TDLAS

December 31, 2023

1 Python translation to the TDLAS MATLAB code with annotation and validation

The script version of code contains individual py files, for function definitions, input definitions and main code. Here, they are combined in systematic order along with anottations on the equations and methods being used in the code

The variable and function names used in this script are the same as the ones used in MATLAB, the variable names may not be appropriate at some locations.

1.1 Input parameters definition

This section contains the definitions of parameters that are used in the computation

```
[1]: # HITRAN database filename
    HITRAN_filename = "01_7000-7500_HITEMP2010_7443.2-7445.5.par"
    # partition filename
                          = "q1.txt"
    PART filename
    # minimum and maximum wave numbers in cm^-1
    nu_1_min
                        = 7443.2
                          = 7445.5
    nu_1_max
     # number of terms between min ans max wave numbers to be considered
    N_nu_range
                          = 3000
     # total radial distance in cm and radial step size
    total_radial_distance = 3.0
    delta x
                          = 0.01
    # parameters for temperature and mole fraction fields construction----
     # maximum flame temperature and ambient temperature
    T mag uni field
                          = A1 = 950.0
    A2
                          = 300.0
    # location where flame temperature reaches half the maximum temperature
    ΧO
                          = 6.0*total_radial_distance/7.0
```

```
# gradient of the transition region
A3 = 0.06

# maximum and ambient mole fraction of water vapour
x_mag_uni_field = B1 = 0.095
B2 = 0.028
```

Following are the constants used in the work

```
[2]: # reference temperature
T0 = 296.0

# Plank's constant
h = 6.62607004e-34

# speed of light in cm/s in vacuum
c = 29979245800

# Boltzmann constant
k = 1.38064852e-23

# molar mass of water
M = 18.01

P = 1 # <----- DOUBT

# Avogadro number
Na = 6.02214076e23
```

1.2 Importing libraries/modules

The needed modules and libraries are imported in this section

```
[3]: import numpy as np
import pandas as pd
import matplotlib.pyplot as plt
from scipy.special import wofz # Voigt function
```

