A SURPRISING POWER LAW RELATIONSHIP PREDICTS TRENDS IN THE TEST ACCURACY FOR VERY DEEP NEURAL NETWORKS

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ABSTRACT

Given two or more Deep Neural Networks (DNNs) with similar architectures, and trained on the same dataset, but trained with different solvers, hyper-parameters, regularization, etc., can we predict which DNN will have the best test accuracy, without peeking at the test data? Solving this question of generalization would have both theoretical impact and great practical importance. In this paper, we show how to use our new theory of Implicit (Heavy Tailed) Self-Regularization for modern Deep Neural Networks to answer this. We examine over 50 different, pre-trained DNNs ranging over 15 different architectures, trained on ImagetNet, with differing test accuracies. We show that, across each architecture (VGG16, VGG19, InceptionV3/V4, ...), the reported test accuracies for each DNN are well correlated with the average of the layer power law exponents, as defined in the Theory of Heavy Tailed Self-Regularization. Our approach requires no changes to the underlying DNN, and does not even require access to the ImageNet data. We present and review these empirical results, and compare and contrast with recent approaches to estimate test performance of DNNs using product norms.

1 Introduction

2 THEORY OF HEAVY TAILED SELF REGULARIZATION

Let us write the Energy Landscape (or optimization function) for a typical DNN with L layers, with activation functions $h_l(\cdot)$, and with weight matrices and b iases \mathbf{W}_l and \mathbf{b}_l , as follows:

$$E_{DNN} = h_L(\mathbf{W}_L \times h_{L-1}(\mathbf{W}_{L-1} \times h_{L-2}(\cdots) + \mathbf{b}_{L-1}) + \mathbf{b}_L). \tag{1}$$

We imagine training this model on some labeled data $\{d_i, y_i\} \in \mathcal{D}$, using Backprop, by minimizing the loss \mathcal{L} For simplicity, we do not indicate the structural details of the layers (e.g., Dense or not, Convolutions or not, Residual/Skip Connections, etc.). Each layer is defined by one or more weight matrices \mathbf{W}_L , or tensors.

In this study, we only need to consider Linear and 2D Convolutional (Conv2D) layers. For the Linear layers, each \mathbf{W}_L is a single $(N\times M)$ (real valued) 2D matrix, where $N\geq M$. These include Dense or Fully Connected (FC) layers, as well as 1D Convolutional (Conv1D) layers, Attention matrices, etc. For the Conv2D layers, with a $c\times d$ kernel, \mathbf{W}_L is a 4-index Tensor, of the form $(N\times M\times c\times d)$, consisting of $c\times d$ 2D feature maps of shape $(N\times M)$. So each Linear layer l gives n=1 2D matrix \mathbf{W}_l , and each Conv2D layer l gives $n=c\times d$ 2D matrices $\mathbf{W}_{l,i}$. A typical modern DNN may have anywhere between 5 and 500 2D $\mathbf{W}_{l,i}$ layer matrices.

Heavy Tailed Universality For any layer weight matrix W, we construct the associated $M \times M$ (uncentered) correlation matrix

$$\mathbf{X} = \frac{1}{N} \mathbf{W}^T \mathbf{W},\tag{2}$$

and form the eigenvalue spectrum of X,

$$\mathbf{X}\mathbf{v}_i = \lambda_i \mathbf{v}_i$$

(where we have dropped the l, i indices).

We call the density of eigenvalues $\rho(\lambda)$ the Empirical Spectral Density (ESD).

Small Neural Networks [charlesG: complete]

The layer matrices in all large, modern DNNs, do not have a scale cut-off, and, instead display scale-invariance. For any modern DNN, the ESD of nearly every **W** layer matrix can be fit to a power law,

$$\rho(\lambda) \sim \lambda^{-\alpha}$$

The power law exponent α lies, 80-90% of the time, in a Universal range between 2 and 4, $\alpha \in [2,4]$. We have verified this empirical Universality on empirical result on over 10,000 layer matrices **W** spanning over 50 pre-trained DNNs. Of course, there are exceptions, and in any real DNN, α may range anywhere from ~ 1.5 to 10 or higher.

The power law exponent α is a complexity metric for a weight matrix; it describes how well that matrix encodes the complex correlations in the training data. So a natural complexity metric for a DNN is to take a weighted average of the power law exponents $\alpha_{l,i}$ for each layer weight matrix $\mathbf{W}_{l,i}$.

$$\hat{\alpha} := \frac{1}{n} \sum_{l,i} b_{l,i} \alpha_{l,i}$$

The smaller $\hat{\alpha}$, the better we expect the DNN to represent the training data. And, presumably, the better the DNN will generalize. The only question is, what are good weights $b_{l,i}$?

