

A new Lorentzian Geometry

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Contents

- 1 The big picture: Regularity & Geometry
- 2 Lorentzian Geometry & General Relativity
- 3 Smooth Lorentzian Geometry is not good enough
- 4 Interlude: Back to the big picture
- 5 Sectional curvature:
From smooth Riemannian to synthetic Lorentzian
- 6 Ricci curvature:
From smooth Riemannian to synthetic Lorentzian
- 7 Summary & Outlook

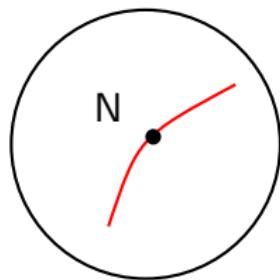
Table of Contents

- 1 The big picture: Regularity & Geometry
- 2 Lorentzian Geometry & General Relativity
- 3 Smooth Lorentzian Geometry is not good enough
- 4 Interlude: Back to the big picture
- 5 Sectional curvature:
From smooth Riemannian to synthetic Lorentzian
- 6 Ricci curvature:
From smooth Riemannian to synthetic Lorentzian
- 7 Summary & Outlook

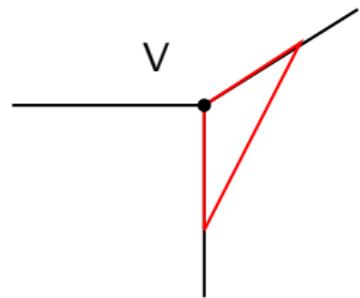
The larger picture: Regularity in geometry

Regularity determines basic geometric features

Example 1: Walking on a sphere vs. walking on a cube.

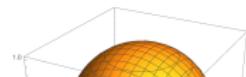


Sphere: Walking along a geodesic/meridian is locally always the shortest path.



Cube: It is always shorter to deviate to the right face than to go along the edge.

Example 2: Squeezing the sphere.



R. Steinbauer



A new Geometry

[Hartman-Wintner, 52]



Mathkoll, Nov. 2025

Table of Contents

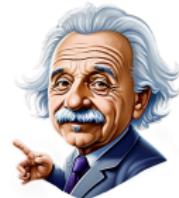
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- 2 Lorentzian Geometry & General Relativity
- 3 Smooth Lorentzian Geometry is not good enough
- 4 Interlude: Back to the big picture
- 5 Sectional curvature:
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- 6 Ricci curvature:
From smooth Riemannian to synthetic Lorentzian
- 7 Summary & Outlook

Lorentzian Geometry is the Language of General Relativity

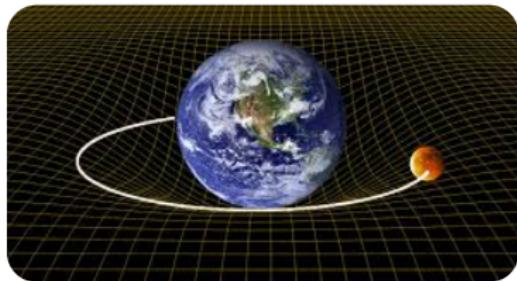
GR, Einstein's theory of space, time & gravity:

Gravity is universal, hence prop. of surrounding spacetime

Curvature proportional to **mass/energy content**



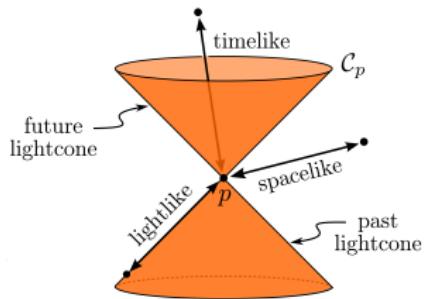
$$Ric - \frac{1}{2} R g = \frac{8\pi G}{c^4} T \quad (\text{E})$$



Matter tells spacetime how to curve,
spacetime tells matter how to move.