1.3 Function definitions

1.3.1 HITRAN database reader function

This function was developed to read the Hi-resolution Transmission database file for the chosen species of H2O. This will return the pandas dataframe with all the columns

```
[4]: def HITRAN_reader(fname):
         # reading all file contents
         fid = open(fname, "r")
         lines = fid.readlines()
         # preparing lists to store line values
         molNo_list = []; isoNo_list = []; transitionWaveNo_list = []
         lineIntensity_list = []; einsteinACoeff_list = []; airBroadenedWidth_list = __
      \hookrightarrow []
         selfBroadenedWidth_list = []; lowEState_list = []; tempCoeff_list = []
         pressureShift_list = []; upperVibQuanta_list = []; lowerVibQuanta_list = []
         upperLocalQuanta_list = []; lowerLocalQuanta_list = []; errorCodes_list = []
         referenceCodes_list = []; flagLineMixing_list = []
         upperStatisticalWeight_list = []; lowerStatisticalWeight_list = []
         # looping through the lines to read individual data
         for line in lines:
             # setting starting read position on the string
             idx = 0
             # reading molecule number
             molNo = int(float(line[idx:idx+2]))
             idx+=2
             # isotopologue number
             isoNo = int(float(line[idx:idx+1]))
             idx+=1
             # transition wavenumber / vaccum wavenumber
             transitionWaveNo = float(line[idx:idx+12])
             idx+=12
             # line intensity
             lineIntensity = float(line[idx:idx+10])
             idx+=10
             # einstein A-coefficient
             einsteinACoeff = float(line[idx:idx+10])
             idx += 10
             # air-broadened half width / air-broadened width
             airBroadenedWidth = float(line[idx:idx+5])
             idx+=5
             # self-broadened half width / self-broadened width
```

```
selfBroadenedWidth = float(line[idx:idx+5])
idx+=5
# lower energy state
lowEState = float(line[idx:idx+10])
idx+=10
# temperature dependence coefficient
tempCoeff = float(line[idx:idx+4])
idx+=4
# pressure shift / air-pressure-induced line shift
pressureShift = float(line[idx:idx+8])
idx += 8
# upper vibrational quanta / upper-state "global" quanta
upperVibQuanta = line[idx:idx+15]
idx+=15
# lower vibrational quanta / lower-state "global" quanta
lowerVibQuanta = line[idx:idx+15]
idx+=15
# upper local quanta / upper-state "local" quanta
upperLocalQuanta = line[idx:idx+15]
idx+=15
# lower local quanta / lower-state "local" quanta
lowerLocalQuanta = line[idx:idx+15]
idx+=15
# error codes / uncertainity indices
errorCodes = line[idx:idx+6]
idx+=6
# reference codes / reference indices
referenceCodes = line[idx:idx+12]
idx+=12
# flag for lineMixing / Flag
flagLineMixing = line[idx:idx+1]
idx+=1
# upper statistical weight / statistical weight for the upper state
upperStatisticalWeight = float(line[idx:idx+7])
idx+=7
```

```
# lower statistical weight / statistical weight for the lower state
      lowerStatisticalWeight = float(line[idx:idx+7])
      idx+=7
      # appending read values to the lists
      molNo_list.append(molNo)
      isoNo_list.append(isoNo)
      transitionWaveNo_list.append(transitionWaveNo)
      lineIntensity list.append(lineIntensity)
      einsteinACoeff_list.append(einsteinACoeff)
      airBroadenedWidth list.append(airBroadenedWidth)
      selfBroadenedWidth_list.append(selfBroadenedWidth)
      lowEState list.append(lowEState)
      tempCoeff_list.append(tempCoeff)
      pressureShift_list.append(pressureShift)
      upperVibQuanta_list.append(upperVibQuanta)
      lowerVibQuanta_list.append(lowerVibQuanta)
      upperLocalQuanta_list.append(upperLocalQuanta)
      lowerLocalQuanta_list.append(lowerLocalQuanta)
      errorCodes_list.append(errorCodes)
      referenceCodes_list.append(referenceCodes)
      flagLineMixing_list.append(flagLineMixing)
      upperStatisticalWeight_list.append(upperStatisticalWeight)
      lowerStatisticalWeight_list.append(lowerStatisticalWeight)
      # break
  # column names for pandas dataframe
  colNames = ["moleculeNumber","isotopologueNumber","transisionWaveNumber",
              "lineIntensity", "einsteinACoefficient", "airBroadenedWidth",
              "selfBroadenedWidth", "lowerStateEnergy", "temperatureDependence",
→ "pressureShift", "upperVibrationalQuanta", "lowerVibrationalQuanta",
              "upperLocalQuanta", "lowerLocalQuanta", "errorCodes",
              "referenceCodes", "flagForLineMixing", "upperStatisticalWeight",
              "lowerStatisticalWeight"]
  # preparing dataframe
  fid = pd.DataFrame(np.
stranspose([molNo_list,isoNo_list,transitionWaveNo_list,
                                    lineIntensity_list,einsteinACoeff_list,
→airBroadenedWidth_list,selfBroadenedWidth_list,
⇔lowEState_list,tempCoeff_list,pressureShift_list,
                                    upperVibQuanta_list,lowerVibQuanta_list,
```

1.3.2 Partition function interpolator

This function interpolates the partition function values over a range of temperatures and obtains the interpolated value for the given temperature

