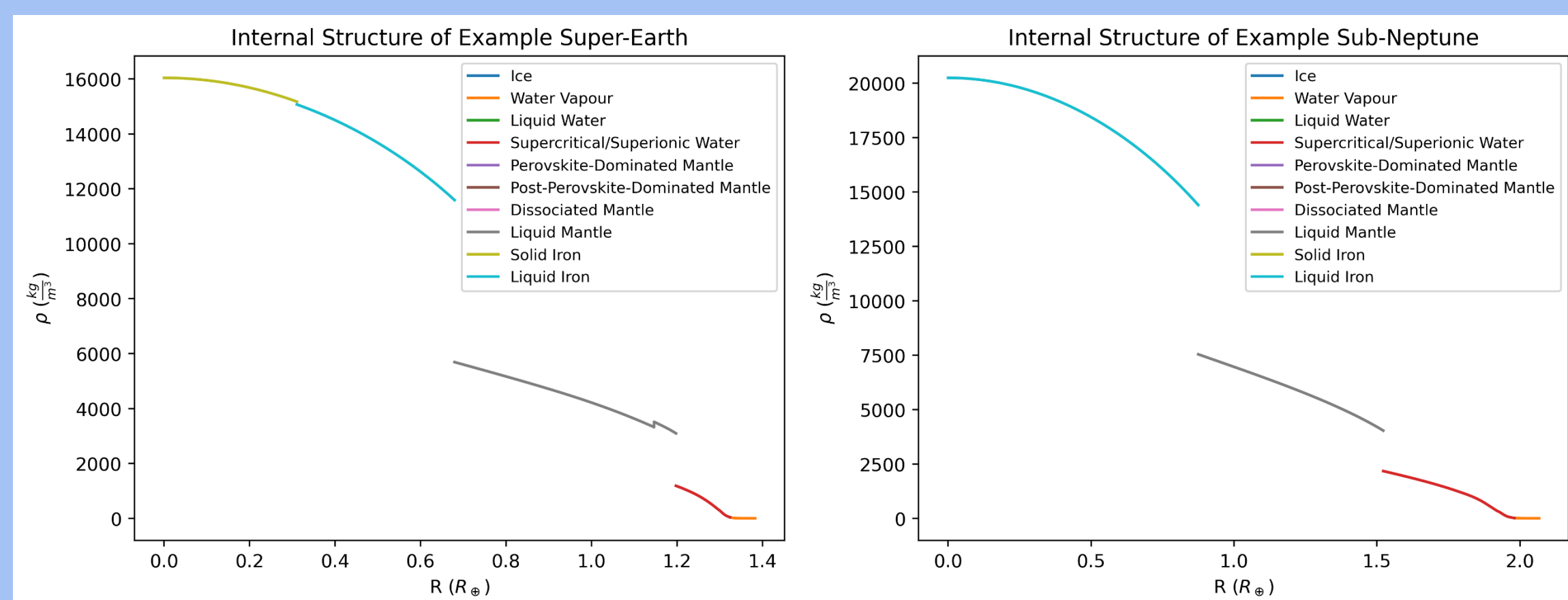
## Radius Valley

- Distribution of planetary radii is bimodal – “Radius Valley”<sup>1</sup>
  - ~1.3  $R_{\oplus}$  super-Earth peak and ~2.4  $R_{\oplus}$  sub-Neptune peak<sup>1</sup>
  - Super-Earth  $\rho$  consistent w/ Earth-like
  - Sub-Neptune  $\rho$  consistent w/ H/He OR water envelope
- Sub-Neptune composition depends on valley formation
  - Atmospheric mass loss  $\rightarrow$  H/He-rich
  - Water-rich formation  $\rightarrow$  water-rich
- FGK slope w/ instellation consistent w/ envelope mass loss<sup>2,3</sup>
- M slope different v. FGK stars<sup>4</sup>, possible water-rich formation
- Both mechanisms could contribute<sup>5</sup>



