COMPASS (SHESHA) Documentation

Release r791

COMPASS team

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CHAPTER

ONE

CONTENTS:

1.1 shesha

1.1.1 bin2d()

shesha.bin2d()

Returns the input 2D array "array", binned with the binning factor "binfact". The input array X and/or Y dimensions needs not to be a multiple of "binfact"; The final/edge pixels are in effect replicated if needed. This routine prepares the parameters and calls the C routine _bin2d. The input array can be of type long, float or double. Last modified: Dec 15, 2003. Author: F.Rigaut SEE ALSO: _bin2d

Parmeters data_in: (np.ndarray): data to binned

binfact: (int): binning factor

1.1.2 indices()

```
shesha.indices (int dim1, int dim2=-1)
```

Return a dimxdimx2 array. First plane is the X indices of the pixels in the dimxdim array. Second plane contains the Y indices.

Inspired by the Python scipy routine of the same name.

New (June 12 2002): dim can either be:

- •a single number N (e.g. 128) in which case the returned array are square (NxN)
- •a Yorick array size, e.g. [#dimension,N1,N2], in which case the returned array are N1xN2
- •a vector [N1,N2], same result as previous case

F.Rigaut 2002/04/03 SEE ALSO: span

Parameters

- dim1 (int) : first dimension
- dim2 (int) : (optional) second dimension

1.1.3 makegaussian()

```
shesha.makegaussian (size, fwhm, xc, yc)
```

Returns a centered gaussian of specified size and fwhm. norm returns normalized 2d gaussian

```
Parameters size: (int):
```

```
fwhm: (float):
```

xc: (int): (optional) center position on x axis yc: (int): (optional) center position on y axis

```
norm: (int): (optional) normalization
```

geom: (Param_geom): geometry settings

loop: (Param_loop): loop settings

1.2 shesha atmos

```
1.2.1 Atmos
class shesha atmos. Atmos
      add_screen()
           Add a screen to the atmos object.
               Parameters size: (float): dimension of the screen (size x size)
                   amplitude: (float): frac
                   altitude: (float): altitude of the screen in meters
                   windspeed: (float): windspeed of the screen [m/s]
                   winddir: (float): wind direction (deg)
                   deltax: (float): extrude deltax pixels in the x-direction at each iteration
                   deltay: (float): extrude deltay pixels in the y-direction at each iteration
                   device: (int): device number
      del_screen()
           Delete a screen from the atmos object
               Parameters alt – (float): altitude of the screen to delete
      disp()
           Display the screen phase at a given altitude
               Parameters alt – (float): altitude of the screen to display
      get_screen()
           Return a numpy array containing the turbulence at a given altitude
               Parameters alt – (float) :altitude of the screen to get
      list alt()
           Display the list of the screens altitude
     move_atmos()
           Move the turbulence in the atmos screen following previous loaded paramters such as windspeed and
           wind direction
1.2.2 atmos init()
shesha_atmos.atmos_init()
     atmos_init(naga_context c, Param_atmos atm, Param_tel tel, Param_geom geom, Param_loop loop,
           wfss=None, Param_target target=None, int rank=0, int clean=1, dict load={})
     Create and initialise an atmos object
           Parameters c: (naga_context) : context
               tel: (Param_tel): telescope settings
```

```
wfss: (list of Param_wfs): (optional) wfs settings
               sensors: (Sensors): (optional) Sensors object on GPU
               target: (Param_target): (optional) target_settings
               overwrite: (int): (optional) overwrite data files if overwite=1 (default 1)
               rank: (int): (optional) rank of the process (default=0)
               clean: (clean): (optional)? (default=1)
               load: (dict) : (optional) ? (default={})
1.3 shesha dms
1.3.1 Dms
class shesha dms.Dms
      add dm()
           Add a dm into a Dms object
               Parameters type_dm: (str): dm type to remove,
                   alt: (float): dm conjugaison altitude to remove,
                   ninflu: (long):,
                   influsize: (long):,
                   ninflupos: (long):,
                   npts: (long):,
                   push4imat: (float):,
                   device: (int): device where the DM will be create (default=-1):
      comp_oneactu()
           Compute the shape of the dm when pushing the nactu actuator
               Parameters type_dm: (str): dm type
                   alt: (float): dm conjugaison altitude
                   nactu: (int): actuator number pushed
                   ampli: (float): amplitude
      computeKLbasis()
           Compute a Karhunen-Loeve basis for the dm:
                 • compute the phase covariance matrix on the actuators using Kolmogorov
                 • compute the geometric covariance matrix
                 • double diagonalisation to obtain KL basis
               Parameters type_dm: (str): dm type
                   alt: (float): dm conjugaison altitude
                   xpos: (np.ndarray[ndim=1,dtype=np.float32_t]): x-position of actuators
                   ypos: (np.ndarray[ndim=1,dtype=np.float32_t]) : y-position of actuators
                   indx_pup: (np.ndarray[ndim=1,dtype=np.int32_t]) : indices of where(pup)
```

1.3. shesha dms 5

```
dim: (long): number of where(pup)
              norm: (float): normalization factor
             ampli: (float): amplitude
getComm()
     Return the voltage command of the sutra_dm
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
         Returns data: (np.ndarray(dims=1,dtype=np.float32)): voltage vector
getInflu()
     Return the influence functions of the DM
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
         Returns data: (np.ndarray(dims=3,dtype=np.float32)): influence functions
get KLbasis()
     Return the klbasis computed by computeKLbasis
         Parameters type_dm: (str): dm type
              alt: (float): dm conjugaison altitude
         Returns KLbasis: (np.ndarray(dims=2,dtype=np.float32)): the KL basis
get dm()
     Return the shape of the dm
         Parameters type_dm: (str): dm type
             alt: (float): dm conjugaison altitude
         Returns data: (np.ndarray(dims=2,dtype=np.float32)): DM shape
load_kl()
     Load all the arrays computed during the initialization for a kl DM in a sutra_dms object
         Parameters alt: (float): dm conjugaison altitude
             rabas: (np.ndarray[ndim=1,dtype=np.float32_t]): TODO
             azbas: (np.ndarray[ndim=1,dtype=np.float32_t]) :
             ord: (np.ndarray[ndim=1,dtype=np.int32_t]):
             cr: (np.ndarray[ndim=1,dtype=np.float32 t]):
             cp: (np.ndarray[ndim=1,dtype=np.float32_t]) :
load_pzt()
     Load all the arrays computed during the initialization for a pzt DM in a sutra_dms object
         Parameters alt: (float): dm conjugaison altitude
             influ: (np.ndarray[ndim=3,dtype=np.float32_t]): influence functions
             influpos: (np.ndarray[ndim=1,dtype=np.int32_t]): positions of the IF
              npoints: (np.ndarray[ndim=1,dtype=np.int32_t]) [for each pixel on the DM screen,]
                the number of IF which impact on this pixel
             istart: (np.ndarray[ndim=1,dtype=np.int32_t]) :
              xoff: (np.ndarray[ndim=1,dtype=np.int32_t]) : x-offset
             yoff: (np.ndarray[ndim=1,dtype=np.int32_t]):y-offset
```

```
kern: (np.ndarray[ndim=1,dtype=np.float32_t]): convoltuon kernel
     load_tt()
          Load all the arrays computed during the initialization for a tt DM in a sutra_dms object
               Parameters alt: (float): dm conjugaison altitude
                   influ: (np.ndarray[ndim=3,dtype=np.float32_t]): influence functions
      oneactu()
          Push on on the nactu actuator of the DM with ampli amplitude and compute the corresponding shape
               Parameters type_dm: (str): dm type
                   alt: (float): dm conjugaison altitude
                   nactu: (int): actuator number
                   ampli: (float): amplitude
      remove_dm()
          Remove a dm from a Dms object
               Parameters type_dm: (str): dm type to remove
                   alt: (float): dm conjugaison altitude to remove
     resetdm()
          Reset the shape of the DM to 0
               Parameters type_dm: (str): dm type
                   alt: (float): dm conjugaison altitude
      set comm()
          Set the voltage command on a sutra_dm
          type_dm: (str): dm type
          alt: (float): dm conjugaison altitude
          comm: (np.ndarray[ndim=1,dtype=np.float32_t]): voltage vector
           shape_dm: (bool): perform the dm_shape after the load (default=False)
      set_full_comm()
          Set the voltage command
          comm: (np.ndarray[ndim=1,dtype=np.float32_t]): voltage vector
          shape_dm: (bool): perform the dm_shape after the load (default=True)
     shape_dm()
          Compute the shape of the DM in a sutra dm object
          type_dm: (str): dm type
           alt: (float): dm conjugaison altitude
1.3.2 comp_dmgeom()
shesha_dms.comp_dmgeom()
     Compute the geometry of a DM: positions of actuators and influence functions
           Parameters dm: (Param_dm): dm settings
               geom: (Param_geom): geom settings
```