1.3.3 Line strength function

This function computes the line strength using the following equation

$$S(T) = S(T_0) \frac{Q(T_0)}{Q(T)} \frac{T_0}{T} exp\left(\frac{-hcE''}{k} \left(\frac{1}{T} - \frac{1}{T_0}\right)\right) \frac{\left[1 - exp\left(\frac{-hc\nu_0}{kT}\right)\right]}{\left[1 - exp\left(\frac{-hc\nu_0}{kT_0}\right)\right]}$$

```
[6]: def lineStrength(SO, nuO, Edd, T, fid_Q):

"""

This function computes the line strength of a particular transition at a given temperature

equation has been split into 4 parts for ease of programming
"""
```

```
part_1 = Q_func(T0,fid_Q["T"],fid_Q["Q"])/Q_func(T,fid_Q["T"],fid_Q["Q"])
part_2 = T0/T*np.exp(-h*c/k*Edd*(1/T - 1/T0))
part_3 = 1 - np.exp(-h*c*nu0/k/T)
part_4 = 1 - np.exp(-h*c*nu0/k/T0)

S = S0*part_1*part_2*part_3/part_4
return S
```

1.3.4 Spectral absorption coefficient function

This function computes the spectral absorption coefficient for a given x over all the wave numbers Here, the a equation is given below, it describes the gausian half-width at half maximum value

$$a = 2\sqrt{\log(2)} \frac{\nu - \nu_0}{\alpha_G}$$

the w equation which describes the lorenztian half-width at half maximum is given below

$$w = \sqrt{\log(2)} \frac{\alpha_L}{\alpha_G}$$

Then the ϕ function value is computed with the equation described below

$$\phi(\nu) = \frac{2}{\alpha_G} \sqrt{\frac{log(2)}{\pi}} V(z) \quad \forall \quad z = a + wi$$

Finally, the value of spectral absorption coefficient is computed by the dot product of (line intensity, mole fraction and presure) and the ϕ function as given below

$$K(x) = (S(T) \times \chi \times P) \cdot \phi(\nu, T)$$

The name of the function is given as per the MATLAB code,

```
[7]: def voigt(dat, nu0, s, delnu_g, delnu_l):
    """
    this function computes value of K (spectral absorption coefficient)
    at given X location for all wave numbers.

The name of this program function is given as per the matlab code
    """

# computing difference between chosen wavenumber ranges and wave number
    # from the HITRAN database
    """ for this, the matrix is obtained by subtracting each element of one
    arrau over each other elements of the other array"""
    vv0 = nu0 - dat.reshape(dat.shape[0],1) # it is (nu-nu0) in the equation
```

```
# computing "a" equation from problem definition
"""

DOUBT:
here a 2 is missing in the code but present in the pblm definition
"""

a = np.sqrt(np.log(2))*vv0/delnu_g

# computing "w" equation from problem definition
w = np.sqrt(np.log(2))*np.ones([dat.shape[0],1])*delnu_l/delnu_g

# combining w and a into a complex function to feed it to voigt function
z = a+w*1j

# computing phi function
"""

DOUBT:
here a 2 is missing in the code but present in the pblm definition
"""

phi = 1/delnu_g*np.sqrt(np.log(2)/np.pi)*wofz(z)

# computing the K function in problem definition for current X value
K = np.matmul(np.real(phi),s)

return K
```

1.4 Computing the Non-uniform Absorption Spectra

Reading the HITRAN database and the partition function table

Extracting individual columns of data from the database and computing a range of wave numbers between the given min and max ranges

```
[9]: # reading data from the database file
    # center wavenumber
    nu0 = fid_HITRAN["transisionWaveNumber"].to_numpy().astype("float64")
    # reference line strength
    s0 = fid_HITRAN["lineIntensity"].to_numpy().astype("float64")
    # air-broadened width
    g_a0 = fid_HITRAN["airBroadenedWidth"].to_numpy().astype("float64")
```

```
# self-broadened width
g_s0 = fid_HITRAN["selfBroadenedWidth"].to_numpy().astype("float64")
# lower state energy
Edd0 = fid_HITRAN["lowerStateEnergy"].to_numpy().astype("float64")
# temperature dependence exponent
n = fid_HITRAN["temperatureDependence"].to_numpy().astype("float64")
# pressure shift coefficient
d0 = fid_HITRAN["pressureShift"].to_numpy().astype("float64")
# computing a range of wave numbers
dat1 = np.linspace(nu_1_min, nu_1_max, N_nu_range)
```

