plyr: divide and conquer

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plyr is a set of tools for a common set of problems: you need to break a big problem down into manageable pieces, operate on each piece and then put all the pieces back together. I call this strategy "split-apply-combine" and the three components form the basis for this paper and the plyr package.

The paper concludes with two case studies using plyr to operate on a 3d array of spatio-temporal dataset, and a large data frame of repeated observations.

1 Introduction

The plyr package provides tools for solving a common class of problems, where you break apart a big complicated data structure into small simple pieces, operate on each piece independently and then put all the pieces back together (possibly in a different format to the original). This paper introduces the ply family of functions which generalise the apply family found in the base package, and include all combinations of input and output of lists, data frames and arrays.

This paper describes version 0.1 of plyr, which requires R 2.7.0 or later and has no run-time dependencies. To install it from within R, run install.packages("plyr"). Information about the latest version of the package can be found online at http://had.co.nz/plyr.

In general, plyr provides a replacement for for loops for a large set of practical problems. The major assumption that the ply functions make is that each piece can be operated on independently of the other pieces, so if there is any dependence between the pieces then you will need to use other tools. The need to provide an alternative to loops does not come about because loops are slow (they are not!), but because they do not clearly express the intent of the algorithm, as important details are mixed in with unimportant book-keeping code. The tools of plyr aim to eliminate this extra code and illuminate the key components of your computations.

Section 3 introduces the plyr family of tools and how to use them. The plyr package also provides a number of helper functions for error recovery, splatting, columnwise processing, and reporting progress, described in Section 4. Section 5 discusses the general strategy that these functions support, including cases studies that explore the performance of veteran baseball players and ozone measured over space and time. Finally, Section 6 maps existing R functions to their plyr counterparts and lists related packages. Section 7 describes future plans.

2 Motivation

Why use plyr? Why not use for loops or the built-in apply functions? This section compares plyr code to base R code for an example that is explained in more detail in Section 5.2.

In this example we are going to remove seasonal affects from satellite measurements of ozone. The ozone was measured on a 24×24 grid, each month for six years, and is stored in a $24 \times 24 \times 72$ 3d array. A single location (ozone[x, y,]) is a vector of 72 values, and we can crudely deasonalise it by looking at the residuals of a robust linear model:

```
one <- ozone[1, 1, ]
month <- factor(rep(1:12, length = 72))
model <- rlm(one ~ month - 1)
deseasf <- resid(model)

deseasf <- function(value) rlm(value ~ month - 1)</pre>
```

The challenge is now to apply this function to each location, assembling the results back into the same form as the input. We also want to keep the intermediate models in a 2d array, so we can reference a local model (model[1, 1]) in a similar way to referencing a local time series (ozone[1, 1,]). In base R, we can tackle this problem with for loops, or with the apply family of functions:

```
Apply functions
For loops
models \leftarrow as.list(rep(NA, 24 * 24))
                                           models <- apply(ozone, 1:2, deseasf)
dim(models) \leftarrow c(24, 24)
                                           resids <- unlist(lapply(models, resid))
deseas \leftarrow array(NA, c(24, 24, 74))
                                           dim(resids) <- c(72, 24, 24)
                                           deseas \leftarrow aperm(resids, c(2, 3, 1))
for (i in seq_len(24)) {
                                           dimnames(deseas) <- dimnames(ozone)</pre>
  for(j in seq_len(24)) {
    mod <- deseasf(ozone[i, j, ])</pre>
    models[[i, j]] <- mod
    deseas[i, j, ] <- resid(mod)</pre>
  }
}
```

The main disadvantage of the for loop is that there is lot of book-keeping code in there. We need to create the output structure before filling them up, and the size of the array is hard coded in multiple places. The apply functions (apply() and lapply()) simplify this task, but there's no straightforward way to go from the 2d array of models to a 3d array of residuals. In plyr, the code is much shorter because these details are taken care of:

```
models <- aaply(ozone, 1:2, deseasf)
deseas <- aaply(models, 1:2, resid)</pre>
```

You may be wondering what those function names mean. All plyr functions have a concise but informative naming scheme: the first two characters describe input and output data types. Both of these functions input and output an array. Other data types are lists and data frames. Because plyr caters for every combination of input and output data types in a consistent way, it is easy to use the data structure that feels most natural for a given problem.

For example, instead of storing the ozone data in a 3d array, we could also store it in a data frame. This type of format is more common if the data is ragged, irregular or incomplete, where we don't have measurements at every possible location for every possible time point. Imagine the data frame is called ozonedf and has columns lat, long, time, month, and value. To repeat the deasonalisation task with this new data format, we first need to tweak our workhorse method:

```
deseasf_df <- function(df) {
  rlm(value ~ month - 1, data = df)
}</pre>
```

Because the data could be ragged, it's much harder to use for loops here and we'll use split(), lapply() and mapply() to complete the task. Here the split-apply-combine strategy maps closely to built-in R functions: we split with split(), apply with lapply() and then combine the pieces into a single result with rbind().

```
pieces <- split(ozonedf, list(ozonedf$lat, ozonedf$long))
models <- lapply(pieces, deseasf)

results <- mapply(function(model, df) {
   cbind(df[rep(1, 72), c("lat", "long")], resid(model))
}, models, pieces)
deseasdf <- do.call("rbind", results)</pre>
```

Much of the complication here is the labelling - we only needed to use mapply() so we could match the original data up with the models. plyr takes care of all the tricky labelling stuff for you, so it only takes two lines:

```
models <- dlply(ozone, .(lat, long), deseas_df)
deseas <- ldply(models, resid)</pre>
```

The following section describes the plyr functions in more detail. If your interest has been whetted by this example, you might want to skip ahead to page 15 to learn more about this data.

