SHESHA Documentation

Release 602

COMPASS team

October 28, 2015

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Please, read the README file to install the Shesha module:

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1.1 Atmos

1.1.1 atmos module

direction

```
class atmos.Atmos
      Bases: object
      add_screen()
           Add a screen to the atmos object.
               Parameters size: (float): dimension of the screen (size x size)
                    amplitude: (float): frac
                    altitude: (float): altitude of the screen in meters
                    windspeed: (float): windspeed of the screen [m/s]
                    winddir: (float): wind direction (°)
                    deltax: (float): extrude deltax pixels in the x-direction at each iteration
                    deltay: (float): extrude deltay pixels in the y-direction at each iteration
                    device: (int): device number
      del_screen()
           Delete a screen from the atmos object
               Parameters alt – (float): altitude of the screen to delete
      disp()
           Display the screen phase at a given altitude
               Parameters alt - (float): altitude of the screen to display
      get_screen()
           Return a numpy array containing the turbulence at a given altitude
               Parameters alt – (float) :altitude of the screen to get
      list_alt()
           Display the list of the screens altitude
      move_atmos()
           Move the turbulence in the atmos screen following previous loaded paramters such as windspeed and wind
```

```
atmos.atmos_init()
Create and initialise an atmos object

Parameters c: (naga_context): context

tel: (Param_tel): telescope settings

geom: (Param_geom): geometry settings

loop: (Param_loop): loop settings

wfss: (list of Param_wfs): (optional) wfs settings

target: (Param_target): (optional) target_settings
```

1.2 Target

1.2.1 target module

```
class target. Target
     Bases: object
     Lambda
           observation wavelength for each target
     add_layer()
           Add a phase screen dm or atmos as layers of turbulence
               Parameters n: (int): index of the target l_type: (str): "atmos" or "dm"
                   alt: (float): altitude
                   xoff: (float): x-offset
                   yoff: (float): y-offset
     apod
           boolean for apodizer
     atmos_trace()
           Raytracing of the target through the atmosphere
               Parameters int: (nTarget): index of the target
                   atm: (atmos): atmos to get through
     dmtrace()
     get_amplipup()
           Return the complex amplitude in the pupil plane of the target.
               Parameters nTarget – (int): index of the target
     get_image()
           Return the image from the target (or long exposure image according to the requested type) :parameters:
               nTarget: (int): index of the target
               type_im: (str): type of the image to get ("se" or "le")
               puponly: (int): if 1, image computed from phase on the pupil only
```

```
Return the phase's screen of the target
               Parameters nTarget – (int): index of the target
     get_phasetele()
          Return the telemetry phase of the target
               Parameters nTarget – (int): index of the target
               Return data (np.ndarray(ndim=2,np.float32)): phase screen
     get_strehl()
          Return the target's strehl
               Parameters nTarget – (int): index of the target
     init strehlmeter()
          Initialise target's strehl
               Parameters nTarget – (int): index of the target
     mag
          magnitude for each target
     ntargets
          number of targets
     xpos
          x positions on sky (in arcsec) for each target
     ypos
          y positions on sky (in arcsec) for each target
target.target_init()
     Create a cython target from parametres structures
          Parameters ctxt: (naga_context):
               atm: (Param_atmos): atmos settings
               geom: (Param_geom): geom settings
               wfs: (Param_wfs): wfs settings
               dm: (Param_dm) : dm settings
1.3 Sensors
1.3.1 sensors module
class sensors. Sensors
     Bases: object
     Constructor: Sensors(nsensors,type_data,npup,nxsub,nvalid,nphase,pdiam,npix,nrebin,nfft,nftota,nphot,lgs,odevice,comm_size,ra
          Parameters nsensors: (int):
               type_data: list of strings):
               npup: (np.ndarray[ndim=1,dtype=np.int64_t]):
```

get_phase()

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nxsub: (np.ndarray[ndim=1,dtype=np.int64_t]) :

```
nvalid: (np.ndarray[ndim=1,dtype=np.int64_t]):
         nphase: (np.ndarray[ndim=1,dtype=np.int64_t]) :
         pdiam: (np.ndarray[ndim=1,dtype=np.float32_t):
         npix: (np.ndarray[ndim=1,dtype=np.int64_t]) :
         nrebin: (np.ndarray[ndim=1,dtype=np.int64 t]):
         nfft: (np.ndarray[ndim=1,dtype=np.int64 t]):
         ntota: (np.ndarray[ndim=1,dtype=np.int64_t]) :
         nphot: (np.ndarray[ndim=1,dtype=np.float32_t]) :
         lgs: (np.ndarray[ndim=1,dtype=np.int32_t]):
         odevice: (int):
         comm_size: (int): MPI communicator size
         rank: (int): process rank
get_amplifoc()
     Return the 'amplifoc' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'amplifoc' from
get bincube()
     Return the 'bincube' array of a given wfs
         Parameters \mathbf{n} – (int): number of the wfs to get the 'bincube' from
get_binimg()
     Return the 'binimg' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'binimg' from
get_camplipup()
     Return the 'camplipup' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'camplipup' from
get imgtele()
     Return the 'image_telemetry' array of a given wfs
         Parameters n – (int): number of the wfs to get the 'image_telemetry' from
get_offsets()
     Return the 'offset' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'offset' from
get_phase()
     Return the phase array of a given wfs
         Parameters \mathbf{n} – (int): number of the wfs to get the phase from
get_pyrimg()
     Return the image of a pyr wfs
         Parameters n - (int): number of the wfs to get the image from
get_rank()
     Return the rank of one of the sensors wfs
         Parameters n - (int): index of the wfs to get the rank for
```

```
get_slopes()
     Return the 'slopes' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'slopes' from
sensors addlayer()
     Call function add layer from the sutra source of a sutra wfs of the Sensors
         Parameters i: (int):
             type_dm: (string):
             alt: (float):
             xoff: (float):
             yoff: (float):
sensors_compimg()
     TODO doc
         Parameters n - (in): index of the wfs
sensors initarr()
     Call the function wfs_initarrays from a sutra_wfs of the Sensors
         Parameters
     n: (int): index of the wfs
     wfs: (Param_wfs):
     geom: (Param_geom):
sensors_initgs()
     Call the function sensors_initgs
         Parameters xpos: (np.ndarray[ndim=1,dtype=np.float32_t]):
             ypos: (np.ndarray[ndim=1,dtype=np.float32_t]) :
             Lambda: (np.ndarray[ndim=1,dtype=np.float32_t]):
             mag: (np.ndarray[ndim=1,dtype=np.float32_t]):
             zerop: (float):
             size: (np.ndarray[ndim=1,dtype=np.int64_t]):
             noise: (np.ndarray[ndim=1,dtype=np.float32_t]):
             seed: (np.ndarray[ndim=1,dtype=np.int64 t]):
sensors_trace()
     Does the raytracing for the wfs phase screen in sutra_wfs
         Parameters n: (int):
             type_trace: (str) ["all"][raytracing across atmos and dms seen] "dm": raytracing across
               dms seen only "atmos": raytracing across atmos only
             atmos: (Atmos) :(optional) Atmos object
             dms: (Dms): (optional) Dms object
             rst: (int): (optional) reset before raytracing if rst = 1
slopes_geom()
     Compute the geometric slopes in a sutra wfs object
```

