TACO outputs		
Parameter	Source module	Description
KIC	pipeline.py	Kepler Input Catalogue identification number
raw_data		Name of the raw input file
git_rev_hash		If git revision is True or not
mean	filter.R	Mean value of the fluxes of in the lightcurve
var		Variance of the fluxes in the lightcurve
start_date		Time of the first point of the lightcurve
end_date		Time of the last point of the lightcurve
fill_factor		Filling factor of the lightcurve
nuNyq	numax_estimate.R	Nyquest frequency
numax0_flag		Flags unreliable values of $\nu_{max,0}$
numax_var		Estimate of ν_{max} derived using the variance of the lightcurve ($\nu_{max} = e^A \cdot \mathrm{Var}^B; A \approx 13.2, B \approx -0.75$)
numax_CWTMexHat		Estimate of ν_{max} derived using a Mexican Hat continuous wavelet transform (CWT).
numax_Morlet		Estimate of ν_{max} derived using a Morlet CWT.
numax0		Selected estimate of ν_{max} based on $\nu_{max,\mathrm{Var}}$, $\nu_{max,\mathrm{CWTMexHat}}$, or $\nu_{max,\mathrm{Morlet}}$.
Hmax	background_fit.py	Total power at ν_{max}
Bmax		Background power at ν_{max}
HBR		Ratio H_{max}/B_{max}
Pn		Level of the white noise
A1		Characteristic amplitude of the first background component
b1		Characteristic frequency of the first background component
A2		Characteristic amplitude of the second background component
b2		Characteristic frequency of the second background component.
A3		Characteristic amplitude of the third background component
b3		Characteristic frequency of the third background component
Pg		Amplitude of the Gaussian envelope encompassing the solar-like oscillations
numax		Central frequency of the Gaussian envelope
sigmaEnv		Width of the Gaussian envelope
lnprob		Median value of $\ln(prob) = \ln(prior) + \ln(L)$; with L being the likelihood function used to evaluate the fit of the background.
npeaks	peak_find.R	Number of peaks found in the PDS with AIC > 2
Deltanu	peak_bag_mode_id02.R	Fitted value of $\Delta\nu$
DeltaNu_sd		Standard deviation of $\Delta\nu$

dNu02		Small frequency separation $\delta\nu_{02}$ between the radial and the quadrupole modes
eps_p		Fitted phase term
eps_sd		Standard deviation of ϵ_p
alpha		Fitted curvature term
alpha_sd		Standard deviation of α
Central_DeltaNu		Fitted central $\Delta\nu$ from the three radial modes closest to ν_{max}
Central_DeltaNu_sd		Standard deviation of Central_DeltaNu
Central_eps_p		Fitted central phase term from the three radial modes closest to ν_{max}
Central_eps_p_sd		Standard deviation of Central_eps_p
Central_alpha		Fitted central curvature term from the three radial mode peaks closest to ν_{max}
Central_alpha_sd		Standard deviation of Central_alpha
gamma0		Global radial mode width determined using the weighted mean of the line widths of the three radial modes closest to ν_{max}. The weights are the mode amplitudes.
modelDFlag		1: npeaks $<$ 3 or none; 2: $\nu_{max} < 10\mu$Hz; 3: number of radial modes $<$ 3; 0: no problems
visibility_ratio	peak_bag_period_spacing.py	Ratio of the total area of the dipole modes over the total area of the radial based on the fitted values
DeltaPi1		Period spacing
coupling		Coupling factor q
eps_g		Phase offset of gravity modes
DeltaPi1_sig		Period spacing error