unyt: Handle, manipulate, and convert data with units in Python

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Summary

Software that processes data with physical units or that models real world quantities with physical units must have some way of managing units. This might be as simple as the convention that all floating point numbers are understood to be in the same physical unit system (for example, the SI MKS units system). While simple approaches like this do work in practice, they also are fraught with possible error, both by programmers modifying the code who unintentionally misinterpret the units, and by users of the software who must take care to supply data in the correct units or who need to infer the units of data returned by the software.

The unyt library is designed both to aid quick calculations at an interactive python prompt and to be tightly integrated into a larger Python application or library. The top-level unyt namespace ships with a large number of predefined units and physical constants to aid setting up quick calculations without needing to look up unit data or the value of a physical constant. Using the unyt library as an interactive calculation aid only requires knowledge of basic Python syntax and awareness of a few of the methods of the unyt_array class - for example, the unyt_array.to() method to convert data to a different unit. As the complexity of the usage increases, unyt provides a number of optional features to aid these cases, including custom unit registries containing both predefined physical units as well as user-defined units, built-in output to disk via the pickle protocol and to HDF5 files using the h5py library (Collette 2013), and round-trip converters for unit objects defined by other popular Python unit libraries.

Physical units in the unyt class are defined in terms of the dimensions of the unit, a string representation, and a floating point scaling to the MKS unit system. Rather than implementing algebra for unit expressions, we rely on the SymPy symbolic algebra library (Meurer et al. 2017) to handle symbolic algebraic manipulation. The unyt.Unit object can represent arbitrary units formed out of base dimensions in the SI unit system: time, length, mass, temperature,

luminance, and electric current. We currently treat units such as mol with the seventh SI base dimension, amount of substance, as dimensionless, although we are open to changing this based on feedback from users. In addition, unyt supports forming quantities defined in other unit systems - in particular CGS Gaussian units common in astrophysics as well as geometrized "natural" units common in relativistic calculations. In addition, unyt ships with a number of other useful predefined unit systems based, including imperial units, Planck units, a unit system for calculations in the solar system, and a "galactic" unit system based on the solar mass, kiloparsecs, and Myr, a convention common in galactic astronomy.

In addition to the unyt.Unit class, unyt also provides a two subclasses of the NumPy (Oliphant 2006) ndarray (Walt, Colbert, and Varoquaux 2011), unyt.unyt_array and unyt.unyt_quantity to represent arrays and scalars with units attached, respectively. The unyt library also provides a unyt.UnitRegistry class to allow custom systems of units, for example to track the internal unit system used in a simulation. These subclasses are tightly integrated with the NumPy ufunc system, which ensures that algebraic calculations that include data with units automatically check to make sure the units are consistent, and allow automatic converting of the final answer of a calculation into a convenient unit.

We direct users interested in usage examples and a guide for integrating unyt into an exiting Python installation to the unyt documentation at hosted at http://unyt.readthedocs.io/en/latest/.

Comparison with Pint and astropy.units

The scientific Python ecosystem has a long history of efforts to develop a library to handle unit conversions and enforce unit consistency. For a relatively recent review of these efforts, see (Bekolay 2013). While we won't exhaustively cover extant Python libraries for handling units in this paper, we will focus on Pint (Grecco 2018) and astropy.units (The Astropy Collaboration et al. 2018), which both provide a robust implementation of an array container with units attached and are commonly used in research software projects. At time of writing a GitHub search for import astropy.units returns approximately 1,500 results and a search for import pint returns approximately 1,500 results.

While unyt provides functionality that overlaps with astropy.units and Pint, there are important differences which we elaborate on below. In addition, it's worth noting that all three codebases had origins at roughly the same time period. Pint initially began development in 2012 according to the git repository logs, and astropy.units began development 2012 and was released as part of astropy 0.2 in 2013, although the initial implementation was adapted from the pynbody library (Pontzen et al. 2013), which started its units implementation

in 2010 according to the git repository logs. In the case of unyt, it originated via the dimensionful library (Stark 2012) in 2012. Later, dimensionful was elaborated on and improved to become vt.units, the unit system for the yt library (Turk et al. 2011) at a yt developer workshop in 2013 and was subsequently released as part of yt 3.0 in 2014. One of the design goals for the yt unit system was the ability to dynamically define "code" units (e.g. units internal to data loaded by yt) as well as units that depend on details of the dataset - in particular cosmological comoving units and the "little h" factor (Croton 2013). For cosmology simulations in particular, this can be tricky because one might want to compare data from multiple outputs in a time series, with each output having a different mapping from internal units to physical units. This despite the fact that each output in the time series represents the same physical system and common workflows involve combining data from multiple outputs with different custom units. This requirement to manage complex custom units and interoperate between custom unit systems drove the yt community to independently develop a custom unit system solution. We have decided to repackage and improve yt.units in the form of unyt to both make it easier to work on and improve the unit system and encourage use of the unit system for scientific python users who do not want to install a heavy-weight dependency like yt.

