

The physics behind chemistry, and the Periodic Table

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1. Introduction

Theoretical chemistry could be seen as a bridge from the real physics of the physicists to the real chemistry of the experimental chemists. We hence expect that any measurable property of any chemical object could, in principle, be calculated to arbitrary accuracy, if the relevant physical laws are known. Moreover, as put by Sidgwick,¹ 'The chemist must resist the temptation to make his own physics; if he does, it will be bad physics—just as the physicist has sometimes been tempted to make his own chemistry, and then it was bad chemistry.'

The first step was the Schrödinger equation since the 1920:ies. Another major step was the inclusion of relativistic effects, using the Dirac equation or approximations to it, basically since the 1970:ies (for some early reviews, see ref.²⁻⁵). These effects are of essential chemical importance, and often explain the differences of the 6th Period elements (Cs-Rn) from their 5th Period counterparts (Rb-Xe).

A classical example on relativistic effects in chemistry is the nobility, trivalency,⁶ and yellow color of gold.^{3,7,8} Another one is the crystal structure of mercury⁹ and probably also the low melting-point of mercury.² No explicit R/NR (relativistic versus non-relativistic) studies on liquid mercury seem to exist yet. A third, new example is the lead-acid battery. It has just been calculated that, of its 2.1 Volts per cell, over 1.7 Volts come from relativistic effects.¹⁰ Without relativity, cars would not start. Numerous further examples exist.

Typical ways of including relativity are the use of pseudopotentials or transformed, approximate Dirac Hamiltonians. Both can be calibrated against full-Dirac benchmarks. For some recent summaries on the methodology we quote Schwerdtfeger,^{11,12} Hess,¹³ Hirao and Ishikawa,¹⁴ Dylla and Faegri,¹⁵ Grant,¹⁶ Reiher and Wolf,¹⁷ or Barysz and Ishikawa.¹⁸

The next physical level brings in the quantum electrodynamical (QED) effects. For light-element problems, such as the hydrogen-atom Lamb shift, precise properties of the hydrogen molecules, or the spectra of the lithium atom, all these effects are already clearly seen, because the accuracy of both theory and experiments is very high. Likewise, the QED effects are conspicuous for highly ionized, heavy, few-electron atoms, such as hydrogenlike gold. For neutral or nearly neutral systems, beyond Li or so, only one order-of-magnitude improvement of the computational accuracy, mainly the treatment of electron correlation with adequate basis sets, is estimated to separate the QED effects from being observed in head-on comparisons of theory and experiment. Examples on such cases are the vibrations of the water molecule,¹⁹ or the ionization potential of the gold atom.^{20–22}

And, that may have been "the last train from physics to chemistry" concerning the fundamental interparticle interactions because, of the possible further terms, parity non-conservation (PNC)^{23,24} splittings are estimated to lie over ten powers of ten further down.²⁵ Like magnetic resonance parameters, the PNC effects can be directly observed. Apart from being a physical challenge, both these effects give new information on molecules, but they are expected to be far too small to influence molecular structures or normal chemical energetics.

2. The levels of theory

2.1. The Dirac-Coulomb-Breit (DCB) Hamiltonian

We use the atomic units (a.u., $e = m_e = \hbar = 4\pi\epsilon_0 = 1$). The Year-2008 standard value of the fine structure constant α is 1/137.035 999 679(94).²⁶ In atomic units the speed of light $c = 1/\alpha$.^{27–29} Please note that, in SI units, c is fixed as 299 792 458 m s⁻¹, but in a.u. it has error limits. The DCB Hamiltonian for electrons in nuclear potential V_n can be written as

$$H = \sum_i h_i + \sum_{i<j} h_{ij}. \quad (1)$$

The one-particle Dirac Hamiltonian

$$h_i = c\boldsymbol{\alpha} \cdot \mathbf{p} + \beta c^2 + V_n, \quad \mathbf{p} = -i\nabla, \quad (2)$$

The two-particle Hamiltonian

$$h_{ij} = h_C + h_B, \quad h_C = 1/r_{ij}. \quad (3)$$

$$h_B = -\frac{1}{2r_{ij}}[\boldsymbol{\alpha}_i \cdot \boldsymbol{\alpha}_j + (\boldsymbol{\alpha}_i \cdot \mathbf{r}_{ij})(\boldsymbol{\alpha}_j \cdot \mathbf{r}_{ij})/r_{ij}^2]. \quad (4)$$

For h_B there are alternative, frequency-dependent forms, see *e.g.* Lindgren.³⁰ In the *Coulomb gauge* used, for a magnetic vector potential \mathbf{A} , one sets $\nabla \cdot \mathbf{A} = 0$. Then the electron-electron interactions can be taken as instantaneous. The first correction, h_B , to the Coulomb interaction h_C in this gauge, physically contains both the interactions between the magnetic moments of the two electrons, and retardation effects. The latter are by some authors already regarded as a QED effect. In correlated calculations (beyond single-Slater-determinant, self-consistent-field ones), electron-like projection operators, P , should be added:

$$h_{ij}^{\text{eff}} = Ph_{ij}P. \quad (5)$$

This is also called the 'No-Virtual-Pair Approximation (NVPA)'. The next term after this H was found by H. Araki³¹ and J. Sucher.³² It corresponds to the exchange of two virtual photons. See also Lindgren et al.³³ This term is clearly visible in the accurate studies on the hydrogen molecule, see below. In Eq. (5), the correlation energy arising from h_B exceeds that arising from h_C beyond $Z = 50$ for He-like systems.³⁴

An example on the level of accuracy that can be reached for the gold atom at DCB CCSD (Coupled cluster singles and doubles) level is given in Table 1.

Table 1: The ionization potential IP and the electron affinity EA of the Au atom from the CCSD calculations of Eliav et al.³⁵ and Landau et al.³⁶ Basis functions up to (spdfghik) were included, and 51 electrons were correlated for Au. The last column gives the calculated, further QED contributions. All contributions in eV.

| Property | Non-rel. | Rel. | Exp. | Exp-Rel | QED |
|----------|----------|-------|------------|---------|--|
| IP | 7.057 | 9.197 | 9.22554(2) | 0.0285 | -0.0255 ²¹ -0.0211 ²² |
| EA | 1.283 | 2.295 | 2.30861(3) | 0.014 | |

We notice that the 'Exp-Rel' and 'QED' terms have a comparable size but, unfortunately, opposite signs. The ratio of the QED to relativistic energies is here $-0.0255/2.14$ or -1.2% , a common result for the ns^1 atoms with $Z \geq 50$.²¹ In that sense the DCB-level relativistic effects were '101 % right'.

Note finally that the relativistic correction to the Au atom IP is $2.14/9.22554$ or 23% of the experimental value. For the experimental EA of the gold atom, the relativistic part is 44% . The Dirac-level relativistic effects are both large and well established.

2.2. The next level: Introducing the QED terms

2.2.1. Qualitative discussion

For the valence energies of the heavier elements, the QED contributions should become discernible in the near future. Because these effects are still unknown for most chemists, a qualitative de-

scription may be helpful. Apart from subsection 2.3., we shall mainly discuss atomic examples. Estimates for molecules can be obtained by adding the monoatomic contributions, as discussed in subsubsection 2.2.2. and subsubsection 2.2.4.

We start by considering the electromagnetic (EM) oscillations of the vacuum. Real oscillations of the EM field can be externally induced by electronic devices, atomic transitions, etc. They also are thermally excited for $h\nu \leq kT$ by thermal, black-body radiation. These are real photons.

However, even at $T = 0$, the zero-point oscillations of the EM field are still there. Very qualitatively, they will shake the point-like Dirac electron, and give it a 'finite size'. This leads to the *vacuum fluctuation* or *self-energy (SE)* contribution. For electrons near a nucleus it is repulsive, because a part of the Coulomb attraction is lost. Parenthetically, if these zero-point oscillations of the EM field are modified, by objects ranging from molecules to macroscopic bodies, this leads to *Casimir forces* between them. A good overview is given by Parsegian.³⁷

Secondly, just as an electric field can polarize a noble-gas atom, by virtual quantum mechanical excitations, the 'empty vacuum' can be electrically polarized by creating virtual electron-positron pairs. This leads to the *vacuum polarization (VP)* contribution. For electrons near a nucleus, it is attractive.