Geometric description via a spacetime manifold (M, g)

Causality is the fundamental structure of Lor. Geometry



$$v \in T_p M, \quad g_p(v, v) \begin{cases} < 0 & \text{timelike} \\ = 0 & \text{lightlike/null} \\ \leq 0 & \text{causal} \\ > 0 & \text{spacelike} \end{cases}$$

extends naturally to
(loc. Lipschitz) curves



Causal relations between points $p, q \in M$

$p \ll q \Leftrightarrow$ connected by future timelike curve

$p \leq q \Leftrightarrow$ connected by future causal curve

Chronological/causal future/past

$I^+(p) := \{q \in M : p \ll q\}$

$J^+(p) := \{q \in M : p \leq q\}$

past versions: $I^-(p)$, $J^-(p)$

for $A \subseteq M$: $I^\pm(A)$, $J^\pm(A)$

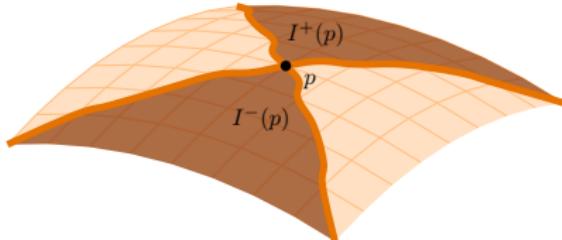


Table of Contents

- 1 The big picture: Regularity & Geometry
- 2 Lorentzian Geometry & General Relativity
- 3 Smooth Lorentzian Geometry is not good enough
- 4 Interlude: Back to the big picture
- 5 Sectional curvature:
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- 6 Ricci curvature:
From smooth Riemannian to synthetic Lorentzian
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Smooth Lorentzian Geometry is not enough

Generally from (E): $Ric - \frac{1}{2}Rg = \frac{8\pi G}{c^4} T$; $Ric \propto \partial^2 g + (\partial g)^2$;
 $T \notin C^0 \rightsquigarrow g \propto C^{1,1}$; below C^2 (smooth f.a.p.p.) non-smooth:=below C^2

Motivations from physics & mathematics

- *PDE/initial value* point-of-view (regularity in hyperbolic PDE!)
- physically relevant *models*: matched spacetimes
[Senovilla, Mars, Sánchez-Pérez, Manzano, Ohanyan, S.,...]
impulsive gravitational waves
[Barrabés, Griffiths, Hogan, Podolský, Sämann, Švarc, S.,...]
- approaches to *Quantum Gravity* (causal set theory, CDT,...)

"The discontinuities of the matter [...] which must be allowed [...] pose enormous problems [...]. To avoid this annoying problem though—despite it being completely fundamental!—we will implicitly assume that g is at least C^2 ."
[Garcia Parrado-Senovilla, CQG, 2005]

But much has happened since!

Table of Contents

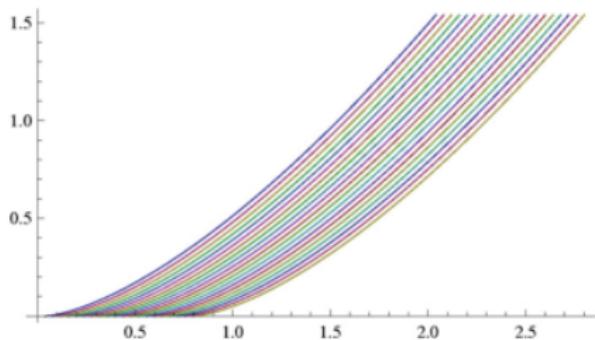
- 1 The big picture: Regularity & Geometry
- 2 Lorentzian Geometry & General Relativity
- 3 Smooth Lorentzian Geometry is not good enough
- 4 Interlude: Back to the big picture
- 5 Sectional curvature:
From smooth Riemannian to synthetic Lorentzian
- 6 Ricci curvature:
From smooth Riemannian to synthetic Lorentzian
- 7 Summary & Outlook

Back to larger picture: Regularity in Lorentzian causality

Regularity determines basic geometric features

Example 3: Lightcones bubble up.

