Applying the Klein-Gordon Theory to Gravitation

Modelling Newtonian gravitation as a classical scalar field theory obeying Klein-Gordon structure

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The Inverse Square Law

Gravitational force:

$$F_m = -G \frac{Mm}{r^2}$$

Electrostatic force:

$$F_e = \frac{1}{4\pi\epsilon_0} \frac{Q_e q_e}{r^2}$$

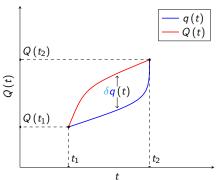
Magnetic force:

$$F_b = \frac{\mu_0}{4\pi} \frac{Q_b q_b}{r^2}$$

Formal Analogies Between the Gravitational and Electrostatic Forces

	Gravitation	Static electricity
	Glavitation	, , ,
Newton's second law	$a' = -\partial' V$	$E' = -\partial' \phi$
	$-\vec{\nabla}V$	$-\vec{ abla}\phi$
	3	3
Gauss' law	$\sum \nabla_i a^i = -4\pi G \rho_m$	$\sum abla_i E^i = rac{1}{\epsilon_0} ho_e$
	i=1	i=1
	$ec{ abla} \cdot ec{a}$	$\vec{ abla}\cdot \vec{a}$
	3	3
Poisson's equation	$\sum \nabla_i \partial^i V = 4\pi G \rho_m$	$\sum abla_i \partial^i \phi = -rac{1}{\epsilon_0} ho_{e}$
	i=1	i=1
	$\nabla^2 V$	$\nabla^2 \phi$

Lagrangian Mechanics



Nature 'selects' the unique on-shell trajectory q(t) given the boundary conditions $(t_1, Q(t_1))$ and $(t_2, Q(t_2))$ for a system.

$$\underbrace{\frac{Q\left(t\right)}{\text{Off-shell}}}_{\text{Off-shell}} = \underbrace{\frac{q\left(t\right)}{\text{On-shell}}}_{\text{Variation}} + \underbrace{\frac{\delta q\left(t\right)}{\text{Variation}}}_{\text{Variation}}$$

Each trajectory Q(t) between the endpoints is associated with a corresponding number called the action.

$$S[Q(t)](t_1,t_2) = \int_{t_1}^{t_2} dt L(Q(t),\dot{Q}(t),t)$$

The integrand $L\left(Q\left(t\right),\dot{Q}\left(t\right),t\right)$ is known as the Lagrangian of the system being modelled and encodes the dynamics of the system.

▶ In general, the action S maps Q(t) to a real number determined by the above integral. Therefore, it is a functional, i.e. a higher-order function which takes in infinite values of the form $\{(t, Q(t)) : t \in \mathbb{R}\}$ and spits out a real.

$$S: \begin{cases} \mathbb{R}^{\mathbb{R}} & \to \mathbb{R} \\ \frac{Q(t)}{Q(t)} & \mapsto \int_{t_1}^{t_2} dt \ L\left(\frac{Q(t)}{Q(t)}, \dot{Q}(t), t\right) \end{cases}$$

Principle of Stationary Action

Lagrange's principle of stationary action

Suppose we vary q(t) about its on-shell evolution as, $q(t) \rightarrow q(t) + \delta q(t)$. Then, the variation in the action satisfies,

$$\delta S \in \mathcal{O}\left(\delta q^2\right)$$

Corollary (First-order approximation)

For very small $\delta q(t)$ i.e.,

$$\forall \, \delta q(t) = \lim_{\epsilon \to 0} \epsilon \eta(t) : \eta(t_1) = \eta(t_2) = 0 :$$

$$\delta S \in \mathcal{O}\left(\epsilon^2 \eta(t)\right) = \{0\}$$

$$\implies \boxed{\delta S = 0}$$

Euler-Lagrange Equation

Lemma (Fundamental lemma of calculus of variations)

The former is possible if and only if the latter is,

$$\forall \, \delta \mathbf{q} : \int_{t_1}^{t_2} dt \, \delta \mathbf{q} \, f\left(\mathbf{q}, \dot{\mathbf{q}}, t\right) = 0$$

$$\iff \forall \, t \in (t_1, t_2) : f\left(\mathbf{q}, \dot{\mathbf{q}}, t\right) = 0$$

Theorem

An on-shell q(t) obeying the principle of stationary action for a given $L(q, \dot{q}, t)$ must also obey the Euler-Lagrange equation of motion:

$$\frac{\partial L}{\partial q} = \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} = \frac{dp}{dt}$$
Generalized force Conjugate momentum

Proof.