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```
Parameters nsensor: (int): wfs number
                       param t (int): method (0 or 1)
sensors.bin2d()
      Returns the input 2D array "array", binned with the binning factor "binfact". The input array X and/or Y
      dimensions needs not to be a multiple of "binfact"; The final/edge pixels are in effect replicated if needed. This
      routine prepares the parameters and calls the C routine bin2d. The input array can be of type long, float or
      double. Last modified: Dec 15, 2003. Author: F.Rigaut SEE ALSO: _bin2d
           Parmeters data_in: (np.ndarray): data to binned
                binfact: (int): binning factor
sensors.fft_goodsize()
      find best size for a fft from size s
           Parameters s: (long) size
sensors.init_wfs_geom()
      Compute the geometry of WFSs: valid subaps, positions of the subaps, flux per subap, etc...
           Parameters wfs: (Param_wfs) : wfs settings
                wfs0: (Param wfs): reference wfs settings
                n: (int): index of the wfs (diplay information purpose only)
                atmos: (Param atmos): atmos settings
                tel: (Param_tel) : telescope settings
                geom: (Param_geom): geom settings
                target: (Param_target): target settings
                loop: (Param_loop): loop settings
                init: (int): (optional)
                verbose: (int): (optional) display informations if 0
sensors.make_lgs_prof1d()
      same as prep_lgs_prof but cpu only. original routine from rico
           Parameters p_tel: (Param_tel) : telescope settings
                prof: (np.ndarray[dtype=np.float32]): Na profile intensity, in arbitrary units
                h: (np.ndarray[dtype=np.float32]): altitude, in meters. h MUST be an array with EQUALLY
                spaced elements.
                beam: (float): size in arcsec of the laser beam
                center: (string): either "image" or "fourier" depending on where the centre should be.
sensors.noise_cov()
                                                                                                       WFS
      Compute
                  the
                          diagonal
                                            the
                                                   noise
                                                           covariance
                                                                          matrix
                                                                                                SH
                                                                                                               (arc-
                                                                                    for
                                      (pi^2/2)*(1/Nphotons)*(d/r0)^2 /
                                                                           (2*pi*d/lambda)<sup>2</sup>
      sec<sup>2</sup>)
              Photon
                         noise:
                                                                                                Electronic
                                                                                                               noise:
      (pi<sup>2</sup>/3)*(wfs.noise<sup>2</sup>/N<sup>2</sup>photons)*wfs.npix<sup>2</sup>*(wfs.npix*wfs.pixsize*d/lambda)<sup>2</sup> / (2*pi*d/lambda)<sup>2</sup>
           Parameters nw: wfs number p_wfs: (Param_wfs) : wfs settings p_atmos: (Param_atmos) : atmos
                settings p_tel: (Param_tel): telescope settings
           Returns cov: (np.ndarray(ndim=1,dtype=np.float64)): noise covariance diagonal
```

```
sensors.prep_lgs_prof()
```

The function returns an image array(double,n,n) of a laser beacon elongated by perpective effect. It is obtaind by convolution of a gaussian of width "lgsWidth" arcseconds, with the line of the sodium profile "prof". The altitude of the profile is the array "h".

```
parameters nsensors: (int) : wfs index
    p_tel: (Param_tel) : telescope settings
    prof: (np.ndarray[dtype=np.float32]) : Na profile intensity, in arbitrary units
    h: (np.ndarray[dtype=np.float32]) : altitude, in meters. h MUST be an array with EQUALLY spaced elements.
    beam: (float) : size in arcsec of the laser beam
    center: (string) : either "image" or "fourier" depending on where the centre should be.
```

Computation of LGS spot from the sodium profile: Everything is done here in 1D, because the Na profile is the result of the convolution of a function $P(x,y) = \operatorname{profile}(x)$. dirac(y) by a gaussian function, for which variables x and y can be split: $\exp(-(x^2+y^2)/2.s^2) = \exp(-x^2/2.s^2) * \exp(-y^2/2.s^2)$ The convolution is (symbol \$ denotes integral) $C(X,Y) = \$ \exp(-x^2/2.s^2) * \exp(-y^2/2.s^2) * \operatorname{profile}(x-X) * \operatorname{dirac}(y-Y) \, dx \, dy$ First one performs the integration along y $C(X,Y) = \exp(-Y^2/2.s^2) * \exp(-x^2/2.s^2) * \operatorname{profile}(x-X) \, dx$ which shows that the profile can be computed by - convolving the 1-D profile - multiplying it in the 2nd dimension by a gaussian function

If one has to undersample the inital profile, then some structures may be "lost". In this case, it's better to try to "save" those structures by re-sampling the integral of the profile, and then derivating it afterwards. Now, if the initial profile is a coarse one, and that one has to oversample it, then a simple re-sampling of the profile is adequate.

```
sensors.type_present()
      Check the present types in a list
           Parameters liste: (list of str): list of types
                pyr: (int): set to 1 if the list contains "pyr" (0 else)
                roof: (int): set to 1 if the list contains "roof" (0 else)
                sh: (int): set to 1 if the list contains "sh" (0 else)
                geo: (int): set to 1 if the list contains "geo" (0 else)
      return 1 if the wfs type is present (0 else)
sensors.wfs init()
      Create and initialise a Sensors object
           Parameters wfs: (list of Param wfs): wfs settings
                p atmos: (Param atmos): atmos settings
                p_tel: (Param_tel) : telescope settings
                p_geom: (Param_geom): geom settings
                p_target: (Param_target) : target settings
                p_loop: (Param_loop) : loop settings
                comm_size: (int): communicator size
                rank: (int): process rank
                dm: (list of Param dm): (optional) dms settings
```

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```
sensors.wheremax()
     return the index of the maximum value of the list
           Parameters liste – (list of values): values to get the index of the maximum from
1.4 RTC
1.4.1 rtc module
class rtc.Rtc
     Bases: object
     add Controller()
           Add a controller in the sutra controller vector of the RTC on the GPU
               Parameters nactu: (int): number of actuators
                   delay: (float): loop delay
                   type_control: (str): controller's type
                   dms: (Dms): sutra_dms object (GPU)
                   type_dmseen: (char**): dms indices controlled by the controller
                   alt: (np.ndarray[ndim=1,dtype=np.float32_t]): altitudes of the dms seen
                   ndm: (int): number of dms controled
                   Nphi: (long): number of pixels in the pupil (used in geo controler case only)
     add_centroider()
           Add a centroider in the sutra centroiders vector of the RTC on the GPU
               Parameters sensor: (Sensors) : sutra_sensors object (GPU)
                   nwfs: (long): number of wfs
                   nvalid: (long): number of valid subaps
                   type_centro: (str): centroider's type
                   offset: (float):
                   scale: (float):
     applycontrol()
           Compute the DMs shapes from the commands computed in a sutra_controller_object. From the command
           vector, it computes the voltage command (adding pertrubation voltages, taking delay into account) and
           then apply it to the dms
               Parameters ncontro: (int): controller index
```

buildcmat()

