

# Optimisation of gas composition and amplification stage for the CYGNO/INITIUM experiment



Giorgio Dho on behalf of CYGNO coll.

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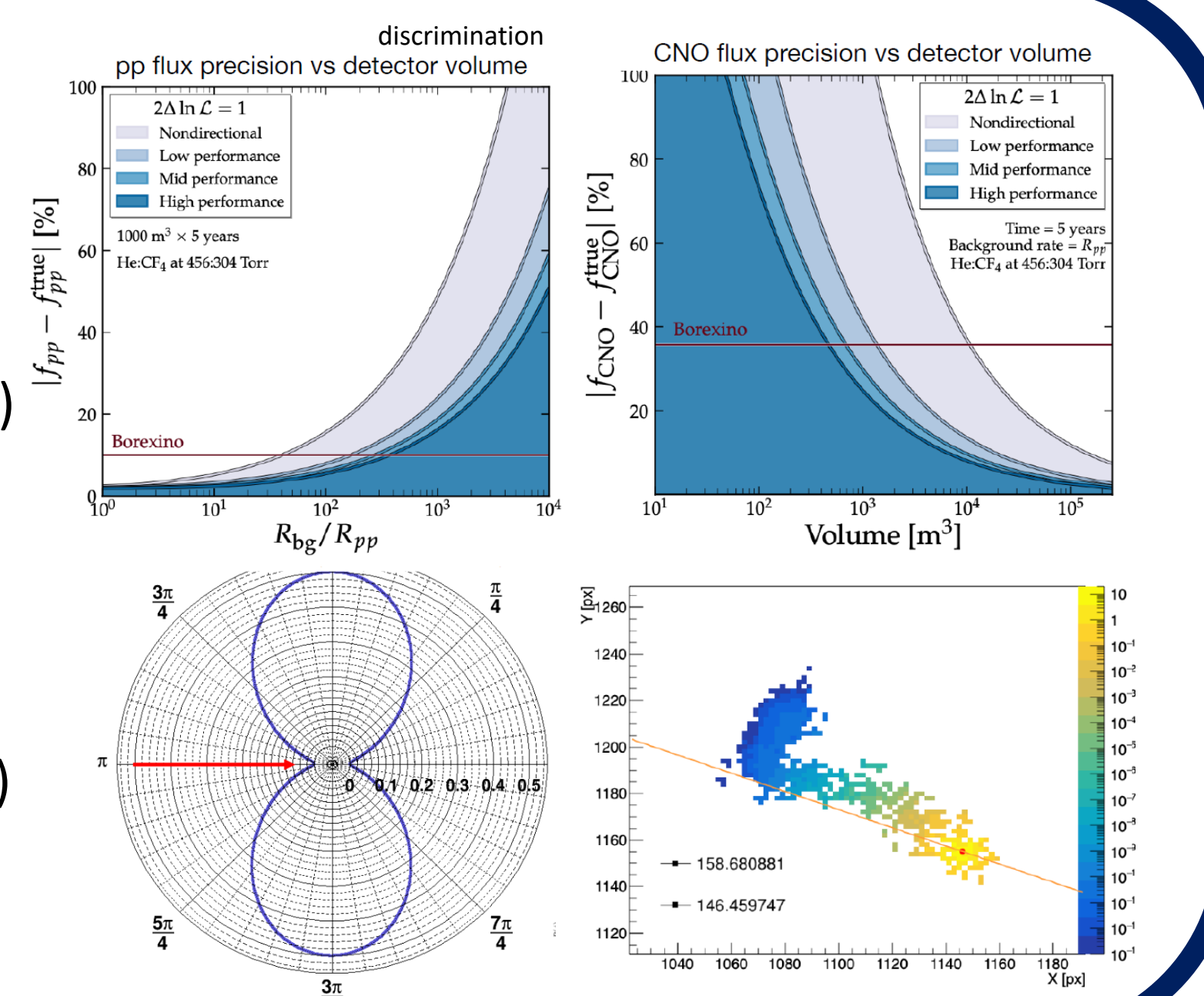
## Physics case

### Directional Dark Matter (DM)

- Assuming DM model detectable by scattering (WIMPs..)
- Directional information crucial for direct detection:
  - Better statistical rejection of background
  - Sidesteps the neutrino fog problem (especially for Solar neutrinos)
  - Will allow 3D astronomy of DM
- Imaging of nuclear recoils (NRs)