It turns out, we can derive the weighted average $\hat{\alpha}$ directly from the more familiar Product Norm.

[charlesG: THE REST OF THIS SECTION is to justify this complexity metric, using the product norm for DNNs. I have sketeched out some of the math,. It just needs to be presented clearly AND TO JUSTIFY computing the average log Norm (which is much faster, much simpler)]

[charles: Maybe should can state a theorem and prove it.]

THEOREM: The data dependent VC-like complexity of a Deep Neural Network can be expressed a weighted average the of power law exponents describing the empirical spectral density of the layer weight matrices

[charles:

PROOF:...

Product Norm Measures of Complexity Recently it has been suggested that the complexity of a DNN, C, can be defined by something akin to a data dependent VC complexity, the product of norms of the layer weight matrices

$$\mathcal{C} \sim \|\mathbf{W}_1\| \times \|\mathbf{W}_2\| \cdots \|\mathbf{W}_L\|$$

where $\|\mathbf{W}\|$ may be the Frobenius norm or even the L1-norm. [michael: what more should we say here?]

To that end, we consider a log complexity

$$\log \mathcal{C} \sim \log \left[\|\mathbf{W}_1\| \times \|\mathbf{W}_2\| \cdots \|\mathbf{W}_L\| \right]$$

$$\sim \left[\, \log \|\mathbf{W}_1\| + \log \|\mathbf{W}_2\| \cdots \log \|\mathbf{W}_L\| \, \right]$$

We define the average log norm of the weight matrices as

$$\langle \log \|\mathbf{W}\| \rangle = \frac{1}{N} \sum_{i=1}^{N} \log_{10} \|\mathbf{W}_i\|$$

which we explicitly define in terms of the base-10 log.

Relation to Heavy Tailed Universality The Frobenius norm is related to the integral of the ESD, over the range the power law is a good fit.

$$\|\mathbf{W}_{l,i}\|^2 \sim \int_{x_{min}}^{x_{max}} \lambda \rho_{l,i}(\lambda) d\lambda$$

Technically, the power only describes the tail of the ESD, for the range $\lambda \in [\lambda_{min}, \lambda_{max}]$. But for most DNN layer matrices, this range covers most the ESD. So this should be a pretty good approximation.

If we take the log matrix norm (and dropping the l, i subscripts) we get

$$\log \|\mathbf{W}\|^{2} \approx \log \int_{\lambda_{min}}^{\lambda_{max}} \lambda \rho(\lambda) d\lambda$$

$$= \log \int_{\lambda_{min}}^{\lambda_{max}} \lambda^{1-\alpha} d\lambda$$

$$= \log \left[\frac{\lambda^{\alpha}}{\alpha} \right]_{\lambda_{min}}^{\lambda_{max}}$$

Since $\lambda_{min} \sim 0$, we have [charles: check this]

$$\log \|\mathbf{W}\|^2 \approx \alpha \log \lambda_{max}$$

Notice that since the power law exponents mostly display Universality empirically, the range $[\lambda_{min}, \lambda_{max}]$ is also roughly the same for all matrices, with the minimum eigenvalue is near zero, $\lambda_{min} \sim 0$, and the maximum eigenvalue is empirically bounded, roughly $\lambda_{max} \sim \mathcal{O}(10^1 - 10^2)$.

This gives

$$\log_{10} \|\mathbf{W}\|^2 \approx (1-2) \times \alpha$$

[charles: FINISH WRITEUP AND and numerical results. In particular, we may be off by some scaling factor in $\rho(\lambda)$] [charlesG: This could all be presented as a theorem–building on work by Hidary and Poggio]

[charles: We need to clean up the math. I think we could use $|\mathbf{W}_F|$ or $|\mathbf{W}|$ or $|\mathbf{W}_F|^2$ and could form the complexity metrix in terms of the product norm of W, the product norm squared of W, or even the product norm of X.].

$$\log \|\mathbf{W}\| \to \sqrt{b\alpha}$$

where the weight factor is b is

$$b = \log \lambda_{max} \sim (1-2)$$

Linear Layer:
$$\log \|\mathbf{W}_l\| \to \sqrt{b_l(\alpha_l)}$$

For the Conv2D layers, we relate the 'Norm' of the 4-index Tensor \mathbf{W}_l to the sum of the integrals of the $n = c \times d$ ESDs for each feature map, giving

Conv2D Layer:
$$\log \|\mathbf{W}_l\| \to -\sum_i \sqrt{b_{l,j}(\alpha_{i,l})}$$

So in the expression for the product norm for $\log C$, let us replace each $\log \|\mathbf{W}_l\|$ for layer l with the sum of the power law exponents $\alpha_{l,i}$ for the n_l $\mathbf{W}_{l,i}$ layer matrices, and take the average over all N_{α} matrices. This gives a new, albeit rather crude complexity metric for the entire DNN

[charles: We don't really use this, but we could:]

$$\hat{\alpha} := \frac{1}{N_{\alpha}} \sum_{i,l} b_{i,j} \alpha_{i,l}$$

[charles: Or we could use]

$$\hat{\alpha} := \frac{1}{N_{\alpha}} \sum_{i,l} \sqrt{b_{i,j} \alpha_{i,l}}$$

[charles: The main complication is if we use log norms or we use a sum of α . Plus we can drop the -1 but we have to deal with the sign change, which I have to work out]

We can now use $\hat{\alpha}$ to analyze numerous pre-trained DNNs...and the results are indeed surprising.

