Simulating Turbulence in HII Regions

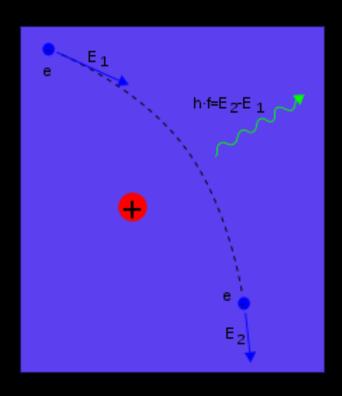
Introduction to HII Regions

- What does the HII mean?
- Physical traits
 - Powered by hot stars
 - Roughly 1 parsec in size
 - A type of nebulae

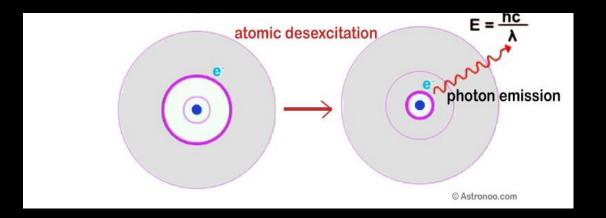


Image attribution: M20 | Trifid Nebula HII Region in Sagittarius 6° from Kaus Borealis (top of the teapot) taken by R Jay GaBany

More on HII Regions

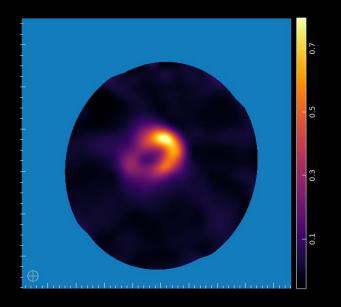


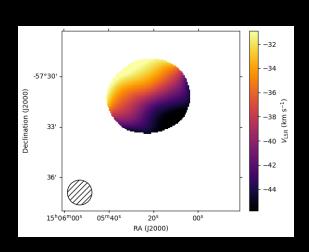
- How do we observe them?
 - o Free-free emission
 - Radio recombination lines (RRLs)

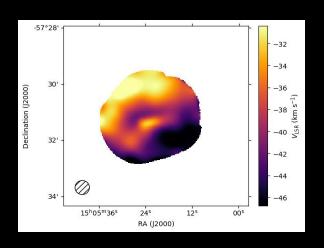


Motivation for Project

- Previous work suggested rotation of HII regions
- New data doesn't match prediction
- Could turbulence explain the differences?







Project Goals

- Simulate turbulence in HII regions
 - Calculating emission for each pixel
 - Overlaying density and velocities
- Test different turbulence parameters
 - Comparing to reality
 - o Use statistical models to compare



Image attribution: https://www.advancedsciencenews.com/wp-content/uploads/2023/07/swirl-g52ac5d4ac_1280.jpg

Exploring Turbulence

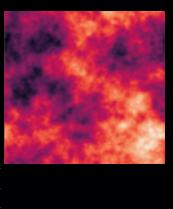
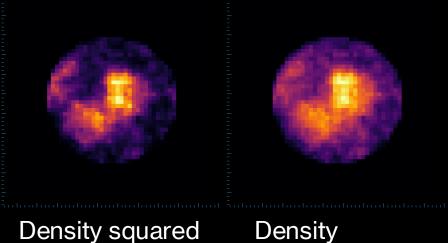


Image
attribution: https://turbustat.re
adthedocs.io/en/latest/genera
ting_test_data.html#threedimensional-fields



- Not all turbulence is made equal
- Using Turbustat
 - Python package
 - o Cubes of density and velocity
- Special case for RRLs