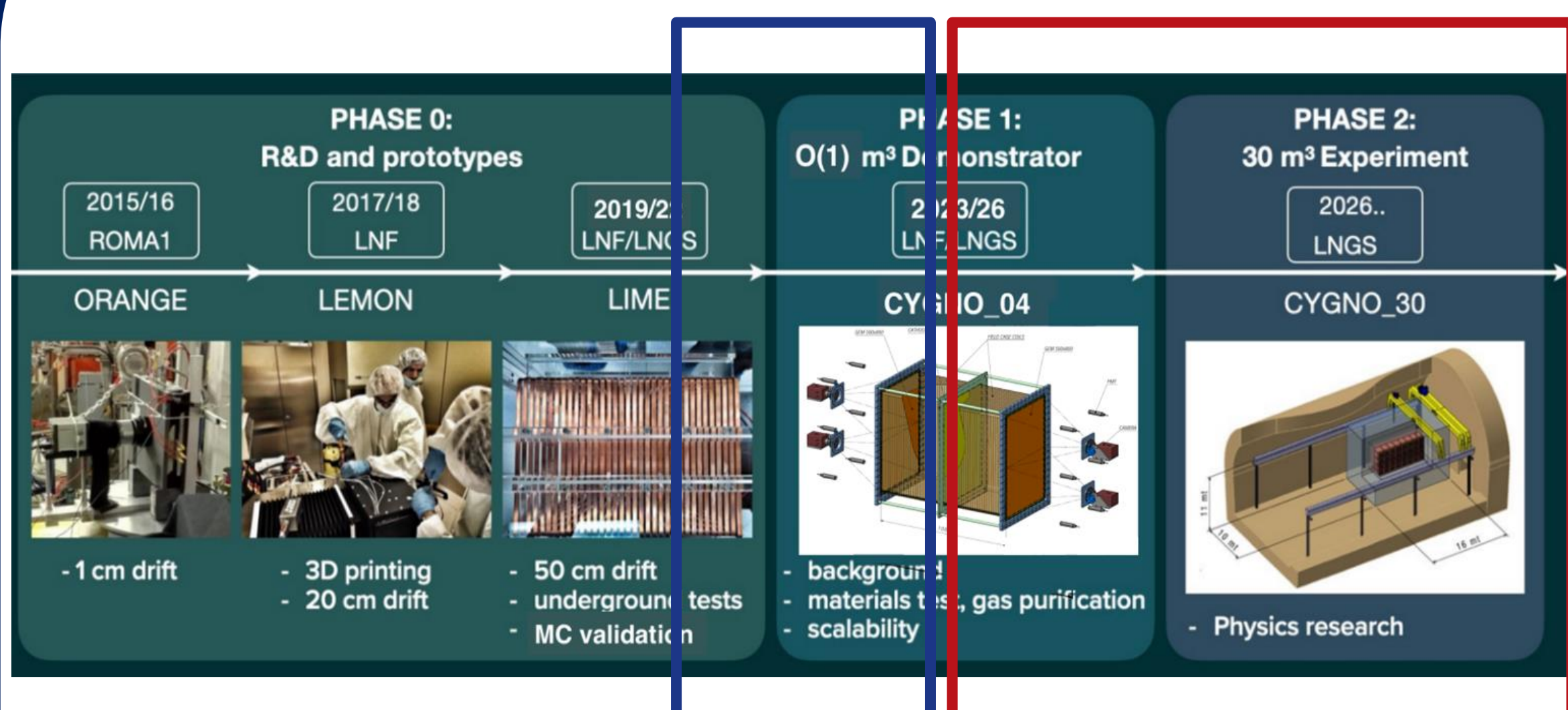
### But not only

- Neutrinos can become signal
  - Imaging of electron recoils (ERs)
- Polarization of X-ray from space
  - Imaging of electron recoils (ERs)
  - Opens new knowledge on large



## The Detector

### Gaseous Time Projection Chamber based detector



See D. Marques presentation

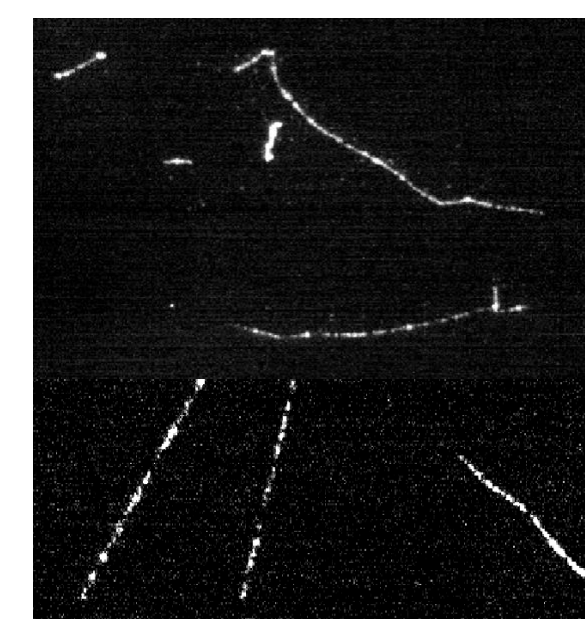
Future

### Main Features

- Imaging detector of low energy Ers and NRs ( $E_{th} \sim 1 \text{ keV}_{ee}$ )
- Operated at atmospheric pressure and room temperature with mixture of  $\text{He:CF}_4$  (60/40)
- Sensitivity to SI and SD for WIMP mass range of  $0.7\text{--}50 \text{ GeV}/c^2$
- Triple Gas Electron Multiplier (GEM) stack for signal amplification
- 3D reconstruction with combined use of the camera and PMTs