Compute the command matrix in a sutra controller ls object

```
Parameters ncontro: (int): controller index nfilt: (int): number of modes to filter filt_tt: (int): (optional) flag to filter TT
```

buildcmatmv()

Compute the command matrix in a sutra_controller_mv object

```
Parameters ncontro: (int): controller index
             cond: (float): conditioning factor for the Cmm inversion
docentroids()
     Compute the centroids with sutra_controller #ncontrol object
         Parameters ncontrol: (optional) controller's index
docentroids geom()
     Compute the geometric centroids with sutra_controller #ncontrol object
         Parameters ncontrol: (optional) controller's index
docontrol()
     Compute the command to apply on the DMs on a sutra_controller object
         Parameters ncontro: (int): controller index
docontrol_geo()
     Compute the command to apply on the DMs on a sutra_controller_geo object
         Parameters ncontro: (int): controller index
doimat()
     Compute the interaction matrix
         Parameters ncontro: (int): controller index
             g dms: (Dms): Dms object
doimat geom()
     Compute the interaction matrix by using a geometric centroiding method
         Parameters ncontro: (int): controller index
             g dms: (Dms): Dms object
             geom: (int): type of geometric method (0 or 1)
getCenbuff()
     Return the centroids buffer from a sutra_controller_ls object. This buffer contains centroids from iteration
     i-delay to current iteration :parameters:
         ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=2,dtype=np.float32_t]): centroids buffer
getCentroids()
     Return the centroids computed by the sutra rtc object
         Parameters ncontrol: (int): controller's index
         Returns data : (np.ndarray[ndim=1,dtype=np.float32_t]) : centroids (arcsec)
getCmmEigenvals()
     Return the eigen values of the Cmm decomposition in a sutra_controller_mv object :parameters:
         ncontro: (int): controller index
         Returns eigenvals: (np.ndarray[ndim=1,dtype=np.float32_t]): eigenvalues
getCom()
     Return the command vector from a sutra controller object
```

```
Parameters ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=1,dtype=np.float32_t]): command vector
getEigenvals()
     Return the eigen values of the imat decomposition in a sutra controller object
         Parameters ncontro: (int): controller index
         Returns eigenvals: (np.ndarray[ndim=1,dtype=np.float32 t]): eigenvalues
getErr()
     Return the command increment (cmat*slopes) from a sutra_controller_ls object
         Parameters ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=1,dtype=np.float32_t]): command increment
getU()
     Return the eigen modes matrix of the imat decomposition from a sutra_controller_ls object
         Parameters ncontro: (int): controller index
         Returns U: (np.ndarray[ndim=2,dtype=np.float32 t]): eigen modes matrix
getVoltage()
     Return the voltage vector that will be effectively applied to the DMs
         Parameters ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=1,dtype=np.float32_t]): voltage vector
get cmat()
     Return the command matrix from a sutra_controller object
         Parameters ncontro: (int): controller index
         Returns cmat: (np.ndarray[ndim=2,dtype=np.float32_t]): command matrix
get_cmm()
     Return the Cmm matrix from a sutra_controller_mv object
         Parameters ncontro: (int): controller index
         Returns cmm: (np.ndarray[ndim=2,dtype=np.float32 t]): Cmm matrix
get cphim()
     Return the Cphim matrix from a sutra_controller_mv object
     :parameters; ncontro: (int): controller index
         Returns cphim: (np.ndarray[ndim=2,dtype=np.float32_t]): Cphim matrix
get_imat()
     Return the interaction matrix of a sutra_controller object
         Parameters ncontro: (int): controller index
         Returns imat: (np.ndarray[ndim=2,dtype=np.float32_t]): interaction matrix
get_mgain()
     Return modal gains from sutra_controller
         Parameters ncontro: (int): controller index
         Returns mgain: (np.ndarray[ndim=1,dtype=np.float32_t]): modal gains
```

```
getcentroids()
     Return the centroids computed by the sutra rtc object If ncontrol <= d control.size, return rtc.d centroids
     Else, compute centroids from wfs[nwfs] with centroider[ncontrol]
         Parameters ncontrol: (int): controller's index
             g wfs: (Sensors): (optional) sutra sensors object
             nwfs: (int): (optional) number of wfs
         Returns data: (np.ndarray[ndim=1,dtype=np.float32_t]): centroids (arcsec)
getolmeas()
     Return the reconstructed open-loop measurement from a sutra_controller_mv object
         Parameters ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=1,dtype=np.float32_t]): reconstructed open-loop
imat_svd()
     Compute the singular value decomposition of the interaction matrix
         Parameters ncontro - controller index
init modalOpti()
     Initialize the modal optimization controller: compute the slopes-to-modes matrix and the transfer func-
     tions
         Parameters ncontro: (int): controller index
             nmodes: (int): number of modes
             nrec: (int): number of recorded open slopes measurements
             M2V: (np.ndarray[ndim=2,dtype=np.float32_t]): modes-to-volt matrix
             gmin: (float): minimum gain for modal optimization
             gmax: (float): maximum gain for modal optimization
             ngain: (int): Number of tested gains
             Fs: (float): sampling frequency
init proj()
     Initialize the projection matrix for sutra controller geo object. The projection matrix is (IFt.IF)**(-1) *
     IFt where IF is the DMs influence functions matrix
         Parameters ncontro: (int): controller index
             dms: (Dms): Dms object
             indx dm: (np.ndarray[ndim=1,dtype=np.int32 t]): indices of where(pup) on DM screen
```

unitpervolt: (np.ndarray[ndim=1,dtype=np.float32_t]): unitpervolt DM parameter

indx_pup: (np.ndarray[ndim=1,dtype=np.int32_t]): indices of where(pup) on ipupil screen

loadOpenLoop()

Load an array of recoded open-loop measurements for modal optimization

Parameters ncontro: (int): controller index

ol_slopes: (np.ndarray[ndim=2, dtype=np.float32_t]): open-loop slopes

loadnoisemat()

