

NAME

swarm — find clusters of nearly-identical nucleotide amplicons

SYNOPSIS

swarm [*options*] [*filename*]

DESCRIPTION

Environmental or clinical molecular studies generate large volumes of amplicons (e.g. SSU-rRNA sequences) that need to be clustered into molecular operational taxonomic units. Traditional clustering methods are based on greedy, input-order dependent algorithms, with arbitrary selection of global cluster size and cluster centroids. To address that problem, we developed **swarm**, a fast and robust method that recursively groups amplicons with *d* or less differences. **swarm** produces stable clusters (or "swarms"), free from centroid selection induced input-order dependency.

Exact clustering is impractical on large data sets when using a naïve all-vs-all approach (i.e. a 2-combination without repetitions), as it implies unrealistic numbers of pairwise comparisons. **swarm** is based on a maximum number of differences, and focuses only on close relationships. An efficient *k*-mer-based filtering and an astute use of comparisons results obtained during the process allows to avoid most of the amplicon comparisons needed in a naïve approach. To speed up the remaining amplicon comparisons, **swarm** implements an extremely fast Needleman-Wunsch algorithm making use of the Streaming SIMD Extensions (SSE2) of modern x86-64 CPUs. If SSE2 instructions are not available, **swarm** exits with an error message.

swarm reads the named input *filename*, a fasta file of nucleotide amplicons. The amplicon identifier is defined as the string comprised between the ">" symbol and the first space or the end of the line, whichever comes first. As **swarm** outputs lists of amplicon identifiers, amplicon identifiers must be unique to avoid ambiguity; **swarm** exits with an error message if identifiers are not unique. If amplicon identifiers end with a "_" followed by a number, that number is used as the amplicon copy number in the statistics output file. The amplicon sequence is defined as a string of [acgt] or [acgu] symbols (case insensitive), starting after the end of the identifier line and ending before the next identifier line or the file end; **swarm** exits with an error message if any other symbol is present. Default is to read from standard input if no file is named, or if the file name is "-".

Options

swarm recognizes the following command-line options:

-b, --break-swarms

output all pairs of nearly-identical amplicons to standard error using a four-columns tab-delimited format: "@@", ampliconA, amplicon B, number of differences. That option, designed to work with the companion script `swarm_breaker.py`, can also be used for swarms post-processing, network analysis and visualization.

-d, --differences *positive integer*

maximum number of differences allowed between two amplicons, meaning that two amplicons will be grouped if they have *integer* (or less) differences. This is **swarm**'s most important parameter. The number of differences is calculated as the number of mismatches (substitutions, insertions or deletions) between the two amplicons once the optimal pairwise global alignment has been found (see "advanced options" for parameters influencing the pairwise alignment). Any *integer* between 1 and 256 can be used, but aligning two very distant amplicons is difficult and results should be considered with caution. Default number of differences is 1.

-h, --help display this help and exit.**-o, --output-file *filename***

output result to *filename*. Result is a list of swarms, one swarm per line. A swarm is a list of amplicon identifiers separated by spaces. Default is to write to standard output.

-s, --statistics-file *filename*

output statistics to the specified file. Default is not to output statistics. The file is a tab-separated table with one swarm per row and seven columns of information: number of

unique amplicons in the swarm, total copy number of amplicons in the swarm, identifier of the initial seed, initial seed copy number (if applicable), number of singletons (amplicons with a copy number of 1), maximum number of generations (i.e. numbers of iterations before the swarm reached its natural limits), and the maximum radius of the swarm (i.e. number of differences between the seed and the furthestmost amplicon in the swarm).

-t, --threads *positive integer*

number of computation threads to use. The number of threads should be lesser or equal to the number of available CPU cores. Default number of threads is 1.

-u, --uclust-file *filename*

output results in uclust-like file format to the specified file. That option does not modify swarm default output format.

-v, --version

output version information and exit.

Advanced options

swarm recognizes advanced command-line options modifying the pairwise global alignment scoring parameters:

-m, --match-reward *positive integer*

reward for a nucleotide match. Default is 5.

-p, --mismatch-penalty *positive integer*

penalty for a nucleotide mismatch. Default is 4.

-g, --gap-opening-penalty *positive integer*

gap open penalty. Default is 12.

-e, --gap-extension-penalty *positive integer*

gap extension penalty. Default is 4.

As **swarm** focuses on close relationships, final results are resilient to model parameters modifications. Modifying model parameters only impacts analysis using a high number of differences.

EXAMPLES

swarm -t 4 -o *myfile.swarms* *myfile.fasta*

Divide the data set *myfile.fasta* into swarms with the finest resolution possible (1 difference) using 4 computation threads, and write the results in the file *myfile.swarms*.

zcat file.fas.gz | **swarm** | awk "{print NF}" | sort -n | uniq -c

Use **swarm** in a pipeline to read a compressed fasta file and to get its swarm size profile (with default parameters).

AUTHORS

Concept by Frédéric Mahé, implementation by Torbjørn Rognes.

REPORTING BUGS

Submit suggestions and bug-reports at <<https://github.com/torognes/swarm/issues>>, send a pull request on <<https://github.com/torognes/swarm>>, or compose a friendly or curmudgeont e-mail to Frédéric Mahé <mahe@rhrk.uni-kl.de> and Torbjørn Rognes <torognes@ifi.uio.no>.

AVAILABILITY

The software is available from <<https://github.com/torognes/swarm>>

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SEE ALSO

swipe, an extremely fast Smith-Waterman database search tool by Torbjørn Rognes (available from <<https://github.com/torognes/swipe>>).

VERSION HISTORY

New features and important modifications of **swarm**:

v1.2.5 released April 11, 2014

Version 1.2.5 removes the need for a POPCNT hardware instruction to be present. Swarm now automatically checks whether POPCNT is available and uses a slightly slower software implementation if not. Only basic SSE2 instructions are now required to run swarm.

v1.2.4 released January 30, 2014

Version 1.2.4 introduces an option `--break-swarms` to output all pairs of amplicons with d differences to standard error. That option is used by the companion script ‘`swarm_breaker.py`’ to refine swarm results. The syntax of the inline assembly code is changed for compatibility with more compilers.

v1.2 released May 16, 2013

Version 1.2 greatly improves speed by using alignment-free comparisons of amplicons based on k -mer word content. For each amplicon, the presence-absence of all possible 5-mers is computed and recorded in a 1024-bits vector. Vector comparisons are extremely fast and drastically reduce the number of costly pairwise alignments performed by swarm. While remaining exact, swarm 1.2 can be more than 100-times faster than swarm 1.1, when using a single thread with a large set of sequences. The minor version 1.1.1, published just before, adds compatibility with Apple computers, and corrects an issue in the pairwise global alignment step that could lead to sub-optimal alignments.

v1.1 released February 26, 2013

Version 1.1 introduces two new important options: the possibility to output swarming results using the uclust output format, and the possibility to output detailed statistics on each swarms. Swarm 1.1 is also faster: new filterings based on pairwise amplicon sequence lengths and composition comparisons reduce the number of pairwise alignments needed and speed up the swarming.

v1.0 released November 10, 2012

First public release