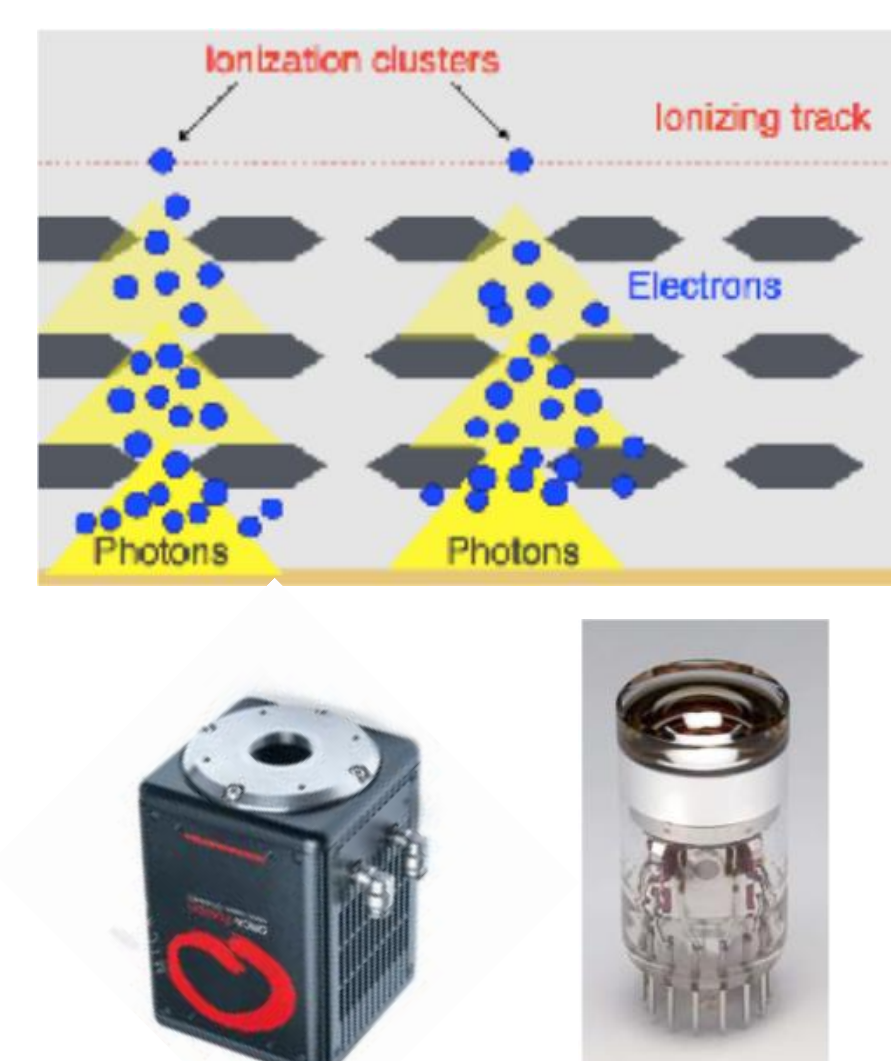
### SCMOS camera

- Single photon sensitivity
- High granularity (2304x2304 pixels)