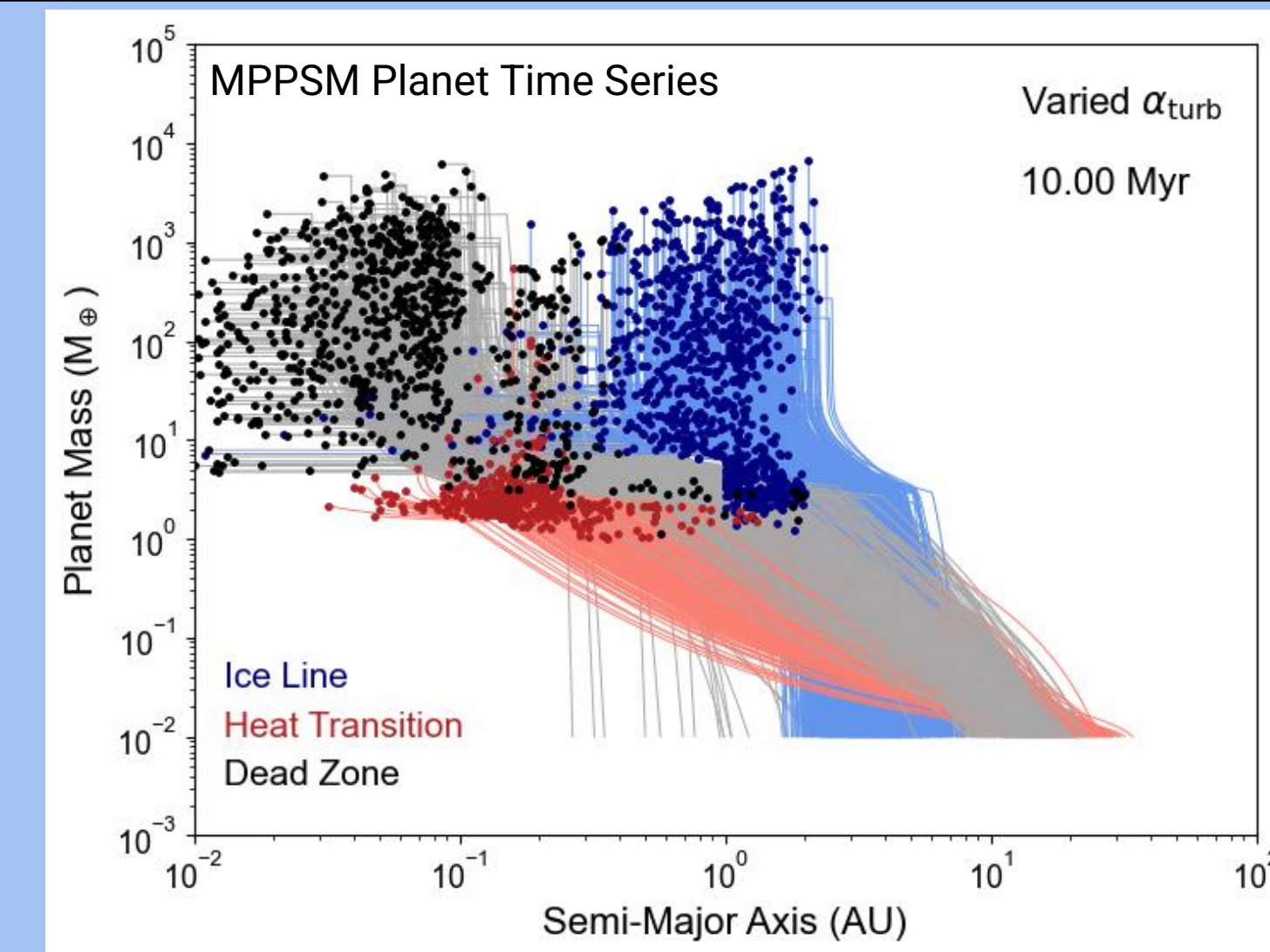
## Interior Structure Updates

- Determine planetary radii from MPPSM output parameters<sup>7</sup>
- Now use most modern Equations of State (EOS)<sup>8-25</sup>
  - Experimental measurements of  $\rho$  using diamond anvils<sup>22</sup>
  - Apply Density Functional Theory (DFT) to high pressures<sup>23</sup>
- More advanced model
  - Mantle composition via Gibbs free energy minimization<sup>13</sup>
  - Non-grey irradiated atmosphere<sup>26-29</sup>
  - Thermal effects (including melting) within mantle, core<sup>8-25</sup>
  - FeS in the core<sup>21,23</sup>
  - Prescription for rotation<sup>30-31</sup>