[Chrūciel-Grant, 12]



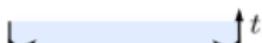
Some null geodesics for $\alpha = 1/3$

$$M = \mathbb{R}^2, g \in C^{0,\alpha} (\alpha < 1)$$

$$g = -(dt + (1 - |t|^\alpha)dx)^2 + dx^2$$

Non-unique null geodesics
~ null cone has full measure
causal bubble

Ex. 4: The future is not always open. [Grant-Kunzinger-Sämann-S, 20]



What can be done: Approaches to non-smooth LG

Spacetime setting: smooth manifold but metric g below C^2 (*analytic*)

- distributional curvature: $g \in H^1 \cap L^\infty$ maximal w stable \mathcal{D}' -curvature
[Geroch-Traschen, 87], [LeFloch-Mardare, 07], [Vickers-S, 09]
- non-linear distributional geometry (algebras of generalised functions)
[Farkas, Grosser, Hanel, Hörmann, Kunzinger, Mayerhofer, Nigsch, Oberguggenberger, Spreitzer, Vickers, S... 01–]
- advanced regularisation techniques $\check{g}_\varepsilon \prec g \prec \hat{g}_\varepsilon$ [Chrūciel-Grant, 12]
- causality for $g \in C^0$ [Sämann, 16], closed cone structures [Minguzzi, 19]
- ~ Singularity theorems in low regularity: $g \in C^{1,1}$, $g \in C^1$, $g \in C^{0,1}$
[Graf, Grant, Kunzinger, Ohayan, Schinnerl, Stojkovic, Vickers, S, 15–]
currently: $g \in W^{1,p}$, $R \in L^p$ [Kunzinger-Reintjes-Vega-S, 25–]
- C^0 -inextendibility [Sbierski, 15–, Ling, 24, Monsani-Solanki-Prados-Sämann, ong.]

A new geometry: beyond the manifold setting (*synthetic*)

This will be the focus for the rest of the talk.

Table of Contents

- 1 The big picture: Regularity & Geometry
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- 4 Interlude: Back to the big picture
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- 6 Ricci curvature:
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How to detect curvature: A glimpse of Riemannian mflds

Riemannian manifold (M, h) ; curvature $R_{XY}Z = [\nabla_X, \nabla_Y]Z - \nabla_{[X,Y]}Z$

$$\text{Sectional curvature } \text{Sec}(X, Y) = \frac{\langle R_{XY}Y, X \rangle}{\|X\|^2\|Y\|^2 - \langle X, Y \rangle^2}$$

Theorem (Toponogov Δ -comparison)

$$\text{Sec} \geq 0 \leq 0 \geq K \leq K \iff$$

For all (small) geodesic triangles $\triangle abc$ in (M, h) consider a comparison triangle $\triangle \bar{a}\bar{b}\bar{c}$ in the Euclidean Plane \mathbb{R}^2 D model space of constant curvature K .

Then for all p, q on its sides and corresponding comparison points \bar{p}, \bar{q}

$$d_h(p, q) \leq d_{\mathbb{R}^2}(\bar{p}, \bar{q}) \leq \bar{d}_K(\bar{p}, \bar{q}).$$

Sectional curvature bounds for metric spaces

Definition (Length space) (X, d) metric space w d intrinsic, i.e.

$$d(x, y) = \inf\{L_d(\gamma) : \gamma \text{ from } x \text{ to } y \text{ continuous}\}$$

Geodesics $\gamma : [0, 1] \rightarrow X$ with $L_d(\gamma) = d(\gamma(0), \gamma(1))$

Definition (Synthetic curvature bounds)

A length space has curvature bounded **below/above** by K if for all (small) *geodesic triangles* $\triangle abc$ and their *comparison triangles* $\triangle \bar{a}\bar{b}\bar{c}$ in the model space of curvature K and all points p, q on its sides and corresponding \bar{p}, \bar{q}

$$d(p, q) \leq \bar{d}_K(\bar{p}, \bar{q}).$$

Curvature bounded below / above: *Alexandrov spaces* / *CAT(K)-spaces*

Metric geometry: rich theory since the 1980-ies: GH-convergence, preservation of curvature bds., Gromov pre-compactness thm.