Below we present a table comparing unyt with astropy.units and Pint. Estimates for lines of code in the library were generated using the cloc tool (Danial 2018); blank and comment lines are excluded from the estimate. Test coverage was estimated using the coveralls output for Pint and astropy.units and using the codecov.io output for unyt.

Library	unyt	astropy.units	Pint
Lines of code	5128	10163	8908
Lines of code excluding tests	3195	5504	4499
Test Coverage	99.91%	93.63%	77.44%

We offer lines of code as a very rough estimate for the "hackability" of the code-base. In general, smaller codebases with higher test coverage have fewer defects (Lipow 1982; Koru, Zhang, and Liu 2007; Gopinath, Jensen, and Groce 2014). This comparison is somewhat unfair in favor of unyt in that astropy.units only depends on NumPy and Pint has no dependencies, while unyt depends on both SymPy and NumPy. Much of the reduction in the size of the unyt library can be attributed to offloading the handling of algebra to SymPy rather than needing to implement the algebra of unit symbols directly in unyt. For potential users who are wary of adding SymPy as a dependency, that might argue in favor of using Pint in favor of unyt.

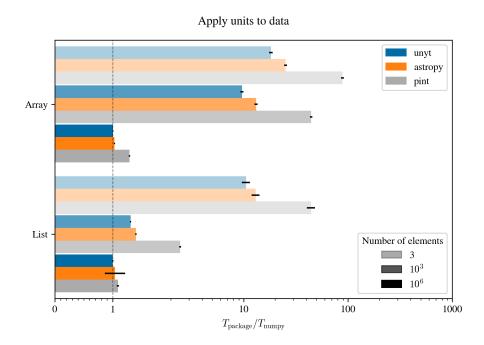


Figure 1: A benchmark comparing the time to apply units to lists and NumPy ndarray instances. Each test is shown for three different sizes of input data, including inputs with size 3, 1,000, and 1,000,000. The black lines at the top of the bars indicate the sample standard deviation. The $T_{\rm numpy}$ time is calculated by benchmarking the time to perform np.asarray(data) where data is either a list of an ndarray.

astropy.units

The astropy units subpackage provides a PrefixUnit class, a Quantity class that represents both scalar and array data with attached units, and a large number of predefined unit symbols. The preferred way to create Quantity instances is via multiplication with a PrefixUnit instance. Similar to unyt, the Quantity class is implemented via a subclass of the NumPy ndarray class. Indeed, in many ways the everyday usage patterns of astropy units and unyt are similar, although unyt is not quite a drop-in replacement for astropy units as there are some API differences. The main functional difference between astropy units and unyt is that astropy units is a subpackage of the larger astropy package. This means that depending on astropy units requires depending on a large collection of astronomically focused software, including a substantial amount of compiled C code. This presents a barrier to usage for potential users of astropy units who are not astronomers or do not need the observational astronomy capabilities provided by astropy.

Pint

The Pint package provides a different API for accessing units compared with unyt and astropy.units. Rather than making units immediately importable from the Pint namespace, instead Pint requires users to instantiate a UnitRegistry instance (unrelated to the unyt.UnitRegistry class), which in turn has Unit instances as attributes. Just like with unyt and astropy.units, creating a Quantity instance requires multiplying an array or scalar by a Unit instance. Exposing the UnitRegistry directly to all users like this does force users of the library to think about which system of units they are working with, which may be beneficial in some cases, however it also means that users have a bit of extra cognitive overhead they need to deal with every time they use Pint.

In addition, the Quantity class provided by Pint is not a subclass of NumPy's ndarray. Instead, it is a wrapper around an internal ndarray buffer. This simplifies the implementation of Pint by avoiding the somewhat arcane process for creating an ndarray subclass, although the Pint Quantity class must also be careful to emulate the full NumPy ndarray API so that it can be a drop-in replacement for ndarray.