3 Usage

Table 1 lists the basic set of plyr functions. Each function is named according to the type of input it accepts and the type of output it produces. The input type determines how the big data structure can be broken down into small pieces, and the output type determines how the pieces are joined back together again. Breaking down input is described in Section 3.1 and piecing together output is described in Section 3.2.

NB: In this paper, the term **array** includes the special cases of vectors (1d arrays) and matrices (2d arrays) as well, and the term **list-array** refers to an list with dimensions (as opposed to an atomic vector, as is more usual). The common atomic vectors

from	array	data frame	list
array	aaply	daply	laply
data frame	aaply adply	ddply	ldply
list	alply	dlply	llply
nothing	a_ply	d_ply	1_ply

Table 1: The 12 key functions of plyr. Arrays include matrices and vectors as special cases.

are logical, character, integer, and numeric. Dimension labels refer to the output of dimnames(), or for 2d structures, the special cases of rownames() and colnames().

The effects of the input and outputs types are orthogonal, so instead of having to learn all 12 functions individually, it is sufficient to learn the three types of input and four types of input. For this reason, it's useful to refer to a complete row (common output type) or column (common input type) of Table 1. The notation we use for this is d*ply to refer an entire row (same input) and *dply for an entire column (same output).

The **ply functions have either two or three main arguments, depending on the type of input:

```
• a*ply(data., margins., fun., ..., progress. = "none")
```

- d*ply(data., variables., fun., ..., progress. = "none")
- l*ply(data., fun., ..., progress. = "none")

The first argument is the data. which will be split up, processed and recombined. The second argument, variables. or margins., describes how to split up the input into pieces. The third argument, fun., is the processing function, and is applied to each piece in turn. All further arguments are passed on to the processing function. If you omit fun. the individual pieces will not be modified, but the entire data structure will be converted from one type to another. The progress. argument controls displaying of a progress bar, and is described at the end of Section 4.

Note that arguments to the **ply functions end in ".". This prevents name clashes with the argument of fun..

3.1 Input

Each type of input has different rules for how to split it up, and are described in detail in the following section. In short:

- a*ply(): Arrays are sliced by dimension in to lower-d pieces.
- d*ply(): Data frames are subsetted by combinations of variables.
- 1*ply(): Each element in a list is a piece.

Technical note: The way the input can be split up is not actually determined by the type of the data structure, but the methods that it responds to. An object split up by a*ply() must respond to dim() and accept multidimensonal indexing; by d*ply(),

must work with split(), and must be coercible to an environment; by list, must work with length() and [[.

The most important result of that is that data frames can also be passed to a*ply(), where they are treated like 2d matrices, and to 1*ply() which will operate column wise on the data frame.

3.1.1 Input: array (a*ply)

The margins argument of a*ply describes which dimensions to slice along (in the same way that apply() does). There are four possible ways to do this for the 2d case is simple, as illustrated Figure 1:

- margins = 1: split into rows
- margins = 2: split into columns
- margins = c(1,2): split into individual cells

The last way is to not split up the matrix at all, and corresponds to margins = c(). (However, there's not much point in using plyr to do this!)

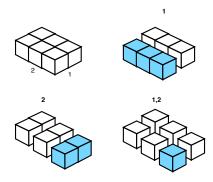


Figure 1: The four ways to split up a 2d matrix, labelled above by the dimensions that they slice up. Original matrix shown at top left, with dimensions labelled. Blue indicates a single piece of the output.

The 3d case is a little more complicated. We have three 2d slices, three 1d slices, and one 0d slice. These are shown in Figure 2. Note how the pieces for the 1d slices correspond to the intersection of the 2d slices. The margins argument works correspondingly for higher dimensions, with a combinatorial explosion in the number of possible ways to slice up the array, choose(slice-d, array-d), to be exact.

These default to working on the first dimension (i.e. row-wise) and automatically splat the function so that function is called not with a single list as input, but each column is passed as a separate argument to the function. Compared to using mapply, for the m*ply functions you will need to cbind the columns together first. This will ensure that each argument has the same length, and allows the m*ply functions to have the same argument order as all the other

Special case: m*ply A special case of operating on arrays corresponds to the mapply function of base R. The plyr equivalents are named maply, mdply, mlply and m_ply. mapply() takes multiple lists of parameters as input, and calls the processing function with a piece from each list as its parameters. The input to m*ply() is a little different: it is a list-array.