1.3. shesha dms

1.3.3 computeDMbasis()

```
\verb|shesha_dms.computeDMbasis|()|
```

Compute a the DM basis []

- · push on each actuator
- get the corresponding dm shape
- · apply pupil mask and store in a column

```
Parameters g_dm: (Dms) : Dms object p_dm: (Param_dm) : dm settings
```

p_geom: (Param_geom) : geom settings

Returns IFbasis = (np.ndarray((indx_valid.size,Nactu),dtype=np.float32)) : DM IF basis

1.3.4 compute_klbasis()

```
shesha_dms.compute_klbasis()
```

Compute a Karhunen-Loeve basis for the dm:

- compute the phase covariance matrix on the actuators using Kolmogorov
- compute the geometric covariance matrix
- double diagonalisation to obtain KL basis

```
Parameters g_dm: (Dms): Dms object

p_dm: (Param_dm): dm settings

p_geom: (Param_geom): geom settings

p_atmos: (Param_atmos): atmos settings

p_tel: (Param_tel): telescope settings
```

1.3.5 dm_init()

```
shesha_dms.dm_init()
Create and initialize a Dms object on the gpu
```

Parameters p_dms: (list of Param_dms) : dms settings p_wfs: (Param_wfs) : wfs settings

p_geom: (Param_geom) : geom settings
p_tel: (Param_tel) : telescope settings

1.3.6 make_pzt_dm()

```
shesha_dms.make_pzt_dm()
```

Compute the actuators positions and the influence functions for a pzt DM

```
Parameters p_dm: (Param_dm): dm settings
geom: (Param_geom): geom settings
p_tel: (Param_tel): tel settings
irc: factor for influence size
```

Returns influ: (np.ndarray(dims=3,dtype=np.float64)): cube of the IF for each actuator

1.4 shesha_param

```
1.4.1 Param loop
class shesha_param.Param_loop
     ittime
          iteration time (in sec)
     niter
          number of iterations
      set ittime()
          Set iteration time
              Parameters t: (float):iteration time
     set niter()
          Set the number of iteration
              Parameters n: (long): number of iteration
1.4.2 Param_tel
class shesha_param.Param_tel
     cobs
          central obstruction ratio.
     diam
          telescope diameter (in meters).
     nbrmissing
          number of missing segments for EELT pupil (max is 20).
     pupangle
          rotation angle of pupil.
     referr
          std of reflectivity errors for EELT segments (fraction).
     set_cobs()
          set the central obstruction ratio
              Parameters c – (float): central obstruction ratio
     set_diam()
          set the telescope diameter
              Parameters d – (float): telescope diameter (in meters)
     set_nbrmissing()
          set the number of missing segments for EELT pupil
              Parameters nb – (long): number of missing segments for EELT pupil (max is 20)
     set_pupangle()
          set the rotation angle of pupil
```

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Parameters p – (float): rotation angle of pupil

```
set_referr()
          set the std of reflectivity errors for EELT segments
               Parameters ref – (float) : std of reflectivity errors for EELT segments (fraction)
     set_spiders_type()
          set the secondary supports type
               Parameters spider – (str): secondary supports type
      set_std_piston()
          set the std of piston errors for EELT segments
               Parameters piston – (float): std of piston errors for EELT segments
     set_std_tt()
          set the std of tip-tilt errors for EELT segments
               Parameters tt – (float): std of tip-tilt errors for EELT segments
     set_t_spiders()
          set the secondary supports ratio
               Parameters spider – (float): secondary supports ratio
     set_type_ap()
          set the EELT aperture type
               Parameters t - (str): EELT aperture type
     spiders_type
          secondary supports type: "four" or "six".
     std piston
          std of piston errors for EELT segments
     std tt
          std of tip-tilt errors for EELT segments
     t_spiders
          secondary supports ratio.
          EELT aperture type: "Nominal", "BP1", "BP3", "BP5" (for back-up plan with 1, 3, or 5 missing
          annulus).
1.4.3 Param_geom
class shesha_param.Param_geom
     cent
          central point of the simulation.
     geom_init()
          Initialize the system geometry
               Parameters tel: (Param_tel): telescope settings
                   pupdiam: (long): linear size of total pupil
                  apod: (int): apodizer
     get_ipupil()
          return the full pupil support
     get_mpupil()
          return the padded pupil
```

```
get_n()
          Return the linear size of the medium pupil
           Return the min(x,y) for valid points for the total pupil
     get_n2()
           Return the max(x,y) for valid points for the total pupil
     get_p1()
          Return the min(x,y) for valid points for the medium pupil
           Return the max(x,y) for valid points for the medium pupil
     get_spupil()
           return the small pupil
     pupdiam
           linear size of total pupil (in pixels).
      set_cent()
           Set the central point of the simulation
               Parameters c - (float): central point of the simulation.
      set_pupdiam()
           Set the linear size of total pupil
               Parameters p – (long) : linear size of total pupil (in pixels).
      set_ssize()
           Set linear size of full image
               Parameters s – (long): linear size of full image (in pixels).
      set_zenithangle()
           Set observations zenith angle
               Parameters z - (float): observations zenith angle (in deg).
      ssize
           linear size of full image (in pixels).
      zenithangle
           observations zenith angle (in deg).
1.4.4 Param wfs
class shesha_param.Param_wfs
     Lambda
           observation wavelength (in µm) for a subap.
      atmos_seen
           1 if the WFS sees the atmosphere layers
     beamsize
           laser beam fwhm on-sky (in arcsec).
     dms seen
          index of dms seen by the WFS
      error_budget
           If True, enable error budget analysis for the simulation
```