Finally, in comparing the output of our benchmarks of Pint, astropy.units, and unyt, we found that in-place operations making use of a NumPy ufunc will unexpectedly strip units in Pint. For example, if a and b are Pint Quantity instances, np.add(a, b, out=out)) will operate on a and b as if neither have units attached. Interestingly, without the out keyword, Pint does get the correct answer, so it's possible that this is a bug in Pint, and we have reported it as such upstream (see https://github.com/hgrecco/pint/issues/644).

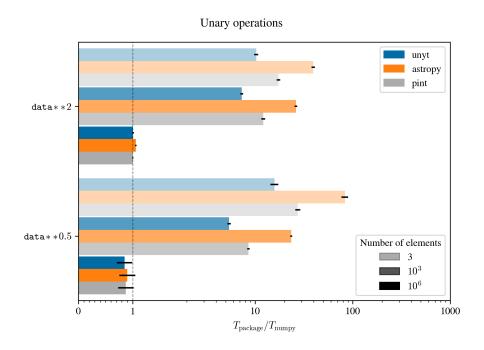


Figure 2: A benchmark comparing the time to square an array and to take the square root of an array. See Figure 1 for a detailed explanation of the plot style.

Binary operations, different units

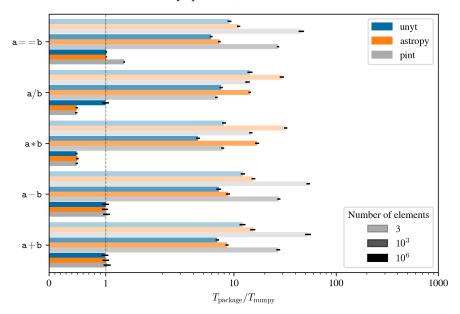


Figure 3: A benchmark comparing the time to perform various binary arithmetic operations on input operans that have different but dimensionally compatible units. See Figure 1 for a detailed explanation of the plot style.

Performance Comparison

Checking units will always add some overhead over using hard-coded unit conversion factors. Thus a library that is entrusted with checking units in an application should incur the minimum possible overhead to avoid triggering performance regressions after integrating unit checking into an application. Optimally, a unit library will add zero overhead regardless of the size of the array. In practice that is not the case for any of the three libraries under consideration, and there is a minimum array size above which the overhead of doing a mathematical operation exceeds the overhead of checking units. It is thus worth benchmarking unit libraries in a fair manner, comparing with the same operation implemented using plain NumPy.

Here we present such a benchmark. We made use of the perf (Stinner 2018) Python benchmarking tool, which not only provides facilities for establishing the statistical significance of a benchmark run, but also can tune a linux system to turn off operating system and hardware features like CPU throttling that might introduce variance in a benchmark. We made use of a Dell Latitude E7270 laptop equipped with an Intel i5-6300U CPU clocked at 2.4 Ghz. The testing environment was based on Python 3.6.3 and had NumPy 1.14.2, sympy 1.1.1, fastcache 1.0.2, Astropy 3.0.1, and Pint 0.8.1 installed. fastcache (Brady 2017) is an optional dependency of SymPy that provides an optimized LRU cache implemented in C that can substantially speed up SymPy. The system was instrumented using perf system tune to turn off CPU features that might interfere with stable benchmarks. We did not make any boot-time Linux kernel parameter changes.

For each of the benchmarks below, we show the ratio of the time to perform an operation with one of unyt, Pint, and astopy.units, $T_{\rm package}$, to the time it takes for NumPy to perform the equivalent operation, $T_{\rm numpy}$. For example, for the comparison of the performance of np.add(a, b) where a and b have different units with the same dimension, the corresponding benchmark to generate $T_{\rm numpy}$ would use the code np.add(a, c*b) where a and b would be ndarray instances and c would be the floating point conversion factor between the units of a and b. Much of the time in $T_{\rm package}$ relative to $T_{\rm numpy}$ is spent in the respective packages calculating the appropriate conversion factor c. Thus the comparisons below depict very directly the overhead for using a unit library over an equivalent operation that uses hard-coded unit-conversion factors.

Applying units to data

In Figure 1 we plot the overhead for applying units to data, showing both Python lists and NumPy ndarray instances as the input to apply data to. Since all three libraries eventually convert input data to a NumPy ndarray, the comparison with array inputs more explicitly shows the overhead for *just* applying units to data. When applying units to a list, all three libraries as well as NumPy need to

Binary operations, same units

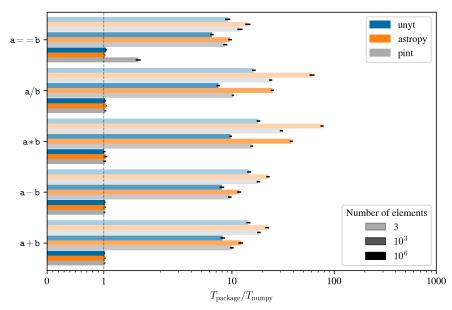


Figure 4: A benchmark comparing the time to perform various binary arithmetic operations on input operands that have the same units. See Figure 1 for a detailed explanation of the plot style.