Until recent times there was almost no information on the expected magnitude of the SE and VP terms for the valence electrons of the heavier, neutral or nearly neutral atoms. Thus the question is, could they be chemically relevant? The first estimates were produced by Dzuba et al. for the Cs³⁸ and Fr³⁹ atoms. They related the ns valence electron Lamb shift of an alkali atom to that of an H-like atom with the same Z by using a quantum-defect formula

$$E^{\text{Lamb}} = \frac{\alpha(Z\alpha)^2}{\pi\nu^3} \left(1 - \frac{d\delta}{dn}\right) F(Z\alpha). \quad (6)$$

The α^3 is the expected behavior for a Lamb shift. The Z^4 behavior of a one-electron atom is changed to Z^2 for the valence electron of a many-electron atom, see eq. (1) of Dzuba et al.³⁹ Here δ is the so-called 'quantum defect'. In a many-electron atom the Rydberg levels are fitted

to a $1/(n - \delta)^2$ behavior, instead of the one-electron $1/n^2$ behavior, and $\nu = n - \delta$ is the effective principal quantum number. The expression in the parentheses yields the electron density at nucleus. It was derived by Fermi and Segrè.⁴⁰ The $F(Z\alpha)$ is defined below in Eq. (7). and already effectively incorporates the relativistic effects on the wave function. This hydrogen-like approach to the electron density at the nucleus is described in Kopfermann⁴¹ and goes back to Fermi and Segrè.⁴⁰

$$E_{n\kappa}^{\text{SE}}(Z\alpha) = \frac{Z^4 \alpha^3}{\pi n^3} F_{n\kappa}(Z\alpha), \quad (7)$$

2.2.2. Vacuum polarization

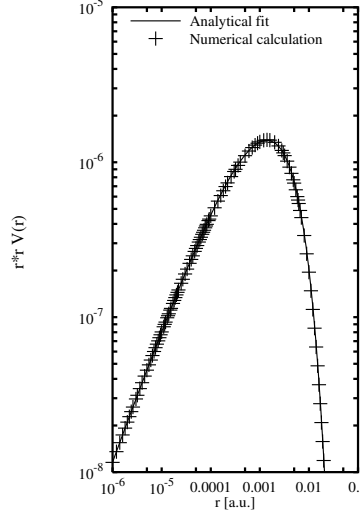
To lowest order, the VP part can be described by the Uehling potential.^{42,43} It is attractive, a local potential, a property of space, and the same for all elements. The analytical expression for a point nucleus is⁴⁴

$$V_n^{\text{eff}}(r) = -\frac{Z}{r}(1 + S(r)) = V_n + V_{\text{Ue}}, \quad (8)$$

$$S(r) = \frac{2\alpha}{3\pi} \int_1^\infty \exp(-2r\chi/\alpha) \left(1 + \frac{1}{2\chi^2}\right) \frac{\sqrt{\chi^2 - 1}}{\chi^2} d\chi \quad (9)$$

The VP effects decay outside ca. 10^{-3} a.u., as seen from Figure 1. Note the factor r^2 from the volume element. The point in the chemical context is, that this term is strongly localized to each nuclear neighborhood.

Figure 1: The enhancing, Uehling potential, V_{Ue} , Eq. (8), multiplied by r^2 from the volume element. The points are given by Eq. (8) - Eq. (9). For the 'fit', see ref.²⁰ (Reproduced from Pyykkö and Zhao.⁴⁵ Copyright IOP.)



A simple way to include the finite-nucleus changes is to replace the $-Z/r$ in Eq. (8) by the finite-nucleus V_n .⁴⁶ The next-order VP terms are the Wichmann-Kroll⁴⁷ and Källen-Sabry⁴⁸ ones.

2.2.3. Self-energy: The benchmarks

The SE part is larger than the VP, and has (for energies) the opposite sign. It can be rigorously treated by first obtaining, for the electrons in question, a complete set of one-particle states at Dirac level, and by then doing the Feynman diagram in Figure 2 (a). This is known as the *Furry picture*. We then have no 'potential' and no 'range' for the SE.

The effective atomic potential for that Dirac problem can, in the simplest case, be taken as a suitably parametrized local model potential.²¹ If it reproduces the Dirac-Fock (DF) (= relativistic Hartree-Fock) valence eigenvalue, it simulates for the QED purpose a DF model. If it reproduces the experimental IP, it simulates a correlated calculation.

More fundamentally, the effective potential for the QED calculations can be obtained by inverting the radial Dirac-Fock equations.⁴⁹ The procedure is as follows: 1) Run first the DF problem

to convergence for the system considered. 2) Then 'invert' the radial DF equations to get an effective local potential $V(r)$ for the occupied state A considered. 3) Then solve the Dirac equations for a complete set of excited states, n , in the same potential, with the same basis of radial spline functions. 4) Finally do the Feynman diagram Figure 2 (a) in Coulomb gauge using the obtained functions, and for instance the 'multiple-commutator method with partial-wave renormalization',⁵⁰

$$\Delta E_{SE}^A = \frac{\alpha}{2\pi i} \sum_n \left[\frac{1 - \vec{\alpha}_1 \cdot \vec{\alpha}_2}{\alpha r_{12}} I_{nA}(r_{12}) \right]_{AnnA} - \delta m_A, \quad (10)$$

as done earlier for the various model potentials by Labzowsky et al.²¹ Here the function

$$I_{nA}(r_{12}) = \int_{-\infty}^{\infty} \frac{d\omega \exp(i|\omega|r_{12})}{E_n(1-i0) - E_A - \omega} \quad (11)$$

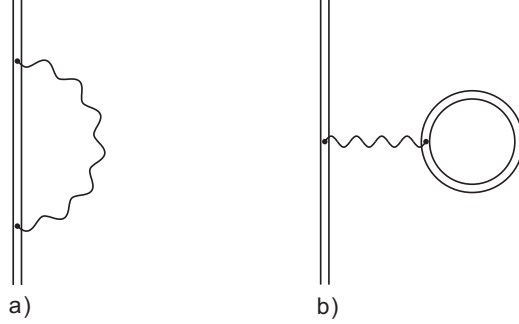
refers to the one-electron Feynman diagram Figure 2 (a) in the Furry picture for state A and intermediate state n , and $-\delta m_A$ arises from renormalization.

Goidenko et al.⁴⁹ found that a 9 % reduction of the electron affinity of the noble gas E118⁵¹ was coming from QED effects, see Table 2. Likewise, the earlier results for the valence electrons of Group 11 and 12 atoms could be confirmed by the inversion method.⁵²

Table 2: Calculated electron affinity of the noble gas E118.

| DCB (av.) | EA/eV | | Reference |
|------------|-----------|-----------|---------------|
| | SE+VP | Total | |
| -0.056(10) | | | ⁵¹ |
| -0.064(2) | 0.0059(5) | -0.058(3) | ⁴⁹ |

Figure 2: The lowest-order Feynman diagrams for self-energy (a) and vacuum polarization (b)). The double solid lines denote the electrons in the atomic potential. The wavy lines are the virtual photons.



In the lowest-order, low- Z formulation of Bethe⁵³ (see⁵⁴) either the self-energy, or the entire Lamb shift, can be expressed in terms of the electron density at the nucleus.

$$E_1^{\text{Lamb}} = \frac{4\alpha^3 Z}{3} \left[-2\ln(\alpha Z) - \ln X + \frac{19}{30} \right] \langle \delta(\mathbf{r}) \rangle. \quad (12)$$

For hydrogen-like atoms $X = 2K_{n0}/(\alpha Z)^2 = 11.77, 16.64, 15.93, 15.64$, and 15.16 for $1s, 2s, 3s, 4s$ and ∞s , respectively, and K_{n0} is the Bethe logarithm (see Labzowsky⁵⁵). For recent reviews on atomic QED calculations, see Beier,⁵⁶ Mohr,⁵⁷ Eides,⁵⁸ Lindgren,^{30,33} or Shabaev et al.⁵⁹ For a benchmark on self-energy screening in two-electron systems, see Indelicato and Mohr.⁶⁰

The proof of the pudding is in the eating. We show two examples on the size of the various contributions for heavy, highly-ionized systems, *viz.* the energies of hydrogen-like Au in Table 3 and of lithium-like uranium in Table 4. Note the agreement between theory and experiment in both cases. We have chosen H-like Au for the availability of all terms. There are calculations for the remaining SESE terms for both H-like and Li-like heavy ions by Yerokhin et al.^{61,62} Concerning the splitting of Li-like U, note the improved nuclear-structure corrections of Kozhedub et al.⁶³ in Table 4.