Load the noise vector on a sutra controller my object

```
Parameters ncontro: (int): controller index
             N: (np.ndarray[ndim=1,dtype=np.float32_t]): noise vector
modalControlOptimization()
     Compute the command matrix with modal control optimization
         Parameter ncontro: controller index
rmcontrol()
     Remove a controller
sensors_compslopes()
     Compute the slopes in a sutra_wfs object. This function is equivalent to docentroids() but the centroids are
     stored in the sutra_wfs object instead of the sutra_rtc object
         Parameters ncentro: (int): centroider index
sensors_initbcube()
     Initialize npix in the sutra_centroider_corr object (useless ?)
         Parameters ncentro: (int): centroider's index
sensors initcorr()
     Initialize sutra_centroider_corr oblect
         Parameters ncentro: (int): centroider's index
             w: (np.ndarray[ndim=1,dtype=np.float32 t]): weight
             corr_norm: (np.ndarray[ndim=2,dtype=np.float32_t]) :
             sizex: (int):
             sizey: (int):
             interpmat: ([ndim=2,dtype=np.float32_t]):
sensors_initweights()
     Load the weight array in sutra_centroider_wcog object
         Parameters ncentro: (int): centroider's index
             w: (np.ndarray[ndim=2, dtype=np.float32_t]): weight
setCom()
     Set the command vector of a sutra controller object to comvec
         Parameters ncontro: (int): controller index
setEigenvals()
     Set the eigen values of the imat decomposition in a sutra controller ls object
         Parameters ncontro: (int): controller index
             eigenvals: (np.ndarray[ndim=1,dtype=np.float32_t]): eigen values
setU()
     Set the eigen modes matrix of the imat decomposition in a sutra_controller_ls object
         Parameters ncontro: (int): controller index
             U: (np.ndarray[ndim=2,dtype=np.float32_t]): eigen modes matrix
set_cmat()
     Set the command matrix on a sutra controller object
```

```
Parameters ncontro: (int): controller index
                   data: (np.ndarray[ndim=2,dtype=np.float32_t]): command matrix to use
     set_cmm()
           Set the Cmm matrix on a sutra controller my object
               Parameters ncontro: (int): controller index
                   data: (np.ndarray[ndim=2,dtype=np.float32 t]): Cmm matrix
     set_decayFactor()
           Set the decay factor on a sutra_controller_generic object
               Parameters ncontro: (int): controller index
                   decay: (np.ndarray[ndim=1,dtype=np.float32_t]): ask to Rico
     set_gain()
           Set the loop gain in sutra_controller object
               Parameters ncontro: (int): controller index
                   gain: (float): loop gain
     set imat()
           Set the interaction matrix on a sutra_controller object
               Parameters ncontro: (int): controller index
                   data: (np.ndarray[ndim=2,dtype=np.float32_t]): interaction matrix to use
     set matE()
           Set the matrix E on a sutra_controller_generic object
               Parameters ncontro: (int): controller index
                   matE: (np.ndarray[ndim=2,dtype=np.float32_t]): ask to Rico
     set_mgain()
           Set modal gains in sutra_controller object
               Parameters ncontro: (int): controller index
                   mgain: (np.ndarray[ndim=1,dtype=np.float32 t]): modal gains
     setnmax()
           set the number of brightest pixels to consider for bpcog centroider
               Parameters ncentro: (int): centroider's index
                   nmax: (int): number of brightest pixels
     setthresh()
           set threshold for the centroider #ncentro
               Parameters ncentro: (int): centroider's index
                   thresh: (float): threshold
rtc.cmat_init()
     Compute the command matrix on the GPU
           Parameters ncontro: (int): g_rtc: (Rtc): p_rtc: (Param_rtc): rtc settings p_wfs: (list of
               Param_wfs): wfs settings p_tel: (Param_tel): telescope settings clean: (int): (optional) clean
               datafiles (imat, U, eigenv) simul name: (str): (optional) simulation's name, use for data files'
               path
```

```
rtc.compute_KL2V()
     Compute the Karhunen-Loeve to Volt matrix (transfer matrix between the KL space and volt space for a pzt dm)
          Parameters p_dms: (list of Param_dm): dms settings
              controller: (Param_controller): controller settings
rtc.correct dm()
     Correct the geometry of the DMs using the imat (filter unseen actuators)
          Parameters p_dms: (list of Param_dm): dms settings
              g_dms: (Dms): Dms object
              p_control: (Param_controller) : controller settings
              p_geom: (Param_geom) : geom settings
              imat: (np.ndarray): interaction matrix
              simul_name: (str): simulation's name, use for data files' path
rtc.create_interp_mat()
     TODO doc
          Parameters dimx: (int):
              dimy: (int):
rtc.create nact geom()
     Compute the DM coupling matrix :param:
          p_dms: (list of Param_dm): dms parameters ndm: (int): dm number
          Returns Nact: (np.array(dtype=np.float64)): the DM coupling matrix
rtc.create_piston_filter()
rtc.doTomoMatrices()
     Compute Cmm and Cphim matrices for the MV controller on GPU
          Parameters g_wfs: (Sensors):
              p wfs: (list of Param wfs): wfs settings
              g_dms: (Dms):
              p_dms: (list of Param_dms): dms settings
              p_geom: (Param_geom) : geom settings
              p_atmos: (Param_atmos): atmos settings
              g_atmos: (Atmos):
              p_tel: (Param_tel) : telescope settings
rtc.get_r0()
     Compute r0 at lambda2 from r0 value at lambda1
          Parameters r0_at_lambda1: (float): r0 value at lambda1
              lambda1: (float): lambda1
              lambda2: (float): lambda2
rtc.imat_geom()
     Compute the interaction matrix with a geometric method
```

```
Parameters g_wfs: (Sensors): Sensors object
               p_wfs: (list of Param_wfs): wfs settings
               p_control: (Param_controller) : controller settings
               g_dms: (Dms): Dms object
               p dms: (list of Param dm): dms settings
               meth: (int): (optional) method type (0 or 1)
rtc.imat_init()
      Initialize and compute the interaction matrix on the GPU
           Parameters ncontro: (int): controller's index
               g_rtc: (Rtc) : Rtc object
               p_rtc: (Param_rtc): rtc settings
               g_dms: (Dms): Dms object
               g wfs: (Sensors): Sensors object
               p_wfs: (list of Param_wfs) : wfs settings
               p_tel: (Param_tel) : telescope settings
               clean: (int): (optional): clean datafiles (imat, U, eigenv)
               simul name: (str): (optional) simulation's name, use for data files' path
rtc.manual imat()
      Compute the interaction matrix 'manually', ie without sutra_rtc doimat method
           Parameters g_rtc: (Rtc): Rtc object
               g_wfs: (Sensors): Sensors object
               g_dms: (Dms): Dms object
               p_dms: (list of Param_dm): dm settings
rtc.openLoopSlp()
      Return a set of recorded open-loop slopes, usefull for modal control optimization
           Parameters g_atm: (Atmos): Atmos object
               g_rtc: (Rtc): Rtc object
               nrec: (int): number of samples to record
               ncontro: (int): controller's index
               g wfs: (Sensors): Sensors object
               p_wfs: (list of Param_wfs) : wfs settings
               p_tar: (Param_target) : target settings
               g_tar: (Target) : Target object
rtc.rtc init()
      Initialize all the sutra_rtc objects: centroiders and controllers
           Parameters g_wfs: (Sensors): Sensors object
               p_wfs: (list of Param_wfs): wfs settings
```

```
g_dms: (Dms): Dms object
               p_dms: (list of Param_dms) : dms settings
               p_geom: (Param_geom) : geom settings
               p_atmos: (Param_atmos): atmos settings
               g atmos: (Atmos): Atmos object
               p_tel: (Param_tel) : telescope settings
               p_loop: (Param_loop) : loop settings
               p_tar: (Param_target) : (optional) target settings
               clean: (int): (optional) clean datafiles (imat, U, eigenv)
               brama: (int): (optional) not implemented yet
               doimat: (int): (optional) force imat computation
               simul_name: (str): (optional) simulation's name, use for path to save data (imat, U...)
           Returns Rtc: (Rtc): Rtc object
rtc.selectDMforLayers()
     For each atmos layer, select the DM which have to handle it in the Cphim computation for MV controller
     :parameters:
           ncontro: (int): controller number p atmos: (Param atmos): atmos parameters p rtc: (Param rtc)
           : rtc parameters p_dms :(list of Param_dm) : dms parameters
           Returns indlayersDM: (np.array(dtype=np.int32)): for each atmos layer, the Dm number corre-
               sponding to it
```