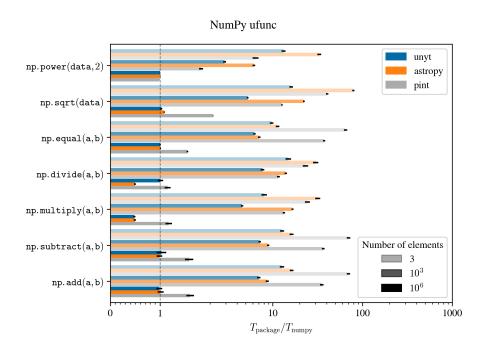


Figure 5: A benchmark comparing the overhead for computing various NumPy ufunc operations. The operands of all binary ufuncs have the same units. See Figure 1 for a detailed explanation of the plot style.

first copy the contents of the list into a NumPy array or a subclass of ndarray. This explains why the overhead is systematically lower when starting with a list.

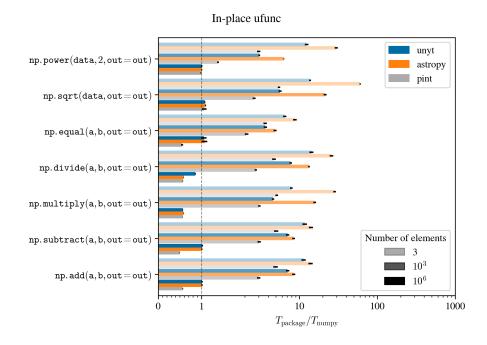


Figure 6: The same as Figure 5, but with in-place ufunc operations. See Figure 1 for a detailed explanation of the plot style.

In all cases, unyt either is fastest by a statistically significant margin, or ties with astropy. Even for large input arrays, Pint still has statistically significant overhead, while both unyt and astropy.units have negligible overhead once the input array size reaches 10^6 elements.

Unary arithmetical operations

Expressions involving powers of data with units, including integer and fractional powers, are very common in the physical sciences. It is therefore very important for a library that handles units to be able to track this case in a performant way. In Figure 2 we present a benchmark comparing Pint, unyt, and astropy.units for the squaring and square root operation. In all cases, unyt has the lowest overhead, with Pint coming in second, and astropy.units trailing. Note that the y-axis is plotted on a log scale, so astropy is as much as 4 times slower than unyt for these operations.

Binary arithmetical operations

Binary operations form the core of arithmetic. It is vital for a library that handles unit manipulation to both transparently convert units when necessary and to ensure that expressions involving quantities with units are dimensionally consistent. In Figure 3 and 4 we present benchmarks for binary arithmetical expressions, both with input data that has the same units and with input data with different units but the same dimensions. In most cases, unyt has less overhead than both astropy and Pint, although there are a few anomalies that are worth explaining in more detail. For comparison operations, Pint exhibits a slowdown even on large input arrays. This is not present for other binary operations, so it's possible that this overhead might be eliminated with a code change in Pint. For multiplication on large arrays, all three libraries have measured overhead of ~ 0.5 that of the "equivalent" Numpy operation. That is because all three libraries produce a results with units given by the product of the input units, that is, there is no need to multiply the result of the multiplication operation by an additional constant, while a pure NumPy implementation would need to multiply the result by a constant to ensure both operands of the operation are in the same units. Finally, for division, both Pint and astropy.units exhibit the same behavior as for multiplication, and for similar reasons: the result of the division operation is output with units given by the ratio of the input units. On the other hand, unyt will automatically cancel the dimensionally compatible units in the ratio and return a result with dimensionless units.

NumPy ufunc performance

Lastly, In Figures 4 and 5, we present benchmarks of NumPy ufunc operations. A NumPy ufunc is a fast C implementation of a basic mathematical operation. This includes arithmetical operators as well as trigonometric and special functions. By using a ufunc directly, one bypasses the Python object protocol and short-circuits directly to the low-level NumPy math kernels. We show both directly using the NumPy ufunc operators (Figure 4) as well as using the same operators with a pre-allocated output array to benchmark in-place operations.