Table 3: Energy contributions (in eV) for H-like Au.⁵⁶ The 'corrections' are counted from the point-nucleus binding energy. The true electron mass is used everywhere.

| Term | Contribution |
|--|------------------|
| Binding energy E_T (point nucleus) | -93459.89 |
| <i>Corrections:</i> | |
| Finite nuclear size | 49.13 |
| Self energy (order α) | 196.68 |
| VP: Uehling contribution | -41.99 |
| VP: Wichmann-Kroll contribution | 1.79 |
| Total vacuum polarization (order α) | -40.20 |
| SESE (2 nd order SE) (a) (b) (c) | uncalculated |
| VPVP (2 nd order VP) (a) (ladder diagrams) | -0.07 |
| VPVP (b) (Källen-Sabry contribution + h.o.) | -0.05 |
| VPVP (c) (Källen-Sabry contribution) | -0.29 |
| SEVP (a) (b) (c) | 0.42 |
| S(VP)E | 0.05 |
| Radiative recoil (estimate) | 0.00 |
| Reduced mass | 0.26 |
| Relativistic recoil | 0.08 |
| Total recoil | 0.34 |
| Nuclear polarization (<i>bottleneck for accuracy!</i>) | -0.02 |
| <i>Sum of corrections</i> | 205.99 |
| Resulting total binding energy | -93253.90 |
| Total shift (theory) | 205.73 |
| (experimental) | 202(8) |

Table 4: The 2s-2p_{1/2} splitting of Li-like U (from ref.⁶⁴).

| | |
|-------------------------------|-------------|
| Exp. ⁶⁴ | 280.645(15) |
| Calc. ⁶³ 2008 | 280.71(10) |
| Calc. ⁶⁵ 2001 | 280.47(7) |
| Calc. ⁶⁶ 2000 | 280.44(10) |
| Calc. ^a | 280.43(7) |
| Inferred two-loop Lamb-shift. | 0.20 |

^a J. Sapirstein and K. T. Cheng, as quoted by Beiersdorfer et al.⁶⁴

An example on a light atomic system is the lithium atom.

Table 5: Properties of the ^7Li atom, calculated by Yan and Drake⁶⁷ and by Puchalski and Pachucki.⁶⁸

| Quantity | Case | Value |
|---------------------------------|--------------------------|--------------------|
| IP/cm ⁻¹ | Exp. | 43 487.159 40(18) |
| | Calc.(tot) ⁶⁷ | 43 487.172 6 (44)) |
| | Lamb (e-n) | -0.305 45(1) |
| | Lamb (other) | +0.059 478 |
| | Calc.(tot) ⁶⁸ | 43 487.159 0(8) |
| $E(2s-2p_{1/2})/\text{cm}^{-1}$ | Exp. | 14 903.648 130(14) |
| | Calc.(tot) ⁶⁷ | 14 903.648 0(30) |
| | Lamb (e-n) | -0.347 95(1) |
| | Lamb (other) | +0.043 4721 |
| | Calc.(tot) ⁶⁸ | 14 903.648 4(10) |
| EA/cm ⁻¹ | Exp. | 4 984.90(17) |
| | Calc.(tot) ⁶⁹ | 4 984.96(18) |

We conclude that these calculations may be a patchwork, but a patchwork that works. Concerning the convergence, the high- Z approach in Table 3 or Table 4 treats the one-electron relativity to all orders, and can treat the virtual-photon exchange (the Feynman diagrams) to an arbitrary order. For the low- Z approach in Table 5 or Table 9, the relativistic effects are treated starting from the Pauli Hamiltonian, which itself only must be used as a first-order perturbation. In the calculations quoted, also the predominant $(\alpha Z)^4$ terms are, however, included.

2.2.4. Approximate self-energy approaches

How to estimate these effects in molecular calculations? We discuss some existing approximate approaches:

1) The Welton potential. Welton⁷⁰ started from the idea of electromagnetic fluctuations, induced by the zero-point oscillations of the vacuum, and obtained an effective SE potential, related to $\nabla^2 V_n$. Here V_n is the nuclear potential. Using the fundamentally calculated hydrogen-like SE for calibration, one obtains

$$E_{\text{SE}} = \frac{\langle ns | \nabla^2 V_n | ns \rangle_{\text{DF}}}{\langle ns | \nabla^2 V_n | ns \rangle_{\text{hyd}}} E_{\text{SE},ns}^{\text{hyd}}. \quad (13)$$

Indelicato and Desclaux⁷¹ thus included electronic screening by taking the ratio between Dirac-Fock (DF) and hydrogenic matrix elements. This method has notably been used by Blundell, Desclaux, Indelicato and coauthors.^{71–74} For its non-relativistic limit, see Dupont-Roc et al.⁷⁵

2) Low- Z approaches. From the Bethe expressions it is not a long step to treat the relativistic effects at the Breit-Pauli level and, concomitantly, to try to model, either the SE part or the entire electron-nuclear Lamb shift, Eq. (12), by slightly renormalizing its Darwin term, as done by Pyykkö et al.¹⁹

$$h^{\text{Pauli}} = -\frac{\alpha^2}{8}\mathbf{p}^4 - \frac{\alpha^2}{8}\nabla^2 V - \frac{\alpha^2}{4}\boldsymbol{\sigma} \cdot (\nabla V) \times \mathbf{p}, \quad (14)$$

with the mass-velocity, Darwin, and spin-orbit contributions, respectively. For a Coulomb potential, $\nabla^2 V = -4Z\pi\delta(\mathbf{r})$. Results were given for the light elements $Z=1-54$. Because the Darwin term is strictly local, and the V_{SE} strongly local, it is a reasonable approximation for a molecule to sum them over all nuclei. Assuming that the Bethe-type Coulomb-field Lamb-shift values can be used for many-electron atoms, we obtain at each nucleus the ratio

$$E_1^{\text{Lamb}}/E_1^{\text{Darwin}} = \frac{8\alpha}{3\pi}[-2\ln(\alpha Z) - \ln X + \frac{19}{30}]. \quad (15)$$

Alternatively one can use the later QED calculations for one-electron atoms, yielding the ratio of one-electron terms

$$E_1^{\text{Lamb}}/E_1^{\text{Darwin}} = 2\alpha F(\alpha Z)/\pi - \frac{8\alpha}{15\pi} = 4.64564 \cdot 10^{-3} F(\alpha Z) - 1.23884 \cdot 10^{-3}. \quad (16)$$

Here the $F(\alpha Z)$ is related to the SE, or the total Lamb shift, by an expression of type

$$E_1^{\text{SE}} = \alpha^3 Z F(\alpha Z) \langle \delta(\mathbf{r}) \rangle \quad (17)$$

The raw data for the function $F(\alpha Z)$ were obtained from the papers of Mohr and coworkers.⁵⁷ This resulted in the "Eq. (6)" ratios $E_1^{\text{Lamb}}/E_1^{\text{Darwin}}$ in Table II of Pyykkö et al.¹⁹ The ratios decrease from 0.04669 for $Z = 1$ to 0.00906 for $Z = 54$, or loosely from 5% to 1%.

An expression, giving the s -state Lamb shift as a renormalized Darwin term was already given by Bjorken and Drell in 1964 in the form⁷⁶

$$E_1^{\text{Lamb}}/E_1^{\text{Darwin}} = (8\alpha/3\pi) \ln(1/Z\alpha). \quad (18)$$

Finally, combining the two-electron Darwin term with the corresponding two-electron Lamb-shift term we get the ratio

$$h_2^{\text{Lamb}}/h_2^{\text{Darwin}} = -\frac{14\alpha}{3\pi} \ln \alpha = 0.053334. \quad (19)$$

Like the Pauli approximation itself, these equations should be used with non-relativistic wave functions, only.⁷⁷

The derivation above referred to a single atom. For molecules, the strongly local character of the SE permits a summation of these renormalized Darwin terms over nuclei. The first such application were the vibrational levels of water in our own original paper.¹⁹ It was estimated that an improvement of the calculations, at the IC-MRCI/aug-cc-pV6Z level for the valence electrons and lower for the core part, by a further order of magnitude would make the QED contributions to certain vibrational lines of H_2O visible. An example is the (501) "bright state" in Table 6. The experimental accuracy is entirely sufficient for seeing the QED effects.

