

# The *spin-transport* documentation

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## Abstract

The *spin-transport* software ([GitHub](#)) is for the dynamic simulation of bulk spin transport—diffusion and separation—in solid media. The project is open-source and still in development.

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## 1 spin-transport: introduction

This repository contains the (developing) open-source code for simulating bulk spin transport—diffusion and separation—in solid media. Multi-spin-species and magnetic resonance simulations are in development.

This is a [Python](#) and [FEniCS](#) project. FEniCS is used to numerically solve the spin transport governing partial differential equations.

End users of this project write Python code to interface with FEniCS.

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## 1.1 Installation

One must first have a working installation of FEniCS. This README assumes the use of [Docker](#) for installation, which is documented [here](#).

Then [clone](#) this repository to the host machine.

## 1.2 Workflow

The FEniCS docs have a section on [workflow](#). There are many ways to instantiate these good practices, but if you're using a nix system, the following may be the easiest.

With the cloned `spin-transport` repository as your working directory, create a link in your path to `spin-transport`'s `fenics` executable bash script.

```
ln fenics /usr/local/bin
```

Now a FEniCS Python script `foo.py` can be started with the command `fenics foo.py` **from the host** instead of manually starting it from a Docker container. This has several advantages, including that there is no need to move scripts into the container and that the complicated syntax need not be remembered.

## 1.3 Testing the installation

To verify that everything is installed correctly, run the Poisson equation demo `ft01_poisson.py` ([source](#)) in your container.

If you installed the `fenics` bash script per the instructions above, you can use the following command (working directory: `spin-transport`).

```
$ fenics ft01_poisson.py
```

If everything is working fine, the output should look something like the following.

```
$ fenics ft01_poisson.py
Calling DOLFIN just-in-time (JIT) compiler, this may take some time.
--- Instant: compiling ---
Calling FFC just-in-time (JIT) compiler, this may take some time.
Calling FFC just-in-time (JIT) compiler, this may take some time.
Calling FFC just-in-time (JIT) compiler, this may take some time.
Solving linear variational problem.
*** Warning: Degree of exact solution may be inadequate for accurate
result in error norm.
Calling FFC just-in-time (JIT) compiler, this may take some time.
Calling FFC just-in-time (JIT) compiler, this may take some time.
Ignoring precision in integral metadata compiled using quadrature
representation. Not implemented.
Calling FFC just-in-time (JIT) compiler, this may take some time.
Calling FFC just-in-time (JIT) compiler, this may take some time.
```

```
Calling FFC just-in-time (JIT) compiler, this may take some time.
error_L2 = 0.00823509807335
error_max = 1.33226762955e-15
```

The directory `spin-transport/poisson` should have been created and should contain two files: `solution.pvd` and `solution000000.vtu`. These files contain the solution data.

## 1.4 Acknowledgement

This work is supported by a grant from the Army Research Office, Materials Science Division under grant proposal **Nanoscale Spin Hyperpolarization and Imaging** with PI [John Marohn, PhD](#).

## 1.5 Contributors

This project stems from a collaboration among three institutions:

- [Cornell University](#),
- [Saint Martin's University](#), and the
- [University of Washington](#).

The lead contributor to this project is [Rico Picone, PhD](#) of Saint Martin's University, co-PI on the ARO grant. Other contributors include [John Marohn, PhD](#) (Cornell, PI), John A. Sidles, PhD (Washington), Joseph L. Garbini, PhD (Washington), and Corinne Isaac (Cornell).

## 2 Short theoretical introduction

The simulation is based on the two spin-species transport equations presented by Picone [3]. These equations present the transport—diffusion and separation—of three conserved quantities represented by the vector-valued function  $\boldsymbol{\rho}$  of time and a single spatial dimension, taken to be the direction of a background magnetic field  $\mathbf{B}$  we call *longitudinal*. The three conserved quantities are represented in the “polarization thermodynamic covector basis”  $\mathbf{e}$  by the components [3]

- $\rho_1$ : a dimensionless dipole-energy density,
- $\rho_2$ : longitudinal nuclear polarization, and
- $\rho_3$ : longitudinal electron polarization.

Several parameters require definition and are summarized in the following list (note that a “ $\sim$ ” over a function denotes that it is a function of the dimensionless spatial coordinate). See [3] for details.

- $B_d$ : maximum dipole-dipole magnetic field.
- $\gamma_2, \gamma_3$ : gyromagnetic ratio for nuclei and electrons, respectively.
- $\bar{\gamma}$ : dimensionless ratio  $\gamma_3/\gamma_2$ .
- $\Gamma_2, \Gamma_3$ : transport coefficients for nuclear and electron polarization, respectively.
- $\bar{\Gamma}$ : dimensionless ratio  $\Gamma_3/\Gamma_2$ .
- $\Delta_2, \Delta_3$ : constant spin densities in solid medium for nuclei and electrons, respectively.
- $\bar{\Delta}$ : dimensionless ratio  $\Delta_3/\Delta_2$ .
- $\bar{r}$ : dimensionless spatial coordinate  $(\partial_r B(r)|_{r=0}/B_d)r$ .
- $\bar{t}$ : dimensionless time  $\Gamma_2(\partial_r B(r)|_{r=0}/B_d)^2 t$ .
- $\bar{B}$ : ratio of magnetic fields  $\partial_{\bar{r}} \tilde{B}/B_d$ .
- $\bar{c}$ : ratio of ratios  $\bar{B}(1 + \bar{\Delta})/(1 + \bar{\gamma}\bar{\Delta})$ .