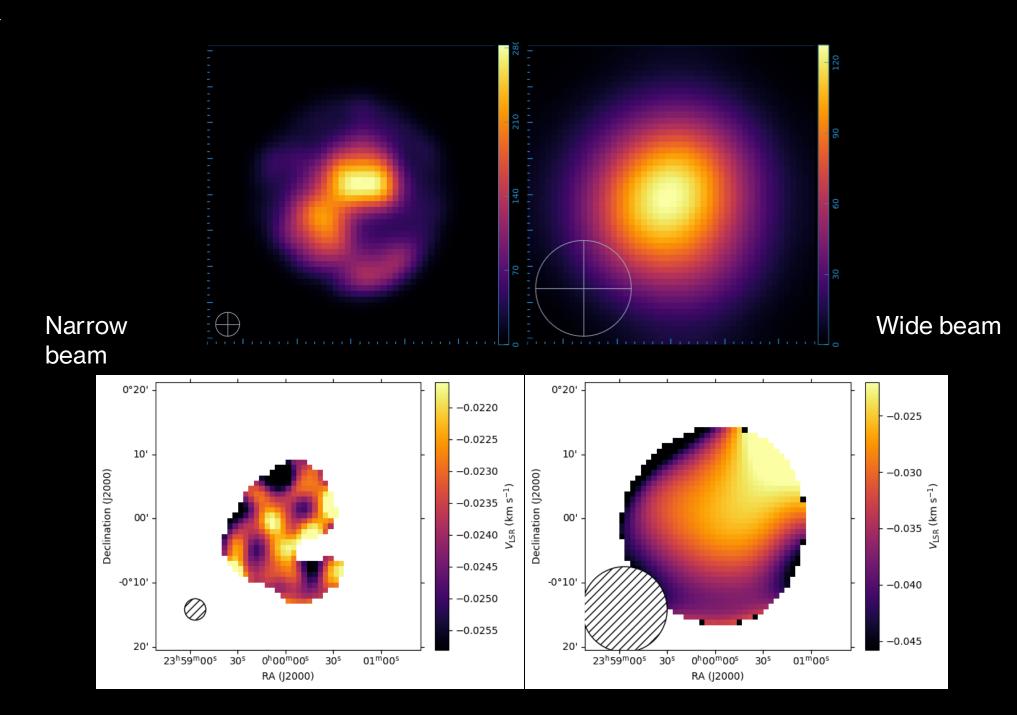
Work Done So Far

- Simulation in working order
 - Turbulence can be added to our region
 - Turbulence currently acting unexpectedly
- Simulation looks similar to data

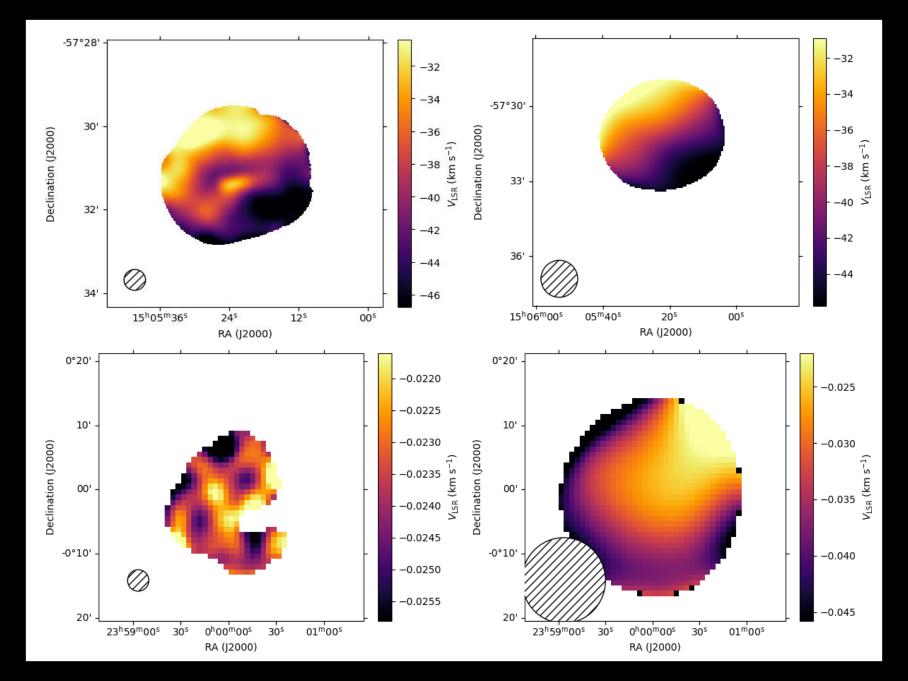
Work Done So Far

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But, what does it look like?

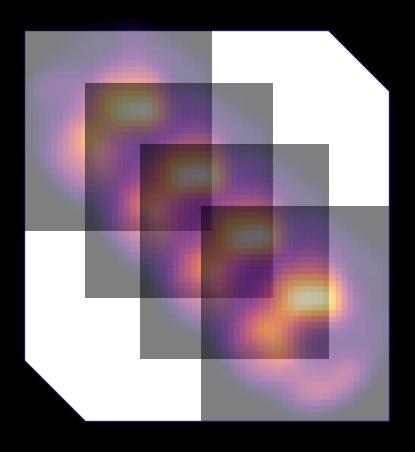






Wide beam

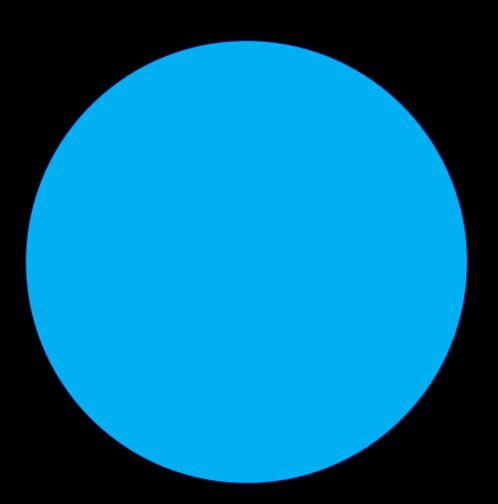
Future Work



- Comparing simulation with reality
 - Using new dataset to test
- Calculating different features of simulation
 - o Finding good test values
 - Making many data cubes

Significance

- We don't understand HII region motion
- Insight into high-mass star formation
- Transfer of energy into surrounding media



Takeaways

- Goal of project
- Progress
- Next steps