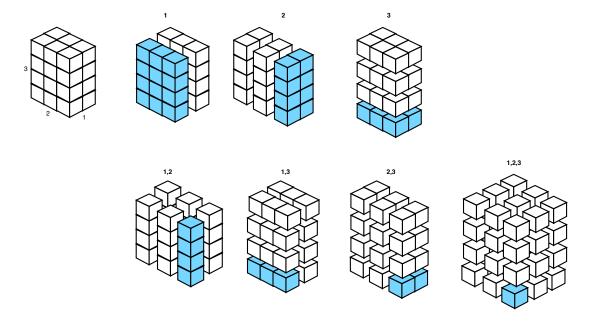


Figure 2: The eight ways to split up a 3d array, labelled above by the dimensions that they slice up. Original array shown at top left, with dimensions labelled. Blue indicates a single piece of the output.

3.1.2 Input: data frame (d*ply)

When operating on a data frame, you usually want to split it up into groups based on combinations variables in the data set. For d*ply you specify which variables (or functions of variables) to use. These variables are specified in a special way to highlight that they are computed first from the data frame, then the global environment (in which case it's your responsibility to ensure that their length is equal to the number of rows in the data frame).

- .(var1) will split the data frame into groups defined by the value of the var1 variable. If you use multiple variables, .(a, b, c), the groups will be formed by the interaction of the variables, and output will be labelled with all three variables.
- You can also use functions of variables: .(round(a)), .(a * b). If you are outputting to a data frame, these will get ugly names (produced by make.names()), but you can override them by specifying names in the call: .(product = a * b)
- By default, plyr will look in the data frame first, and then in the global environment .(anothervar). However, you are encouraged to keep all related variables in the same data frame: this makes things much easier in the long run.

Figure 3 shows two examples of splitting up up a simple data frame. Splitting up data frames is easier to understand (and to draw!) than splitting up arrays, because they're only 2 dimensional.

3.1.3 Input: list (1*ply)

Lists are the simplest type of input to deal with because they are already naturally divided into pieces: the elements of the list. For this reason, the 1*ply functions don't

			.(sex)			.(age)		
name	age	sex	name	age	sex	name	age	sex
John	13	Male	John	13	Male	John	13	Male
Mary	15	Female	Peter	13	Male	Peter	13	Male
Alice	14	Female	Roger	14	Male	Phyllis	13	Female
Peter	13	Male						
			name	age	sex	name	age	sex
Roger	14	Male	Mary	15	Female	Alice	14	Female
Phyllis	13	Female	Alice	14	Female	Roger	14	Male
			Phyllis	13	Female			
			,			name	age	sex
						Mary	15	Female

Figure 3: Two examples of splitting up a data frame by variables. If the data frame was split up by both sex and age, there would only be one subset with more than one row: 13-year-old males.

need an argument that describes how to break up the data structure.

Special case: r*ply A special case of operating on lists corresponds to replicate() in base R, and is useful for drawing distributions of random numbers. This is a little bit different to the other plyr methods. Instead of the data. argument, it has n. the number of replications to run, and instead of a function it accepts a expression.

3.2 Output

The output type defines how the pieces will be joined back together again, and how they will be labelled. The labels are particularly important to allow you to match up the input with the output.

The input and output types are the same, except there is an additional output option, which discard the output. This is useful for functions with side effects that make changes outside of R

The output type also places some restrictions on what type of results the processing function should return. Generally, the processing function should return the same type of data as the eventual output, (i.e. vectors, matrices and arrays for *aply and data frames for *dply) but some other formats are accepted for convenience and are described in Table 2. These are explained in more detail in the individual output type sections.

3.2.1 Output: array (*aply)

With array output the shape of the output array is determined by the input splits and the dimensionality of each individual result. Figures 4 and 5 illustrate this pictorially for simple 1d and 2d cases. For arrays, the pieces contribute to the output in the expected way; lists are related like a 1d array; and data frames get a dimension for each variable in the split. The dimnames of the array will be the same as the input, if an array; or extracted from the subsets, if a data frame.

Output	Processing function restrictions	Null output
*aply *dply frame *lply *_ply	atomic array, or list data frame, or atomic vector none none	<pre>logical() data.frame() list()</pre>

Table 2: Summary of processing function restrictions and null output values for all output types. Explained in more detail in each output section.

The processing function should return an atomic (i.e. logical, character, numeric or integer) array of fixed size/shape, or a list. If atomic, the extra dimensions will added perpendicular to the original dimensions. If a list, the output will be a list-array. If there are no results, *aply will return a logical vector of length 0.

All *aply functions have a drop. argument. When this is true, the default, any dimensions of length one will be dropped. This is useful because in R, a vector of length three is not equivalent to a 3×1 matrix or a $3 \times 1 \times 1$ array.

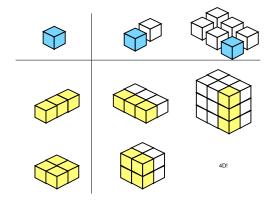


Figure 4: Results from outputs of various dimensionalty from a **single** value, shown top left. Columns indicate input: (left) a vector of length two, and (right) a 2×2 matrix. Rows indicate the shape of a single processed piece: (top) a vector of length 3, (bottom) a 2×2 matrix. Extra dimensions are added perpendicular to existing ones. The array in the bottom-right cell is 4d and so is not shown.

3.2.2 Output: data frames (*dply)

When the output is a data frame, it will contain the results and additional columns that identify where in the original data each row came from. These columns make it possible to merge the old and new data if you need to. If the input was a data frame, there will be a column for variables used to split up the original data; if it was a list, a column for the names of the list; if an array, a column for the names of each splitting dimension. Figure 6 illustrates this for data frame input.