$$\delta S = 0 \quad \text{[Principle of stationary action]}$$

$$\delta \int_{t_1}^{t_2} dt \, L\left(q, \dot{q}, t\right) = 0$$

$$\int_{t_1}^{t_2} dt \, \delta L\left(q, \dot{q}, t\right) = 0 \quad \text{[Additivity of variations]}$$

$$\int_{t_1}^{t_2} dt \left[\delta q \frac{\partial L}{\partial q} + \delta \dot{q} \frac{\partial L}{\partial \dot{q}} + \mathcal{M} \frac{\partial L}{\partial t} \right] = 0 \quad \text{[Chain rule for variations]}$$

$$\int_{t_1}^{t_2} dt \left[\delta q \frac{\partial L}{\partial q} + (\dot{\delta q}) \frac{\partial L}{\partial \dot{q}} \right] = 0 \quad \text{[Commutativity of derivatives]}$$

$$\int_{t_1}^{t_2} dt \, \delta q \frac{\partial L}{\partial q} + \int_{t_1}^{t_2} dt \, (\dot{\delta q}) \frac{\partial L}{\partial \dot{q}} = 0$$

Proof.

$$\int_{t_{1}}^{t_{2}} dt \, \delta q \frac{\partial L}{\partial q} + \frac{\partial L}{\partial \dot{q}} \int_{t_{1}}^{t_{2}} dt \, (\dot{\delta q}) - \int_{t_{1}}^{t_{2}} dt \, \left[\int dt \, (\dot{\delta q}) \right] \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} = 0$$
[Integration by parts]
$$\int_{t_{1}}^{t_{2}} dt \, \delta q \frac{\partial L}{\partial q} + \frac{\partial L}{\partial \dot{q}} [\delta q]_{t_{1}}^{t_{2}} - \int_{t_{1}}^{t_{2}} dt \, \delta q \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} = 0$$

$$[\delta q \, (t_{1}) = \delta q \, (t_{2})]$$

$$\forall \, \delta q : \int_{t_{1}}^{t_{2}} dt \, \delta q \, \left(\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \right) = 0$$

$$\iff \frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} = 0$$

$$\iff \frac{\partial L}{\partial q} - \frac{dp}{dt} = 0 \quad \Box$$

[Fundamental lemma of the calculus of variations]

Noether's Theorem

Theorem (Noether's first theorem)

If the action S[q(t)] remains invariant under perturbations of the following form,

$$q \rightarrow q + \delta q$$

then the following quantity is conserved,

$$j = p \, \delta q$$
$$\frac{dj}{dt} = 0$$

Proof.

$$\begin{split} \delta L &= \frac{\partial L}{\partial q} \delta q + \frac{\partial L}{\partial \dot{q}} \delta \dot{q} \\ &= \dot{p} \delta q + p \delta \dot{q} \qquad \text{[Euler-Lagrange equation]} \\ &= \frac{d}{dt} \left(p \delta q \right) \\ \text{But } \delta L &= 0 \\ \implies \frac{d}{dt} \left(p \delta q \right) &= 0 \end{split}$$

Example

If $S\left[q\left(t\right)\right]$ is symmetric (i.e. conserved) under a small time-independent translation $q \to q + \epsilon$, we obtain the invariant $j = p\epsilon$. Since $\frac{dj}{dt} = 0$, $\frac{d\epsilon}{dt} = 0$, we get $\frac{dp}{dt} = 0$.

Classical Mechanics

The Lagrangian for classical mechanics is of the form,

$$L(q, \dot{q}, t) = T(\dot{q}) - V(q)$$

$$= \frac{1}{2} mg \dot{q}^2 - V(q)$$

$$= \frac{1}{2} mv^2 - V(q)$$

► The equation of motion obtained by applying the Euler-Lagrange equation to the above Lagrangian is,

$$\frac{d}{dt}(mv) + \frac{\partial V}{\partial q} = 0$$

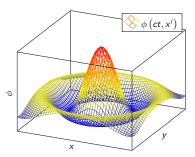
This is Newton's second law. If the entire system concerned is symmetric under small translations on q, we have $\frac{\partial V}{\partial q}=0$ implying $\frac{d}{dt}(mv)=0$. This is Newton's third law.



Classical Field Theory

- A classical field is a tensor field on spacetime (which is a pseudo-Riemannian manifold obeying dynamical field equations such as the Einstein field equations). Therefore, a classical field is some rank (p,q) tensor $\phi^{\mu_1\dots\mu_p}_{(1,\dots,\mu_q)}(x^\alpha)$ at each point x^α in space and time with $\alpha\in(0,1,2,3)$.
- ► A classical field obeys the following principles:
 - 1. Principle of stationary action
 - 2. Local Lorentz invariance
 - 3. Locality
 - 4. Gauge invariance
- ▶ The simplest classical field theory is that of rank (0,0) tensor fields i.e. scalar fields $\phi(x^{\alpha})$, in a flat spacetime \mathcal{M} . We will study such fields in the following slides.