Table 6: Calculated and observed energies (in cm^{-1}) for the vibrational ($v_1 v_2 v_3$) states of water¹⁹

| State | Calc. | +Lamb | Observed |
|-------|----------|-------|----------|
| (010) | 1598.19 | -0.09 | 1594.75 |
| (100) | 3657.68 | 0.18 | 3657.05 |
| (501) | 19776.00 | 1.01 | 19781.10 |

There are numerous later tests on water, from that of Polyansky et al.⁷⁸ to that of Kahn et al.,⁷⁹ and Császár et al.⁸⁰ Both the accuracy of the BO energies, and the non-adiabatic corrections still present obstacles for seeing the QED corrections. Further molecules where this approach has been tested are NH_3 ,⁸¹ EF_3 ; $\text{E}=\text{B-Ga}$,⁸² H_2S ,^{83,84} OH , FO , HOF and F_2O ,⁸⁵ H_3^+ ,⁸⁶ or CH_2 .⁸⁷ For more general reviews on high-precision molecular calculations, see Tarczay et al.,⁸⁸ Helgaker et al.,⁸⁹ or Lodi and Tennyson.⁹⁰

3) The 'ratio method'. Pyykkö et al.²⁰ noted that the ratio $E_{\text{SE}}/E_{\text{VP}}$ was fairly constant for given Z , as function of n . Thus in the *ratio method* one could multiply the $\langle V_{\text{Ue}} \rangle$ by that ratio to get an estimate for the E_{SE} . For heavy elements, the E_{SE} was evaluated from the $2s$ SE/VP ratio of Johnson and Soff.⁹¹ The total valence-electron Lamb shift became

$$E_L = \langle V_{\text{Ue}} \rangle (E_{\text{SE}} + E_{\text{VP}})/E_{\text{VP}}. \quad (20)$$

4) Effective local potentials. a) A further way would be to simulate the SE contribution by local potentials. The first such potential was the modified electron-proton potential for a hydrogen atom, introduced by Pais⁹² to account for the hydrogen Lamb shift:

$$V_{\text{Pais}}(r) = -\frac{e^2}{r} [1 - 2e^{-kr}], \quad (21)$$

with k^{-1} of the order of the classical electron radius $r_0 = e^2/mc^2$. This reproduced the observed upwards $2s$ shift of about 0.03 cm^{-1} . Note the transition from $-1/r$ at large r to $+1/r$ at small r .

b) Another potential was proposed by Fricke,⁹³ who folded the nuclear potential with a Gaus-

sian function $\exp(-\kappa r^2)$ with $\kappa = 1/\langle(\delta r)^2\rangle$ and

$$\langle(\delta r)^2\rangle = \frac{2\alpha^3}{\pi} \log(1/Z\alpha). \quad (22)$$

c) If one only wants to reproduce the energy, the 'width' of the chosen SE potential is arbitrary and could range from nuclear dimensions to much more diffuse values. Tulub et al.^{94,95} introduced a very compact repulsive excess potential, of the same shape as that of a homogeneously charged spherical nucleus with radius R_n , for energy levels or for magnetic dipole (M1) hyperfine splittings, respectively. V_0 is a fitting constant.

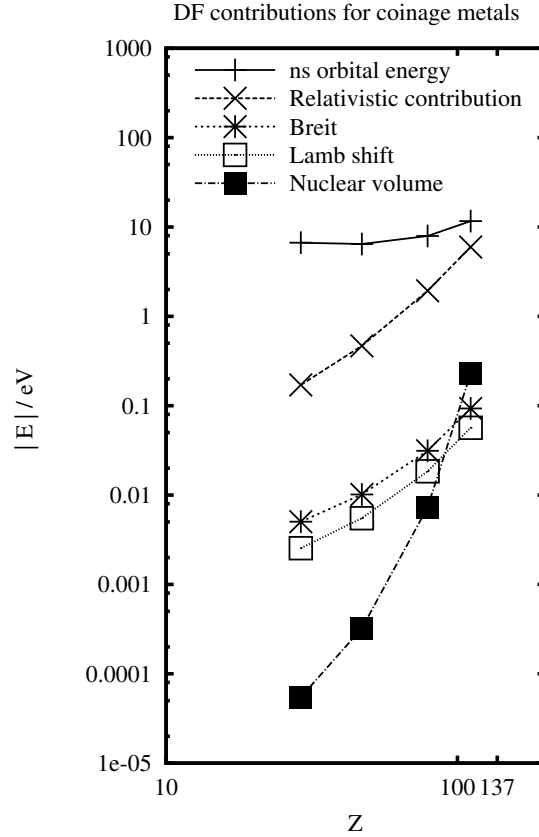
$$\begin{aligned} V_{\text{QED}} &= V_0 \left[1 - \left(\frac{r}{R_n} \right)^2 \right], r < R_n, \\ &= 0, r > R_n. \end{aligned} \quad (23)$$

d) In the *A-model* of Pyykkö et al.,²⁰ an extended size was given, not to the electron as in the Welton model, but to the nucleus. An inflated mass number, A , reproduced the total H-like $2s$ Lamb shift with

$$A = a \exp(-bZ), \quad a = 2.36 \cdot 10^5, \quad b = 0.0555. \quad (24)$$

For SE only, $a = 2.09 \cdot 10^5$, $b = 0.05001$. The corresponding radius of a homogeneously charged nucleus is $r_A = 2.2677 \cdot 10^{-5} A^{1/3}$. A comparison of the Dirac-level, Breit, QED, and finite-nuclear-volume effects for the coinage metals Cu, Ag, Au, and Rg is shown in Figure 3.

Figure 3: The total Dirac-Koopmans level ionization potentials and their relativistic, Breit, QED and nuclear-volume contributions²⁰ for the atoms Cu - Rg. Copyright APS.



e) Eides et al.⁵⁸ give a non-relativistic momentum-space potential whose Fourier-transform to r -space yields the term

$$V_{SE}(r) = \frac{8\alpha^4 Z (2\pi)^{1/2}}{3r^3}. \quad (25)$$

Note that this potential is strongly singular near origin, and it has not been applied to atomic or molecular calculations. As seen from Figure 4, at moderate distances it cuts through many of the alternative potentials.

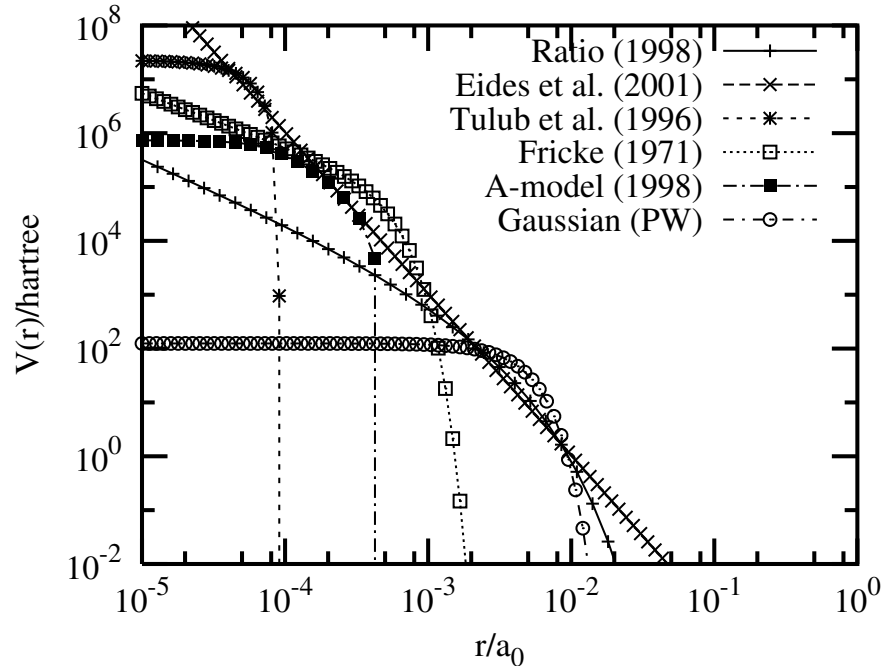
f) As mentioned above, the SE energy shifts could be simulated by a potential of any width, from a δ function to atomic dimensions by choosing a suitable pre-coefficient. If we add other physical properties, also the 'width' or shape of the effective SE potential V_{SE} could be semiem-

pirically fitted. Pyykkö and Zhao⁴⁵ used the H-like 2s-state Lamb shift and magnetic dipole (M1) hyperfine data of Boucard and Indelicato⁹⁶ or Yerokhin et al.⁹⁷ to determine the B and β parameters of a two-parameter Gaussian effective potential for all atoms

$$V_{SE}(r) = B \exp(-\beta r^2). \quad (26)$$

A quadratic polynomial fit, done for both B and β at $29 \leq Z \leq 83$, was still meaningful in the superheavy domain. For a comparison of the different SE potentials for the higher ($> 1s$) s-electrons of Cs, see Figure 4.

Figure 4: Effective local SE potentials for the ns ($n > 1$) electrons of Cs. Reproduced from Pyykkö and Zhao.⁴⁵ Copyright IOP.



No molecular applications of this potential have yet been reported.

g) Flambaum and Ginges⁹⁸ derived an effective SE potential from first principles.⁹⁹ It has been tested on a number of atoms by Thierfelder and Schwerdtfeger.²² It contains an SE part with both

an electric and magnetic, momentum-dependent potential,

$$V_{\text{QED}} = V_{\text{VP}} + V_{\text{SE}}, \quad V_{\text{SE}}(p) = \left[\frac{g(-p^2)}{2m} \vec{\gamma} \vec{p} + f(-p^2) - 1 \right] \phi(p), \quad (27)$$

whence it cannot be directly compared with the alternative purely-electric SE potentials. The numerical agreement with other calculations is good.

h) Finally we note that hydrogenic estimates, scaled with an effective Z_{eff} , may be useful for inner shells but are not reliable for the valence shells, whose effective field is far from Coulombic. The E_{SE} values, produced by the earlier versions of the Grasp atomic code¹⁰⁰ are of such scaled hydrogenic type.