Energy x-y coordinate

### Optical Readout



### PMT

- Fast detector

straight track

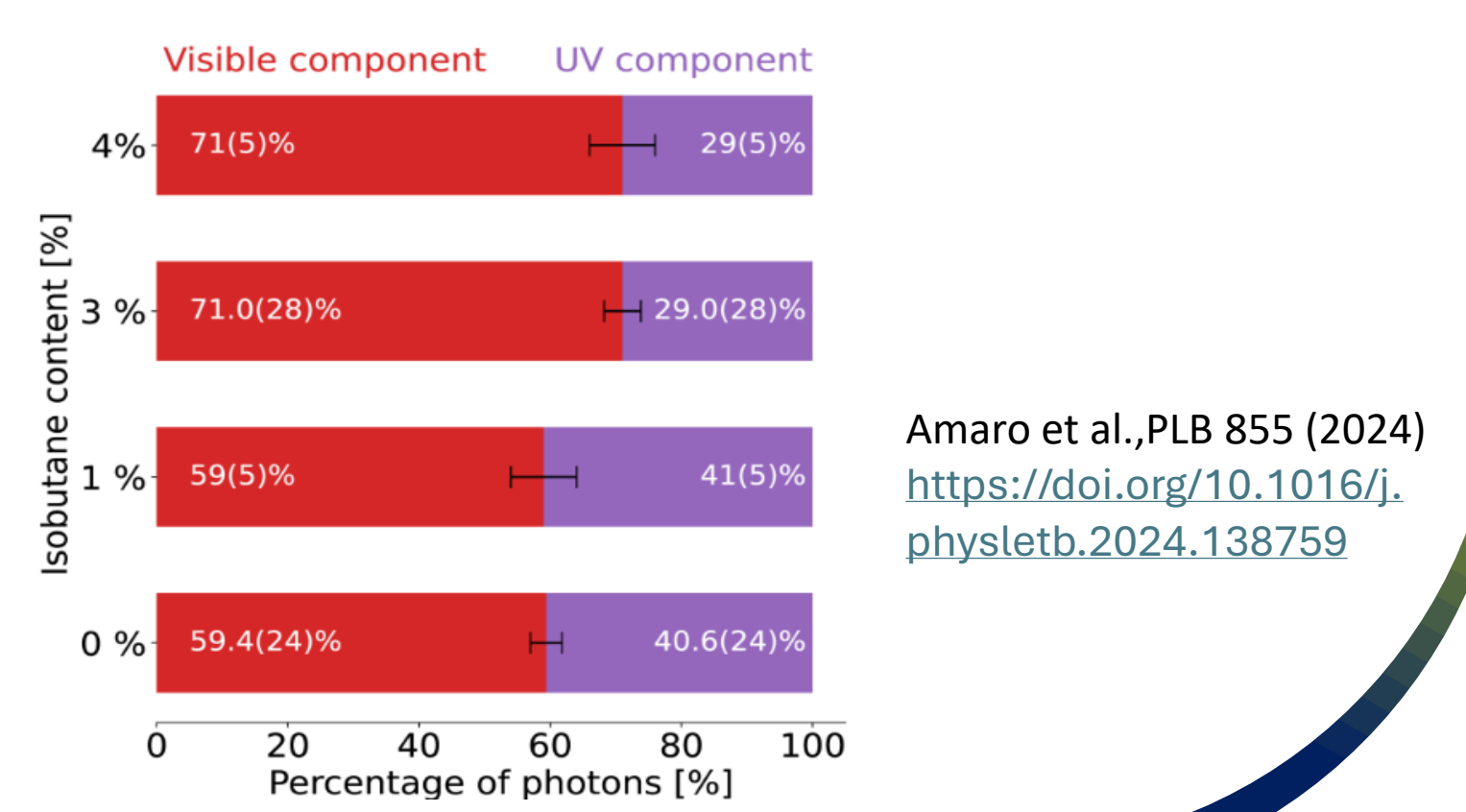
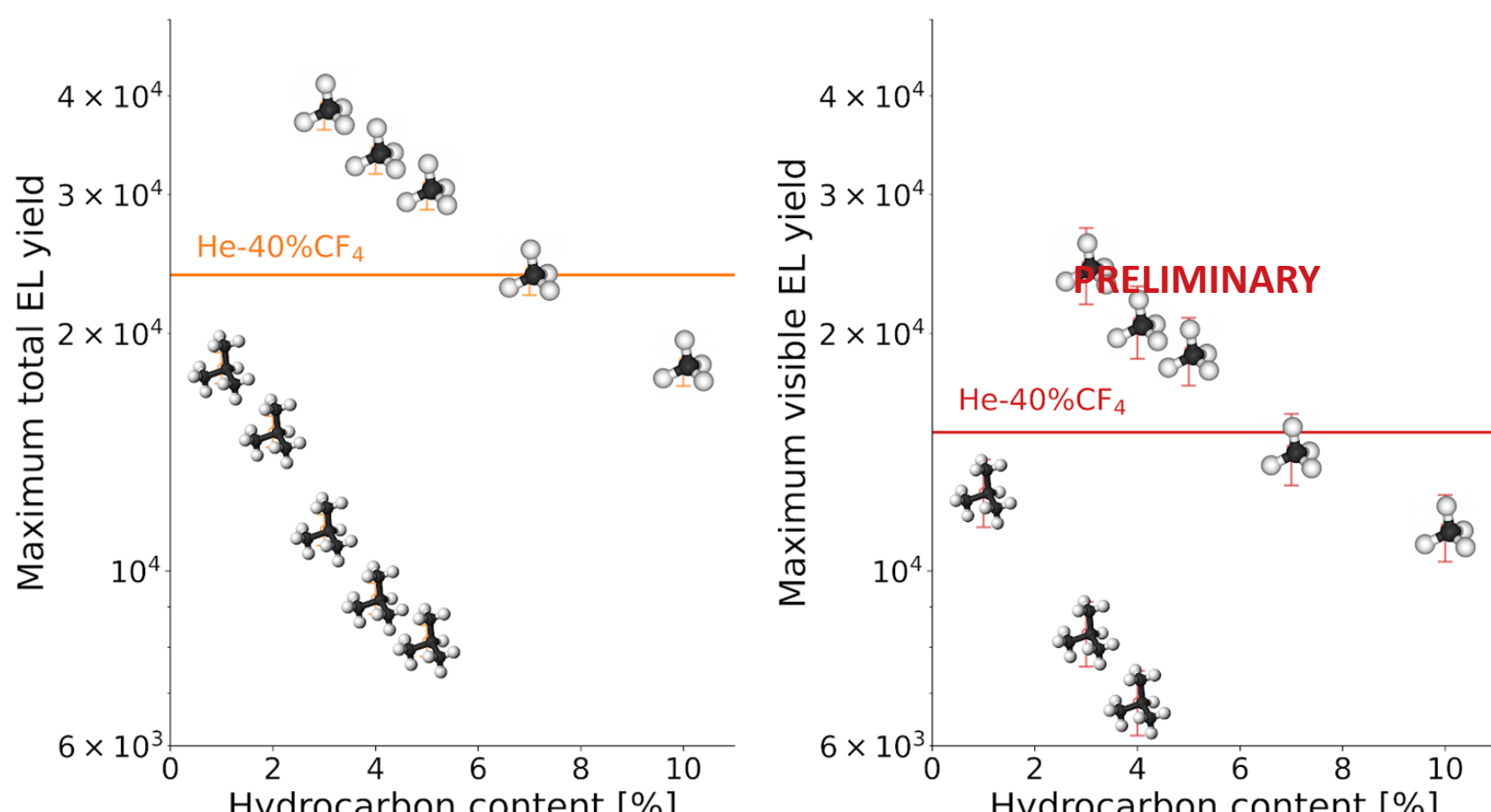
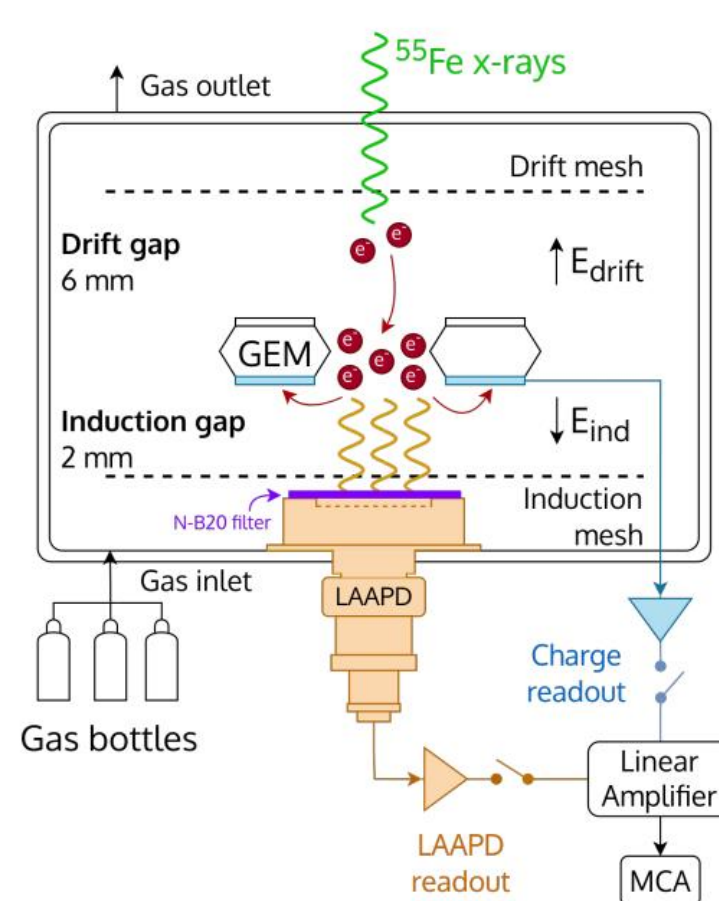
tilted track

Energy z coordinate

Full 3D reconstruction

## H-rich gases

- The possibility of adding hydrogen rich gas is under study to gain sensitivity to lower DM masses
- Limit estimation with 2% isobutane or 5% methane
- Both isobutane and methane in <10% concentration were tested