Principle of Stationary Action for Classical Fields



To construct the action for a particle, we integrated its Lagrangian between endpoints in time. A field such as $\phi(x^{\alpha})$, however, lives in space and time. Therefore, its action is a volume integral of a Lagrangian density \mathcal{L} , in a 4-dimensional region of spacetime $\Omega \subset \mathcal{M}$,

$$oxed{S\left[\phi\left(x^{lpha}
ight)
ight] = \int_{\Omega} d^4x \, \mathcal{L}\left(\phi\left(x^{lpha}
ight), \partial_{\mu}\phi\left(x^{lpha}
ight), x^{
u}
ight)}$$

The Lagrangian density is so-called as it looks like a Lagrangian (integrable over some time interval $\Omega^{(1)}$) when integrated over a region of space $\Omega^{(3)}$:

$$L(\phi(x^{\alpha}), \partial_{\mu}\phi(x^{\alpha}), x^{\nu}) = \int_{\Omega^{(3)}} d^{3}x \, \mathcal{L}(\phi(x^{\alpha}), \partial_{\mu}\phi(x^{\alpha}), x^{\nu})$$

$$S[\phi(x^{\alpha})] = \int_{\Omega} d^{4}x \, \mathcal{L}(\phi(x^{\alpha}), \partial_{\mu}\phi(x^{\alpha}), x^{\nu})$$

$$= \int_{\Omega^{(1)}} cdt \, L(\phi(x^{\alpha}), \partial_{\mu}\phi(x^{\alpha}), x^{\nu})$$

► The principle of stationary action for fields states that for small variations $\delta \phi$ of a field ϕ in its on-shell configuration, the action remains stationary,

$$\delta S = 0$$

Euler-Lagrange Equation for Classical Fields

Lemma (Fundamental lemma of multivariable calculus of variations)

$$\forall \, \delta \phi : \int_{\Omega} d^4 x \delta \phi \, f \left(\phi, \partial_{\mu} \phi, x^{\nu} \right) = 0$$

$$\iff \forall \, x^{\alpha} \in \Omega \backslash \partial \Omega : f \left(\phi, \partial_{\mu} \phi, x^{\nu} \right) = 0$$

Einstein summation convention

Dummy indices, i.e. pairs of upper and lower tensor indices, are implicitly summed over.

Example

$$A_\mu B^\mu = \sum_{\mu=0}^3 A_\mu B^\mu$$

Theorem

A field ϕ obeys the principle of stationary action if and only if it also satisfies,

$$\frac{\partial \mathcal{L}}{\partial \phi} = \nabla_{\mu} \underbrace{\frac{\partial \mathcal{L}}{\partial \left(\partial_{\mu} \phi\right)}}_{\text{Conjugate momentum tensor}} = \nabla_{\mu} \pi^{\mu}$$

Proof.

$$\delta S = 0 \qquad \qquad \text{[Principle of stationary action]}$$

$$\delta \int_\Omega d^4x \, \mathcal{L} = 0$$

$$\int_\Omega d^4x \, \delta \mathcal{L} = 0 \qquad \qquad \text{[Additivity of variations]}$$

$$\int_{\Omega} d^4 x \left[\delta \phi \frac{\partial \mathcal{L}}{\partial \phi} + \delta \left(\partial_{\mu} \phi \right) \underbrace{\frac{\partial \mathcal{L}}{\partial \left(\partial_{\mu} \phi \right)}}_{\pi^{\mu}} + \delta x^{\mu} \partial_{\mu} \mathcal{L} \right] = 0$$

[Multivariable chain rule for variations]

$$\int_{\Omega} d^4x \left[\delta \phi \frac{\partial \mathcal{L}}{\partial \phi} + (\partial_{\mu} \delta \phi) \pi^{\mu} \right] = 0$$

[Commutativity of variations and covariant derivatives]

$$\int_{\Omega} d^4x \, \delta\phi \frac{\partial \mathcal{L}}{\partial \phi} + \pi^{\mu} \underbrace{\int_{\Omega} d^4x \, \partial_{\mu} \delta\phi}_{\text{Constant surface term}} - \int_{\Omega} d^4x \, \left[\int d^4x \, \partial_{\mu} \delta\phi \right] \nabla_{\mu} \pi^{\mu} = 0$$

[Volume integration by parts]

Using Stokes' theorem, the constant surface term can be set to 0. We then find,

$$\begin{split} \int_{\Omega} d^4 x \, \delta \phi \frac{\partial \mathcal{L}}{\partial \phi} - \int_{\Omega} d^4 x \, \delta \phi \, \nabla_{\mu} \pi^{\mu} &= 0 \\ \int_{\Omega} d^4 x \, \delta \phi \left(\frac{\partial \mathcal{L}}{\partial \phi} - \nabla_{\mu} \pi^{\mu} \right) &= 0 \\ \frac{\partial \mathcal{L}}{\partial \phi} - \nabla_{\mu} \pi^{\mu} &= 0 \\ \iff \frac{\partial \mathcal{L}}{\partial \phi} - \nabla_{\mu} \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi)} &= 0 \quad \Box \end{split}$$

[Fundamental lemma of multivariable calculus of variations]

Noether's Theorem for Classical Fields