Numerical Example of Power Law / Norm Relation But first, some numerical examples of the relation between $log||\mathbf{W}||$ and α BLAH

BLAH

BLAH

3 EMPIRICAL RESULTS ON PRETRAINED DNNS

VGG and VGG BN Models We start by looking at the VGG class of models, including VGG11, VGG13, VGG16, and VGG19, and their counterparts with Batch Normalization, VGG11_BN, VGG13_BN, VGG16_BN and VGG19_BN. Figure 1

ResNet PyTorch Models

More PreTrained Models More Pretrained Models

| Architecture | Model | Top1 | Top5 | L | N_{α} | $\hat{\alpha}$ | $\hat{\alpha}*$ |
|--------------|----------|------|------|---|--------------|----------------|-----------------|
| VGG11 | VGG11 | | | | | | |
| | VGG11 BN | | | | | | ' |
| VGG13 | VGG13 | | | | | | |
| | VGG13 BN | | | | | | |
| VGG16 | VGG16 | | | | | | |
| | VGG16 BN | | | | | | |
| VGG19 | VGG19 | | | | | | |
| | VGG19 BN | | | | | | |

Table 1: VGG Architectures and DNN Models

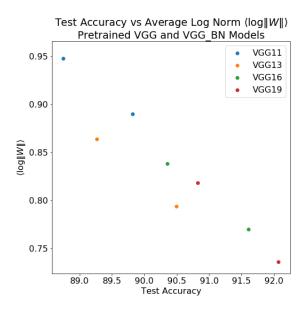


Figure 1: Pretrained VGG and VGG BN Architectures and DNNs. Test Accuracy and average log norm $\hat{\alpha*}$ for VGG11 vs VGG11_BN (blue), VGG13 vs VGG13_BN (orange), VGG16 vs VGG16_BN (green), and VGG19 vs VGG19_BN (red). Note that $\hat{\alpha*}$ does not include the last layer, connecting the model to the labels.

| Architecture | Model | Test Accuracy |
|-------------------|-------|---------------|
| ResNet (larger) | | |
| ResNet (extended) | | |
| ResNet (b) | | |

Table 2: ResNet Architectures and DNN Models

| Architecture | Model | Test Accuracy |
|--------------|-------|---------------|
| GoogLeNet | | |
| ResNeXt | | |
| SqueezeNet | | |

Table 3: Other Models

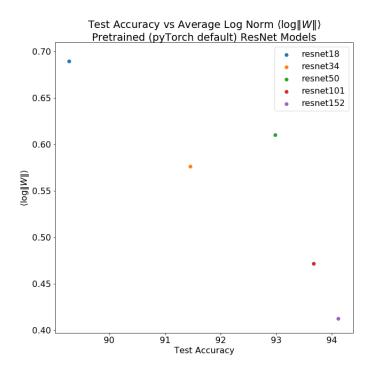


Figure 2: Pretrained ResNet models available in PyTorch

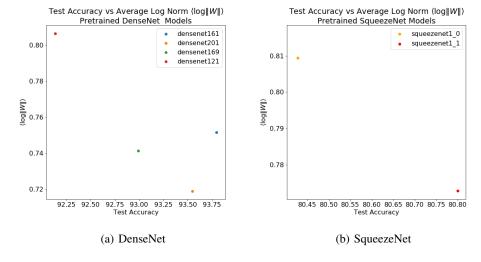


Figure 3: Densenet and SqueezeNet PyTorch Models

4 DISCUSSION AND CONCLUSION

Clearly, our theory opens the door to address numerous very practical questions. One of the most obvious is whether our RMT-based theory is applicable to other types of layers such as convolutional layers. Initial results suggest yes, but the situation is more complex than the relatively simple picture we have described here. These and related directions are promising avenues to explore.