Filtering the entries from database based on the following criteria - line intensity should be greater than 10^{-30} - dropping extreme wave numbers

```
[10]: # selecting the indices of entries with nu0 falling b/w extreme nu values and
    # with line intensity <= 10^-30
    index1 = (nu0 > nu_1_min) & (nu0 < nu_1_max)
    index2 = s0 > 1e-30
    index = index1 & index2

nu1 = nu0[index]
    s01 = s0[index]
    g_a01 = g_a0[index]
    g_s01 = g_s0[index]
    Edd1 = Edd0[index]
    n1 = n[index]
    d01 = d0[index]
```

pressure is taken to be contstant, hence, multiplying the line strength with pressure and some conversion factor

```
[11]: SO = s01*2.479371939e19*P # <----- DOUBT
```

Constructing the temperature and mole fraction profiles along with discretizing the x-locations

```
[12]: # temperature and mole fraction fields construction------

# discretized radial distance
X = np.arange(0,total_radial_distance+delta_x,delta_x)

# non uniform temperature field
Tnon = A2+(A1-A2)/(1+np.exp((X-X0)/A3))

# non uniform mole fraction field
xnon = B2+(B1-B2)/(1+np.exp((X-X0)/A3))
```

Creating 2D matrices with size being the number of line strength entities and the number of x-locations, and computing the spatial step size

Looping through all the filtered database entries and the x-locations and computing the following

- Pressure shift coefficient

$$\delta_p = \delta_0 \left(\frac{T_0}{T}\right)^{0.96}$$

- Air broadened width

$$\gamma_{air}(T) = \gamma_{air}(T_0) \left(\frac{T_0}{T}\right)^{n_{air}}$$

- Self broadened width

$$\gamma_{self}(T) = \gamma_{self}(T_0) \left(\frac{T_0}{T}\right)^{0.75}$$

- Shifted center wave number

$$\nu_0 = \nu + P(1-\chi)\delta_p$$

- Line strength with product of mole fraction and spatial step size

$$\begin{split} S(T) = S(T_0) \frac{Q(T_0)}{Q(T)} \frac{T_0}{T} exp\left(\frac{-hcE''}{k} \left(\frac{1}{T} - \frac{1}{T_0}\right)\right) \frac{\left[1 - exp\left(\frac{-hc\nu_0}{kT}\right)\right]}{\left[1 - exp\left(\frac{-hc\nu_0}{kT_0}\right)\right]} \\ S(T)\chi dx \end{split}$$

- Gaussian half width at half maximum (1000 is a unit conversion factor for N_a and 10^4 for c in cm/s)

$$\alpha_G = \nu_0 \left(\frac{2N_akTlog(2)\times 1000\times 10^4}{mc^2}\right)^{\frac{1}{2}}$$

- Lorentzian half width at half maximum

$$\alpha_L = P\left(\chi\gamma_{self} + (1-\chi)\gamma_{air}\right)$$

```
[14]: # looping through all x-locations and given line intensities
for i in range(ns):
    for j in range(nt):
        # getting current mole fraction value
```

```
x = xnon[j]
       # computing the pressure shift coefficient
      d[j,i] = d01[i]*(T0/Tnon[j])**0.96
       # air broadened width
      g_a[j,i] = g_a01[i]*(T0/Tnon[j])**n1[i]
       # self broadened width
      g_s[j,i] = g_s01[i]*(T0/Tnon[j])**0.75
       # shifted center wave nuber
      nu0_s[j,i] = nu1[i] + P*(1-x)*d[j,i]
       # line strength
      s[j,i] = lineStrength(S0[i], nu0_s[j,i], Edd1[i], Tnon[j],
\hookrightarrowfid_Q)*dx[j]*x
       # qaussian HWHM
      delnu_g[j,i] = np.sqrt(2*k*np.log(2)*Na/c**2*1000*10**4)*nu0_s[j,i]*np.
⇔sqrt(Tnon[j]/M)
       # delnu_g[j,i] = 0.5*7.162242257e-7*nu0_s[j,i]*np.sqrt(Tnon[j]/M)
       # lorentian HWHM
      delnu_1[j,i] = P*(x*g_s[j,i]+(1-x)*g_a[j,i])
```