fracsub

```
minimal illumination fraction for valid subaps.
fssize
     size of field stop in arcsec.
fstop
     Fields of the wfs diaphragm shape: "square" or "none"
     altitude of guide star (in m) 0 if ngs.
     magnitude of guide star.
laserpower
     laser power in W.
lgsreturnperwatt
     return per watt factor (high season: 10 ph/cm2/s/W).
lltx
     x position (in meters) of llt.
llty
     y position (in meters) of llt.
     desired noise : < 0 = \text{no noise} / 0 = \text{photon only} / > 0 \text{ photon} + \text{ron}.
nphotons4imat
     number of photons per subap used for doing imat
npix
     number of pixels per subap.
nxsub
     linear number of subaps.
openloop
     1 if in "open-loop" mode (i.e. does not see dm).
optthroughput
     wfs global throughput.
pixsize
     pixel size (in arcsec) for a subap.
proftype
     type of sodium profile "gauss", "exp", etc ...
pyr_ampl
     pyramid wfs modulation amplitude radius [arcsec].
pyr_loc
     Location of modulation, before/after the field stop. valid value are "before" or "after" (default "after").
pyr_npts
     total number of point along modulation circle [unitless].
pyrtype
     Type of pyramid, either 0 for "Pyramid" or 1 for "RoofPrism".
set_Lambda()
     Set the observation wavelength
          Parameters L – (float) : observation wavelength (in \mum) for a subap
set_altna()
     Set the corresponding altitude
```

```
Parameters a – (np.ndarray[ndim=1,dtype=np.float32]) : corresponding altitude
set_atmos_seen()
     Tells if the wfs sees the atmosphere layers
         Parameters i - (int) : 1 if the WFS sees the atmosphere layers
set_beamsize()
     Set the laser beam fwhm on-sky
         Parameters b – (float) : laser beam fwhm on-sky (in arcsec)
set dms seen()
    Set the index of dms seen by the WFS
         Parameters dms_seen - (np.ndarray[ndim=1,dtype=np.int32_t): index of dms seen by
             the WFS
set_errorBudget()
     Set the error budget flag: if True, enable error budget analysis for this simulation
         Parameters error_budget - (bool): error budget flag
set_fracsub()
     Set the minimal illumination fraction for valid subaps
         Parameters f – (float): minimal illumination fraction for valid subaps
set fssize()
     Set the size of field stop
         Parameters f – (float): size of field stop in arcsec
set fstop()
    Set the size of field stop
         Parameters f - (str): size of field stop in arcsec
set_gsalt()
     Set the altitude of guide star
         Parameters g – (float) : altitude of guide star (in m) 0 if ngs
set_gsmag()
     Set the magnitude of guide star
         Parameters g – (float) : magnitude of guide star
set_kernel()
     Set the attribute kernel
         Parameters k - (float):
set_laserpower()
     Set the laser power
         Parameters 1 - (float): laser power in W
set_lgsreturnperwatt()
     Set the return per watt factor
         Parameters 1pw – (float): return per watt factor (high season: 10 ph/cm2/s/W)
set_lltx()
     Set the x position of llt
         Parameters 1 - (float) : x position (in meters) of llt
set_llty()
    Set the y position of llt
         Parameters 1 - (float): y position (in meters) of llt
```

```
set_noise()
     Set the desired noise
         Parameters \mathbf{n} – (float): desired noise: < 0 = \text{no noise} / 0 = \text{photon only} / > 0 \text{ photon} + \text{ron}
set_nphotons4imat()
     Set the desired numner of photons used for doing imat
         Parameters nphot – (float) : desired number of photons
set_npix()
     Set the number of pixels per subap
         Parameters n - (long): number of pixels per subap
set nxsub()
     Set the linear number of subaps
         Parameters n - (long): linear number of subaps
set_openloop()
     Set the loop state (open or closed)
         Parameters o – (long): 1 if in "open-loop" mode (i.e. does not see dm)
set_optthroughput()
     Set the wfs global throughput
         Parameters o – (float): wfs global throughput
set_pixsize()
     Set the pixel size
         Parameters p - (float): pixel size (in arcsec) for a subap
set_profna()
     Set the sodium profile
         Parameters p – (np.ndarray[ndim=1,dtype=np.float32]) : sodium profile
set_proftype()
     Set the type of sodium profile
         Parameters p – (str): type of sodium profile "gauss", "exp", etc ...
set_pyr_ampl()
     Set the pyramid wfs modulation amplitude radius
         Parameters p – (float) : pyramid wfs modulation amplitude radius (in arsec)
set_pyr_loc()
     Set the location of modulation
         Parameters p - (str): location of modulation, before/after the field stop. valid value are
             "before" or "after" (default "after")
set_pyr_npts()
     Set the total number of point along modulation circle
         Parameters p - (long): total number of point along modulation circle
set_pyrtype()
     Set the type of pyramid,
         Parameters p – (str): type of pyramid, either 0 for "Pyramid" or 1 for "RoofPrism"
set_type()
     Set the type of wfs
         Parameters t - (str): type of wfs ("sh" or "pyr")
```

```
set_xpos()
           Set the guide star x position on sky
               Parameters \mathbf{x} – (float): guide star x position on sky (in arcsec)
      set_ypos()
           Set the guide star y position on sky
               Parameters y - (float): guide star y position on sky (in arcsec)
      set_zerop()
           Set the detector zero point
               Parameters z - (float): detector zero point
      type_wfs
           type of wfs: "sh" or "pyr".
           guide star x position on sky (in arcsec).
     ypos
           guide star x position on sky (in arcsec).
      zerop
           detector zero point.
1.4.5 Param_atmos
class shesha_param.Param_atmos
     L0
           L0 per layers in meters.
      alt
           altitudes of each layer.
      deltax
           x translation speed (in pix / iteration) for each layer.
           y translation speed (in pix / iteration) for each layer.
      dim_screens
           linear size of phase screens.
      frac
           fraction of r0 for each layer.
      nscreens
           number of turbulent layers.
     pupixsize
           pupil pixel size (in meters).
           global r0.
      set_L0()
           Set the L0 per layers
               Parameters 1 – (lit of float) : L0 for each layers
      set_alt()
           Set the altitudes of each layer
               Parameters 1 – (lit of float) : altitudes
```

```
set_deltax()
           Set the translation speed on axis x for each layer
               Parameters 1 – (lit of float) : translation speed
      set_deltay()
           Set the translation speed on axis y for each layer
               Parameters 1 – (lit of float) : translation speed
      set_dim_screens()
           Set the size of the phase screens
               Parameters 1 – (lit of float) : phase screens sizes
      set_frac()
           Set the fraction of r0 for each layers
               Parameters 1 - (lit of float) : fraction of r0
      set_nscreens()
           Set the number of turbulent layers
               Parameters n - (long) number of screens.
      set_pupixsize()
           Set the pupil pixel size
               Parameters xsize – (float): pupil pixel size
      set_r0()
           Set the global r0
               Parameters \mathbf{r} – (float) : global r0
      set_seeds()
           Set the seed for each layer
               Parameters 1 – (lit of int) : seed
      set_winddir()
           Set the wind direction for each layer
               Parameters 1 – (lit of float) : wind directions
      set_windspeed()
           Set the the wind speed for each layer
               Parameters 1 – (lit of float) : wind speeds
     winddir
           wind directions of each layer.
     windspeed
           wind speeds of each layer.
1.4.6 Param dm
class shesha_param.Param_dm
      alt
           conjugaison altitude (im m)
      center_name
           filename for influ hdf5 file
      coupling
           actuators coupling (<0.3)
```

```
cube_name
     name for influence cube in hdf5
file influ hdf5
     filename for influ hdf5 file
hyst
     actuators hysteresis (<1.)
margin
     outside margin for actuator select
     inside margin for actuator select
margin_out
     outside margin for actuator select
nact
     number of actuators in the diameter
nkl
     number of kl modes
pupoffset
       2.
push4imat
     nominal voltage for imat
set_alt()
     set the conjugaison altitude
         Parameters a – (float) : conjugaison altitude (im m)
set_center_name()
     set the name of hdf5 influence file
         Parameters filename – (str): Hdf5 file influence name
set_coupling()
     set the actuators coupling
         Parameters c - (float): actuators coupling (<0.3)
set_cube_name()
     set the name of influence cube in hdf5
         Parameters cubename - (str): name of influence cube
set file influ hdf5()
     set the name of hdf5 influence file
         Parameters filename – (str): Hdf5 file influence name
set i1()
     TODO doc
         Parameters i1 – (np.ndarray[ndim=1,dtype=np.int32_t]):
set_influ()
     Set the influence function
         Parameters influ – (np.ndarray[ndim=3,dtype=np.float32_t]): influence function
set_j1()
     TODO doc
         Parameters j1 – (np.ndarray[ndim=1,dtype=np.int32_t]):
```

```
set_margin()
     set the margin for outside actuator select
         Parameters \mathbf{n} – (float): pupille diametre ratio for actuator select
set_margin_in()
     set the margin for inside actuator select (central obstruction)
         Parameters \mathbf{n} – (float): pupille diametre ratio for actuator select
set_margin_out()
     set the margin for outside actuator select
         Parameters n – (float): pupille diametre ratio for actuator select
set nact()
     set the number of actuator
         Parameters n - (long): number of actuators in the diameter
set_ntotact()
     set the total number of actuators
         Parameters n - (long): total number of actuators
set_pattern()
     set the pattern type
         Parameters t - (str): type of pattern
set_push4imat()
     set the nominal voltage for imat
         Parameters p - (float): nominal voltage for imat
set_thresh()
     set the threshold on response for selection
         Parameters t - (float): threshold on response for selection (<1)
set_type()
     set the dm type
         Parameters t - (str): type of dm
set_unitpervolt()
     set the Influence function sensitivity
         Parameters \mathbf{u} – (float): Influence function sensitivity in unit/volt
set_x_name()
     set the name of x coord of influence fonction in file
         Parameters t - (str): name of x coord of influence
set_xpos()
     Set the x positions of influ functions (lower left corner)
         Parameters xpos - (np.ndarray[ndim=1,dtype=np.float32_t]): x positions of influ func-
             tions
set_y_name()
     set the name of y coord of influence fonction in file
         Parameters yname – (str): name of y coord of influence
set_ypos()
     Set the y positions of influ functions (lower left corner)
         Parameters ypos – (np.ndarray[ndim=1,dtype=np.float32_t]): y positions of influ func-
             tions
```

```
threshold on response for selection (<1)
     type_dm
           type of dm
     type_pattern
           type of pattern
     unitpervolt
          Influence function sensitivity in unit/volt. Optional [0.01] Stackarray: mic/volt, Tip-tilt: arcsec/volt.
          name for x coord of influence
     y_name
           name for y coord of influence
1.4.7 Param_target
class shesha_param.Param_target
     Lambda
          observation wavelength for each target
      apod
           boolean for apodizer
     dms_seen
          index of dms seen by the target
     mag
           magnitude for each target
     ntargets
          number of targets
      set_Lambda()
          Set the observation wavelength
               Parameters 1 – (list of float) : observation wavelength for each target
      set_apod()
          Tells if the apodizer is used
          The apodizer is used if a is not 0 :param a: (int) boolean for apodizer
      set_dms_seen()
          set the dms seen by the target
               Parameters 1 – (list of int): index for each dm
      set_mag()
           set the magnitude
               Parameters 1 – (list of float) : magnitude for each target
      set_nTargets()
           Set the number of targets
               Parameters n - (int): number of targets
      set_xpos()
           Set the x positions on sky (in arcsec)
               Parameters 1 - (list of float) : x positions on sky for each target
```