How to detect curvature: smooth Lorentzian world

$$\text{Sectional curvature } \text{Sec}(X, Y) = \frac{\langle R(X, Y)Y, X \rangle}{\langle X, X \rangle \langle Y, Y \rangle - \langle X, Y \rangle^2}$$

Kulkarni: If Sec is bounded below (above), then it is constant.

Definition (“Correct” curvature bounds) [Andersson-Howard, 98]

A smooth Lorentzian manifold has $\text{Sec} \geq K$ if

spacelike sectional curvatures $\geq K$ and *timelike* sectional curvatures $\leq K$.

How to go beyond Lorentzian manifolds?

	smooth	synthetic
pos. def.	Riemannian mflds	metric length spaces
Lorentzian	Lorentzian manifolds	??? Lor. length spaces

What is the analogue of metric (length) spaces in the *Lorentzian setting*?

Serious issue

No natural distance function \leadsto no direct entry into world of metric spaces

Lorentzian (pre-)length spaces [Kunzinger-Sämann, 18] based on causal spaces [Kronheimer-Penrose, 67] & the use of the **time separation**

Null distance [Sormani-Vega, 16] uses a time function to define distance function; contributions by [Allen, Burtscher, Galloway, Garcia-Heveling, Sakovich, Kunzinger, S...]

Lorentzian time separation vs. Riemannian distance

Length of causal curves $\gamma : I \rightarrow M$: $L_g(\gamma) := \int_I \sqrt{-g(\dot{\gamma}, \dot{\gamma})}$

Null curves have $L_g(\gamma) = 0$, timelike can be arbitrarily short!

Time separation function

$\tau(p, q) := \sup\{L_g(\gamma) : \gamma \text{ future causal from } p \text{ to } q\}$

Properties:

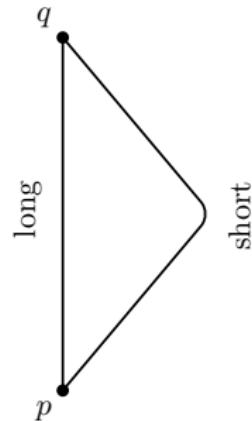
- lower semi-continuous,
- satisfies reverse Δ -inequality:

$$\tau(p, q) + \tau(q, r) \leq \tau(p, r)$$

- τ is not a distance function

Why: implements basic causality already of SRT (twin paradox)

Consequences: No cheap entry into the world of metric geometry!



Causality w/o metric: Lorentzian (pre-)length spaces

Causal space: X (metrizable) topological space with *abstract causality*:

\leq preorder on X , \ll transitive relation contained in \leq

Abstract time separation: $\tau: X \times X \rightarrow [0, \infty]$ lower semicontinuous

Definition

[Kunzinger-Sämann, 18]

(X, \ll, \leq, τ) is a *Lorentzian pre-length space* if for $p \leq q \leq r$

$$\tau(p, r) \geq \tau(p, q) + \tau(q, r) \quad \text{and} \quad \tau(p, q) \begin{cases} = 0 & \text{if } x \not\leq y \\ > 0 & \Leftrightarrow x \ll y \end{cases}$$

It is a *Lorentzian length space* if τ is intrinsic.

Examples

- *smooth spacetimes* (M, g) with usual time separation τ
- *Lorentz(-Finsler) spacetimes*, of *low regularity* ($g \in C^0 + \dots$)
- *finite directed graphs* (causal sets)

Lorentzian *causality theory*

variants by [Braun-McCann, 23-], [Minguzzi-Suhr, 24-], [Müller, 24-]