**two-species magnetization transport equations.** The preferred form of the continuity equation is in the polarization basis with the dimensionless spatial coordinate and time variable, which we call the *two-species magnetization transport equations*:

$$\begin{aligned} \partial_{\bar{t}} \rho_1 = & -\frac{\bar{c}^2}{1 + \bar{\Delta}} \left( (1 - \rho_2^2) + \bar{\Gamma} \bar{\Delta} \bar{\gamma}^2 (1 - \rho_3^2) \right) \operatorname{arctanh} \rho_1 + \\ & -\frac{\bar{c}}{1 + \bar{\Delta}} (\partial_{\bar{r}} \rho_2 - \bar{\Gamma} \bar{\Delta} \bar{\gamma} \partial_{\bar{r}} \rho_3) + (1 + \bar{\Gamma}) \partial_{\bar{r}}^2 \rho_1 \end{aligned} \quad (1)$$

$$\partial_{\bar{t}} \rho_2 = \partial_{\bar{r}} \left( \bar{c} (1 - \rho_2^2) \operatorname{arctanh} \rho_1 \right) + \partial_{\bar{r}}^2 \rho_2 \quad (2)$$

$$\partial_{\bar{t}} \rho_3 = -\bar{\Gamma} \bar{\gamma} \partial_{\bar{r}} \left( \bar{c} (1 - \rho_3^2) \operatorname{arctanh} \rho_1 \right) + \bar{\Gamma} \partial_{\bar{r}}^2 \rho_3. \quad (3)$$

Bloch equation dynamics for a each species and  $T_1 \gg T_2$  can be approximated without the transverse dynamics as [2, 1]

$$\partial_{\bar{t}} \rho_2 = \frac{\rho_{20}}{\bar{T}_{1p}} - \frac{\rho_2}{\bar{\tau}_p} \quad (4)$$

$$\partial_{\bar{t}} \rho_3 = \frac{\rho_{30}}{\bar{T}_{1e}} - \frac{\rho_3}{\bar{\tau}_e}. \quad (5)$$

where, for each species,

$$\bar{\tau} = \frac{1}{\frac{1}{\bar{\tau}_1} + \frac{\bar{\tau}_2 \omega_1^2}{1 + \bar{\tau}_2^2 \bar{\delta}^2}} \quad (6)$$

where it is important to use dimensionless time constants and where we define (again, for each species)

$$\omega_1 = -\gamma B_1 \quad \text{and} \quad \bar{\delta} = \gamma B_d \bar{\tau}. \quad (7)$$

The Bloch equation dynamics can be combined with the transport dynamics of Equations 1 – 3 such that the *two-species Block-transport* equations can be written

$$\begin{aligned} \partial_{\bar{t}} \rho_1 = & -\frac{\bar{c}^2}{1 + \bar{\Delta}} \left( (1 - \rho_2^2) + \bar{\Gamma} \bar{\Delta} \bar{\gamma}^2 (1 - \rho_3^2) \right) \operatorname{arctanh} \rho_1 + \\ & -\frac{\bar{c}}{1 + \bar{\Delta}} (\partial_{\bar{r}} \rho_2 - \bar{\Gamma} \bar{\Delta} \bar{\gamma} \partial_{\bar{r}} \rho_3) + (1 + \bar{\Gamma}) \partial_{\bar{r}}^2 \rho_1 \end{aligned} \quad (8)$$

$$\partial_{\bar{t}} \rho_2 = \partial_{\bar{r}} \left( \bar{c} (1 - \rho_2^2) \operatorname{arctanh} \rho_1 \right) + \partial_{\bar{r}}^2 \rho_2 + \frac{\rho_{20}}{\bar{\Gamma}_{1p}} - \frac{\rho_2}{\bar{\tau}_p} \quad (9)$$

$$\partial_{\bar{t}} \rho_3 = -\bar{\Gamma} \bar{\gamma} \partial_{\bar{r}} \left( \bar{c} (1 - \rho_3^2) \operatorname{arctanh} \rho_1 \right) + \bar{\Gamma} \partial_{\bar{r}}^2 \rho_3 + \frac{\rho_{30}}{\bar{\Gamma}_{1e}} - \frac{\rho_3}{\bar{\tau}_e}. \quad (10)$$

## References

- [1] Jean-Philippe Grivet. Simulation of Magnetic Resonance Experiments. *American Journal of Physics*, 61(12):1133–1139, 1993.
- [2] Rico A. R. Picone. Investigating the Effects of Polarizing Diffusion in Magnetic Resonance Force Microscopy, 2010.
- [3] Rico A. R. Picone. Separative Magnetization Transport: Theory, Model, and Experiment, 2014.