thresh

```
set_ypos()
          Set the y positions on sky (in arcsec)
               Parameters 1 – (list of float): y positions on sky for each target
     set_zerop()
          Set the detector zero point
               Parameters z – (float) : detector zero point
     xpos
          x positions on sky (in arcsec) for each target
          y positions on sky (in arcsec) for each target
     zerop
          target flux for magnitude 0
1.4.8 Param rtc
class shesha_param.Param_rtc
     set centroiders()
          Set the centroiders
               Parameters 1 – (list of Param_centroider) : centroiders settings
     set_controllers()
          Set the controller
               Parameters 1 – (list of Param controller) : controllers settings
     set_nwfs()
          Set the number of wfs
               Parameters n - (int) number of wfs
1.4.9 Param_centroider
class shesha_param.Param_centroider
     interpmat
          optional reference function(s) used for corr centroiding
          number of brightest pixels
     nwfs
          index of wfs in y wfs structure on which we want to do centroiding
     set_nmax()
          Set the number of brightest pixels to use for bpcog
               Parameters n: (int): number of brightest pixels
     set nwfs()
          Set the index of wfs
               Parameters n - (int): index of wfs
      set_sizex()
          Set sizex parameters for corr centroider (interp_mat size)
               Parameters s: (long): x size
```

```
set_sizey()
          Set sizey parameters for corr centroider (interp_mat size)
               Parameters s: (long): y size
      set_thresh()
          Set the threshold for tcog
               Parameters t: (float): threshold
      set_type()
          Set the centroider type :param t: (str) : centroider type
      set_type_fct()
          Set the type of ref function
               Parameters \mathbf{f} – (str) : type of ref function
      set_weights()
           Set the weights to use with wcog or corr
               Parameters w: (np.ndarray[ndim=3 ,dtype=np.float32_t]): weights
      set_width()
           Set the width of the Gaussian
               Parameters \mathbf{w} – (float): width of the gaussian
      sizex
          x-size for inter mat (corr)
      sizey
          x-size for inter mat (corr)
     thresh
          Threshold
     type_centro
          type of centroiding cog, tcog, bpcog, wcog, corr
     type_fct
           type of ref function gauss, file, model
     weights
          optional reference function(s) used for centroiding
     width
           width of the Gaussian
1.4.10 Param_controller
class shesha_param.Param_controller
      TTcond
           tiptilt condition number for cmat filtering with mv controller
      cmat
          full control matrix
      cured ndivs
          subdivision levels in cured
     delay
          loop delay [frames]
     gain
          loop gain
```

```
gmax
     Maximum gain for modal optimization
gmin
     Minimum gain for modal optimization
imat
     full interaction matrix
maxcond
     max condition number
modopti
     Flag for modal optimization
nactu
     number of controled actuator per dm
ndm
     index of dms in controller
ngain
     Number of tested gains
nmodes
     Number of modes for M2V matrix (modal optimization)
     Number of sample of open loop slopes for modal optimization computation
nvalid
     number of valid subaps per wfs
nwfs
     index of wfss in controller
set TTcond()
     Set the tiptilt condition number for cmat filtering with mv controller
     :param : (float) : tiptilt condition number
set_cmat()
     Set the full control matrix
         Parameters cmat – (np.ndarray[ndim=2,dtype=np.float32_t]): full control matrix
set cured ndivs()
     Set the subdivision levels in cured
         Parameters c - (long): subdivision levels in cured
set delay()
     Set the loop delay expressed in frames
         Parameters d: (float) :delay [frames]
set_gain()
     Set the loop gain
         Parameters g: (float): loop gain
set_gmax()
     Set the maximum gain for modal optimization
         Parameters g – (flaot) : maximum gain for modal optimization
set_gmin()
     Set the minimum gain for modal optimization
```

Parameters g - (float): minimum gain for modal optimization

Chapter 1. Contents:

```
set_imat()
          Set the full interaction matrix
              Parameters imat – (np.ndarray[ndim=2,dtype=np.float32_t]): full interaction matrix
      set maxcond()
          Set the max condition number
          :param: (float): max condition number
      set modopti()
          Set the flag for modal optimization
              Parameters m - (int): flag for modal optimization
      set nactu()
          Set the number of controled actuator
              Parameters 1 – (list of int): number of controlled actuator per dm
      set_ndm()
          Set the indices of dms
              Parameters 1 – (list of int) : indices of dms
      set_ngain()
          Set the number of tested gains
              Parameters n - (int): number of tested gains
     set_nkl()
          Set the number of KL modes used for computation of covmat in case of minimum variance controller
              Parameters n – (long): number of KL modes
     set nmodes()
          Set the number of modes for M2V matrix (modal optimization)
              Parameters n - (int): number of modes
      set_nrec()
          Set the number of sample of open loop slopes for modal optimization computation
              Parameters n - (int): number of sample
     set_nvalid()
          Set the number of valid subaps
              Parameters 1 – (list of int): number of valid subaps per wfs
      set nwfs()
          Set the indices of wfs
              Parameters 1 – (list of int): indices of wfs
     type_control
          type of controller
1.4.11 get_classAttributes()
shesha_param.get_classAttributes(Param_class)
     Return all the attribute names of the given Param_class
          Parameters Param_class - shesha parameters class
          Returns list of strings (attributes names)
          Example import shesha as ao get_classAttributes(ao.Param_wfs)
```

1.4.12 indices()

```
shesha_param.indices()
DOCUMENT indices(dim)
Return a dimxdimx2 array.
contains the Y indices. Inspecan either be:
```

Return a dimxdimx2 array. First plane is the X indices of the pixels in the dimxdim array. Second plane contains the Y indices. Inspired by the Python scipy routine of the same name. New (June 12 2002), dim can either be:

- •a single number N (e.g. 128) in which case the returned array are square (NxN)
- •a Yorick array size, e.g. [#dimension,N1,N2], in which case the returned array are N1xN2
- •a vector [N1,N2], same result as previous case

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```
Parameters dim1: (int): first dimension dim2: (int): (optional) second dimension
```

1.4.13 make_apodizer()

1.4.14 makegaussian()

```
shesha\_param.makegaussian(size, fwhm, xc, yc)
```

Returns a centered gaussian of specified size and fwhm. norm returns normalized 2d gaussian

```
Parameters size: (int):

fwhm: (float):

xc: (int): (optional) center position on x axis

yc: (int): (optional) center position on y axis

norm: (int): (optional) normalization
```

1.4.15 rotate()

```
shesha_param.rotate()
```

Rotates an image of an angle "ang" (in DEGREES).

The center of rotation is cx,cy. A zoom factor can be applied.

(cx,cy) can be omitted :one will assume one rotates around the center of the image. If zoom is not specified, the default value of 1.0 is taken.

```
Parameters im: (np.ndarray[ndim=3,dtype=np.float32_t]) : array to rotate ang: (float) : rotation angle (in degrees)
cx: (int) : (optional) rotation center on x axis (default: image center)
cy: (int) : (optional) rotation center on x axis (default: image center)
```

```
zoom: (float): (opional) zoom factor (default =1.0)
```

1.4.16 rotate3d()

```
shesha_param.rotate3d()
     Rotates an image of an angle "ang" (in DEGREES).
     The center of rotation is cx,cy. A zoom factor can be applied.
      (cx,cy) can be omitted :one will assume one rotates around the center of the image. If zoom is not specified,
     the default value of 1.0 is taken.
     modif dg: allow to rotate a cube of images with one angle per image
           Parameters im: (np.ndarray[ndim=3,dtype=np.float32_t]) : array to rotate
               ang: (np.ndarray[ndim=1,dtype=np.float32_t]): rotation angle (in degrees)
               cx: (int): (optional) rotation center on x axis (default: image center)
               cy: (int): (optional) rotation center on x axis (default: image center)
               zoom: (float): (opional) zoom factor (default =1.0)
1.5 shesha rtc
1.5.1 Rtc
class shesha_rtc.Rtc
      add Controller()
           Add a controller in the sutra_controller vector of the RTC on the GPU
               Parameters nactu: (int): number of actuators
                   delay: (float): loop delay
                   type_control: (str): controller's type
                   dms: (Dms): sutra_dms object (GPU)
                   type dmseen: (char**): dms indices controlled by the controller
                   alt: (np.ndarray[ndim=1,dtype=np.float32_t]): altitudes of the dms seen
                   ndm: (int): number of dms controled
                   Nphi: (long): number of pixels in the pupil (used in geo controler case only)
      add_centroider()
           Add a centroider in the sutra_centroiders vector of the RTC on the GPU
               Parameters sensor: (Sensors): sutra_sensors object (GPU)
                   nwfs: (long): number of wfs
                   nvalid: (long): number of valid subaps
                   type_centro: (str): centroider's type
```

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offset: (float): scale: (float):

```
applycontrol()
     Compute the DMs shapes from the commands computed in a sutra_controller_object. From the com-
     mand vector, it computes the voltage command (adding pertrubation voltages, taking delay into ac-
     count) and then apply it to the dms
         Parameters ncontro: (int): controller index
buildcmat()
     Compute the command matrix in a sutra controller ls object
         Parameters ncontro: (int): controller index
             nfilt: (int): number of modes to filter
             filt_tt: (int): (optional) flag to filter TT
buildcmatmv()
     Compute the command matrix in a sutra_controller_mv object
         Parameters ncontro: (int): controller index
             cond: (float): conditioning factor for the Cmm inversion
docentroids()
     Compute the centroids with sutra_controller #ncontrol object
         Parameters ncontrol: (optional) controller's index
docentroids geom()
     Compute the geometric centroids with sutra_controller #ncontrol object
         Parameters ncontrol: (optional) controller's index
docontrol()
     Compute the command to apply on the DMs on a sutra_controller object
         Parameters ncontro: (int): controller index
docontrol_geo()
     Compute the command to apply on the DMs on a sutra_controller_geo object for the target direction
         Parameters ncontro: (int): controller index
docontrol_geo_onwfs()
     Compute the command to apply on the DMs on a sutra_controller_geo object for the wfs direction
         Parameters ncontro: (int): controller index
doimat()
     Compute the interaction matrix
         Parameters ncontro: (int): controller index
             g dms: (Dms): Dms object
doimat_geom()
     Compute the interaction matrix by using a geometric centroiding method
         Parameters ncontro: (int): controller index
              g dms: (Dms): Dms object
             geom: (int): type of geometric method (0 or 1)
getCenbuff()
     Return the centroids buffer from a sutra_controller_ls object. This buffer contains centroids from
     iteration i-delay to current iteration.
         Parameters ncontro: (int): controller index
```

Returns data: (np.ndarray[ndim=2,dtype=np.float32_t]): centroids buffer

```
getCentroids()
     Return the centroids computed by the sutra_rtc object
         Parameters ncontrol: (int): controller's index
         Returns data : (np.ndarray[ndim=1,dtype=np.float32_t]) : centroids (arcsec)
getCmmEigenvals()
     Return the eigen values of the Cmm decomposition in a sutra_controller_mv object
         Parameters ncontro: (int): controller index
         Returns eigenvals: (np.ndarray[ndim=1,dtype=np.float32 t]): eigenvalues
getCom()
     Return the command vector from a sutra_controller object
         Parameters ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=1,dtype=np.float32_t]): command vector
getEigenvals()
     Return the eigen values of the imat decomposition in a sutra_controller object
         Parameters ncontro: (int): controller index
         Returns eigenvals: (np.ndarray[ndim=1,dtype=np.float32_t]): eigenvalues
getErr()
     Return the command increment (cmat*slopes) from a sutra_controller_ls object
         Parameters ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=1,dtype=np.float32 t]): command increment
getU()
     Return the eigen modes matrix of the imat decomposition from a sutra_controller_ls object
         Parameters ncontro: (int): controller index
         Returns U: (np.ndarray[ndim=2,dtype=np.float32_t]): eigen modes matrix
getVoltage()
     Return the voltage vector that will be effectively applied to the DMs
         Parameters ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=1,dtype=np.float32_t]): voltage vector
get_IFsparse()
     Get the influence functions matrix computed by the geo controller Return a scipy.sparse object which
     shape is (nactus, Npts in the pupil)
         Parameters ncontro: (int): controller index
         Returns IF: (scipy.sparse): influence functions matrix
get cmat()
     Return the command matrix from a sutra_controller object
         Parameters ncontro: (int): controller index
         Returns cmat: (np.ndarray[ndim=2,dtype=np.float32_t]): command matrix
get_cmm()
     Return the Cmm matrix from a sutra_controller_mv object
         Parameters ncontro: (int): controller index
         Returns cmm: (np.ndarray[ndim=2,dtype=np.float32_t]): Cmm matrix
get_cphim()
     Return the Cphim matrix from a sutra_controller_mv object
```

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```
:parameters; ncontro: (int): controller index
         Returns cphim: (np.ndarray[ndim=2,dtype=np.float32_t]): Cphim matrix
get imat()
     Return the interaction matrix of a sutra controller object
         Parameters ncontro: (int): controller index
         Returns imat : (np.ndarray[ndim=2,dtype=np.float32_t]) : interaction matrix
get mgain()
     Return modal gains from sutra controller
         Parameters ncontro: (int): controller index
         Returns mgain: (np.ndarray[ndim=1,dtype=np.float32_t]): modal gains
get_nfiltered()
     Get the number of filtered modes for cmat computation
         Parameters ncontro: (int): controller index p_rtc: (Param_rtc): rtc parameters
getcentroids()
     Return the centroids computed by the sutra_rtc object If ncontrol <= d_control.size, return
     rtc.d_centroids Else, compute centroids from wfs[nwfs] with centroider[ncontrol]
         Parameters ncontrol: (int): controller's index
             g_wfs: (Sensors): (optional) sutra_sensors object
             nwfs: (int): (optional) number of wfs
         Returns data: (np.ndarray[ndim=1,dtype=np.float32 t]): centroids (arcsec)
getolmeas()
     Return the reconstructed open-loop measurement from a sutra_controller_mv object
         Parameters ncontro: (int): controller index
         Returns data: (np.ndarray[ndim=1,dtype=np.float32_t]): reconstructed open-loop
     Compute the singular value decomposition of the interaction matrix
         Parameters ncontro - controller index
init modalOpti()
     Initialize the modal optimization controller: compute the slopes-to-modes matrix and the transfer
     functions
         Parameters ncontro: (int): controller index
             nmodes: (int): number of modes
             nrec: (int): number of recorded open slopes measurements
             M2V: (np.ndarray[ndim=2,dtype=np.float32_t]): modes-to-volt matrix
             gmin: (float): minimum gain for modal optimization
             gmax: (float): maximum gain for modal optimization
             ngain: (int): Number of tested gains
             Fs: (float): sampling frequency
init_proj()
     Initialize the projection matrix for sutra_controller_geo object. The projection matrix is (IFt.IF)**(-1)
     * IFt where IF is the DMs influence functions matrix
```

```
Parameters ncontro: (int): controller index
             dms: (Dms): Dms object
             indx_dm: (np.ndarray[ndim=1,dtype=np.int32_t]): indices of where(pup) on DM screen
             unitpervolt: (np.ndarray[ndim=1,dtype=np.float32_t]): unitpervolt DM parameter
             indx_pup: (np.ndarray[ndim=1,dtype=np.int32_t]) : indices of where(pup) on ipupil
              screen
loadOpenLoop()
     Load an array of recoded open-loop measurements for modal optimization
         Parameters ncontro: (int): controller index
             ol_slopes: (np.ndarray[ndim=2, dtype=np.float32_t]): open-loop slopes
loadnoisemat()
     Load the noise vector on a sutra_controller_mv object
         Parameters ncontro: (int): controller index
             N: (np.ndarray[ndim=1,dtype=np.float32_t]): noise vector
modalControlOptimization()
     Compute the command matrix with modal control optimization
         Parameter ncontro: controller index
rmcontrol()
     Remove a controller
sensors compslopes()
     Compute the slopes in a sutra wfs object. This function is equivalent to docentroids() but the centroids
     are stored in the sutra_wfs object instead of the sutra_rtc object
         Parameters ncentro: (int): centroider index
sensors initbcube()
     Initialize npix in the sutra_centroider_corr object (useless ?)
         Parameters ncentro: (int): centroider's index
sensors_initcorr()
     Initialize sutra_centroider_corr oblect
         Parameters ncentro: (int): centroider's index
              w: (np.ndarray[ndim=1,dtype=np.float32_t]): weight
             corr\_norm: (np.ndarray[ndim=2,dtype=np.float32\_t]):
             sizex: (int):
             sizey: (int):
             interpmat: ([ndim=2,dtype=np.float32_t]) :
sensors_initweights()
     Load the weight array in sutra_centroider_wcog object
         Parameters ncentro: (int): centroider's index
              w: (np.ndarray[ndim=2, dtype=np.float32_t]): weight
setCentroids()
     Set the centroids vector of a sutra_controller object to centro
         Parameters ncontro:
                                         (int)
                                                               controller
                                                                              index
                                                                                          centro:
             (np.ndarray[ndim=1,dtype=np.float32_t]): centroids vector
```