- Transit radii calculated to compare to observations<sup>26</sup>



## Planet Population Synthesis

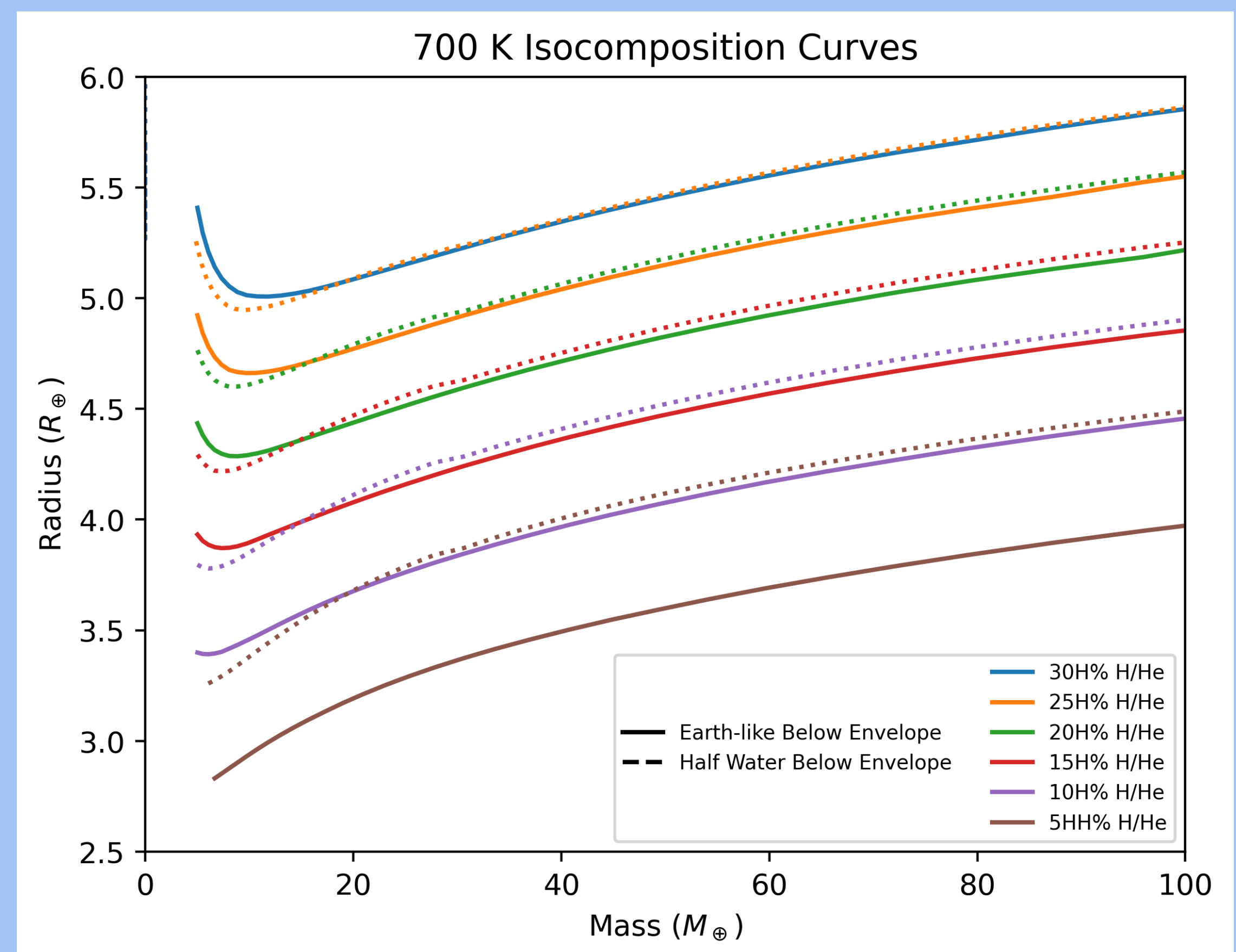
- Simulate planet formation, create synthetic population
- McMaster Planet Population Synthesis Model (MPPSM)<sup>6</sup>
  - Planetesimal accretion in disk around sunlike star
  - One lunar-mass embryo/disk
  - Power-law disk evolved via turbulent viscosity
  - Disk wind-driven advection and mass loss
  - Planets form in planet traps where disk conditions change<sup>7</sup>
    - Ice Line: Water vapour  $\rightarrow$  condensed
    - Heat Transition: Heat via viscous dissipation  $\rightarrow$  irradiation
    - Dead-Active Zone: MRI inactive  $\rightarrow$  MRI active