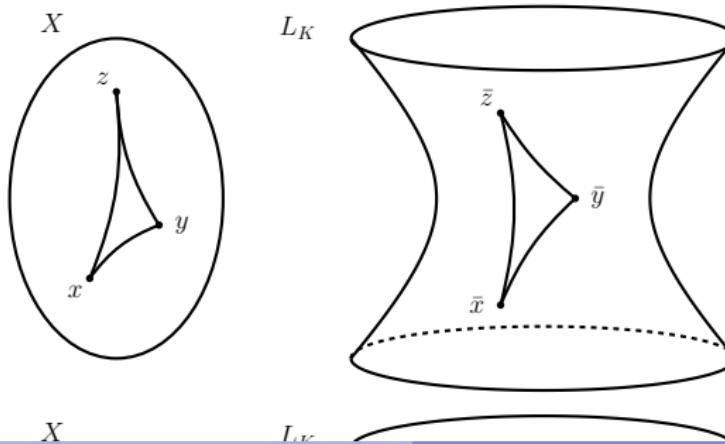
Timelike curvature via triangle comparison

Definition (Synthetic curvature bounds)

(X, \ll, \leq, τ) has *timelike curvature $\geq K \leq \bar{K}$* if

- ① some technical conditions hold
- ② for all *small timelike triangles* Δxyz and their comparison $\Delta \bar{x}\bar{y}\bar{z}$ in L_K and all p, q resp. \bar{p}, \bar{q}

$$\tau(p, q) \leq \bar{\tau}_K(\bar{p}, \bar{q}).$$



Faithful extension of
sectional curvature bounds

Timelike curvature comparison: The whole truth

Theorem (Equivalence of curvature conditions)

[Beran-Kunzinger-Rott, 24]

Let X be a chronological LpLS. Recall curvature bds. below/above by K in the following senses:

- 1 Timelike triangle comp.
- ii One-sided timelike
- iii Causal triangle co
- iv One-sided causal t
- v Strict causal trian
- vi One-sided strict ca
- vii Monotonicity com
- viii One-sided monoto

If X is strongly causal, r
bounds additionally is lo
(8), then *all notions of c*

- ix Angle comp.

In general the following relations hold:

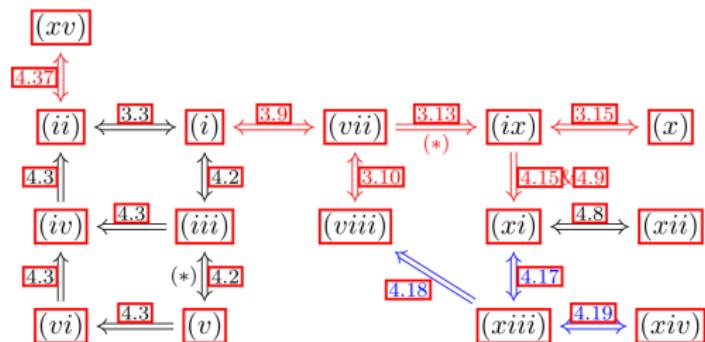


Figure 4: All relations between different formulations of curvature bounds for Lorentzian pre-length space. Black arrows are always valid, red arrows require X to be regular, and blue arrows require X to be strongly causal, regular and locally D_K -geodesic. The two instances of additional assumptions for one direction of curvature bounds are decorated by (*) in the figure.

Selected results (1/3)

Theorem

[Kunzinger-Sämann 19, Beran-Sämann, 22]

In a strongly causal Lorentzian pre-length space with *timelike curvature bounded below* timelike geodesics do *not branch*.

Theorem

[Grant-Kunzinger-Sämann, 19]

A timelike geodesically complete spacetime (or LLS) is *inextendible as a regular LLS*, i.e., any LLS-extension necessarily has unbounded curvature.

Complements class. [Beem-Ehrlich, 80s], C^0 -results [Galloway-Ling-Sbierski, 18] [Minguzzi-Suhr, 19] and relates to *curvature blow up*!

Splitting theorem

[Beran-Ohanyan-Rott-Solis, 23]

Let X be a globally hyperbolic LLS with global timelike $K \geq 0$. If X contains a complete timelike line (& some technical conditions) then it splits into a product $\mathbb{R} \times S$ with S a metric length space with $K \geq 0$.

Generalises smooth Lorentzian & synthetic Riemannian results.