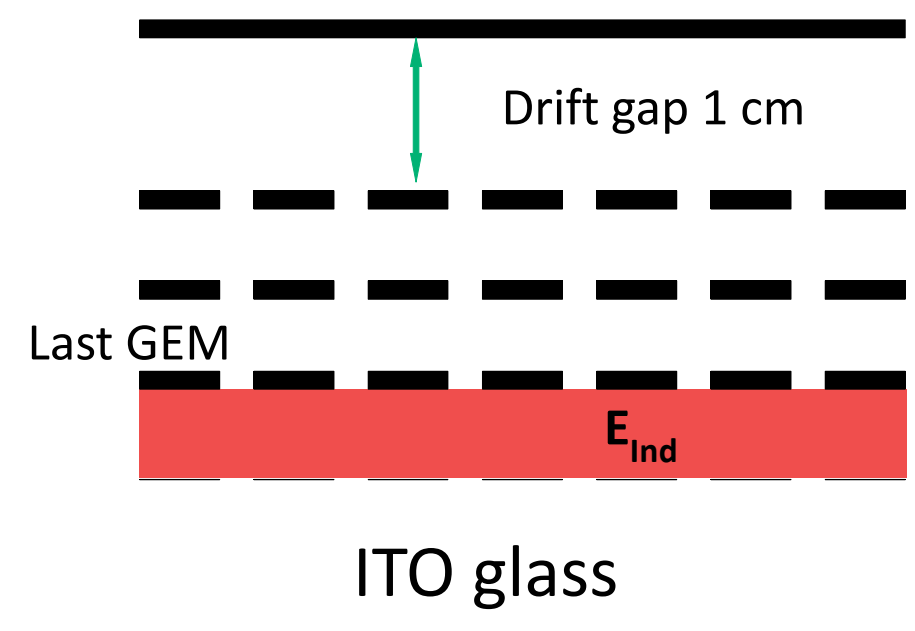
Amaro et al., PLB 855 (2024) <https://doi.org/10.1016/j.physletb.2024.138759>

- Methane gave larger stability
- Larger gain achieved than  $\text{He:CF}_4$  alone
- Level of visible to UV light quenching estimated with filters (pass >300 nm)

## Enhanced light yield

- Increasing the light yield produced in the amplification stage allows to:
  - lower the energy threshold
  - Improve signal to noise ratio