5 APPENDIX

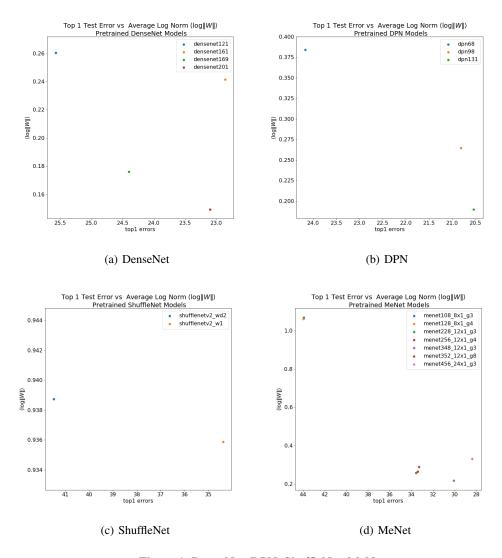


Figure 4: DenseNet, DPN, ShuffleNet, MeNet

| Architecture | Model | Top 1 Error | |
|--------------|--------------------|-------------|--|
| DenseNet | densenet121 | 25.57 | |
| | densenet161 | 22.86 | |
| | densenet169 | 24.4 | |
| | densenet201 | 23.1 | |
| DPN | dpn68 | 24.17 | |
| | dpn98 | 20.81 | |
| | dpn131 | 20.54 | |
| MeNet | menet108_8x1_g3 | 43.92 | |
| | menet128_8x1_g4 | 43.95 | |
| | menet228_12x1_g3 | 33.57 | |
| | menet256_12x1_g4 | 33.41 | |
| | menet348_12x1_g3 | 30.1 | |
| | menet352_12x1_g8 | 33.31 | |
| | menet456_24x1_g3 | 28.4 | |
| MobileNet | mobilenet_wd4 | 46.26 | |
| | mobilenet_wd2 | 36.3 | |
| | mobilenet_w3d4 | 33.54 | |
| | mobilenet_w1 | 29.86 | |
| MobileNetV2 | mobilenetv2_wd4 | 49.72 | |
| | mobilenetv2_wd2 | 36.54 | |
| | mobilenetv2_w3d4 | 31.89 | |
| | mobilenetv2_w1 | 29.31 | |
| FDMobileNet | fdmobilenet_wd4 | 55.77 | |
| | fdmobilenet_wd2 | 43.85 | |
| | fdmobilenet_w1 | 34.7 | |
| SE-ResNet | seresnet50 | 22.47 | |
| | seresnet101 | 21.88 | |
| | seresnet152 | 21.48 | |
| SE-ResNeXt | seresnext50_32x4d | 21.0 | |
| | seresnext101_32x4d | 19.96 | |
| ShuffleNet | shufflenetv2_wd2 | 41.48 | |
| | shufflenetv2_w1 | 34.39 | |
| | | | |

Table 4: Even more pretrained DNN models.

| Model | Top 1 Error | $\hat{\alpha}$ | |
|---------------------|--|---|---|
| resnet10 | 37.46 | | |
| resnet12 | 36.18 | | |
| resnet14 | 33.17 | | |
| resnet16 | 30.9 | | |
| resnet18 | 29.52 | | |
| resnet34 | 25.66 | | |
| resnet50 | 23.79 | | |
| condensenet74_c4_g4 | 26.25 | | |
| condensenet74_c8_g8 | 28.93 | | |
| | resnet10 resnet12 resnet14 resnet16 resnet18 resnet34 resnet50 condensenet74_c4_g4 | resnet10 37.46 resnet12 36.18 resnet14 33.17 resnet16 30.9 resnet18 29.52 resnet34 25.66 resnet50 23.79 condensenet74_c4_g4 26.25 | resnet10 37.46 resnet12 36.18 resnet14 33.17 resnet16 30.9 resnet18 29.52 resnet34 25.66 resnet50 23.79 condensenet74_c4_g4 26.25 |

Table 5: Counter Examples

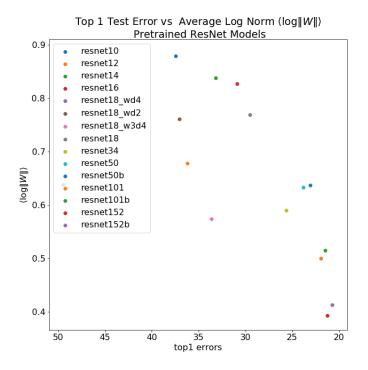


Figure 5: Pretrained ResNet models available in OSMR package

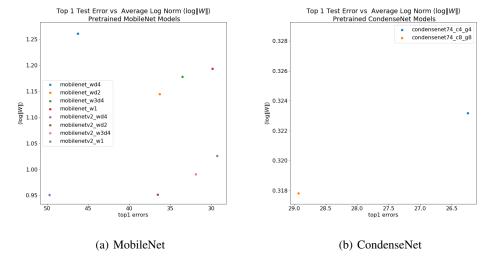


Figure 6: CounterExamples