Riemannian Comparison Theorem

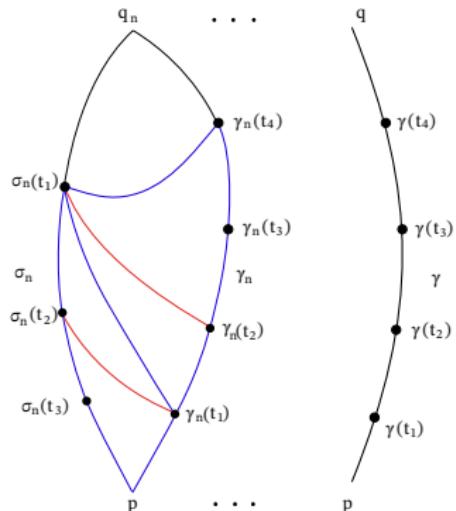
Classical

$$0 <$$

C^∞

tl. S

Thm



Let (X, \ll, \geq, τ) be a chronological, regular, local and finitely τ -measurable Lorentzian pre-length space with curvature $\leq K$ and let $\gamma \in TGeo(X)$ be such that $q := \gamma(1)$ are symmetric conjugate along γ . Then

Closely related to

Rauch comparison Theorem

[Grant-Kunzing]

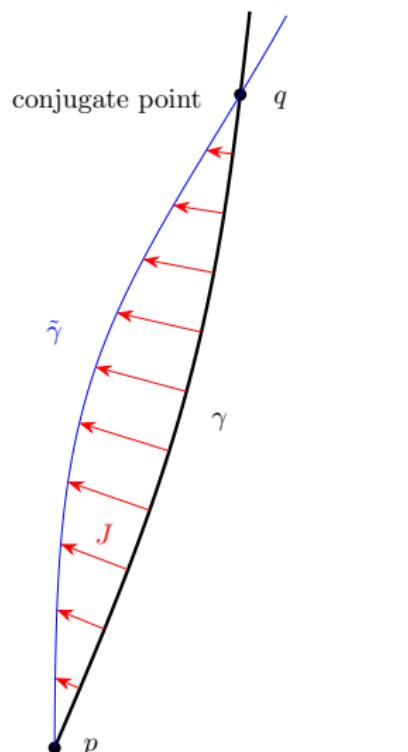
conjugate pt.

Synth. Riem.

$$\text{CAT}(K) \Rightarrow$$

Synthetic L

tl. Sec. $\leq K$



Selected results (3/3)

- *Time functions* [Burtscher-García-Heveling, 21]
- *Null distance & LLS* [Kunzinger-S., 22]
- compatibility w *distr. curv. bds* [Erös-Kunzinger-Ohanyan-Vardabasso, 25]
- Lorentzian Hausdorff *dimension, measure* [McCann-Sämann, 22]
- *Gluing* of Lorentzian length spaces [Beran-Rott, 24, Rott, 23]
- *Causal/ideal bdries* [Ake Hau-Burgos-Solis, 23–25], [Burgos-Che-Prados, ong.]
- *symp./cont. geo.* [Abbondandolo-Benedetti-Polterovich, 22, Hedicke, 24]
- *Sub-Lorentzian* Heisenberg group as LLS [Borza-Rigoni-Zoghłami, 25]

Lorentzian Gromov–Hausdorff convergence

Stability of lower timelike, Geometric precompactness [Mondino-Sämann, 25]
variants by [Minguzzi-Suhr, 22], [Bykov-Minguzzi-Suhr, 24], [Müller, 24–]
[Sakovich–Sormani, 24], [Che-Perales-Sormani, 25]
[Perales-Prados-Sormani-Zoglami, ong.]