Alessi & Pudritz (2022) Fig. 7<sup>6</sup>

## New Mass-Radius Curves

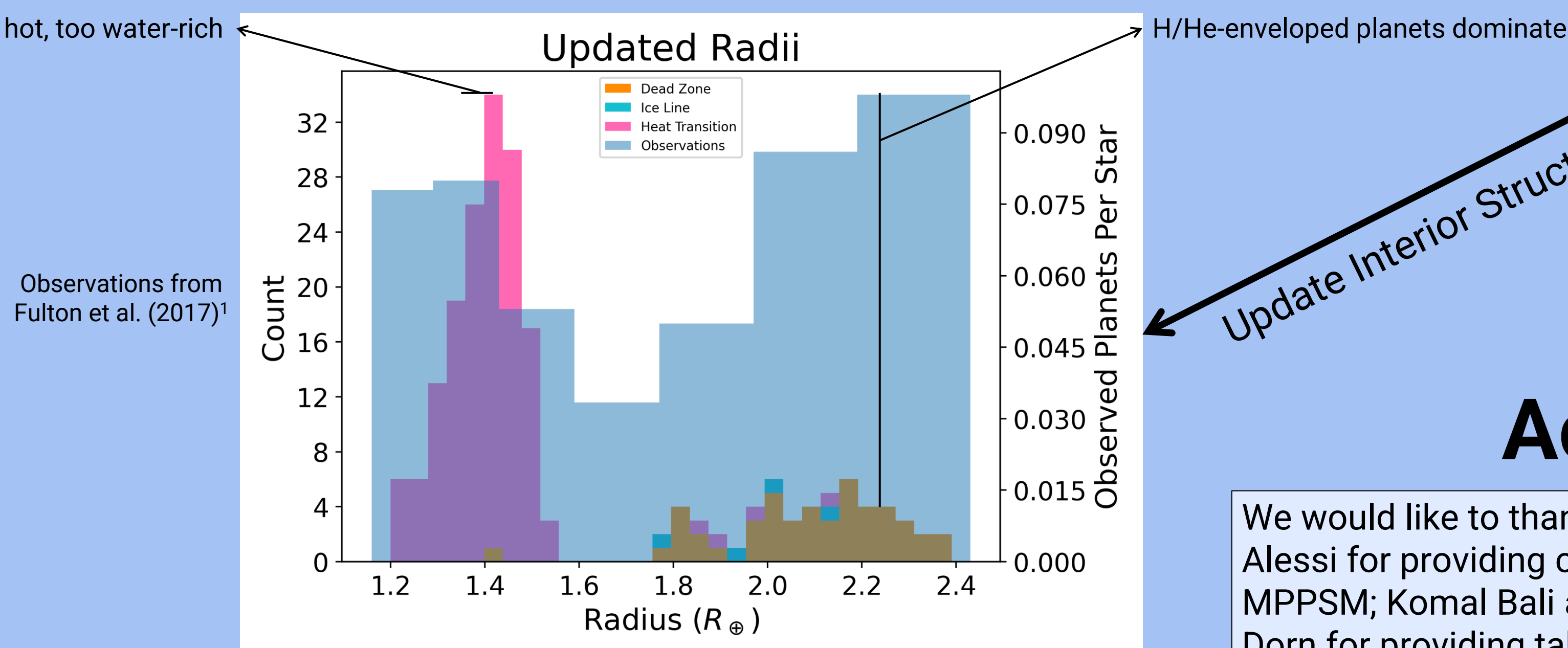
- Radii calculated for varying masses, compositions
- Current mass-radius curves old EOS or not made for high T



## Synthetic Radius Valley

- Calculated radii of MPPSM planets w/ new interior structure
  - Is radius valley replicated *without* H/He-rich planets?
- Peak radii approximated but relative occurrence incorrect
  - Water worlds could be *secondary* FGK sub-Neptune source
- Updated interior structure model *required* for radius valley
- W/o winds, no radius valley b/c less migration

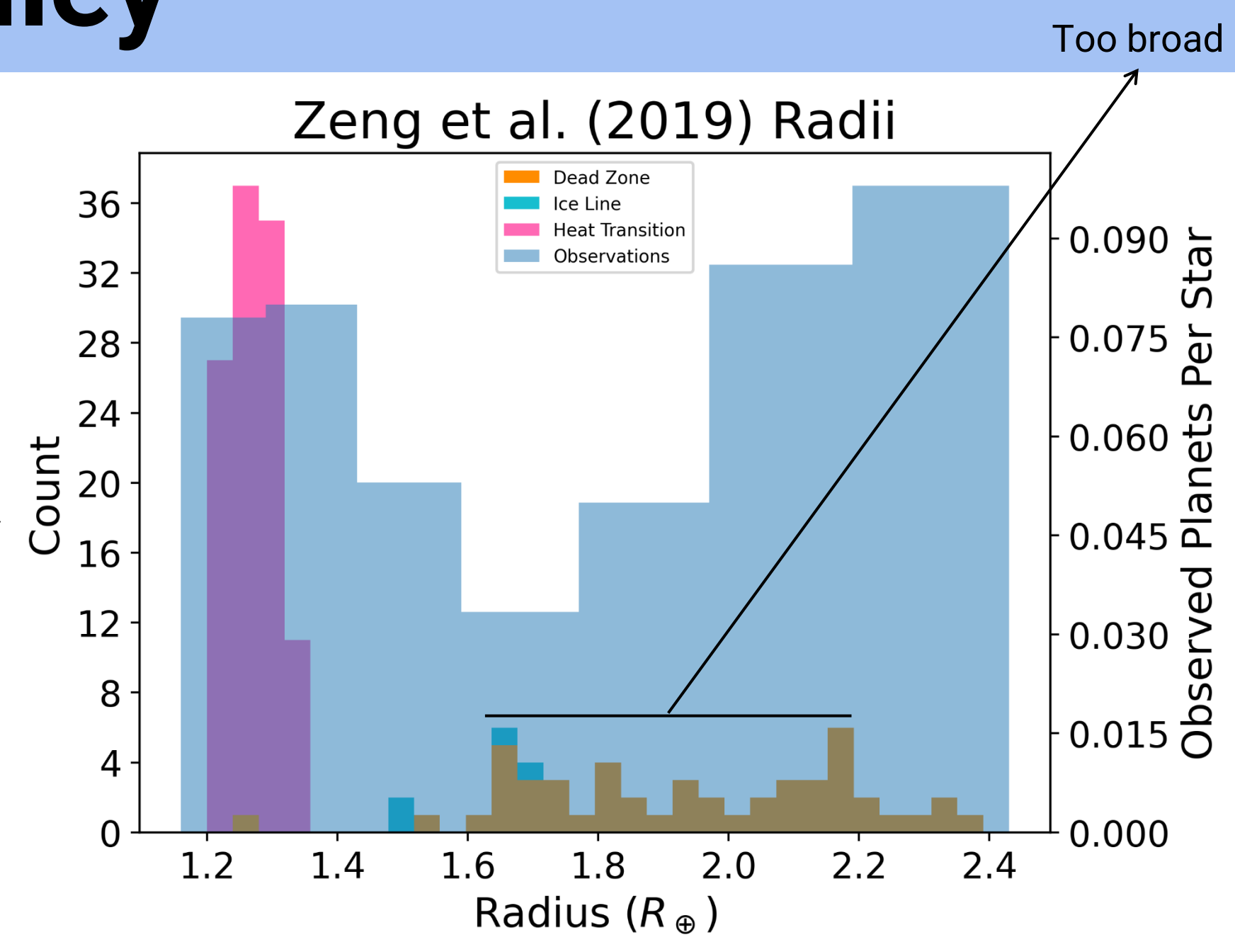
Planets too hot, too water-rich



## Future Work

- Incorporate metals in H/He envelope
- Non-radiogenic heating
- Implement active-dead zone boundary
- Create and run MPPSM for M dwarf stars
- Statistically quantify synthetic-observational discrepancy
- Apply improved interior structure to MPPSM advances
  - Pebble accretion, new wind prescription, N-body, etc.

We would like to thank Matthew Alessi for providing code used in MPPSM; Komal Bali and Caroline Dorn for providing tabulated MR relationships from their interior structure model; and Jonas Haldemann for insights about water equations of state. We would also like to thank the creators of the BICEPS model, Jonas Haldemann, Caroline Dorn, Julia Venturini, Yann Alibert, and Willy Benz, for coagulating many of the EOS we use in our model.



Observations from Fulton et al. (2017)<sup>1</sup>

## Acknowledgements & References

- Fulton, B. J., Petigura, E. A., Howard, A. W., Isaacson, H., Marcy, G. W., Cargile, P. A., Hebb, L. Weiss, L. M., Johnson, J. A., Morton, T. D., Sinukoff, E., Crossfield, I. J. M., and Hirsch, L. A. (2017). The California-Kepler Survey. III. A Gap in the Radius Distribution of Small Planets. *154*(3):109.
- Lopez, E. D. and Rice, K. (2018). How formation time-scales affect the period dependence of the transition between rocky super-Earths and gaseous sub-Neptunes and implications for  $\rho_p$ .
- Gupta, A. and Schlichting, H. E. (2019). Sculpting the valley in the radius distribution of small exoplanets as a by-product of planet formation: the core-powered mass-loss mechanism.
- Cloutier, R. and Menou, K. (2020). Evolution of the Radius Valley around Low-mass Stars from Kepler and K2.
- Burn, C., Moras, C., Mishra, L., Haldemann, J., Venturini, J., Enshuhuber, A., and Henning, T. (2024). A radius valley between migrated steam worlds and evaporated rocky cores.
- Alessi, M. and Pudritz, R. E. (2022). Combined effects of disc winds and turbulence-driven accretion on planet populations.
- Alessi, M., Pudritz, R. E., and Omland, A. J. (2020). Formation of planetary populations. II. Effects of initial disc size and radial dust drift.
- Chabrier, G., Mazevet, S., and Soubiran, F. (2019). A New Equation of State for Dense Hydrogen-Helium Mixtures. II. Taking into Account Hydrogen-Helium Interactions.
- Chabrier, G. and Debras, F. (2021). A New Equation of State for Dense Hydrogen-Helium Mixtures. II. Taking into Account Hydrogen-Helium Interactions.
- Howard, S. and Guillot, T. (2023). Accounting for non-ideal mixing effects in the hydrogen-helium equation of state.
- Haldemann, J., Alibert, Y., Mordant, C., and Benz, W. (2020). AQUA: a collection of H<sub>2</sub>O equations of state for planetary models.
- Sturruel, L. and Lithgow-Bertelloni, C. (2024). Thermodynamics of mantle minerals - III: the role of iron.
- Connolly, J. A. D. (2009). The geodynamic equation of state: What and how.
- Ichikawa, H. and Tsuchiya, T. (2020). Ab initio Thermodynamic Properties of Liquid Iron-Nickel-Light Element Alloys.
- Sakai, T., Dekura, H., and Hirao, N. (2016). Experimental and theoretical thermal equations of state of MgSiO<sub>3</sub> post-perovskite at multi-megabar pressures.
- Fischer, R. A., Campbell, A. J., Shorner, G. A., Lord, O. T., Dera, P., and Prakapenka, V. B. (2011). Equation of state and phase diagram of FeO.
- Musielak, R., Mazevet, S., and Guyot, F. (2019). Physical properties of MgO at deep planetary conditions.
- Melosh, H. J. (2007). A hydrocode equation of state for SiO<sub>2</sub>.
- Falk, S., Tauschwitz, A., and Iosilevsky, I. (2018). The equation of state package FEOS for high energy density matter.
- Stewart, S., Davies, E., Duncan, M., Lock, S., Root, S., Townsend, J., Kraus, R., Caracac, R., and Jacobsen, S. (2020). The shock physics of giant impacts: Key requirements for the equations of state.
- Ichikawa, H. and Tsuchiya, T. (2020). Ab initio Thermodynamic Properties of Liquid Iron-Nickel-Light Element Alloys.
- Kawayama, Y., Morad, G., Nakajima, Y., Hirose, K., Baron, A. Q. R., Kawaguchi, S. I., Tsuchiya, T., Ishikawa, D., Hirao, N., and Ohishi, Y. (2020). Equation of State of Liquid Iron under Extreme Conditions.
- Hakim, B., Rivoldini, A., Van Hoolst, T., Cottenier, S., Jaeken, J., Chust, T., and Steine-Neumann, G. (2018). A new ab initio equation of state of Fe-Fe and its implication on the interior structure and mass-radius relations of rocky super-Earths.
- Fei, Y., Murphy, C., Shibasaki, Y., Shahar, A., and Huang, H. (2016). Thermal equation of state of hcp-iron: Constraint on the density deficit of Earth's solid inner core.
- Dorokopets, P. I., Dymshts, A. M., Litakov, K. D., and Sokolova, T. S. (2017). Thermodynamics and Equations of State of Iron to 350 GPa and 6000 K.
- Guillot, T. (2010). On the radiative equilibrium of irradiated planetary atmospheres.
- Parmentier, V. and Guillot, T. (2014). A non-grey analytical model for irradiated atmospheres. I. Derivation.
- Parmentier, V., Guillot, T., Fortney, J. J., and Marley, M. S. (2015). A non-grey analytical model for irradiated atmospheres. II. Analytical vs. numerical solutions.
- Friedman, R. S., Lustig-Yeager, J., Fortney, J. J., Lupu, R. E., Marley, M. S., and Lodders, K. (2014). Gaseous Mean Opacities for Giant Planet and Ultracool Dwarf Atmospheres over a Range of Metallicities and Temperatures.
- Paxton, B., Cantiello, M., Arras, P., Bildsten, L., Brown, E. F., Dotter, A., Mankovich, C., Montgomery, M. H., Stello, D., Timmes, F. X., and Townsend, R. (2013). Modules for Experiments in Stellar Astrophysics (MESA): Planets, Oscillations, Rotation, and Massive Stars.
- Paxton, B., Smolec, R., Schwab, J., Gaitis, A., Bildsten, L., Dotter, A., Farmer, R., Goldberg, J. A., Jermyn, A. S., Kanbur, S. M., Marchant, P., Thoul, A., Townsend, R. H. D., Wolf, W. M., Zhang, M., and Timmes, F. X. (2019). Modules for Experiments in Stellar Astrophysics (MESA): Pulsating Variable Stars, Rotation, Convective Boundaries, and Energy Conservation.