1.5.1 param module

```
class param.Param_atmos
Bases: object

LO

LO per layers in meters.

alt

altitudes of each layer.

deltax

x translation speed (in pix / iteration) for each layer.

deltay

y translation speed (in pix / iteration) for each layer.

dim_screens

linear size of phase screens.

frac

fraction of rO for each layer.
```

```
nscreens
     number of turbulent layers.
pupixsize
     pupil pixel size (in meters).
r0
     global r0.
seeds
set_L0()
     Set the L0 per layers
         Parameters 1 – (lit of float) : L0 for each layers
set_alt()
     Set the altitudes of each layer
         Parameters 1 – (lit of float) : altitudes
set deltax()
     Set the translation speed on axis x for each layer
         Parameters 1 – (lit of float) : translation speed
set deltay()
     Set the translation speed on axis y for each layer
         Parameters 1 – (lit of float) : translation speed
set_dim_screens()
     Set the size of the phase screens
         Parameters 1 – (lit of float): phase screens sizes
set_frac()
     Set the fraction of r0 for each layers
         Parameters 1 – (lit of float): fraction of r0
set_nscreens()
     Set the number of turbulent layers
         Parameters \mathbf{n} – (long) number of screens.
set_pupixsize()
     Set the pupil pixel size
         Parameters xsize – (float): pupil pixel size
set_r0()
     Set the global r0
         Parameters r – (float) : global r0
set_seeds()
     Set the seed for each layer
         Parameters 1 – (lit of float) : seed
set_winddir()
     Set the wind direction for each layer
         Parameters 1 - (\text{lit of float}): wind directions
```

```
set windspeed()
           Set the the wind speed for each layer
               Parameters 1 – (lit of float) : wind speeds
     winddir
           wind directions of each layer.
     windspeed
           wind speeds of each layer.
class param.Param_centroider
     Bases: object
     interpmat
           optional reference function(s) used for corr centroiding
     nmax
           number of brightest pixels
     nwfs
           index of wfs in y_wfs structure on which we want to do centroiding
     set nmax()
           Set the number of brightest pixels to use for bpcog
               Parameters n: (int): number of brightest pixels
     set nwfs()
           Set the index of wfs
               Parameters n - (int): index of wfs
     set_sizex()
           Set sizex parameters for corr centroider (interp_mat size)
               Parameters s: (long): x size
     set_sizey()
           Set sizey parameters for corr centroider (interp_mat size)
               Parameters s: (long): y size
     set thresh()
           Set the threshold for tcog
               Parameters t: (float): threshold
     set_type()
           Set the centroider type :param t: (str): centroider type
     set_type_fct()
           Set the type of ref function
               Parameters \mathbf{f} – (str): type of ref function
     set_weights()
           Set the weights to use with wcog or corr
               Parameters w: (np.ndarray[ndim=3 ,dtype=np.float32_t]) : weights
     set width()
           Set the width of the Gaussian
               Parameters \mathbf{w} – (float) : width of the gaussian
```

```
sizex
          x-size for inter mat (corr)
     sizey
          x-size for inter mat (corr)
     thresh
          Threshold
     type_centro
          type of centroiding cog, tcog, bpcog, wcog, corr
     type_fct
          type of ref function gauss, file, model
     weights
          optional reference function(s) used for centroiding
     width
          width of the Gaussian
class param.Param_controller
     Bases: object
     TTcond
          tiptilt condition number for cmat filtering with mv controller
     cmat
          full control matrix
     cured ndivs
          subdivision levels in cured
     delay
          loop delay [frames]
     gain
          loop gain
     gmax
          Maximum gain for modal optimization
     gmin
          Minimum gain for modal optimization
     imat
          full interaction matrix
     maxcond
          max condition number
     modopti
          Flag for modal optimization
     nactu
          number of controled actuator per dm
     ndm
          index of dms in controller
     ngain
          Number of tested gains
     nkl
```

```
nmodes
     Number of modes for M2V matrix (modal optimization)
nrec
     Number of sample of open loop slopes for modal optimization computation
nvalid
     number of valid subaps per wfs
nwfs
     index of wfss in controller
set TTcond()
     Set the tiptilt condition number for cmat filtering with my controller
     :param : (float) : tiptilt condition number
set_cmat()
     Set the full control matrix
         Parameters cmat - (np.ndarray[ndim=2,dtype=np.float32_t]): full control matrix
set_cured_ndivs()
     Set the subdivision levels in cured
         Parameters c – (long): subdivision levels in cured
set_delay()
     Set the loop delay expressed in frames
         Parameters d: (float) :delay [frames]
set_gain()
     Set the loop gain
         Parameters g: (float): loop gain
set_gmax()
     Set the maximum gain for modal optimization
         Parameters g – (flaot): maximum gain for modal optimization
set qmin()
     Set the minimum gain for modal optimization
         Parameters g – (float) : minimum gain for modal optimization
set imat()
     Set the full interaction matrix
         Parameters imat – (np.ndarray[ndim=2,dtype=np.float32_t]): full interaction matrix
set maxcond()
     Set the max condition number
     :param : (float) : max condition number
set_modopti()
     Set the flag for modal optimization
         Parameters m – (int): flag for modal optimization
set_nactu()
     Set the number of controled actuator
         Parameters 1 – (list of int): number of controlled actuator per dm
```

```
set ndm()
          Set the indices of dms
              Parameters 1 – (list of int): indices of dms
     set ngain()
          Set the number of tested gains
              Parameters n - (int): number of tested gains
     set_nkl()
          Set the number of KL modes used for computation of covmat in case of minimum variance controller
              Parameters n – (long): number of KL modes
     set_nmodes()
          Set the number of modes for M2V matrix (modal optimization)
              Parameters n - (int): number of modes
     set_nrec()
          Set the number of sample of open loop slopes for modal optimization computation
              Parameters n - (int): number of sample
     set_nvalid()
          Set the number of valid subaps
              Parameters 1 – (list of int): number of valid subaps per wfs
     set nwfs()
          Set the indices of wfs
              Parameters 1 - (list of int) : indices of wfs
     set_type()
     type_control
          type of controller
class param.Param_dm
     Bases: object
     _com
          current command
     _i1
     influ
          influence functions
     influkernel
     _influpos
     influsize
          total number of actuators
     _influstart
     _j1
     _klbas
          np.ndarray to a kl struct
     _n1
          position of leftmost pixel in largest support
```

```
_n2
     position of rightmost pixel in largest support
_ninflu
     Influence functions
ntotact
     total number of actuators
pitch
     inter-actuator space in pixels
_pupil
     pupil mask for this dm
_puppixoffset
_xpos
     x positions of influ functions
_ypos
     y positions of influ functions
alt
     conjugaison altitude (im m)
coupling
     actuators coupling (<0.3)
hyst
     actuators hysteresis (<1.)
margin
nact
     number of actuators in the diameter
nkl
     number of kl modes
pupoffset
       2.
push4imat
     nominal voltage for imat
set_alt()
     set the conjugaison altitude
         Parameters a – (float) : conjugaison altitude (im m)
set_coupling()
     set the actuators coupling
         Parameters c - (float): actuators coupling (<0.3)
set_i1()
     TODO doc
         Parameters i1 – (np.ndarray[ndim=1,dtype=np.int32_t]):
set_influ()
     Set the influence function
         Parameters influ – (np.ndarray[ndim=3,dtype=np.float32_t]): influence function
```

```
set_j1()
           TODO doc