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```
setCom()
     Set the command vector of a sutra_controller object to comvec
         Parameters ncontro: (int): controller index
setEigenvals()
     Set the eigen values of the imat decomposition in a sutra controller ls object
         Parameters ncontro: (int): controller index
             eigenvals: (np.ndarray[ndim=1,dtype=np.float32 t]): eigen values
setU()
     Set the eigen modes matrix of the imat decomposition in a sutra controller ls object
         Parameters ncontro: (int): controller index
             U: (np.ndarray[ndim=2,dtype=np.float32_t]): eigen modes matrix
set_cmat()
     Set the command matrix on a sutra_controller object
         Parameters ncontro: (int): controller index
             data: (np.ndarray[ndim=2,dtype=np.float32_t]): command matrix to use
set cmm()
     Set the Cmm matrix on a sutra controller my object
         Parameters ncontro: (int): controller index
             data: (np.ndarray[ndim=2,dtype=np.float32_t]): Cmm matrix
set decayFactor()
     Set the decay factor on a sutra_controller_generic object
         Parameters ncontro: (int): controller index
             decay: (np.ndarray[ndim=1,dtype=np.float32_t]): ask to Rico
set_gain()
     Set the loop gain in sutra_controller object
         Parameters ncontro: (int): controller index
             gain: (float): loop gain
set imat()
     Set the interaction matrix on a sutra controller object
         Parameters ncontro: (int): controller index
             data: (np.ndarray[ndim=2,dtype=np.float32_t]): interaction matrix to use
set matE()
     Set the matrix E on a sutra_controller_generic object
         Parameters ncontro: (int): controller index
             matE: (np.ndarray[ndim=2,dtype=np.float32_t]): ask to Rico
set mgain()
     Set modal gains in sutra_controller object
         Parameters ncontro: (int): controller index
             mgain: (np.ndarray[ndim=1,dtype=np.float32_t]): modal gains
set openloop()
     Set the openloop state to a sutra_controller object
         Parameters ncontro: (int): controller index
             openloop: state of the controller
```

```
setnmax()
           set the number of brightest pixels to consider for bpcog centroider
               Parameters ncentro: (int): centroider's index
                   nmax: (int): number of brightest pixels
      setthresh()
           set threshold for the centroider #ncentro
               Parameters ncentro: (int): centroider's index
                   thresh: (float): threshold
1.5.2 cmat init()
shesha_rtc.cmat_init()
     Compute the command matrix on the GPU
           Parameters ncontro: (int):
               g_rtc: (Rtc):
               p_rtc: (Param_rtc) : rtc settings
               p_wfs: (list of Param_wfs): wfs settings
               p_tel : (Param_tel) : telescope settings
               clean: (int): (optional) clean datafiles (imat, U, eigenv)
               simul_name: (str): (optional) simulation's name, use for data files' path
               load: (dict): (optional) dictionary of matrices to load and their path
1.5.3 compute_KL2V()
shesha_rtc.compute_KL2V()
     Compute the Karhunen-Loeve to Volt matrix (transfer matrix between the KL space and volt space for a pzt
     dm)
           Parameters p_dms: (list of Param_dm): dms settings
               controller: (Param_controller): controller settings
1.5.4 correct_dm()
shesha_rtc.correct_dm()
     Correct the geometry of the DMs using the imat (filter unseen actuators)
           Parameters p_dms: (list of Param_dm): dms settings
               g_dms: (Dms): Dms object
               p_control: (Param_controller): controller settings
               p_geom: (Param_geom): geom settings
               imat: (np.ndarray): interaction matrix
               simul_name: (str): simulation's name, use for data files' path
               load: (dict): (optional) dictionary of matrices to load and their path
```

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```
1.5.5 create_interp_mat()
shesha_rtc.create_interp_mat()
     TODO doc
         Parameters dimx: (int):
             dimy: (int):
1.5.6 create nact geom()
shesha_rtc.create_nact_geom()
     Compute the DM coupling matrix
         Param p_dms: (list of Param_dm): dms parameters ndm: (int): dm number
         Returns Nact: (np.array(dtype=np.float64)): the DM coupling matrix
1.5.7 create_piston_filter()
shesha_rtc.create_piston_filter()
1.5.8 doTomoMatrices()
shesha_rtc.doTomoMatrices()
     Compute Cmm and Cphim matrices for the MV controller on GPU
         Parameters g_wfs: (Sensors):
             p_wfs: (list of Param_wfs): wfs settings
             g_dms: (Dms):
             p_dms: (list of Param_dms): dms settings
             p_geom: (Param_geom) : geom settings
             p_atmos: (Param_atmos): atmos settings
             g_atmos: (Atmos):
             p_tel: (Param_tel) : telescope settings
1.5.9 get_r0()
shesha_rtc.get_r0()
     Compute r0 at lambda2 from r0 value at lambda1
         Parameters r0_at_lambda1: (float): r0 value at lambda1
             lambda1: (float): lambda1
             lambda2: (float): lambda2
1.5.10 imat_geom()
shesha_rtc.imat_geom()
     Compute the interaction matrix with a geometric method
```

```
Parameters g_wfs: (Sensors): Sensors object
               p_wfs: (list of Param_wfs): wfs settings
               p_control: (Param_controller): controller settings
               g_dms: (Dms): Dms object
               p_dms: (list of Param_dm): dms settings
               meth: (int): (optional) method type (0 or 1)
1.5.11 imat_init()
shesha_rtc.imat_init()
     Initialize and compute the interaction matrix on the GPU
          Parameters ncontro: (int): controller's index
               g_rtc: (Rtc): Rtc object
               p_rtc: (Param_rtc) : rtc settings
               g_dms: (Dms): Dms object
               g_wfs: (Sensors): Sensors object
               p_wfs: (list of Param_wfs): wfs settings
               p_tel: (Param_tel) : telescope settings
               clean: (int): (optional): clean datafiles (imat, U, eigenv)
               simul_name: (str): (optional) simulation's name, use for data files' path
               load: (dict): (optional) dictionary of matrices to load and their path
1.5.12 manual imat()
shesha_rtc.manual_imat()
     Compute the interaction matrix 'manually', ie without sutra_rtc doimat method
          Parameters g_rtc: (Rtc): Rtc object
               g_wfs: (Sensors): Sensors object
               g_dms: (Dms): Dms object
               p_dms: (list of Param_dm): dm settings
1.5.13 openLoopSlp()
shesha_rtc.openLoopSlp()
     Return a set of recorded open-loop slopes, usefull for modal control optimization
           Parameters g_tel: (Telescope): Telescope object
               g_atm: (Atmos): Atmos object
               g_rtc: (Rtc): Rtc object
               nrec: (int): number of samples to record
               ncontro: (int): controller's index
               g_wfs: (Sensors): Sensors object
               p_wfs: (list of Param_wfs): wfs settings
```

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```
p_tar: (Param_target) : target settings
               g_tar: (Target): Target object
1.5.14 rtc_init()
shesha_rtc.rtc_init()
      Initialize all the sutra_rtc objects : centroiders and controllers
           Parameters g_tel: (Telescope): Telescope object
               g_wfs: (Sensors): Sensors object
               p_wfs: (list of Param_wfs) : wfs settings
               g_dms: (Dms): Dms object
               p_dms: (list of Param_dms): dms settings
               p_geom: (Param_geom) : geom settings
               p_atmos: (Param_atmos): atmos settings
               g_atmos: (Atmos): Atmos object
               p_tel: (Param_tel) : telescope settings
               p_loop: (Param_loop) : loop settings
               p_tar: (Param_target) : (optional) target settings
               clean: (int): (optional) clean datafiles (imat, U, eigenv, pztok, pztnok)
               brama: (int): (optional) not implemented yet
               doimat: (int): (optional) force imat computation
               simul_name: (str): (optional) simulation's name, use for path to save data (imat, U...)
               load: (dict): (optional) dictionary of matrices to load and their path
           Returns Rtc: (Rtc): Rtc object
```

1.6 shesha sensors

1.6.1 Sensors

```
class shesha sensors. Sensors
```

Constructor: Sensors(nsensors,type_data,npup,nxsub,nvalid,nphase,pdiam,npix,nrebin,nfft,nftota,nphot,lgs,odevice,comm_si

```
Parameters nsensors: (int):
    type_data: list of strings):
    npup: (np.ndarray[ndim=1,dtype=np.int64_t]):
    nxsub: (np.ndarray[ndim=1,dtype=np.int64_t]):
    nvalid: (np.ndarray[ndim=1,dtype=np.int64_t]):
    nphase: (np.ndarray[ndim=1,dtype=np.int64_t]):
    pdiam: (np.ndarray[ndim=1,dtype=np.int64_t]):
    npix: (np.ndarray[ndim=1,dtype=np.int64_t]):
    nrebin: (np.ndarray[ndim=1,dtype=np.int64_t]):
    nfft: (np.ndarray[ndim=1,dtype=np.int64_t]):
```

```
ntota: (np.ndarray[ndim=1,dtype=np.int64_t]):
         nphot: (np.ndarray[ndim=1,dtype=np.float32_t]) :
         nphot4imat: (np.ndarray[ndim=1,dtype=np.float32_t]) :
         lgs: (np.ndarray[ndim=1,dtype=np.int32_t]):
         odevice: (int):
         comm size: (int): MPI communicator size
         rank: (int): process rank
get amplifoc()
     Return the 'amplifoc' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'amplifoc' from
get_amplifoc_pyr()
     Return the 'amplifoc' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'amplifoc' from
get_bincube()
     Return the 'bincube' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'bincube' from
get bincubeNotNoisy()
    Return the 'bincube_not_noisy' array of a given wfs. It's the bincube before noise has been added
         Parameters n – (int): number of the wfs to get the 'bincube_not_noisy' from
get binimg()
    Return the 'binimg' array of a given wfs
    :param n: (int) :number of the wfs to get the 'binimg' from
    :options for raw image computation tel (Telescope): shesha telescope atmos (Atmos): shesha at-
         mos dms (Dms): shesha dms
get_camplipup()
     Return the 'camplipup' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'camplipup' from
get_camplipup_pyr()
     Return the 'camplipup' array of a given wfs in the pyr case
         Parameters n - (int): number of the wfs to get the 'camplipup' from
get_ftlgskern()
    Return the ftlgskern array of a given wfs
         Parameters n - (int): number of the wfs to get the phase from
get_fttotim_pyr()
    Return the 'fttotim' array of a given wfs
         Parameters \mathbf{n} – (int): number of the wfs to get the 'amplifoc' from
get_hrimg_pyr()
     Return the phase array of a given wfs
         Parameters n - (int): number of the wfs to get the phase from
get_imgtele()
     Return the 'image_telemetry' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'image_telemetry' from
```

```
:options for raw image computation tel (Telescope): shesha telescope atmos (Atmos): shesha at-
         mos dms (Dms): shesha dms
get_lgskern()
     Return the lgskern array of a given wfs
         Parameters n - (int): number of the wfs to get the phase from
get offsets()
     Return the 'offset' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'offset' from
get_phase()
     Return the phase array of a given wfs
         Parameters \mathbf{n} – (int): number of the wfs to get the phase from
get_pyrimg()
     Return the image of a pyr wfs
         Parameters \mathbf{n} – (int): number of the wfs to get the image from
get_pyrimghr()
     Return the high res image of a pyr wfs
         Parameters n - (int): number of the wfs to get the image from
get rank()
     Return the rank of one of the sensors wfs
         Parameters n - (int): index of the wfs to get the rank for
get slopes()
     Return the 'slopes' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'slopes' from
get_subsum()
     Return the 'subsum' array of a given wfs
         Parameters n - (int): number of the wfs to get the 'subsum' from
reset_phase()
     Reset the phase's array of a given wfs
         Parameters \mathbf{n} – (int) : index of the given wfs
sensors_addlayer()
     Call function add_layer from the sutra_source of a sutra_wfs of the Sensors
         Parameters i: (int):
             type_dm: (string):
             alt: (float):
             xoff: (float):
             yoff: (float):
sensors_compimg()
     TODO doc
         Parameters n - (in): index of the wfs
sensors_initarr()
     Call the function wfs_initarrays from a sutra_wfs of the Sensors
         Parameters n: (int): index of the wfs
              wfs: (Param_wfs):
```

```
sensors_initgs()
           Call the function sensors_initgs
               Parameters xpos: (np.ndarray[ndim=1,dtype=np.float32_t]):
                   ypos: (np.ndarray[ndim=1,dtype=np.float32_t]) :
                   Lambda: (np.ndarray[ndim=1,dtype=np.float32_t]):
                   mag: (np.ndarray[ndim=1,dtype=np.float32 t]):
                   zerop: (float):
                   size: (np.ndarray[ndim=1,dtype=np.int64_t]):
                   noise: (np.ndarray[ndim=1,dtype=np.float32_t]):
                   seed: (np.ndarray[ndim=1,dtype=np.int64_t ]) :
      sensors_trace()
           Does the raytracing for the wfs phase screen in sutra_wfs
               Parameters n: (int):
                   type_trace: (str) ["all"][raytracing across atmos and dms seen] "dm": raytracing
                      across dms seen only "atmos": raytracing across atmos only
                   tel: (Telescope):(optional) Telescope object
                   atmos: (Atmos):(optional) Atmos object
                   dms: (Dms): (optional) Dms object
                   rst: (int): (optional) reset before raytracing if rst = 1
      set bincube()
           Set the bincube of the WFS numner n
               Parameters n: (int): WFS number data: (np.ndarray[ndim=3,dtype=np.float32_t]): bin-
                   cube to use
      set_phase()
           Set the phase array of a given wfs
               Parameters
                   • \mathbf{n} – (int): number of the wfs to get the phase from
                    • data – (np.ndarray) : the phase to set
      slopes_geom()
           Compute the geometric slopes in a sutra_wfs object
               Parameters nsensor: (int): wfs number
                      param t (int): method (0 or 1)
1.6.2 bin2d()
shesha_sensors.bin2d()
     Returns the input 2D array "array", binned with the binning factor "binfact". The input array X and/or Y
     dimensions needs not to be a multiple of "binfact"; The final/edge pixels are in effect replicated if needed.
     This routine prepares the parameters and calls the C routine _bin2d. The input array can be of type long,
      float or double. Last modified: Dec 15, 2003. Author: F.Rigaut SEE ALSO: _bin2d
           Parmeters data_in: (np.ndarray): data to binned
```

binfact: (int): binning factor

```
1.6.3 fft_goodsize()
shesha_sensors.fft_goodsize()
     find best size for a fft from size s
          Parameters s: (long) size
1.6.4 init_wfs_geom()
shesha_sensors.init_wfs_geom()
     Compute the geometry of WFSs: valid subaps, positions of the subaps, flux per subap, etc...
          Parameters wfs: (Param_wfs) : wfs settings
              wfs0: (Param wfs): reference wfs settings
              n: (int): index of the wfs (diplay information purpose only)
              atmos: (Param_atmos): atmos settings
              tel: (Param_tel): telescope settings
              geom: (Param_geom): geom settings
              target: (Param_target): target settings
              loop: (Param_loop): loop settings
              init: (int): (optional)
              verbose: (int): (optional) display informations if 0
1.6.5 make_lgs_prof1d()
shesha_sensors.make_lgs_prof1d()
     same as prep_lgs_prof but cpu only. original routine from rico
          Parameters p_tel: (Param_tel) : telescope settings
              prof: (np.ndarray[dtype=np.float32]): Na profile intensity, in arbitrary units
              h: (np.ndarray[dtype=np.float32]) : altitude, in meters. h MUST be an array with
              EQUALLY spaced elements.
              beam: (float): size in arcsec of the laser beam
              center: (string): either "image" or "fourier" depending on where the centre should be.
1.6.6 noise cov()
shesha_sensors.noise_cov()
     Compute the diagonal of
                                   the noise covariance matrix for a SH WFS
                           (pi^2/2)*(1/Nphotons)*(d/r0)^2
                                                         / (2*pi*d/lambda)^2
                                                                                   Electronic
     (pi^2/3)*(wfs.noise^2/N^2photons)*wfs.npix^2*(wfs.npix*wfs.pixsize*d/lambda)^2 / (2*pi*d/lambda)^2
          Parameters nw: wfs number p_wfs: (Param_wfs) : wfs settings p_atmos: (Param_atmos) :
              atmos settings p_tel: (Param_tel): telescope settings
          Returns cov: (np.ndarray(ndim=1,dtype=np.float64)): noise covariance diagonal
```

1.6.7 prep_lgs_prof()

```
shesha_sensors.prep_lgs_prof()
```

The function returns an image array(double,n,n) of a laser beacon elongated by perpective effect. It is obtaind by convolution of a gaussian of width "lgsWidth" arcseconds, with the line of the sodium profile "prof". The altitude of the profile is the array "h".

```
parameters nsensors: (int) : wfs index
    p_tel: (Param_tel) : telescope settings
    prof: (np.ndarray[dtype=np.float32]) : Na profile intensity, in arbitrary units
    h: (np.ndarray[dtype=np.float32]) : altitude, in meters. h MUST be an array with
    EQUALLY spaced elements.
    beam: (float) : size in arcsec of the laser beam
    center: (string) : either "image" or "fourier" depending on where the centre should
    be.
```

Computation of LGS spot from the sodium profile: Everything is done here in 1D, because the Na profile is the result of the convolution of a function $P(x,y) = \operatorname{profile}(x)$. dirac(y) by a gaussian function, for which variables x and y can be split: $\exp(-(x^2+y^2)/2.s^2) = \exp(-x^2/2.s^2) * \exp(-y^2/2.s^2)$ The convolution is (symbol \$ denotes integral) $C(X,Y) = \$ \exp(-x^2/2.s^2) * \exp(-y^2/2.s^2) * \operatorname{profile}(x-X) * \operatorname{dirac}(y-Y) dx dy$ First one performs the integration along y $C(X,Y) = \exp(-Y^2/2.s^2) * \exp(-x^2/2.s^2) * \operatorname{profile}(x-X) dx$ which shows that the profile can be computed by - convolving the 1-D profile - multiplying it in the 2nd dimension by a gaussian function

If one has to undersample the inital profile, then some structures may be "lost". In this case, it's better to try to "save" those structures by re-sampling the integral of the profile, and then derivating it afterwards. Now, if the initial profile is a coarse one, and that one has to oversample it, then a simple re-sampling of the profile is adequate.

1.6.8 type_present()

```
shesha_sensors.type_present()
      Check the present types in a list
           Parameters liste: (list of str): list of types
                pyr: (int): set to 1 if the list contains "pyr" (0 else)
                roof: (int): set to 1 if the list contains "roof" (0 else)
                sh: (int): set to 1 if the list contains "sh" (0 else)
                geo: (int): set to 1 if the list contains "geo" (0 else)
      return 1 if the wfs type is present (0 else)
1.6.9 wfs init()
shesha_sensors.wfs_init()
      Create and initialise a Sensors object
           Parameters wfs: (list of Param_wfs) : wfs settings
                p_atmos: (Param_atmos): atmos settings
                p_tel: (Param_tel) : telescope settings
                p_geom: (Param_geom) : geom settings
                p target: (Param target): target settings
```

```
p_loop: (Param_loop) : loop settings
               dm: (list of Param_dm): (optional) dms settings [=None]
               comm_size: (int) : (optional) communicator size [=1]
               rank: (int): (optional) process rank [=0]
1.6.10 wheremax()
shesha_sensors.wheremax()
     return the index of the maximum value of the list
           Parameters liste – (list of values): values to get the index of the maximum from
1.7 shesha target
1.7.1 Target
class shesha_target.Target
     Lambda
           observation wavelength for each target
     add_layer()
           Add a phase screen dm or atmos as layers of turbulence
               Parameters n: (int): index of the target
                   1_type: (str): "atmos" or "dm"
                   alt: (float): altitude
                   xoff: (float): x-offset
                   yoff: (float): y-offset
     apod
          boolean for apodizer
     atmos_trace()
           Raytracing of the target through the atmosphere
               Parameters nTarget: (int): index of the target
                   atm: (atmos): atmos to get through
                   tel: (Telescope): telescope
     dmtrace()
          Raytracing of the target through thedms
               Parameters ntar: (int): index of the target
                   dms: (Dms): dms to go through
                   reset: (int): if >0, reset the screen before raytracing
                   do_phase_var: (int): if 0, doesn't take the screen into account in the phase average
                   (unused)
     get_amplipup()
           Return the complex amplitude in the pupil plane of the target.
```

Parameters nTarget - (int): index of the target

```
get_image()
     Return the image from the target (or long exposure image according to the requested type)
         Parameters nTarget: (int): index of the target
             type_im: (str): type of the image to get ("se" or "le")
             puponly: (int): if 1, image computed from phase on the pupil only
              comp le: (bool): if False (default), the computed image is not taken into account in the
             LE image
get_phase()
     Return the phase's screen of the target
         Parameters nTarget - (int): index of the target
get_phasetele()
     Return the telemetry phase of the target
         Parameters nTarget - (int): index of the target
         Return data (np.ndarray(ndim=2,np.float32)): phase screen
get_strehl()
     Compute and return the target's strehl
         Parameters nTarget – (int): index of the target
         Return strehl (np.array(4,dtype=np.float32)) : [Strehl SE, Strehl LE, instantaneous phase
              variance over the pupil, average phase variance over the pupil]
init_strehlmeter()
     Initialise target's strehl
         Parameters nTarget - (int): index of the target
mag
     magnitude for each target
ntargets
     number of targets
reset_phase()
     Reset the phase's screen of the target
         Parameters nTarget - (int): index of the target
reset_strehl()
     Reset the target's strehl
         Parameters nTarget – (int): index of the target
set phase()
     Set the phase's screen of the target
         Parameters
              • nTarget – (int) : index of the target
              • data – (np.ndarray[ndim=2,dtype=np.float32_t]) : phase screen
xpos
     x positions on sky (in arcsec) for each target
ypos
     y positions on sky (in arcsec) for each target
```

1.7.2 target_init()

```
shesha_target.target_init()
Create a cython target from parametres structures

Parameters ctxt: (naga_context):

atm: (Param_atmos): atmos settings
geom: (Param_geom): geom settings
wfs: (Param_wfs): wfs settings
dm: (Param_dm): dm settings
```

CHAPTER

TWO

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