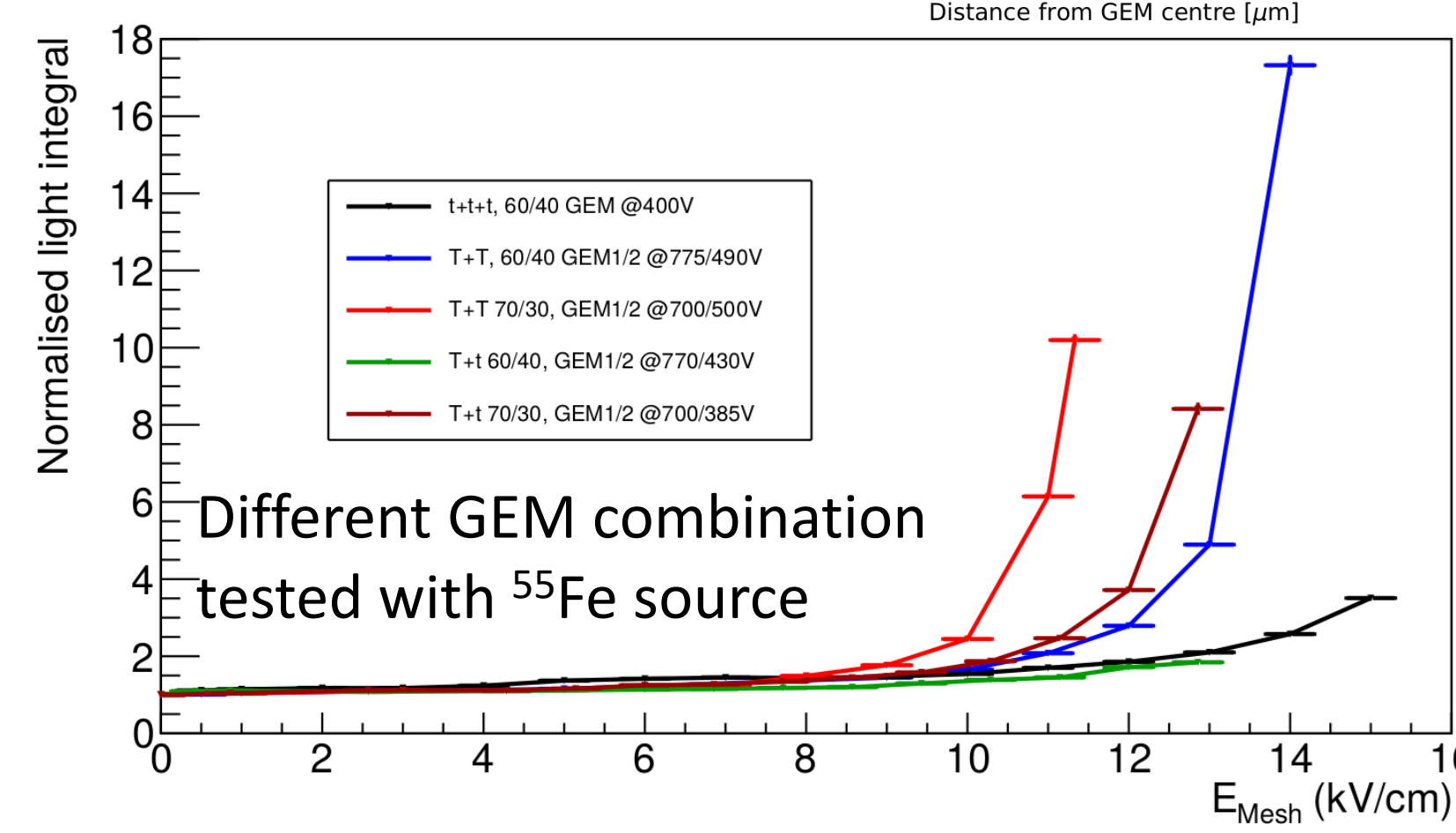
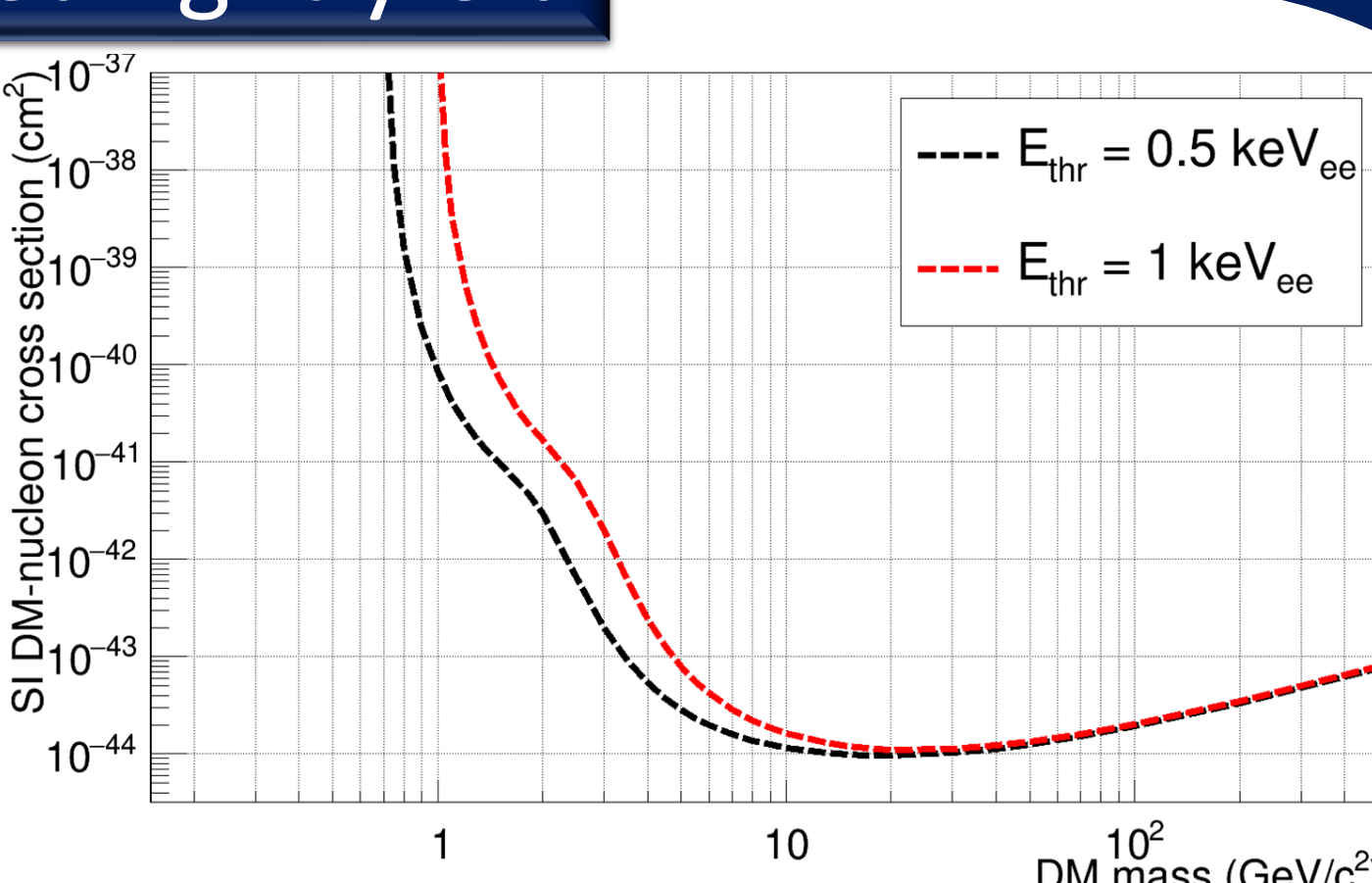
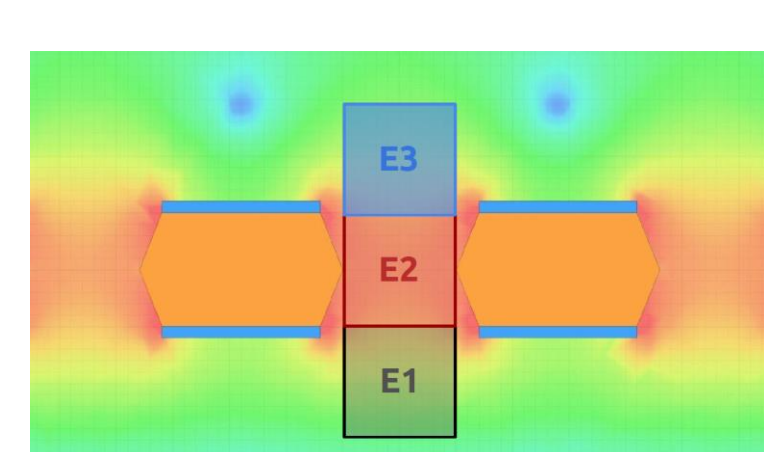
- Concept: add a strong electric field (10- 15 kV/cm) below the last GEM of amplification



### Advantages:

- Large increment in light yield
- Light yield surpassed regular GEM operation
- Intrinsic diffusion stable

Allows the use of 2 GEM stack with reduced intrinsic diffusion (~ 25% less) and comparable light yield (~ factor 2 less)