               Parameters j1 – (np.ndarray[ndim=1,dtype=np.int32_t]):
     set nact()
          set the number of actuator
               Parameters n - (long): number of actuators in the diameter
     set ntotact()
           set the total number of actuators
               Parameters n - (long): total number of actuators
     set_push4imat()
           set the nominal voltage for imat
               Parameters p – (float): nominal voltage for imat
     set_thresh()
           set the threshold on response for selection
               Parameters t - (float): threshold on response for selection (<1)
     set_type()
           set the dm type
               Parameters t - (str): type of dm
     set_unitpervolt()
           set the Influence function sensitivity
               Parameters \mathbf{u} – (float): Influence function sensitivity in unit/volt
     set_xpos()
           Set the x positions of influ functions
               Parameters xpos – (np.ndarray[ndim=1,dtype=np.float32_t]): x positions of influ functions
     set_ypos()
          Set the y positions of influ functions
               Parameters ypos – (np.ndarray[ndim=1,dtype=np.float32 t]): y positions of influ functions
     thresh
          threshold on response for selection (<1)
     type_dm
           type of dm
     unitpervolt
          Influence function sensitivity in unit/volt. Optional [0.01] Stackarray: mic/volt, Tip-tilt: arcsec/volt.
class param.Param_geom
     Bases: object
     _mpupil
     _n
     _n1
     _n2
      _p1
```

```
_p2
cent
     central point of the simulation.
geom_init()
     Initialize the system geometry
         Parameters tel: (Param_tel) : telescope settings
              pupdiam: (long): linear size of total pupil
              apod: (int): apodizer
get_ipupil()
     return the full pupil support
get_mpupil()
     return the padded pupil
get_n()
     Return the linear size of the medium pupil
get_n1()
     Return the min(x,y) for valid points for the total pupil
     Return the max(x,y) for valid points for the total pupil
get_p1()
     Return the min(x,y) for valid points for the medium pupil
     Return the max(x,y) for valid points for the medium pupil
get_spupil()
     return the small pupil
pupdiam
     linear size of total pupil (in pixels).
set_cent()
     Set the central point of the simulation
         Parameters \mathbf{c} – (float) : central point of the simulation.
set_pupdiam()
     Set the linear size of total pupil
         Parameters p - (long): linear size of total pupil (in pixels).
set ssize()
     Set linear size of full image
         Parameters s – (long): linear size of full image (in pixels).
set_zenithangle()
     Set observations zenith angle
         Parameters z - (float): observations zenith angle (in deg).
     linear size of full image (in pixels).
zenithangle
     observations zenith angle (in deg).
```

```
class param.Param_loop
     Bases: object
     ittime
          iteration time (in sec)
     niter
          number of iterations
     set ittime()
          Set iteration time
              Parameters t: (float):iteration time
     set niter()
          Set the number of iteration
              Parameters n: (long): number of iteration
class param.Param_rtc
     Bases: object
     centroiders
     controllers
     nwfs
     set centroiders()
          Set the centroiders
              Parameters 1 – (list of Param_centroider) : centroiders settings
     set_controllers()
          Set the controller
              Parameters 1 – (list of Param_controller) : controllers settings
     set_nwfs()
          Set the number of wfs
              Parameters n - (int) number of wfs
class param.Param_target
     Bases: object
     Lambda
          observation wavelength for each target
     apod
          boolean for apodizer
     dms seen
          index of dms seen by the target
     mag
          magnitude for each target
     ntargets
          number of targets
     set_Lambda()
          Set the observation wavelength
              Parameters 1 – (list of float) : observation wavelength for each target
```

```
set_apod()
           Tells if the apodizer is used
           The apodizer is used if a is not 0 :param a: (int) boolean for apodizer
     set mag()
           set the magnitude
               Parameters 1 – (list of float) : magnitude for each target
     set_nTargets()
           Set the number of targets
               Parameters n - (int): number of targets
     set_xpos()
           Set the x positions on sky (in arcsec)
               Parameters 1 - (list of float) : x positions on sky for each target
     set_ypos()
           Set the y positions on sky (in arcsec)
               Parameters 1 – (list of float): y positions on sky for each target
     xpos
           x positions on sky (in arcsec) for each target
     ypos
           y positions on sky (in arcsec) for each target
     zerop
           target flux for magnitude 0
class param.Param_tel
     Bases: object
     cobs
           central obstruction ratio.
     diam
           telescope diameter (in meters).
     nbrmissing
           number of missing segments for EELT pupil (max is 20).
     pupangle
           rotation angle of pupil.
     referr
           std of reflectivity errors for EELT segments (fraction).
     set_cobs()
           set the central obstruction ratio
               Parameters c – (float): central obstruction ratio
     set_diam()
           set the telescope diameter
               Parameters d – (float): telescope diameter (in meters)
     set_nbrmissing()
           set the number of missing segments for EELT pupil
               Parameters nb – (long): number of missing segments for EELT pupil (max is 20)
```

```
set_pupangle()
           set the rotation angle of pupil
               Parameters p – (float) : rotation angle of pupil
     set referr()
          set the std of reflectivity errors for EELT segments
               Parameters ref – (float): std of reflectivity errors for EELT segments (fraction)
     set_spiders_type()
           set the secondary supports type
               Parameters spider – (str): secondary supports type
     set_t_spiders()
           set the secondary supports ratio
               Parameters spider - (float): secondary supports ratio
     set_type_ap()
           set the EELT aperture type
               Parameters t - (str): EELT aperture type
     spiders_type
          secondary supports type: "four" or "six".
     t spiders
          secondary supports ratio.
     type_ap
          EELT aperture type: "Nominal", "BP1", "BP3", "BP5" (for back-up plan with 1, 3, or 5 missing annulus).
class param.Param_wfs
     Bases: object
     Lambda
           observation wavelength (in µm) for a subap.
     _Nfft
           array size for fft for a subap (in pixel)
      Ntot
           total size of hr image for a subap (in pixel)
     _altna
          corresponding altitude
     _azimuth
           angles of rotation for each spot
     beam
           1d beam function
     _binmap
          (int*) array of pixels transform from full FoV hr images to binned images
     fluxPerSub
           fraction of nphotons per subap
      ftbeam
           1d beam function fft
```

```
halfxy
     (float*) phase offset for 1/2 pixel shift in (x,y)
     (int*) array of pixels transform from minimal FoV image to (in case type is sh or geo)
istart
     (int*) x start indexes for cutting phase screens
isvalid
     (int*) array of 0/1 for valid subaps
_jstart
     (int*) y start indexes for cutting phase screens
_lgskern
     lgs kernels for each subap
_nphotons
     number of photons per subap
     rebin factor from hr to binned image for a subap
_nvalid
     number of valid subaps
pdiam
     pupil diam for a subap (in pixel)
_phasemap
     (int*) array of pixels transform from phase screen into subaps phase screens
_prof1d
     hr profile
_profcum
     hr profile cumulated
_profna
     sodium profile
_pyr_cx
_pyr_cy
_pyr_offsets
qpixsize
     quantum pixel size for the simulation
subapd
     subap diameter (m)
_{\tt submask}
     (float*) fieldstop for each subap
_validsubsx
     (int*) indices of valid subaps along axis x
_validsubsy
     (int*) indices of valid subaps along axis y
atmos_seen
```