- *Machine learning* in spacetimes [Law-Lucas, 23]

Table of Contents

- 1 The big picture: Regularity & Geometry
- 2 Lorentzian Geometry & General Relativity
- 3 Smooth Lorentzian Geometry is not good enough
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From smooth Riemannian to synthetic Lorentzian
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Ricci bounds via optimal transport: the basic idea

- *Optimal Transport*: [Monge, Kantorovich]

move matter (distribution μ_1) in the cheapest/optimal way (to μ_2)

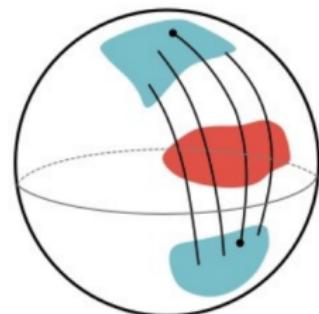
- *Minimize*

$$\int_{X \times Y} c(x, y) d\pi(x, y) \quad (c \dots \text{cost function})$$

over transport plans $\pi \in \mathcal{P}(X \times Y)$

w given marginals $(\text{pr}_X)_\sharp \pi = \mu_1, (\text{pr}_Y)_\sharp \pi = \mu_2$

- What is *optimal* depends on cost,
but also on *distances* & *geometry*!
- Turn this on its head:
define *curvature* by requiring that
OT behaves as in model spaces
 - Riemannian case: cost $c = d$
 - Lorentzian case: cost $c = \tau$



Transporting *clouds* of
points on the sphere

Ricci Bounds via Optimal Transport: Riemannian case

Thm. (Displacement convexity)

[Lott-Villani, Sturm 06-09]

(M, h) complete Riemannian manifold

$\text{Ric} \geq 0 \iff (M, d_h, \text{vol}_h) \text{ is an } \text{CD}(0, \infty)\text{-space}$

Definitions. On a metric measure space (X, d, \mathfrak{m}) define

- *Wasserstein distance:* $W_2(\mu_1, \mu_2) = \left(\inf_{\pi \in \Pi} \int_{X \times X} d(x, y)^2 d\pi(x, y) \right)^{\frac{1}{2}}$
- *Wasserstein geodesic:* continuous curve $(\mu_t)_{0 \leq t \leq 1}$

$$W_2(\mu_s, \mu_t) = |t - s| \cdot W_2(\mu_1, \mu_2)$$

- *Entropy functional:* $E(\mu|\mathfrak{m}) = \int_X \rho \log \rho d\mathfrak{m}$ for $\mu = \rho \mathfrak{m}$
- *$\text{CD}(0, \infty)$ -space:* E convex along Wasserstein geodesics, i.e.,

$$E(\mu_t|\mathfrak{m}) \leq (1 - t)E(\mu_1|\mathfrak{m}) + tE(\mu_2|\mathfrak{m})$$

Again turn this into definition of synthetic curvature bounds... $\text{CD}(K, N)$
~*(R)CD-spaces:* stability u. measured GH-conv. [Ambrosio-Gigli-Savare, 14]

Ricci Bounds via Optimal Transport: Lorentzian case

Thm. (Lor. displacement convexity) [McCann, Mondino-Suhr, 20]

(M, g) globally hyperbolic n -dim. spacetime

$\text{Ric}(X, X) \geq 0$ for X timelike $\iff (M, \tau, \text{vol}_g)$ is TCD($0, n$)-space

Definitions. Measured Lorentzian pre-length space $(X, \mathfrak{m}, \ll, \leq, \tau)$

- OT & causality [Eckstein-Miller, 17]: $\pi \in \Pi_{\leq}$ (concentrated on \ll)

- *p-Lorentz Wasserstein distance:* $(0 < p \leq 1)$ (for reverse \triangle -ineq.)

$$l_p(\mu_1, \mu_2) = \left(\sup_{\pi \in \Pi_{\leq}} \int_{X \times X} \tau(x, y)^p d\pi(x, y) \right)^{1/p}$$

- *Entropy functional:* $\text{Ent}(\mu | \mathfrak{m}) = \int \rho \log(\rho) d\mathfrak{m}$ for $\mu = \rho \mathfrak{m}$

- *TCD(K, N):* along l_p -geos μ_t we have for $e(t) := \text{Ent}(\mu_t | \mathfrak{m})$

$$e''(t) - 1/N e'(t)^2 \geq K \int_{X \times X} \tau(x, y)^2 \pi(dx dx)$$

Turn into definition: measured Lorentzian pre-length space w TCD(K, N)

Selected results

- *Ricci curvature bounds* via optimal transport on *null hypersurfaces* [McCann, 24, Ketterer, 24, Cavalletti-Manini-Mondino, 24]
- OT for general Lorentzian cost [Calisti-Ohanyan-Sálamo, 25]
- Compatibility with distributional bounds [Braun-Sálamo, ongoing]
- (vacuum) *Einstein equations* [Mondino-Suhr, 23]
- Hawking singularity theorem in TCD/TMCP [Cavalletti-Mondino 24]
- Penrose singularity theorem [Cavalletti-Manini-Mondino, 24]

First order Sobolev calculus on metric measure spacetimes

Infinitesimal Minkowskianity, nonlinear elliptic p -d'Alembertian comparison
[Beran-Braun-Calisti-Gigli-McCann-Ohanyan-Rott-Sämann, 24]

- Lorentzian splitting
 - ▶ Elliptic proof of class. result [Braun-Gigli-McCann-Ohanyan-Sämann, 24]
 - ▶ C^1 (weighted) spacetimes [Braun-Gigli-McCann-Ohanyan-Sämann, 25]

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Summary

A new Lorentzian geometry

(Measured) Lorentzian Length Spaces $(X, \mathfrak{m} \ll, \leq, \tau)$

- provide a general mathematical setting for
 - ▶ *sectional* curvature and
 - ▶ *Ricci* curvature (bounds)
- that contains
 - ▶ *low regularity spacetimes* but also
 - ▶ *discrete spaces*

Gives framework for

- approaches to non-smooth spacetime geometry & beyond
- fundamentally discrete approaches to QG
 - causal set theory, CDT, causal fermion systems, ...

Outlook: Causal set theory [Bombelli-Lee-Meyer-Sorkin, 87]

ingredients: causal set (X, \leq) , partial order

locally finite: $J(x, y) = \{z : x \leq z \leq y\}$ finite

CS hypothesis: QT of causal sets X ;

(M, g) approximation of X

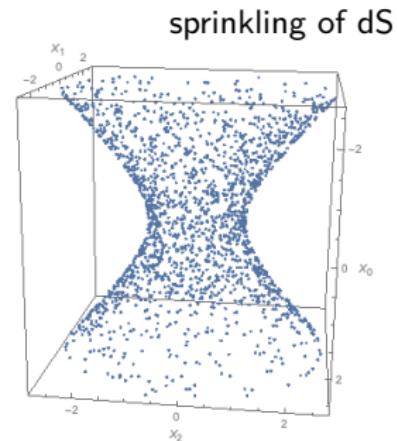
$$\mathcal{C}(M, \rho) \ni X \longleftrightarrow (M, g)$$

Hauptvermutung of CST

X can be embedded at density ρ_C into distinct spacetimes iff they are “close”.

(X, \ll, \leq, τ) is a LpLS

- chain: $C := (x_i)_{i=1}^n : x_i < x_{i+1}$
- length: $L(C) = n$; $\ll := <$
- $\tau(x, y) := \sup\{L(C) : C \text{ chain from } x \text{ to } y\}$



Hauptvermutung translates into statement on convergence of LpLS.

Olivier Ricci curvature and stability

[Barton-Borza-Röhrig, ongoing]

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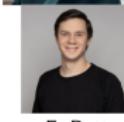
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Thank you for your attention

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- slide 3, Cartoon-Einstein: <https://wallpapers.com/albert-einstein-png> (freely downloadable)
- slide 3, Spacetime-curvature: Einstein relatively easy, <https://einsteinrelativelyeasy.com/>
- slide 4, lightcone : arXiv:1607.04202[gr-qc], Le Tiec, Alexandre and Novak, Jerome, Theory of Gravitational Waves. arXiv:1607.04202, <https://doi.org/10.48550/arXiv.1607.04202>
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