The processing functions should either return a data frame, or a (named) atomic vector of fixed length, which will form the columns of the output. If there are no results, *dply will return an empty data frame. plyr provides an as.data.frame method for functions which can be handy: as.data.frame(mean) will create a new function which outputs a data frame.

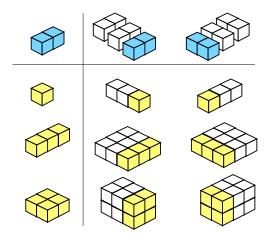


Figure 5: Results from outputs of various dimensionalty from a 1d vector, shown top left. Columns indicate input: (left) a 2×3 matrix split by rows and (right) and 3×2 matrix split by columns. Rows indicate the shape of a single processed piece: (top) a single value, (middle) a vector of length 3, and (bottom) a 2×2 matrix.

.(s	sex)	.(age)		.(sex, age)			
sex	value	age	value		sex	age	value
Male	3	13	3		Male	13	2
Female	3	14	2		Male	14	1
		15	2		Female	13	1
					Female	14	1
					Female	15	1

Figure 6: Illustrating the output from using ddply() on the example from Figure 3 with nrow(). Splitting variables shown above each example. Note how the extra labelling columns are added so that you can identify to which subset the results apply.

3.2.3 Output: list (*lply)

This is the simplest output format, where each processed piece is joined together in a list. The list also stores the labels associated with each pieces, so that if you use ldply or laply to further process the list the labels will appear as if you had used aaply, adply, daply or ddply directly. llply is convenient for calculating complex objects once (e.g. models), from which pieces of interest are later extracted into arrays and data frames.

There are no restrictions on the output of the processing function. If there are no results, *lply will return a list of length 0.

3.3 Output: nothing (*_ply)

Sometimes you are operating on a list purely for the side effects (e.g. plots, caching, output to screen/file). This is a little more efficient than abandoning the output of *lply because it doesn't store the intermediate results.

4 Helpers

The plyr package also provides a number of helper function which take a function (or functions) as input and return a modified function as output.

• splat() converts a function to use. This is useful when you want to pass a function a row of data frame or array, and don't want to manually pull it apart in your function. For example:

```
hp_per_cyl <- function(hp, cyl, ...) hp / cyl
splat(hp_per_cyl)(mtcars[1,])
splat(hp_per_cyl)(mtcars)</pre>
```

Generally, splatted functions should have ... as an argument, so you only need to specify the variables that you are interested in. For more information on how splat works, see do.call.

splat() is applied to functions used in m*ply by default.

- each() takes a list of functions and produces a function that runs each function on the inputs and returns a named vector of outputs. For example, each(min, max) is short hand for function(x) c(min = min(x), max = max(x)). Using each with a single function is useful if you want a named vector as output.
- colwise() converts a function that works on vectors, to one that operates columnwise of data frame, returning a data fram. For example, colwise(median) is a function that computes the median of each column of a data frame.
 - The optional .if argument specialises the function to only run on certain types of vector, e.g. .if = is.factor or .if = is.numeric. These two restrictions are provided in the premade calcolwise and numcolwise.
- failwith() sets a default value to return if the function throws an error. For example, failwith(NA, f) will return an NA whenever f throws an error.

The optional quiet argument suppresses any notice of the error when it is TRUE.

• Given a function, as.data.frame.function() creates a new function which coerces the output of the input function to a data frame. This is useful when you are using *dply() and the default column-wise output is not what you want.

Each plyr function also has a **progress**. argument which allows you to monitor the progress of long running operations. There are four difference progress bars:

- "none", the default. No progress bar is displayed.
- "text" provides a textual progress bar which.
- "win" and "tk" provide graphical progress bars for Windows and systems with the tcl/tk package (the mac and most linux platforms).

The progress bars assume that processing each piece takes the same amount of time, so will not be 100% accurate.

5 Strategy

Having learned the basic structure and operation of the plyr family of functions, you will now see some examples of using them in practice. The following two case studies explore two data sets: a data frame of batting records from long-term baseball players, and a 3d array recording ozone measurements that vary over space and time. Neither of these data studies do more than scratch the surface of possible analyses, but do show of a number of different ways to use plyr.

Both cases follow a similar process:

- 1. Extract a subset of the data for which it is easy to solve the problem
- 2. Solve the problem by hand, checking results as you go.
- 3. Write a function that encapsulates the solution.
- 4. Use the appropriate plyr function to split up the original data, apply the function to each piece and join the pieces back together.

The code shown in this paper is necessarily abbreviated. The data sets are large, and often only small subsets of the data are shown. The code focuses on data manipulation, and much of the graphics code is omitted. You are encouraged to experiment with the full code yourself

5.1 Case study: baseball

The baseball data set contains the batting records for all professional US players with 15 or more years of data. The complete list of variables are described fully ?baseball, but for this example we will focus on just four: id, which identifies the player, year the year of the record and rbi the number of runs that the player made in the season, and at bat, the number of times the player had an opportunity to hit the ball.

(This is a rather crude analysis, as it doesn't take into account the people that might already be on the other plates)

What we'll explore is the performance of a batter over his career. To get started, we need to calculate the careeryear, i.e. the number of years since the player started playing. This is easy to do if we have a single player:

```
> baberuth <- subset(baseball, id == "ruthba01")
> baberuth$cyear <- baberuth$year - min(baberuth$year) + 1

To do this for all players, we first make a function:</pre>
```

```
> calculate_cyear <- function(df) {
+   transform(df,
+     cyear = year - min(year),
+     cpercent = (year - min(year)) / (max(year) - min(year))
+   )
+ }</pre>
```

and then split up the whole data frame into people, run the function on each piece and join them back together into a data frame:

```
> baseball <- ddply(baseball, .(id), calculate_cyear)
> baseball <- subset(baseball, ab >= 25)
```

To summarise the pattern across all players, we first need to figure out what the common patterns are. A time series plot of rbi/ab, runs per bat, is a good place to start. We do this for Babe Ruth, as shown in Figure 5.1, then write a function to do it for any player (taking care to ensure common scale limits) and then use d_ply to save a plot for every player to a pdf. We use two tricks here: reorder to sort the players in order of average rbi / ab, and failwith to ensure that even if a single plot doesn't work we will still get output for the others.

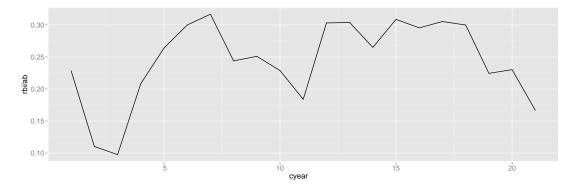


Figure 7: Runs per bat for Babe Ruth.

```
xlim <- range(baseball\$cyear, na.rm=TRUE)
ylim <- range(baseball\$rbi / baseball\$ab, na.rm=TRUE)
plotpattern <- function(df) {
    qplot(cyear, rbi / ab, data = df, geom="line", xlim = xlim, ylim = ylim)
}

pdf("paths.pdf", width=8, height=4)
d_ply(baseball, .(reorder(id, rbi / ab)), failwith(NA, plotpattern),
    print. = TRUE)
dev.off()</pre>
```

Flicking through the 1145 plots reveals few common patterns, although many players do seem to have a roughly linear trend with quite a bit of noise. We'll start by fitting a linear model to each player and then exploring the results. This time we'll skip doing it by hand and go directly to the function. (Not recommended in practice!)

Now we have a list of 1145 models, one for each player. To do something interesting with these, we need to extract some summary statistics. We'll extract the coefficients of the model (the slope and intercept), and a measure of model fit so we can ensure we're not drawing conclusions based on models that fit the data very poorly, the R-squared. The first few rows of coef are shown in Table 3.

```
> rsq <- function(x) summary(x)$r.squared
> coef <- ldply(models, function(x) c(coef(x), rsq(x)))
> names(coef) <- c("id", "intercept", "slope", "rsquare")</pre>
```

id	intercept	slope	rsquare
aaronha01	0.18	0.00	0.00
abernte02	0.00		0.00
adairje01	0.09	-0.00	0.01
adamsba01	0.06	0.00	0.03
adamsbo03	0.09	-0.00	0.11
adcocjo01	0.15	0.00	0.23

Table 3: The first few rows of the coef data frame. Note that the player ids from the original data have been preserved

Figure 5.1 displays the distribution of r-squared across the models. The models generally do a very bad job of fitting the data! Figure 5.1 summarises these bad models. These plots show a negative correlation between slope and intercept, and the particularly bad models have estimates for both values close to 0. Reassuringly, there are no players in the bottom left quadrant with both negative slope and intercept.

This concludes the baseball player case study, which used used ddply, d_ply, dlply and ldply. Our statistical analysis was not very sophisticated, but the tools of plyr made it very easy to work at the player level, and then combine results into a single summary.

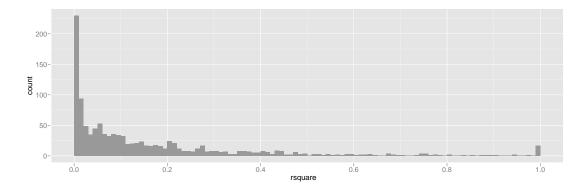


Figure 8: Histogram of model r-squared with bin width of 0.05. Most models fit very poorly! The spike of models with a r-squared of 1 are players with only two data points, found by inspecting ldply(models[coef\$rsquare == 1], "[[", "model")

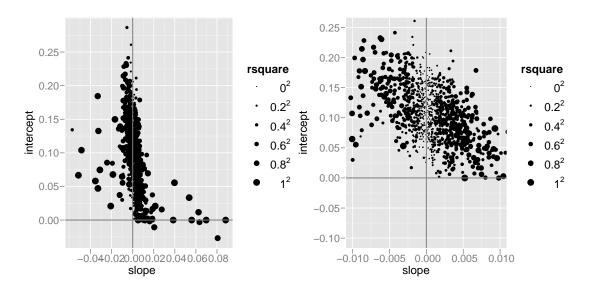


Figure 9: A scatterplot of model intercept and slope, with one point for each model (player). The size of the points is proportion to the R-square of the model. Vertical and horizontal lines emphasis the x and y origins.

5.2 Case study: ozone

In this case study we will analyse a 3d array that records ozone levels over a 24×24 spatial grid at 72 time points (Hobbs et al., To appear). This produces a $24 \times 24 \times 72$ 3d array, containing a total of 41 472 data points. Figure 5.2 is one way of displaying this data. Conditional on spatial location, each glyph shows the evolution of ozone levels for each of the 72 months (6 years). The striking seasonal patterns make it difficult to see if there are any long-term changes. In this case study, we will explore how to separate out and visualise the seasonal effects.

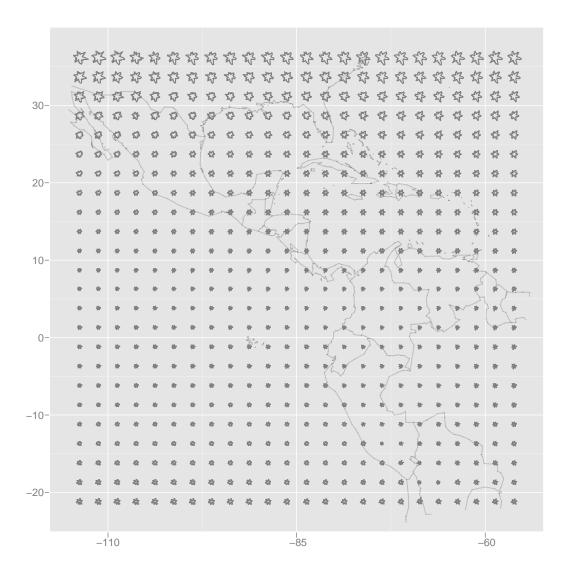


Figure 10: Star glyphs showing variation in ozone over time at each spatial location.

Again we will start with the simplest case: a single time point, from location (1, 1). Figure 5.2 displays this in two ways: as a single line over time, or a line for each year over the months. This second plot illustrates the striking seasonal variation at this time point. The following code sets up some useful variables.

```
> value <- ozone[1, 1, ]
> time <- 1:72 / 12</pre>
```

```
> month.abbr <- c("Jan", "Feb", "Mar", "Apr", "May",
+ "Jun", "Jul", "Aug", "Sep", "Oct", "Nov", "Dec")
> month <- factor(rep(month.abbr, length = 72), levels = month.abbr)
> year <- rep(1:6, each = 12)</pre>
```

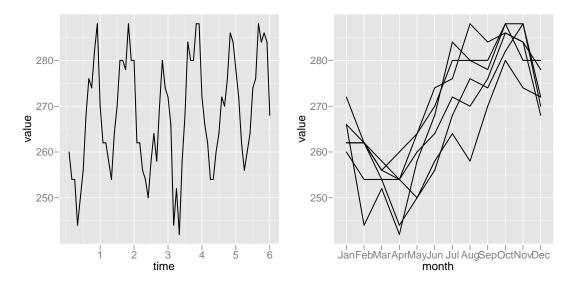


Figure 11: Two ways of displaying the seasonal changes. (Left) A single time series over all six years and (right) a line for each year.

We are going to use a quick and dirty method to remove the seasonal variation: residuals from a robust linear model, predicting amount of ozone by month. We could use a regular linear model, but then our seasonal estimates might be thrown off by an very unusual month. Figure 5.2 shows the deseasonalised trend from location (1, 1)

```
> library(MASS)
+
> deseas1 <- rlm(value ~ month - 1)
> summary(deseas1)

Call: rlm(formula = value ~ month - 1)
Residuals:
    Min    1Q Median    3Q    Max
```

1.0

3.0

11.3

Coefficients:

-3.3

-18.7

	Value	Std. Error	t value
${\tt monthJan}$	264.40	2.75	96.19
${\tt monthFeb}$	259.20	2.75	94.30
${\tt monthMar}$	255.00	2.75	92.77
${\tt monthApr}$	252.00	2.75	91.68
${\tt monthMay}$	258.51	2.75	94.05
${\tt monthJun}$	265.34	2.75	96.53
${\tt monthJul}$	274.00	2.75	99.68
monthAug	276.67	2.75	100.66

```
      monthSep
      277.00
      2.75
      100.78

      monthOct
      285.00
      2.75
      103.69

      monthNov
      283.60
      2.75
      103.18

      monthDec
      273.20
      2.75
      99.39
```

Residual standard error: 4.45 on 60 degrees of freedom

> coef(deseas1)

```
monthJan monthFeb monthMar monthApr monthMay monthJun monthJul monthAug 264 259 255 252 259 265 274 277 monthSep monthOct monthNov monthDec 277 285 284 273
```

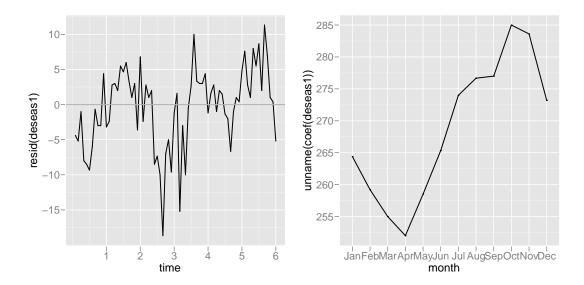


Figure 12: Deasonalised ozone trends. (Left) deasonalised trend over six years. (Right) Estimates of seasonal effects. Compare to Figure 5.2

We next turn this into a function and fit the model to each spatial location. This does take a little while, but we are fitting 576 models! From those models we extract the deasonalised values (the residuals) and the seasonal coefficients. Looking at the dimensionality we see that they're in the same format as the original data. We also carefully label the new dimensions. This is important: just as data frames should have descriptive variable names, arrays should always have descriptive dimension labels.

```
> deseasf <- function(value) rlm(value ~ month - 1)
> models <- alply(ozone, 1:2, deseasf)
+
ERROR: attempt to apply non-function
> coefs <- laply(models, coef)
ERROR: could not find function "fun."
> dimnames(coefs)[[3]] <- month.abbr
> names(dimnames(coefs))[3] <- "month"
+
> deseas <- laply(models, resid)
ERROR: Results must have the same dimensions.</pre>
```

```
> dimnames(deseas)[[3]] <- 1:72
> names(dimnames(deseas))[3] <- "time"
+
> dim(coefs)
[1] 24 24 12
> dim(deseas)
[1] 24 24 72
```

We now have a lot of data to try and understand: for each of the 576 locations we have 12 estimates of monthly effects, and 72 residuals. There are many different ways we could visualise this data. Figures 5.2 and 5.2 visualise these results with star glyph plots. For plotting, it's more convenient to have the data in data frames. There are few different ways to do this: we can convert from the 3d array to a data frame with melt() from the reshape package, or use ldply() instead of laply(). For this example, we'll use a combination of these techniques. We'll convert the original array to a data frame, add on some useful columns, and then perform the same steps as above with this new format. Notice how our effect labelling the dimensions pays off with useful columns in the data frame.

```
> coefs_df <- melt(coefs)
> coefs_df <- ddply(coefs_df, .(lat, long), transform,
   avg = mean(value),
   std = value / max(value)
+ )
> levels(coefs_df$month) <- month.abbr</pre>
> head(coefs_df)
    lat long month value avg
1 -21.2 -114
                May
                      264 269 0.928
2 -21.2 -114
                Apr
                      259 269 0.909
3 -21.2 -114
                      255 269 0.895
                Aug
4 -21.2 -114
                      252 269 0.884
                Jan
5 -21.2 -114
                Sep
                      259 269 0.907
6 -21.2 -114
                Jul
                      265 269 0.931
> deseas_df <- melt(deseas)</pre>
> head(deseas_df)
    lat long time value
1 -21.2 -114
                 1 - 4.40
2 -18.7 -114
                 1 - 3.33
3 -16.2 -114
                 1 - 2.96
4 -13.7 -114
                 1 - 5.00
5 -11.2 -114
                 1 -4.00
  -8.7 -114
                 1 - 3.00
```

The star glyphs show temporal patterns conditioned on location. We can also look at spatial pattern conditional on time. One way to do this is to draw tile plots where each cell of the 24×24 grid is coloured according to its value. The following code sets up a function with constant scales to do that. Figure 5.2 shows the spatial variation of seasonal coefficients for January and July.

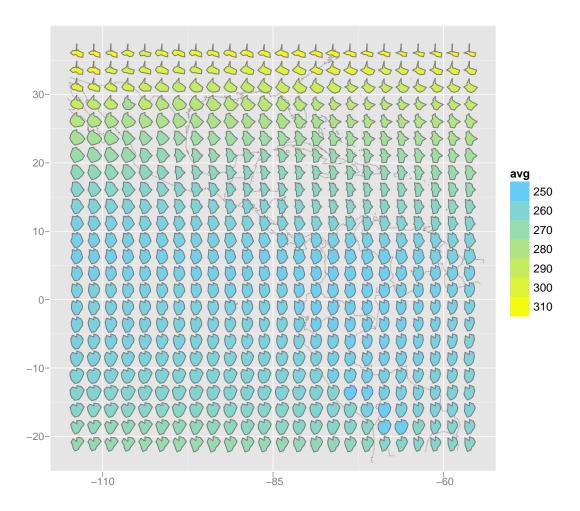


Figure 13: Star glyphs showing seasonal variation. Estimates of seasonal effects are standardised to have the same maximum at each location to make it easier to compare the general pattern. The glyph colours give the overall average ozone measurement. Note the strong spatial correlation: nearby glyphs have similar shapes.

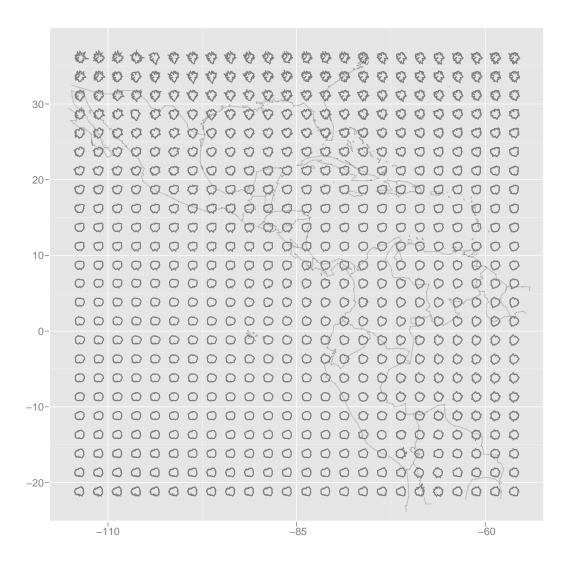


Figure 14: Star glyphs showing deasonalised trends. This plot contains a lot of data—over 40,000 observations—and rewards detailed study. Looking at a printed version also helps as the resolution of a printer (600 dpi) is much higher than that of the screen (100 dpi).

```
> coef_limits <- range(coefs_df$value)
> coef_mid <- mean(coefs_df$value)
> monthsurface <- function(mon) {
+    df <- subset(coefs_df, month == mon)
+    qplot(long, lat, data = df, fill = value, geom="tile") +
+    scale_fill_gradient(limits = coef_limits,
+    low=brightblue, high="yellow") + map + opts(aspect.ratio = 1)
+ }</pre>
```

We could do the same thing for the values themselves, but we'd probably want to make an animation rather than looking at all 72 plots individually. The *_ply functions are useful for making animations because we are only calling the plotting function for its side effects, not because we're interested in its value. For lattice and ggplot graphics we need to use the print. argument to ensure that the plots are rendered.

```
pdf("ozone-animation.pdf", width=8, height=8)
l_ply(month.abbr, monthsurface, print. = TRUE)
dev.off()
```

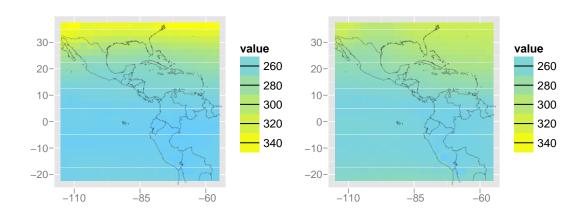


Figure 15: Tile plots of coefficients for January (left) and July (right).

5.3 Other uses

The transform() and subset() functions work will in combination with plyr. Transform makes it very easy to perform randomisation within groups. For example, if we wanted to break the dependence between ozone values and time within each spatial location, we could do:

```
ddply(coefs_df, .(lat, long), transform, time = sample(time))
```

This technique is useful for performing block bootstrapping and other related permutation tests. Scaling variables within a group is also trivial:

```
ddply(coefs_df, .(lat, long), transform, value = scale(value))
```

If we wanted to extract the observation in each group with the lowest value ozone of ozone, it's just as easy:

```
ddply(coefs_df, .(lat, long), subset, value == min(value))
```

For simulations, mdply() can be very useful, because it's easy to generate a grid of parameter values and then evaluate them. This can also be useful when testing many possible combinations to input to a function.

```
mdply(expand.grid(mean = 1:5, sd = 1:5), as.data.frame(rnorm), n = 10)
```

6 Related worked

There are a number of other approaches to solving the problems that plyr solves. You can always use loops, but loops create a lot of book-keeping code that obscures the intent of your algorithm. This section describes other high-level approaches similar to plyr.

Table 4 describes the equivalent between functions in base R and the functions provided by plyr. The built-in R functions focus mainly on arrays and lists, not data frames, and most attempt to return an atomic data structure if possible, and if not, a list. This ambiguity of the output type can make programming against these functions tricky. Compared to aaply, aaply returns the new dimensions first, rather than last, which means it is not idempotent when used with the identity function. For aaply, aaply(x, a, identity) == aperm(x, a) regardless of the value of a.

Base function	Input	Output	plyr function
aggregate	d	d	ddply + colwise
apply	a	a/l	$\mathtt{aaply} \; / \; \mathtt{alply}$
by	d	1	dlply
lapply	1	1	llply
mapply	a	a/l	$\mathtt{maply} \ / \ \mathtt{mlply}$
replicate	\mathbf{r}	a/l	$\mathtt{raply} \; / \; \mathtt{rlply}$
sapply	1	a	laply

Table 4: Mapping between apply functions and plyr functions.

Related functions tapply, are and sweep have no corresponding function in plyr, and remain useful. merge is useful for combining summaries with the original data. The cast function in the reshape package (Wickham, 2005) is closely related to aaply. There are also packages that attempt to simplify this class of problem:

The doBy (Højsgaard, 2008) package provides versions of order, sample, split, subset, summary and transform that make it easy to perform each of these operations on subsets of data frames, joining the results back into a data frame. These functions are rather like specialised version of ddply but they use a formula based interface, which, particularly for summary, makes it easy to only operate on selected columns.

- The gdata (Warnes and Gorjanc., 2008) package contains a bundle of helpful data manipulation functions, including frameApply which works like ddply or dlply depending on its arguments.
- The scope (Bergsma, 2007) package provides scope, scoop, skim, score and probe which provide a composable set of functions for operating symbolically on subsets of data frames.

You might also find Spector (2008) to be useful. It is an excellent introduction to many of the thorny problems encountered when manipulating data in R.

7 Future plans

The current major shortcoming of plyr is speed and memory usage. While theoretically it should be possible to do most of the operations without duplicating the input data set, currently plyr makes at least one copy. It is my aim to eventually implement these functions in C for maximum speed and memory efficiency, so that they are competitive with the built in operations.

I am also interested in connecting with papply (Currie, 2005) and related packages to make it easy to split up large tasks across multiple cores and multiple computers.

References

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