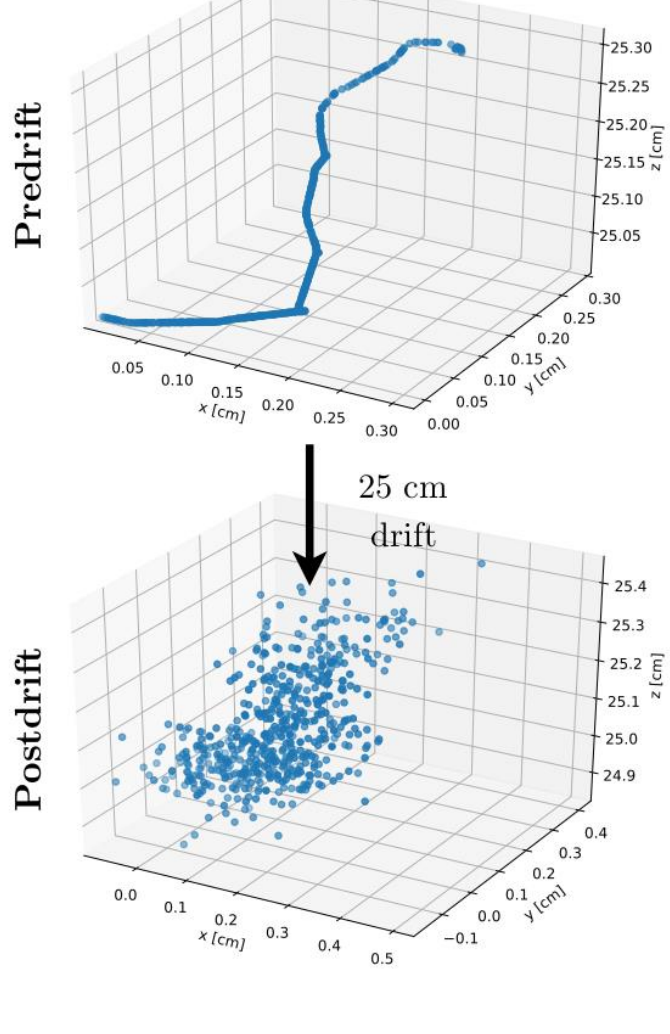
Amaro et al., preprint (2024) [arXiv:2406.05713](https://arxiv.org/abs/2406.05713)

## Negative ion Drift

### Principle

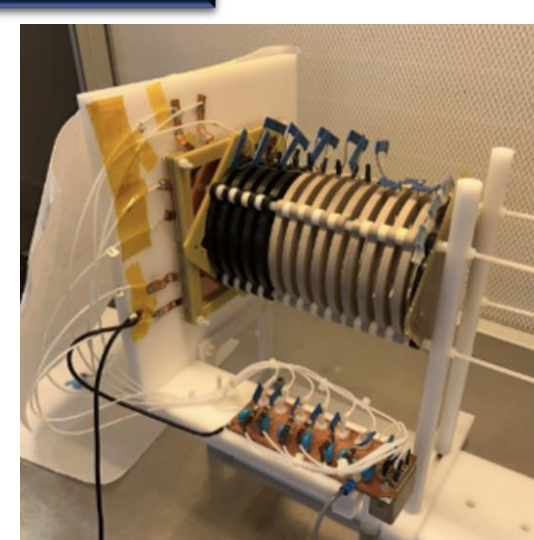
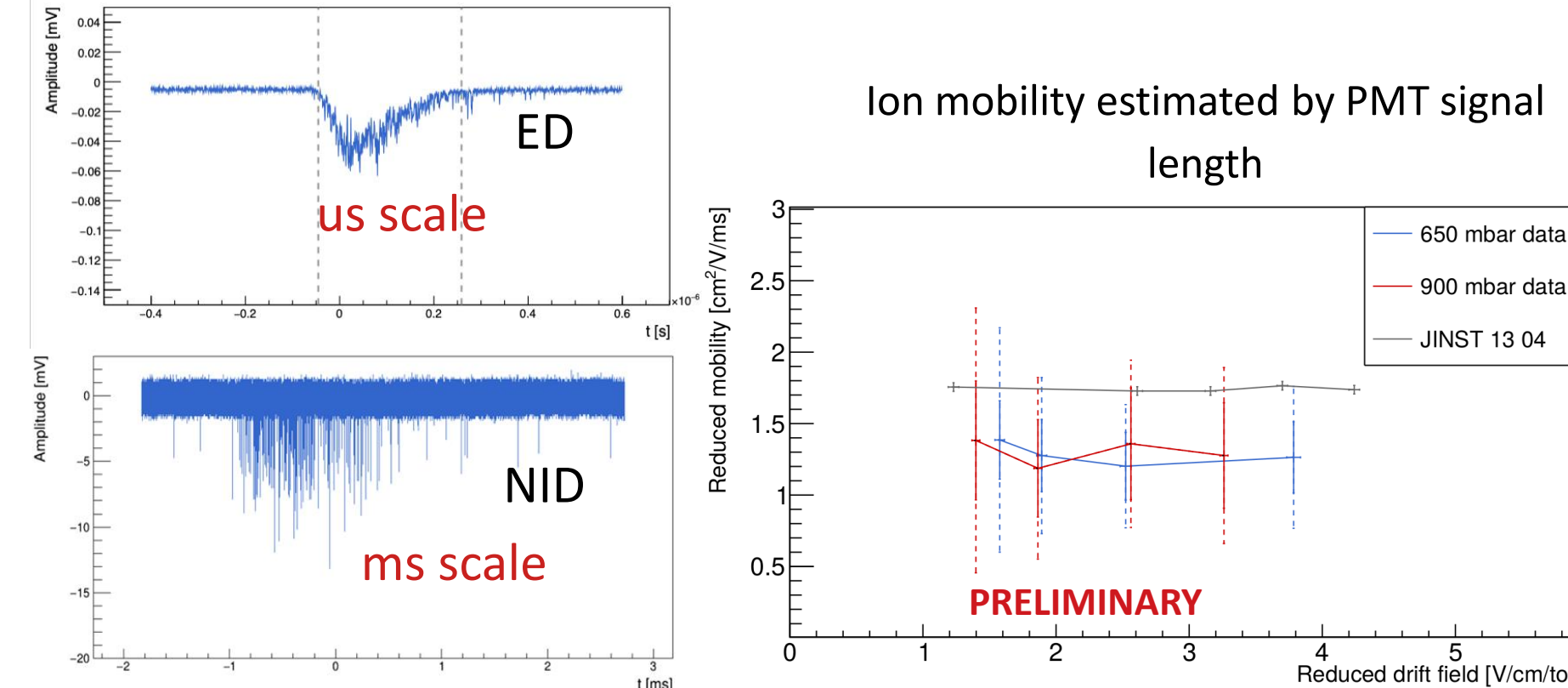
- Diffusion in gas limits the amount of information retrievable to characterise the topology of the tracks
- Electronegative gases can capture primary electron and generate negative ions
- Diffusion can be reduced to thermal limit (maybe below?)