30

1 if the WFS sees the atmosphere layers

beamsize laser beam fwhm on-sky (in arcsec). dms seen index of dms seen by the WFS fracsub minimal illumination fraction for valid subaps. fssize size of field stop in arcsec. fstop size of field stop in arcsec. gsalt altitude of guide star (in m) 0 if ngs. gsmag magnitude of guide star. kernel laserpower laser power in W. lgsreturnperwatt return per watt factor (high season: 10 ph/cm2/s/W). 11tx x position (in meters) of llt. llty y position (in meters) of llt. noise desired noise : < 0 = no noise / 0 = photon only / > 0 photon + ron.npix number of pixels per subap. nxsub linear number of subaps. openloop 1 if in "open-loop" mode (i.e. does not see dm). optthroughput wfs global throughput. pixsize pixel size (in arcsec) for a subap. proftype type of sodium profile "gauss", "exp", etc ... pyr_ampl pyramid wfs modulation amplitude radius [arcsec]. pyr_loc Location of modulation, before/after the field stop. valid value are "before" or "after" (default "after").

1.5. Param 31

total number of point along modulation circle [unitless].

pyr_npts

```
pyrtype
     Type of pyramid, either 0 for "Pyramid" or 1 for "RoofPrism".
set Lambda()
     Set the observation wavelength
         Parameters L - (float): observation wavelength (in \mum) for a subap
set altna()
     Set the corresponding altitude
         Parameters a – (np.ndarray[ndim=1,dtype=np.float32]): corresponding altitude
set_atmos_seen()
     Tells if the wfs sees the atmosphere layers
         Parameters i - (int) : 1 if the WFS sees the atmosphere layers
set_beamsize()
     Set the laser beam fwhm on-sky
         Parameters b - (float): laser beam fwhm on-sky (in arcsec)
set_dms_seen()
     Set the index of dms seen by the WFS
         Parameters dms_seen - (np.ndarray[ndim=1,dtype=np.int32_t): index of dms seen by the
             WFS
set fracsub()
     Set the minimal illumination fraction for valid subaps
         Parameters f – (float) : minimal illumination fraction for valid subaps
set_fssize()
     Set the size of field stop
         Parameters f – (float): size of field stop in arcsec
set_fstop()
     Set the size of field stop
         Parameters \mathbf{f} – (str): size of field stop in arcsec
set qsalt()
     Set the altitude of guide star
         Parameters g – (float) : altitude of guide star (in m) 0 if ngs
set gsmag()
     Set the magnitude of guide star
         Parameters g – (float) : magnitude of guide star
set_kernel()
     Set the attribute kernel
         Parameters k – (float):
set_laserpower()
     Set the laser power
         Parameters 1 – (float) : laser power in W
set lgsreturnperwatt()
     Set the return per watt factor
```

```
Parameters 1pw – (float): return per watt factor (high season: 10 ph/cm2/s/W)
set lltx()
     Set the x position of llt
         Parameters 1 - (float) : x position (in meters) of llt
set llty()
     Set the y position of llt
         Parameters 1 – (float) : y position (in meters) of llt
set_noise()
     Set the desired noise
         Parameters \mathbf{n} – (float): desired noise: < 0 = \text{no noise} / 0 = \text{photon only} / > 0 \text{ photon} + \text{ron}
set_npix()
     Set the number of pixels per subap
         Parameters n - (long): number of pixels per subap
set nxsub()
     Set the linear number of subaps
         Parameters n – (long): linear number of subaps
set_openloop()
     Set the loop state (open or closed)
         Parameters o – (long): 1 if in "open-loop" mode (i.e. does not see dm)
set_optthroughput()
     Set the wfs global throughput
         Parameters o – (float): wfs global throughput
set_pixsize()
     Set the pixel size
         Parameters p - (float): pixel size (in arcsec) for a subap
set_profna()
     Set the sodium profile
         Parameters p – (np.ndarray[ndim=1,dtype=np.float32]) : sodium profile
set_proftype()
     Set the type of sodium profile
         Parameters p – (str): type of sodium profile "gauss", "exp", etc ...
set pyr ampl()
     Set the pyramid wfs modulation amplitude radius
         Parameters p – (float) : pyramid wfs modulation amplitude radius (in arsec)
set_pyr_loc()
     Set the location of modulation
         Parameters \mathbf{p} – (str): location of modulation, before/after the field stop. valid value are "be-
              fore" or "after" (default "after")
set_pyr_npts()
     Set the total number of point along modulation circle
         Parameters p - (long): total number of point along modulation circle
```