2.2.5. Summary of numerical results for atoms

Some valence-shell atomic results from various QED approaches are collected for the full Lamb shift to Table 7 and for the SE effects on magnetic-dipole hyperfine interactions in Table 8. Further results on the latter are reported for E119 and E120⁺ by Dinh et al.¹⁰¹ Calibration results for M1 hyperfine splittings and g factors of 1s to 3s and 2p states of hydrogenlike ions with $Z = 1-12$ are given by Yerokhin and Jentschura.¹⁰² For further results on the individual SE contributions, see.⁴⁵ The QED corrections to the $p_{1/2}$ states of Li - Cs are discussed by Sapirstein and Cheng.¹⁰³ The SE term for E111 (Rg) is reported by Indelicato et al.¹⁰⁴

For the IP of Be, the estimates by Chung et al.,¹⁰⁵ and earlier work, using a Z_{eff} , estimated from the relativistic energy shift, and a Bethe-type $\alpha^3 Z_{\text{eff}}^4$ formula, give a 2s Lamb shift of 0.126 meV, rather larger than the later results in Table 7. Estimates for other Be-like systems are also given by them.

Discussing the trends, as seen from Figure 3 for Group 11, the valence-ns-electron Lamb shifts follow a similar trend as the Dirac-level relativistic effects. It is roughly Z^2 , where Z is the full nuclear charge. The sign is a destabilization of the valence ns levels. For the outermost, $np_{1/2}$ valence electron of the Group-13 elements B–Tl, the sign is negative. For the heaviest member,

E113, the sign becomes positive again.

The QED contributions for the discrete 2s-3s transition of a ${}^9\text{Be}$ atom are given to order α^4 by Stanke et al.¹⁰⁶ For the 2s-ns transitions of the isoelectronic B^+ , see Bubin et al.¹⁰⁷ The contribution to the electron affinity of Li is an increase of $0.007(0) \text{ cm}^{-1}$.⁶⁹

Pachucki and Sapirstein¹⁰⁸ calculated the dipole polarizability of helium. Of the total 1.383 191(2) a.u., the QED contribution was 0.000 030 a.u. Łach et al.¹⁰⁹ calculated the full α^3 term, and obtained 0.000 030 666(3) a.u. Their total value is 1.383 760 79(23) a.u.

Table 7: The total Lamb shift $E_L = E_{VP} + E_{SE}$ (in meV) for the valence shells of various Groups, G, and Periods (for their numbers, see Figure 5) of the Periodic Table, for atomic systems. The 'approach', A, is either self-consistent field ('DF' = Dirac-Fock) or includes some estimate of electron correlation ('EC'). Positive numbers indicate net destabilization. For quick reference a shortened name of the first author is given.

| G | A | Period | | | | | | | Year | Ref. | Eq. |
|-----------------|-----------------|--------|--------|-------|-------------------|--------------------|------------------|-------------------|------|-----------------------|-----------------------|
| | | 2 | 3 | 4 | 5 | 6 | 7 | 8 | | | |
| 1 | | Li | Na | K | Rb | Cs | Fr | E119 | | | |
| | DF | 0.040 | 0.28 | 0.47 | 1.03 | 1.92 | 4.75 | 17.5 ^a | 1998 | Pyy ²⁰ | Eq. (20) |
| | DF | 0.037 | 0.27 | 0.46 | 1.10 | 1.85 | 4.60 | 17.3 | 1999 | Lab ²¹ | Eq. (10) ^b |
| | DF ^c | | | | 1.28 | 2.23 | 3.57 | 10.32 | 2005 | Eli ¹¹⁰ | Eq. (13) |
| | DF | | | | | 2.0 | 4.5 | 8.3 | 2005 | Fla ^{98,111} | Eq. (27) |
| | DF ^c | 0.033 | 0.288 | 0.511 | 1.298 | 2.05 | 4.68 | 8.96 | 2010 | Thi ²² | Eq. (27) |
| | EC | | | | | 3.5 | 9.5 | | 1983 | Dzu ^{38,39} | Eq. (6) |
| | EC | 0.051 | 0.43 | 0.81 | 1.99 | 3.30 | 7.58 | | 1999 | Lab ²¹ | Eq. (10) ^b |
| | EC ^f | | | | | 2.9 | 6.5 | | 2002 | Sap ¹¹² | Figure 2 |
| | EC | | | | | 2.7 | 5.8 | 10.6 | 2005 | Fla ⁹⁸ | Eq. (27) |
| 2 | | Be | Mg | Ca | Sr | Ba | Ra | E120 | | | |
| | DF ^c | 0.087 | 0.419 | 0.65 | 1.53 ^a | 2.35 | 5.19 | 9.49 | 2010 | Thi ²² | Eq. (27) |
| | EC | 0.0722 | | | | | | | 2007 | Pac ¹¹³ | |
| | EC | | | | | 4.6 ⁺ | 9.5 ⁺ | 14.9 ⁺ | 2008 | Din ¹¹¹ | Eq. (27) |
| 11 | | | | Cu | Ag | Au | Rg | | | | |
| | DF | | | 2.54 | 5.51 | 18.42 | 56.56 | | 1998 | Pyy ²⁰ | Eq. (20) |
| | DF | | | 2.42 | 5.40 | 17.5 | 54.7 | | 1999 | Lab ²¹ | Eq. (10) ^b |
| | DF | | | | 5.50 | | 56.3 | | 2009 | Goi ⁵² | Eq. (10) ^d |
| | DF ^c | | | | 6.52 | | 62.6 | | 2009 | Goi ⁵² | Eq. (10) ^d |
| | DF ^c | | | 3.05 | 6.48 | 21.1 | 52.9 | | 2010 | Thi ²² | Eq. (27) |
| | EC | | | 4.61 | 9.32 | 25.5 | | | 1999 | Lab ²¹ | Eq. (10) ^b |
| 12 | | | | Zn | Cd | Hg | Cn | | | | |
| | DF | | | | 6.17 | | 65.3 | | 2009 | Goi ⁵² | Eq. (10) ^d |
| | DF ^c | | | | 6.60 | | 69.1 | | 2009 | Goi ⁵² | Eq. (10) ^d |
| | DF ^c | | | 3.08 | 6.43 | 20.5 | 52.6 | | 2010 | Thi ²² | Eq. (27) |
| | DF | | | | | 26.1 ⁺ | | | 1999 | Lab ²¹ | Eq. (10) ^b |
| | EC | | | | | 32.5 ⁺ | | | 1999 | Lab ²¹ | Eq. (10) ^b |
| 13 | | B | Al | Ga | In | Tl | E113 | | | | |
| | DF ^c | -0.23 | -0.545 | -1.85 | -3.07 | -5.00 | 32.4 | | 2010 | Thi ²² | Eq. (27) |
| | DF | | | | | 37.1 ⁺⁺ | | | 1999 | Lab ²¹ | Eq. (10) ^b |
| | EC | | | | | 42.0 ⁺⁺ | | | 1999 | Lab ²¹ | Eq. (10) ^b |
| 18 ^e | | Ne | Ar | Kr | Xe | Rn | E118 | | | | |
| | DF ^c | -1.012 | -1.16 | -2.05 | -2.47 | -3.33 | -0.62 | | 2010 | Thi ²² | Eq. (27) |

^a A printing error in original paper is corrected.

^b Full SE calculation in model potentials, simulating DF or IP(exp).

^c Calculated as total-energy differences.

^d Full SE calculation in inverted DF potential.

^e In Group 18, Period 1, the DF^c value for He using Eq. (27) is 0.172 meV Thi.²²

^f The largest value in various Dirac-Slater potentials chosen.

⁺ Monocation. ⁺⁺ Dication.

Table 8: SE-induced changes of magnetic M1 hyperfine integrals for the valence orbitals of ns¹ metals.

| Atom | $\delta/\%$ | |
|------------------|-------------|---------------------|
| | PW(DF) | Ref. ¹¹⁴ |
| Rb | -0.53 | -0.44 |
| Cs | -0.87 | -0.75 |
| Fr | -1.77 | -1.45 |
| Cu | -0.36 | |
| Ag | -0.78 | |
| Au | -1.58 | |
| Hg ⁺ | -1.44 | |
| Tl ²⁺ | -1.38 | |

A further application area for the QED terms are the inner-shell electronic transitions of neutral or nearly neutral atoms. An example for the superheavy elements E112 to E118 was published by Gaston et al.¹¹⁵

We conclude by mentioning the approach by Lindgren et al.^{30,116–119} to attack the combined electronic many-body and QED problem from the beginning.

2.3. Accurate calculations on diatomics

An extraordinary example on the accuracy of present quantum chemistry are the calculations on H₂ isotopologues,¹²⁰ see Table 9. A slight deviation between theory and experiment for D₂ was resolved by a later experiment by Liu et al.¹²¹ The later work includes a measurement¹²² and a calculation¹²³ on HD, see the same table. A finite-nuclear-volume contribution to $D_0(\text{D}_2)$ of -0.0002 cm⁻¹ is included. For H₂ this correction is estimated to lie below 0.0001 cm⁻¹.¹²⁰ Some other species treated are H₂⁺ isotopologues,^{124–127} ³He⁴He⁺,¹²⁸ and He₂.¹²⁹

The background of the H₂ work is well described by Piszczatowski et al.¹²⁰ The QED corrections were probably first evaluated by Ladik.¹³⁰

Table 9: Dissociation energies for H₂ and D₂ (in cm⁻¹) from Piszczatowski et al.¹²⁰ The terms are classified by powers of the fine-structure constant, α .

| Order | Term | H ₂ | D ₂ | HD |
|---------------------|--------------------|----------------|----------------------------|-----------------------------|
| α^0 | Born-Oppenheimer | 36112.5927(1) | 36746.1623(1) | |
| | Adiabatic | 5.7711(1) | 2.7725(1) | |
| | Non-adiabatic | 0.4339(2) | 0.1563(2) | |
| | Total α^0 | 36118.7978(2) | 36749.0910(2) | |
| α^2 | Mass-velocity | 4.4273(2) | 4.5125(2) | |
| | One-el. Darwin | -4.9082(2) | -4.9873(2) | |
| | Two-el. Darwin | -0.5932(1) | -0.5993(1) | |
| | Breit | 0.5422(1) | 0.5465(1) | |
| | Total α^2 | -0.5319(3) | -0.5276(3) | |
| $\alpha^2 m_e/m_p$ | Estimate | 0.0000(4) | 0.0000(2) | |
| α^3 | One-el. Lamb shift | -0.2241(1) | -0.2278(1) | |
| | Two-el. Lamb shift | 0.0166(1) | 0.0167(1) | |
| | Araki-Sucher | 0.0127(1) | 0.0128(1) | |
| | Total α^3 | -0.1948(2) | -0.1983(2) | |
| $\alpha^3 m_e/m_p$ | Estimate | 0.0000(2) | 0.0000(1) | |
| α^4 | One-loop term | -0.0016(8) | -0.0016(8) | |
| Total theory | | 36118.0695(10) | 36748.3633(9) ^a | 36405.7828(10) ^b |
| Exp. ¹³¹ | | 36118.062(10) | 36748.343(10) | |
| Exp. ¹³² | | 36118.0696(4) | | |
| Exp. ¹²¹ | | | 36748.36286(68) | |
| Exp. ¹²² | | | | 36405.78366(36) |

^a Includes -0.0002 cm⁻¹ from the finite deuteron size. ^b Pachucki and Komasa.¹²³

Calculations for the individual IR lines of hydrogen molecules using NBO-level ('Non-Born-Oppenheimer') methods are reported for HD by Stanke et al.,¹³³ and for D₂ and T₂ by Bubin et al.¹³⁴ HeH⁺ was treated in ref.¹³⁵ Relativistic, α^2 corrections were included.

For the finite nuclear mass corrections to electric and magnetic interactions in diatomic molecules, see Pachucki.¹³⁶

2.4. Further small terms and curiosities

2.4.1. The finite nuclear size

The nuclear charge distribution can be taken as a Fermi one¹³⁷ with the parameters

$$\rho(r) = \rho_0 / [1 + \exp((r - c)(4 \ln 3)/t)], \quad (28)$$

where ρ_0 is a normalization constant to obtain a charge Z and the surface thickness $t = 2.3$ fm (Fermi) for $Z > 45$. Using

$$A = 0.00733Z^2 + 1.3Z + 63.6. \quad (29)$$

for the atomic mass, the RMS nuclear radius c (in fm) is extrapolated in the program from known values of c as a function of $A(Z)$ for large Z .

For recent reviews on the finite nuclear charge distributions, and their inclusion in Quantum Chemistry, see Andrae.^{138,139} Ultimately one needs an explicit, quantum-mechanical description of both the nucleus and the electrons. In refs.^{140,141} one treated the exchange of virtual photons between a ²⁰⁹Bi nucleus 'valence proton' and a single valence electron.

2.4.2. Nuclear electric polarizability

Because the nucleus itself has an electric polarizability, α_n , an electron at distance r will enjoy a further attraction

$$V = -\alpha_n/2r^4. \quad (30)$$

This term is actually thought to limit the accuracy of the calculation on H-like Au in Table 3. A novel application of this polarizability would be a van der Waals-bound dineutron, the ultimate noble-gas molecule.¹⁴²

2.4.3. 'Nuclear relativity'

Could the relativistic dynamics of the nuclei become relevant? The question is of principal interest in the NBO calculations (see Chapter 2.3.) where the electronic and nuclear motions are handled on equal footing. For a spin- $\frac{1}{2}$ nucleus with an anomalous magnetic moment κ ,

$$h_{BP} = -\frac{p^4}{8m^3c^2} + \frac{1+2\kappa}{8m^2c^2} \nabla^2 V + \frac{1+2\kappa}{4m^2c^2} \frac{V'}{r} \mathbf{L} \cdot \boldsymbol{\sigma}. \quad (31)$$

For a proton, $m = m_p$, the anomalous magnetic moment

$$\kappa = 1.792\,84734. \quad (32)$$

The expression is adapted from^{143,144} for a spin-zero, infinite-mass potential source.

In their first NBO study, Adamowicz' group¹⁴⁵ neglected this correction. In later work (for a summary, see their own review¹⁴⁶ in the present issue) they included it. In the practice this does not matter. Consider as an order-of-magnitude estimate vibrations of frequency ν . Then the critical parameter is $h\nu/mc^2$. It therefore seems unlikely that 'nuclear relativity' could be seen in many molecular spectra.

For the case of H_2 , with the lowest reduced mass of $m = m_p/2 = 918m_e$, we can make the following rough estimate for the relativistic lowering of the various vibrational states, n , using the mass-velocity Hamiltonian h_m , only, and the harmonic estimate

$\langle T \rangle = \langle V \rangle = E_n/2$, where the vibrational energy, $E_n = (n + \frac{1}{2})h\nu$,

$$\langle h_m \rangle = \langle n | -\frac{p^4}{8m^3c^2} | n \rangle = -\frac{1}{2mc^2} \langle T^2 \rangle \approx -\frac{1}{2mc^2} \langle T \rangle^2 = -\frac{1}{8mc^2} [(n + \frac{1}{2})h\nu]^2. \quad (33)$$

The corresponding relativistic change of the transition energy

$$\Delta_r(E_{n+1} - E_n) = -\frac{1}{4mc^2} (n+1)(h\nu)^2. \quad (34)$$

For the lowest, $n = 0$, vibrational transition of H_2 this gives $-1.28 \cdot 10^{-6} \text{ cm}^{-1}$, over two orders of magnitude below the estimated inaccuracies of the theoretical¹²⁰ and experimental (see Stanke et al.¹⁴⁷) values of 4161.1661(5) and 4161.1660(3) cm^{-1} , respectively.

If a similar harmonic argument were stretched to the dissociation limit, the highest vibrational levels of H_2 would descend by

$$\langle h_m \rangle = \langle -\frac{p^4}{8m^3c^2} \rangle = -\frac{1}{2mc^2} \langle T^2 \rangle \approx -\frac{1}{8mc^2} (D_0)^2 = -1.96 \cdot 10^{-10} \text{ au} = -4.31 \cdot 10^{-5} \text{ cm}^{-1}. \quad (35)$$

This contribution is less than two orders of magnitude beyond the precision of $1 \cdot 10^{-3} \text{ cm}^{-1}$ in Table 9.

It should be added that, as done by Piszczatowski et al.¹²⁰ (p. 3045), before the small contributions here, one should consider the electron-nucleus Breit interaction and the fact that the accurate "nonadiabatic" wave function depends on the reduced rather than the true mass of the electron. These 'nonadiabatic' contributions to the wave function give an $\alpha^2(m_e/m_p)$ contribution to the conventional mass-velocity and Darwin energy.

2.4.4. Magnetic and hyperfine effects

At one-electron Dirac level one includes these effects via the Hamiltonian

$$h = c\boldsymbol{\alpha} \cdot \mathbf{A} \quad (36)$$

where \mathbf{A} is the vector potential of the magnetic, external and/or nuclear fields.

Beyond Dirac level, the most conspicuous QED effects are those on the g -factor of the electron, see Table 10. The leading Schwinger¹⁴⁸ term a_{10} exceeds one part per thousand. The a_{20} term is known as the Karplus-Kroll¹⁴⁹ one. We give in the table the latest available standard value for g . The g calculation by Gabrielse et al.¹⁵⁰ could be inverted to yield $\alpha^{-1} = 137.035\,999\,710(96)$.

Another example is the magnetic dipole hyperfine splitting of the hydrogen-atom ground state,

see Table 10. Here the QED terms a_{10} and a_{21} actually override the leading Dirac term a_{22} .

Table 10: Relativistic and QED corrections to 1s-state hyperfine splitting of H-like atoms^a.

| Term | Origin | a_{10} | a_{20} | a_{21} | a_{22} | Next |
|-----------|--------------|------------------|---------------------|------------------------|----------|-------------------|
| Dirac eq. | Dynamics | - | - | - | 3/2 | $O(\alpha^4 Z^4)$ |
| QED | El. g-factor | $\frac{1}{2\pi}$ | $-.328478966/\pi^2$ | | | |
| QED | Vac. pol. | - | - | $\frac{3}{4}$ | - | $O(\alpha^3 Z^2)$ |
| QED | Self en. | - | - | $\ln 2 - \frac{13}{4}$ | - | $O(\alpha^3 Z^2)$ |
| Total/ppm | | 1161.4 | -1.8 | -96.2 | 79.9 | |

$$\Delta E = \Delta E^{\text{NR}} [1 + a_{10}(\alpha) + a_{20}(\alpha^2) + a_{21}(\alpha^2 Z) + a_{22}(\alpha^2 Z^2) + \dots]$$

^a See Sapirstein¹⁵¹ and Sunnergren et al.¹⁵²

Electron $g = 2 \cdot 1.001\,159\,652\,181\,11(74)$.

Fine str. const. $\alpha = 1/137.035\,999\,679(94)$.²⁶

For the terms arising in the relativistic theory of ESR and NMR variables, see the recent summaries by Aucar et al.,¹⁵³ Autschbach,^{154–158} Kutzelnigg and Liu,^{159,160} or Vaara et al.¹⁶¹ For all terms at the Breit-Pauli level, see Manninen et al.¹⁶²

Returning to QED effects, for valence ns -state hyperfine interactions near Au, Hg or Tl, the SE-induced decrease is estimated to be ca. -1.5%, per atom,⁴⁵ see Table 8. This is comparable with other small effects, such as many examples on solvation.

2.4.5. Retardation at large distances

At large R , retardation will change the R^{-6} dispersion forces to R^{-7} ones. This is of direct importance in a case like He₂, and it is often known as the Casimir effect.¹⁶³ For detailed studies, see Przybytek et al.¹²⁹ They conclude that 95% of these effects can be included by using the Breit and Araki-Sucher terms.

3. The Periodic Table

Chemistry is about the chemical elements.¹⁶⁴ These chemical elements can be ordered in a Periodic Table. The currently experimentally known 118 elements snugly fit to the PT in Figure 5. One

case where a chemical property has sizable QED contributions, is the electron affinity of the last element, the noble gas E118, see Table 2. Another, potentially observable property are the K- and L-shell ionization potentials of E112 to E118.¹¹⁵

For the 172 first elements, the PT in Figure 6 was recently proposed on basis of Dirac-Fock calculations on both atoms and ions.

Figure 5: A Periodic Table for $Z = 1 - 118$. Reproduced by permission of the PCCP Owner Societies from Pyykkö.¹⁶⁵ The IUPAC PT¹⁶⁶ coincides with this table, but so far only includes the elements, up to roentgenium (E111).

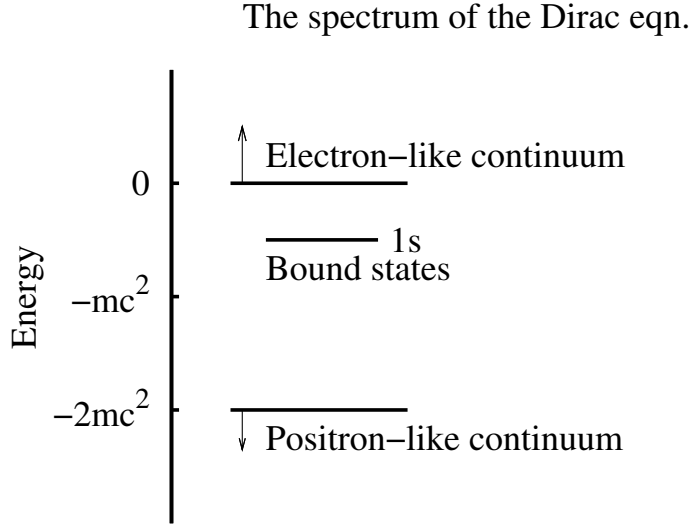
| Periodic Table 1-118 | | | | | | | | | | | | | | | | | | 18 Orbitals | | | | | | |
|----------------------|----------|----------|---|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|----------|-------------|----------|----------|---------|----------|----|------|
| 1 | 1 H | | 2 | | | | | | | | | | | | | 13 | 14 | 15 | 16 | 17 | 2 He | | 1s | |
| 2 | 3 Li | | 4 Be | | | | | | | | | | | | | | 5 | 6 | 7 | 8 | 9 | 10 Ne | | 2s2p |
| 3 | 11 Na | | 12 Mg | | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 | 11 | 12 | 13 Al | 14 Si | 15 P | 16 S | 17 Cl | 18 Ar | 3s3p | | | |
| 4 | 19 K | 20 Ca | 21 Sc | 22 Ti | 23 V | 24 Cr | 25 Mn | 26 Fe | 27 Co | 28 Ni | 29 Cu | 30 Zn | 31 Ga | 32 Ge | 33 As | 34 Se | 35 Br | 36 Kr | 4s3d4p | | | | | |
| 5 | 37 Rb | 38 Sr | 39 Y | 40 Zr | 41 Nb | 42 Mo | 43 Tc | 44 Ru | 45 Rh | 46 Pd | 47 Ag | 48 Cd | 49 In | 50 Sn | 51 Sb | 52 Te | 53 I | 54 Xe | 5s4d5p | | | | | |
| 6 | 55 Cs | 56 Ba | 57-71 La Ce Pr Nd Pm Sm Eu Gd Tb Dy Ho Er Tm Yb Lu | 72 Hf | 73 Ta | 74 W | 75 Re | 76 Os | 77 Ir | 78 Pt | 79 Au | 80 Hg | 81 Tl | 82 Pb | 83 Bi | 84 Po | 85 At | 86 Rn | 6s5d6p | | | | | |
| 7 | 87 Fr | 88 Ra | 89-103 Ac Th Pa U Np Pu Am Cm Bk Cf Es Fm Md No Lr | 104 Rf | 105 Db | 106 Sg | 107 Bh | 108 Hs | 109 Mt | 110 Ds | 111 Rg | 112 Cn | 113 | 114 | 115 | 116 | 117 | 118 | 7s6d7p | | | | | |
| | | | | | | | | | | | | | | | | | | | | | | | | |
| 6 | 57 La | | 58 Ce | 59 Pr | 60 Nd | 61 Pm | 62 Sm | 63 Eu | 64 Gd | 65 Tb | 66 Dy | 67 Ho | 68 Er | 69 Tm | 70 Yb | 71 Lu | 4f | | | | | | | |
| 7 | 89 Ac | | 90 Th | 91 Pa | 92 U | 93 Np | 94 Pu | 95 Am | 96 Cm | 97 Bk | 98 Cf | 99 Es | 100 Fm | 101 Md | 102 No | 103 Lr | 5f | | | | | | | |

Figure 6: A suggested Periodic Table for $Z = 1 - 172$. Reproduced by permission of the PCCP Owner Societies from Pyykkö.¹⁶⁵