CYGNUS white paper: <https://arxiv.org/abs/2008.12587>  
Helium in 20 Torr  $\text{SF}_6$



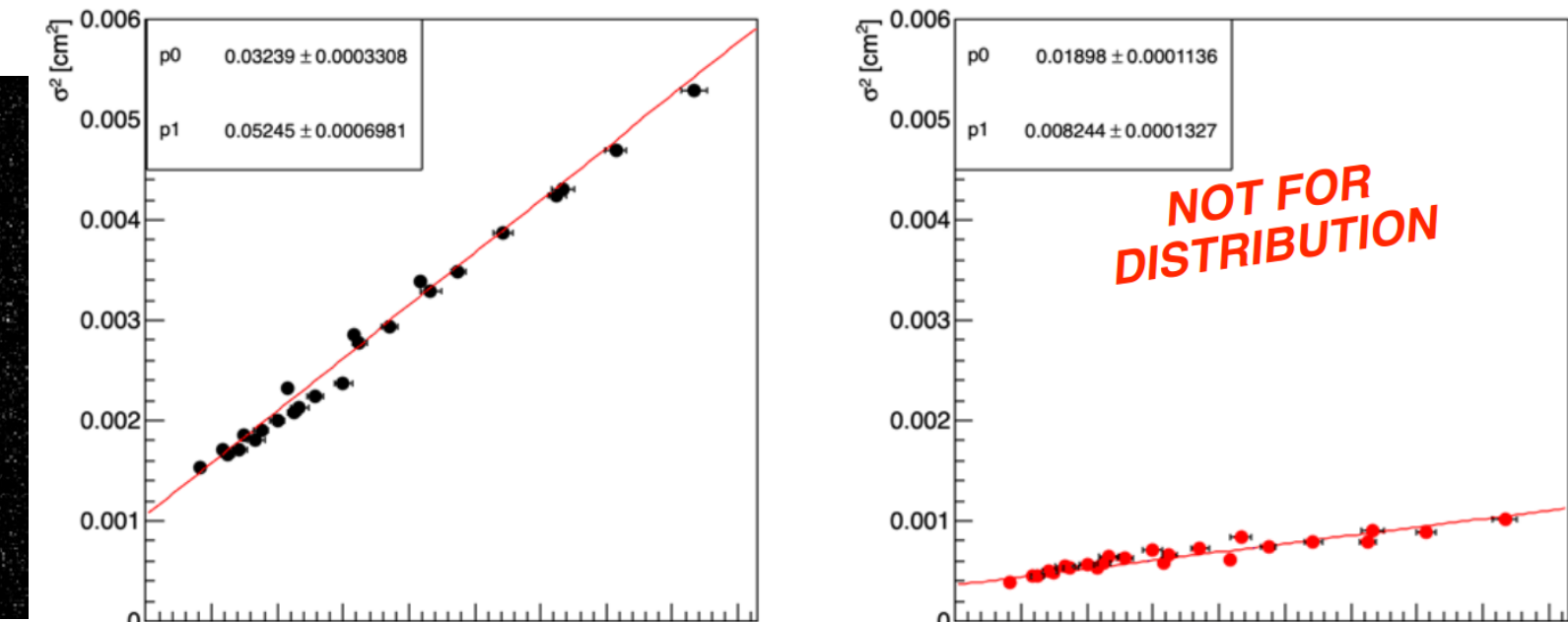
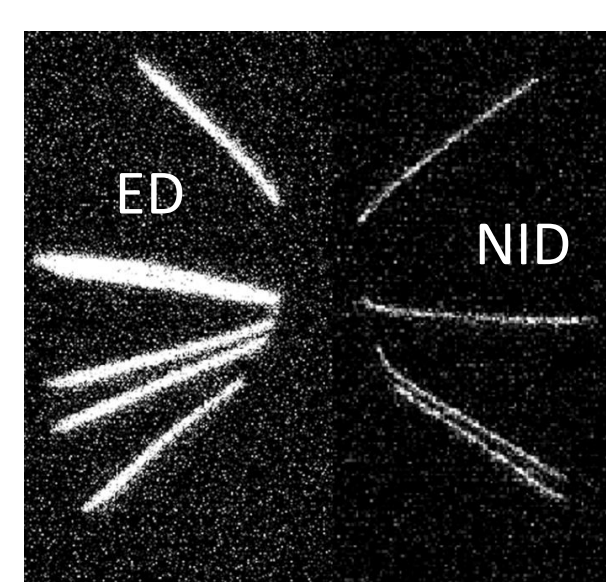
### Experimental realisation

- TPC  $10 \times 10 \text{ cm}^2$  15 cm drift length,  $^{241}\text{Am}$  source, 900 and 650 mbar
- Gas mixture ED:  $\text{He:CF}_4$  (60/40)  
NID:  $\text{He:CF}_4:\text{SF}_6$  (59/39.4/1.6)



### Diffusion measurement

- Exploiting the central width of collimated alpha tracks



Extremely low transverse diffusion achieved

Gas	Diffusion @1 kV/cm [ $\mu\text{m}/\sqrt{\text{cm}}$ ]
Ar:CH <sub>4</sub> (90/10):	600
He:CF <sub>4</sub> (60/40):	140
He:CF <sub>4</sub> :SF <sub>6</sub> (59:39.4:1.6):	35