1.5. Param 33

```
set_pyrtype()
           Set the type of pyramid,
               Parameters p – (str): type of pyramid, either 0 for "Pyramid" or 1 for "RoofPrism"
     set_type()
           Set the type of wfs
               Parameters t - (str): type of wfs ("sh" or "pyr")
     set_xpos()
           Set the guide star x position on sky
               Parameters \mathbf{x} – (float): guide star x position on sky (in arcsec)
     set_ypos()
           Set the guide star y position on sky
               Parameters y - (float): guide star y position on sky (in arcsec)
     set_zerop()
           Set the detector zero point
               Parameters z – (float) : detector zero point
     type_wfs
           type of wfs: "sh" or "pyr".
     xpos
           guide star x position on sky (in arcsec).
     ypos
           guide star x position on sky (in arcsec).
     zerop
           detector zero point.
param.indices()
     DOCUMENT indices(dim) Return a dimxdimx2 array. First plane is the X indices of the pixels in the dimxdim
     array. Second plane contains the Y indices. Inspired by the Python scipy routine of the same name. New (June
     12 2002): dim can either be: - a single number N (e.g. 128) in which case the returned array are
           square (NxN)
          •a Yorick array size, e.g. [#dimension,N1,N2], in which case the returned array are N1xN2
          •a vector [N1,N2], same result as previous case
     F.Rigaut 2002/04/03 SEE ALSO: span
           Parameters dim1: (int): first dimension dim2: (int): (optional) second dimension
param.make_apodizer()
     TODO doc
           Parameters (int): im:
               (int): pupd:
               (str): filename:
               (float): angle:
param.makeqaussian (size, fwhm, xc, yc)
     Returns a centered gaussian of specified size and fwhm. norm returns normalized 2d gaussian
```

```
Parameters size: (int):
                fwhm: (float):
                xc: (int): (optional) center position on x axis
                yc: (int): (optional) center position on y axis
                norm: (int): (optional) normalization
param.rotate()
      Rotates an image of an angle "ang" (in DEGREES).
      The center of rotation is cx,cy. A zoom factor can be applied.
      (cx,cy) can be omitted :one will assume one rotates around the center of the image. If zoom is not specified, the
      default value of 1.0 is taken.
           Parameters im: (np.ndarray[ndim=3,dtype=np.float32_t]): array to rotate
                ang: (float): rotation angle (in degrees)
                cx: (int): (optional) rotation center on x axis (default: image center)
                cy: (int): (optional) rotation center on x axis (default: image center)
                zoom: (float): (opional) zoom factor (default =1.0)
param.rotate3d()
      Rotates an image of an angle "ang" (in DEGREES).
      The center of rotation is cx,cy. A zoom factor can be applied.
      (cx,cy) can be omitted :one will assume one rotates around the center of the image. If zoom is not specified, the
      default value of 1.0 is taken.
      modif dg: allow to rotate a cube of images with one angle per image
           Parameters im: (np.ndarray[ndim=3,dtype=np.float32_t]) : array to rotate
                ang: (np.ndarray[ndim=1,dtype=np.float32_t]): rotation angle (in degrees)
                cx: (int): (optional) rotation center on x axis (default: image center)
                cy: (int): (optional) rotation center on x axis (default: image center)
                zoom: (float): (opional) zoom factor (default =1.0)
```

1.6 Dms

1.6.1 dms module

```
class dms.Dms
    Bases: object
    add_dm()
    comp_oneactu()
        Compute the shape of the dm when pushing the nactu actuator
        Parameters type_dm: (str): dm type
        alt: (float): dm conjugaison altitude
        nactu: (int): actuator number pushed
```

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ampli: (float): amplitude

```
computeKLbasis()
     Compute a Karhunen-Loeve basis for the dm:
           · compute the phase covariance matrix on the actuators using Kolmogorov
           • compute the geometric covariance matrix

    double diagonalisation to obtain KL basis

         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
             xpos: (np.ndarray[ndim=1,dtype=np.float32_t]): x-position of actuators
             ypos: (np.ndarray[ndim=1,dtype=np.float32_t]): y-position of actuators
             indx_pup: (np.ndarray[ndim=1,dtype=np.int32_t]): indices of where(pup)
             dim: (long): number of where(pup)
             norm: (float): normalization factor
             ampli: (float): amplitude
getComm()
     Return the voltage command of the sutra dm
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
         Returns data: (np.ndarray(dims=1,dtype=np.float32)): voltage vector
getInflu()
     Return the influence functions of the DM
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
         Returns data: (np.ndarray(dims=3,dtype=np.float32)): influence functions
qet KLbasis()
     Return the klbasis computed by computeKLbasis
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
         Returns KLbasis: (np.ndarray(dims=2,dtype=np.float32)): the KL basis
get_dm()
     Return the shape of the dm
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
         Returns data: (np.ndarray(dims=2,dtype=np.float32)): DM shape
load_kl()
     Load all the arrays computed during the initialization for a kl DM in a sutra_dms object
```

```
Parameters alt: (float): dm conjugaison altitude
             rabas: (np.ndarray[ndim=1,dtype=np.float32_t]): TODO
             azbas: (np.ndarray[ndim=1,dtype=np.float32_t]) :
             ord: (np.ndarray[ndim=1,dtype=np.int32_t]):
             cr: (np.ndarray[ndim=1,dtype=np.float32 t]):
             cp: (np.ndarray[ndim=1,dtype=np.float32 t]):
load_pzt()
     Load all the arrays computed during the initialization for a pzt DM in a sutra_dms object
         Parameters alt: (float): dm conjugaison altitude
             influ: (np.ndarray[ndim=3,dtype=np.float32_t]): influence functions
             influpos: (np.ndarray[ndim=1,dtype=np.int32_t]): positions of the IF
             npoints: (np.ndarray[ndim=1,dtype=np.int32_t]) [for each pixel on the DM screen,] the
                number of IF which impact on this pixel
             istart: (np.ndarray[ndim=1,dtype=np.int32_t]) :
             xoff: (np.ndarray[ndim=1,dtype=np.int32_t]) : x-offset
             yoff: (np.ndarray[ndim=1,dtype=np.int32_t]) :y-offset
             kern: (np.ndarray[ndim=1,dtype=np.float32 t]): convoltuon kernel
load tt()
     Load all the arrays computed during the initialization for a tt DM in a sutra_dms object
         Parameters alt: (float): dm conjugaison altitude
             influ: (np.ndarray[ndim=3,dtype=np.float32_t]): influence functions
oneactu()
     Push on on the nactu actuator of the DM with ampli amplitude and compute the corresponding shape
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
             nactu: (int): actuator number
             ampli: (float): amplitude
remove_dm()
     Remove a dm from a Dms object
         Parameters type_dm: (str): dm type to remove
             alt: (float): dm conjugaison altitude to remove
resetdm()
     Reset the shape of the DM to 0
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
set_comm()
     Set the voltage command on a sutra_dm
     type_dm: (str): dm type
```

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```
alt: (float): dm conjugaison altitude
           comm: (np.ndarray[ndim=1,dtype=np.float32_t]): voltage vector
     shape_dm()
           Compute the shape of the DM in a sutra_dm object
           type_dm: (str): dm type
           alt: (float): dm conjugaison altitude
dms.comp_dmgeom()
     Compute the geometry of a DM: positions of actuators and influence functions
           Parameters dm: (Param_dm): dm settings
               geom: (Param_geom): geom settings
dms.compute_klbasis()
     Compute a Karhunen-Loeve basis for the dm:

    compute the phase covariance matrix on the actuators using Kolmogorov

             • compute the geometric covariance matrix
             • double diagonalisation to obtain KL basis
           Parameters g_dm: (Dms): Dms object
               p_dm: (Param_dm): dm settings
               p_geom: (Param_geom): geom settings
               p_atmos: (Param_atmos): atmos settings
               p_tel: (Param_tel) : telescope settings
dms.dm_init()
     Create and initialize a Dms object on the gpu
           Parameters p_dms: (list of Param_dms): dms settings
               p_wfs: (Param_wfs): wfs settings
               p_geom: (Param_geom) : geom settings
               p_tel: (Param_tel) : telescope settings
```

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