Computing the absorption spectra

$$K(x) = P\chi(x) \sum_{i=1}^{N_x} S(T_i) \phi(\nu_i, T_i)$$

```
[15]: # pre-allocating absorbance array for all wave numbers
"""

it is -ln(I(nu)/I_0)
"""

Absorbance = np.zeros([N_nu_range])

# computing absorbance coefficient K for each x and suming it up for total
for i in range(X.shape[0]-1): # <------ DOUBT

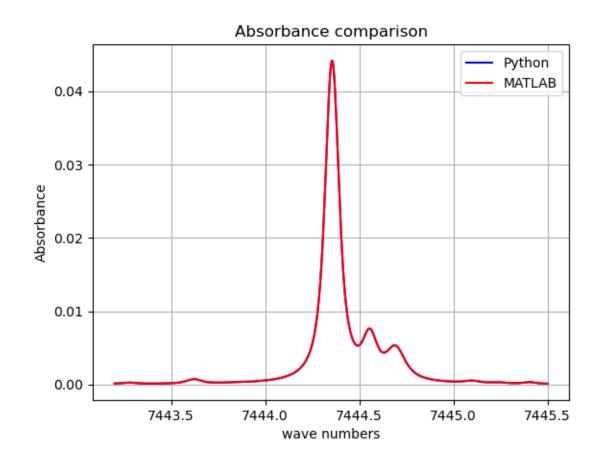
Absorbance += voigt(dat1, nu0_s[i,:], s[i,:], delnu_g[i,:], delnu_l[i,:])</pre>
```

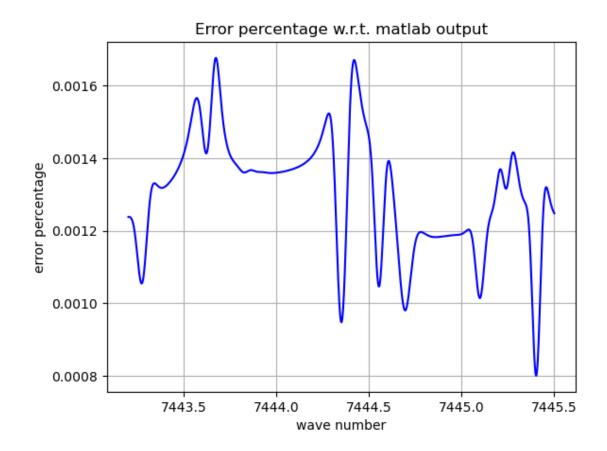
1.5 Post-processing

Preparing pandas dataframe with computed values, and compute error percentage with MATLAB output, then write computed values to a csv file

Plotting comparison and error graphs

```
[17]: # plotting absorbance graph
      plt.figure()
      plt.plot(dat1, Absorbance, '-b', label = "Python")
      plt.plot(fid_mat["wavenumber"], fid_mat["absorbance"], '-r', label = "MATLAB")
      plt.grid()
      plt.legend()
      plt.xlabel("wave numbers")
      plt.ylabel("Absorbance")
      plt.title("Absorbance comparison")
      plt.savefig("absorbance.png", dpi = 150)
      # plotting error percentage between matlab and python
      plt.figure()
      plt.plot(fid_ab["wavenumber"],fid_ab["error_percentage"], '-b')
      plt.grid()
      plt.xlabel("wave number")
      plt.ylabel("error percentage")
      plt.title("Error percentage w.r.t. matlab output")
      plt.savefig("error.png", dpi = 150)
      plt.show()
```





[]: