

Water as a Potential Sculptor of the M Dwarf Radius Valley



0.014

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Background

- Distribution of planetary radii is bimodal "Radius Valley" 1
 - Slope w/ instellation implies atmospheric escape^{2,3}
- Slope different around M v. FGK stars⁴
 - Different formation mechanism? Water worlds?⁵

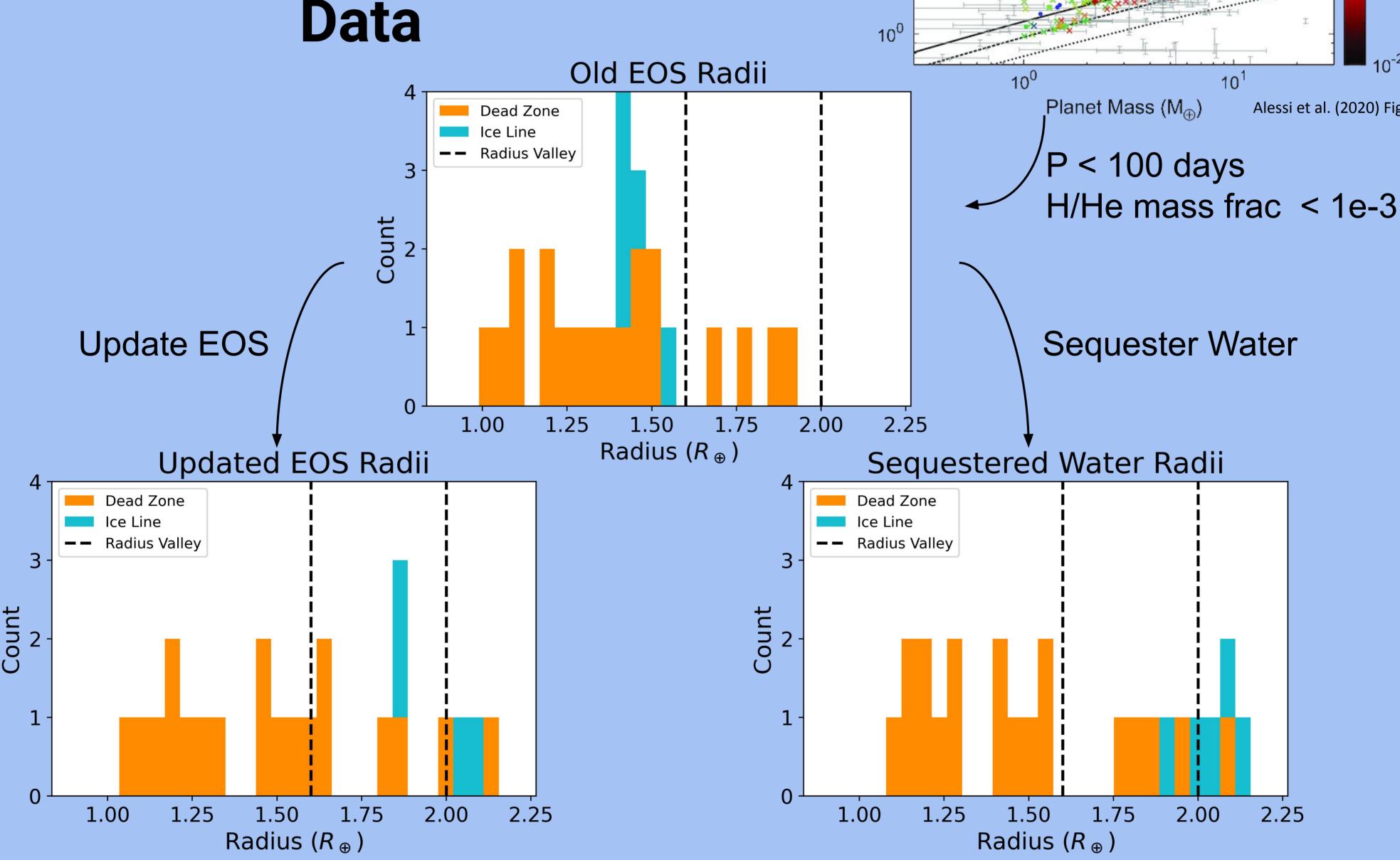
The Project

- McMaster Planet Population Synthesis model⁶
 - Planetesimal accretion in disk around FGK star
 - Planets form in planet traps at dead zone, ice line, and heat transition
 - Disks chemically evolve
- Recalculate planet radii w/ new advances
 - New Equations of State (EOS) for water⁷, iron⁸, silicates⁹, opacities¹⁰
 - Sequestration of water into planetary interior¹¹
- Can radius valley be replicated solely w/ water?

Impact erosion (W19) 0.012 Measured slope (this work) 0.010 0.008 2.0 0.006 0.0041.0 0.002 0.000 0.5^{1} Cloutier & Menou (2020) Fig. 114 Insolation $[F_{\oplus}]$ I.L. Composition Core After Evaporation ($t_{evap} = 1 \text{ Gyr}$) --- D.Z. Composition Core 10⁰ Pure Iron Core Stripped H.J. Semi-Majo 10⁰ Planet Mass (M_⊕) Alessi et al. (2020) Fig. 7⁶

■ Gas-poor formation (LR18) ■ • Photoevaporation (LR18)

Core-powered mass loss (GS19)



Updated EOS and sequestration separate water and dry worlds

Water does NOT solely replicate radius valley, but could contribute

Results

Some water worlds in the valley

Future Work

- Update McMaster Planet Population Synthesis model for M stars
- Increase sample size by running more simulations
- Take advantage of disk chemistry tracking to improve mantle model

Acknowledgements & References

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2.00

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