| Period | Periodic Table 1-172 | | | | | | | | | | | | | | | | | 18 | Orbitals |
|--------|----------------------|----------|----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|----------|----------|----------|----------|
| 1 | 1 H | 2 | | | | | | | | | | | 13 | 14 | 15 | 16 | 17 | 2 He | 1s |
| 2 | 3 Li | 4 Be | | | | | | | | | | | 5 B | 6 C | 7 N | 8 O | 9 F | 10 Ne | 2s2p |
| 3 | 11 Na | 12 Mg | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 | 11 | 12 | 13 Al | 14 Si | 15 P | 16 S | 17 Cl | 18 Ar | 3s3p |
| 4 | 19 K | 20 Ca | 21 Sc | 22 Ti | 23 V | 24 Cr | 25 Mn | 26 Fe | 27 Co | 28 Ni | 29 Cu | 30 Zn | 31 Ga | 32 Ge | 33 As | 34 Se | 35 Br | 36 Kr | 4s3d4p |
| 5 | 37 Rb | 38 Sr | 39 Y | 40 Zr | 41 Nb | 42 Mo | 43 Tc | 44 Ru | 45 Rh | 46 Pd | 47 Ag | 48 Cd | 49 In | 50 Sn | 51 Sb | 52 Te | 53 I | 54 Xe | 5s4d5p |
| 6 | 55 Cs | 56 Ba | 57-71 | 72 Hf | 73 Ta | 74 W | 75 Re | 76 Os | 77 Ir | 78 Pt | 79 Au | 80 Hg | 81 Tl | 82 Pb | 83 Bi | 84 Po | 85 At | 86 Rn | 6s5d6p |
| 7 | 87 Fr | 88 Ra | 89-103 | 104 Rf | 105 Db | 106 Sg | 107 Bh | 108 Hs | 109 Mt | 110 Ds | 111 Rg | 112 Cn | 113 | 114 | 115 | 116 | 117 | 118 | 7s6d7p |
| 8 | 119 | 120 | 121- | 156 | 157 | 158 | 159 | 160 | 161 | 162 | 163 | 164 | 139 | 140 | 169 | 170 | 171 | 172 | 8s7d8p |
| 9 | 165 | 166 | | | | | | | | | | | 167 | 168 | | | | | 9s9p |
| 6 | 57 La | 58 Ce | 59 Pr | 60 Nd | 61 Pm | 62 Sm | 63 Eu | 64 Gd | 65 Tb | 66 Dy | 67 Ho | 68 Er | 69 Tm | 70 Yb | 71 Lu | 4f | | | |
| 7 | 89 Ac | 90 Th | 91 Pa | 92 U | 93 Np | 94 Pu | 95 Am | 96 Cm | 97 Bk | 98 Cf | 99 Es | 100 Fm | 101 Md | 102 No | 103 Lr | 5f | | | |
| 8 | 141 | 142 | 143 | 144 | 145 | 146 | 147 | 148 | 149 | 150 | 151 | 152 | 153 | 154 | 155 | 6f | | | |
| 8 | 121 | 122 | 123 | 124 | 125 | 126 | 127 | 128 | 129 | 130 | 131 | 132 | 133 | 134 | 135 | 136 | 137 | 138 | 5g |

One reason to discuss the Periodic Table in the present context, are the limits imposed by the spectrum of the Dirac equation in a nuclear (or atomic) field, see Figure 7.

Figure 7: The schematic spectrum of a Dirac electron in an atomic field.



P. Pyykkö, University of Helsinki, 2011.

There actually are three special Z values to consider, near 118, 137, and 172. Already Gordon¹⁶⁷ noticed that a unique solution of the Dirac-Coulomb problem (for a point-like nucleus) exists up to $\alpha Z = \sqrt{3}/2$, or $Z \approx 118.7$. Above that, there is another, irregular solution that should be avoided.^{168,169} Beyond $\alpha Z = 1$, or $Z \approx 137$, for the electron total angular momentum $j = \frac{1}{2}$, the dE/dZ would become infinite and the energy E imaginary.¹⁷⁰ Note that the energy in Figure 7 then only reached $-mc^2$. For a finite nucleus a normalizable solution always exists.¹⁷¹ With realistic nuclear dimensions, one can go to about $Z = 172$, where the $1s$ eigenvalue would dive to the lower continuum at $-2mc^2$ (for references, see¹⁶⁵). No detailed studies on the actual physical implications for stationary, supercritical systems appear to exist. They may or may not be serious. For the situation in the late 1970:ies, see Reinhardt and Greiner¹⁷² or Rafelski et al.¹⁷³ Anyway, it is reasonable to terminate Figure 6 at $Z = 172$. If the overcritical situation is reached during an atomic collision, a vacancy in the resulting $1s$ state would fly out as a real positron.

As emphasized by Wang and Schwarz,¹⁷⁴ the periodicity is driven by the noble-gas-like closed-shell structures. The filling-order for the 118 first elements is shown in Figure 8.

Figure 8: The schematic Aufbau principle for the 118 first elements.

Aufbau Principle, $Z = 1-118$

| | | | | | |
|--|-----------------|-----------------|-----------------|-----------------|--------|
| | | | | | Period |
| | $\overline{7s}$ | $\overline{7p}$ | $\overline{6d}$ | $\overline{5f}$ | 7 |
| | $\overline{6s}$ | $\overline{6p}$ | $\overline{5d}$ | $\overline{4f}$ | 6 |
| | $\overline{5s}$ | $\overline{5p}$ | $\overline{4d}$ | | 5 |
| | $\overline{4s}$ | $\overline{4p}$ | $\overline{3d}$ | | 4 |
| | $\overline{3s}$ | $\overline{3p}$ | | | 3 |
| | $\overline{2s}$ | $\overline{2p}$ | | | 2 |
| | $\overline{1s}$ | | | | 1 |

P. Pyykkö, University of Helsinki, 2010.

The next thing to notice is that the first shell of every quantum number l (1s, 2p, 3d, 4f, 5g) is anomalously compact, not having any radial nodes (for details and references, see ref.^{165,175} This makes the second-period elements anomalous, the 2s and 2p shells having similar sizes, despite of different energies. The following point is the possibility of partial-screening effects. An example is that selenium is only slightly larger than sulfur, because the 3d¹⁰ shell is filled before it.¹⁷⁶ Another example is the lanthanide contraction which partially explains the large 6s electron binding energy of Au or Hg.^{176,177} The other partial explanation is relativity, which stabilizes the s and p shells, and destabilizes the d and f shells, both valence-shell effects roughly increasing as Z^2 down a column, and having a local 'gold maximum' in Group 11³ along a given Period. Indeed, when passing from Period 5 to Period 6, the main new factor is relativistic effects.^{2,178} As an example, the only 'normal' coinage metal is silver. Copper is anomalous in having a very compact, nodeless 3d shell. Gold is anomalous due to its large relativistic effects. These mechanisms suffice to $Z=118$, and beyond.

Beyond $Z = 118$, the next two elements E119 and E120 have $8s^1$ and $8s^2$ electron configurations. Beyond them, the 8p, 7d, 6f, and 5g shells all have a chance to be occupied in a single atom or atomic ion, for earlier literature see Pyykkö.¹⁶⁵ The placement of the 5g elements in the new Periodic Table in Figure 6 was fixed by considering ions. For instance, E125(VI) was found

to have a $5g^1$ electron configuration, placing E125 in Group 7, and the nominal 5g series at $Z = 121-138$. It should, however, be emphasized that considerable overlap may occur between filling the 5g, 8p, 6f, and 7d shells. The broad, general order of atomic levels for the 118 first elements in Figure 8 is followed by

$$8s < 5g \leq 8p_{1/2} < 6f < 7d < 9s < 9p_{1/2} < 8p_{3/2}, \quad (37)$$

as discussed using Dirac-Fock calculations on atoms and ions,¹⁶⁵ and already found in the Dirac-Slater atomic work by Fricke et al.¹³⁷

Very few molecular calculations exist yet in this superheavy domain. An early piece of insight was the quasirelativistic multiple-scattering calculation on $[(E125)F_6]$ by Makhyoun¹⁷⁹ finding, indeed, that it was a $5g^1$ system.

Finally we note that low-lying atomic orbitals which are empty in the atomic single-configuration ground state, can participate in chemical bonding. Examples are 1) the 8s of E118, 2) the 2p of Be, 3p of Mg or 4p of Zn, 3) the (n-1)d of Ca, Sr and Ba, or 4) the 5f of Th. In this sense these four cases could be called *pre-s*, *pre-p*, *pre-d*, and *pre-f* elements, respectively.

4. Conclusion

These are all the terms of which news have come to Helsinki. The importance of relativistic (Dirac) effects in heavy-element chemistry is no longer new, but it is useful both to occasionally remind the broad chemical audience about them, and to check the soundness of the methods used.

The next physical level, quantum electrodynamics, has a double significance. On one hand it is about two orders of magnitude below the Dirac-level relativistic effects, and, being small, thus indirectly verifies the soundness of the latter. On the other hand, quantum chemical methods are becoming increasingly accurate, and it is therefore expected that even these QED terms will soon be needed for fully understanding the chemistry of the heavier elements. For the lightest elements, up to Li or Be, they already have been clearly visible for a long time. Some extraordinarily accurate

work on the H₂ isotopologues has just been reported. Isolated examples on QED effects on the potentially observable properties of the superheavy elements are starting to appear. The relativistic and QED effects together determine many of the chemical trends in, and possibly the prescribed upper limit of the Periodic Table.

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Biography

Pekka Pyykkö was born in Hinnerjoki, Finland in 1941, and received his education in the nearby city of Turku with a Ph. D. in 1967. His two latest employers were Åbo Akademi University in 1974-84, and the University of Helsinki in 1984-2009. Since November 2009 he continues research in Helsinki as Professor Emeritus. He now has about 300 papers. He led in 1993-98 the program 'Relativistic Effects in Heavy-Element Chemistry and Physics (REHE)' of the European Science Foundation (ESF), and in 2006-08 the Finnish Centre of Excellence in Computational Molecular Science (CMS). In addition to his own research, he currently chairs two Academies, and one Editorial Board.

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