As for the other benchmarks, unyt tends to have the lowest amount of overhead, although there are some significant exceptions. For np.power, Pint has the lowest overhead, except for very large input arrays, where the overhead for all three libraries is negligible. On the other hand, for np.sqrt, np.equal, np.add, and np.subtract, Pint still has statistically significant overhead for large input arrays. Finally, for the in-place ufunc comparison, Pint has the lowest overhead for all operations. However, as discussed above, this is because of a bug in Pint which causes the library to ignore units when calling a ufunc with the out keyword set.

Conclusions

In this paper we present the unyt library, giving background on the reasons for its existence and some historical context for its origin. We also present a set of benchmarks for common arithmetic operations, comparing the performance of unyt with Pint and astropy.units. In general, we find that unyt either outperforms or matches the performance astropy.units and Pint, depending on the operation and size of the input data. We also demonstrate that the unyt library constitutes a smaller codebase with higher test coverage than both Pint and astropy.units.

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References

Bekolay, Trevor. 2013. "A Comprehensive Look at Representing Physical Quantities in Python." https://www.youtube.com/watch?v=N-edLdxiM40.

Brady, Peter. 2017. "Fastcache." *GitHub Repository*. https://github.com/pbrady/fastcache; GitHub.

Collette, Andrew. 2013. Python and Hdf5. O'Reilly.

Croton, D. J. 2013. "Damn You, Little h! (Or, Real-World Applications of the Hubble Constant Using Observed and Simulated Data)" 30 (October): e052. https://doi.org/10.1017/pasa.2013.31.

Danial, Al. 2018. "Cloc." GitHub Repository. https://github.com/AlDanial/cloc; GitHub.

Gopinath, Rahul, Carlos Jensen, and Alex Groce. 2014. "Code Coverage for Suite Evaluation by Developers." In *Proceedings of the 36th International Conference on Software Engineering*, 72–82. ICSE 2014. New York, NY, USA: ACM. https://doi.org/10.1145/2568225.2568278.

Grecco, Hernan E. 2018. "Pint." GitHub Repository. https://github.com/hgrecco/pint; GitHub.

Koru, A. G., D. Zhang, and H. Liu. 2007. "Modeling the Effect of Size on Defect Proneness for Open-Source Software." In *Predictor Models in Software Engineering, 2007. PROMISE'07: ICSE Workshops 2007. International Workshop on*, 10–10. https://doi.org/10.1109/PROMISE.2007.9.

Lipow, M. 1982. "Number of Faults Per Line of Code." *IEEE Trans. Softw. Eng.* 8 (4). Piscataway, NJ, USA: IEEE Press: 437–39. https://doi.org/10.1109/TSE.1982.235579.

Meurer, Aaron, Christopher P. Smith, Mateusz Paprocki, Ondřej Čertík, Sergey B. Kirpichev, Matthew Rocklin, AMiT Kumar, et al. 2017. "SymPy: Symbolic Computing in Python." *PeerJ Computer Science* 3 (January): e103. https://doi.org/10.7717/peerj-cs.103.

Oliphant, Travis E. 2006. A Guide to Numpy. Trelgol Publishing.

Pontzen, A., R. Roškar, G. S. Stinson, R. Woods, D. M. Reed, J. Coles, and T. R. Quinn. 2013. "pynbody: Astrophysics Simulation Analysis for Python."

Stark, Casey W. 2012. "Dimensionful." *GitHub Repository*. https://github.com/caseywstark/dimensionful; GitHub.

Stinner, Victor. 2018. "Perf." GitHub Repository. https://github.com/vstinner/perf; GitHub.

The Astropy Collaboration, A. M. Price-Whelan, B. M. Sipőcz, H. M. Günther, P. L. Lim, S. M. Crawford, S. Conseil, et al. 2018. "The Astropy Project: Building an inclusive, open-science project and status of the v2.0 core package." *ArXiv E-Prints*, January.

Turk, M. J., B. D. Smith, J. S. Oishi, S. Skory, S. W. Skillman, T. Abel, and M. L. Norman. 2011. "yt: A Multi-code Analysis Toolkit for Astrophysical Simulation Data." *The Astrophysical Journal Supplement Series* 192 (January): 9. https://doi.org/10.1088/0067-0049/192/1/9.

Walt, S. van der, S. C. Colbert, and G. Varoquaux. 2011. "The Numpy Array: A Structure for Efficient Numerical Computation." *Computing in Science Engineering* 13 (2): 22–30. https://doi.org/10.1109/MCSE.2011.37.