

Hartree-Fock methods

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Why Hartree-Fock? Derivation of Hartree-Fock equations in coordinate space

Hartree-Fock (HF) theory is an algorithm for finding an approximative expression for the ground state of a given Hamiltonian. The basic ingredients are

- Define a single-particle basis $\{\psi_\alpha\}$ so that

$$\hat{h}^{\text{HF}} \psi_\alpha = \varepsilon_\alpha \psi_\alpha$$

with the Hartree-Fock Hamiltonian defined as

$$\hat{h}^{\text{HF}} = \hat{t} + \hat{u}_{\text{ext}} + \hat{u}^{\text{HF}}$$

- The term \hat{u}^{HF} is a single-particle potential to be determined by the HF algorithm.
- The HF algorithm means to choose \hat{u}^{HF} in order to have

$$\langle \hat{H} \rangle = E^{\text{HF}} = \langle \Phi_0 | \hat{H} | \Phi_0 \rangle$$

that is to find a local minimum with a Slater determinant Φ_0 being

Why Hartree-Fock? Derivation of Hartree-Fock equations in coordinate space

We will show that the Hartree-Fock Hamiltonian \hat{h}^{HF} equals our definition of the operator \hat{f} discussed in connection with the new definition of the normal-ordered Hamiltonian (see later lectures), that is we have, for a specific matrix element

$$\langle p | \hat{h}^{\text{HF}} | q \rangle = \langle p | \hat{f} | q \rangle = \langle p | \hat{t} + \hat{u}_{\text{ext}} | q \rangle + \sum_{i \leq F} \langle p i | \hat{V} | q i \rangle_{AS},$$

meaning that

$$\langle p | \hat{u}^{\text{HF}} | q \rangle = \sum_{i \leq F} \langle p i | \hat{V} | q i \rangle_{AS}.$$

The so-called Hartree-Fock potential \hat{u}^{HF} brings an explicit medium dependence due to the summation over all single-particle states below the Fermi level F . It brings also in an explicit dependence on the two-body interaction (in nuclear physics we can also have complicated three- or higher-body forces). The two-body interaction, with its contribution from the other bystanding

Variational Calculus and Lagrangian Multipliers

The calculus of variations involves problems where the quantity to be minimized or maximized is an integral.

In the general case we have an integral of the type

$$E[\Phi] = \int_a^b f(\Phi(x), \frac{\partial \Phi}{\partial x}, x) dx,$$

where E is the quantity which is sought minimized or maximized. The problem is that although f is a function of the variables Φ , $\partial \Phi / \partial x$ and x , the exact dependence of Φ on x is not known. This means again that even though the integral has fixed limits a and b , the path of integration is not known. In our case the unknown quantities are the single-particle wave functions and we wish to choose an integration path which makes the functional $E[\Phi]$ stationary. This means that we want to find minima, or maxima or saddle points. In physics we search normally for minima. Our task is therefore to find the minimum of $E[\Phi]$ so that its variation δE is zero subject to specific constraints. In our case the constraints appear as the integral which expresses the orthogonality of the

Variational Calculus and Lagrangian Multipliers

Let us specialize to the expectation value of the energy for one particle in three-dimensions. This expectation value reads

$$E = \int dx dy dz \psi^*(x, y, z) \hat{H} \psi(x, y, z),$$

with the constraint

$$\int dx dy dz \psi^*(x, y, z) \psi(x, y, z) = 1,$$

and a Hamiltonian

$$\hat{H} = -\frac{1}{2} \nabla^2 + V(x, y, z).$$

We will, for the sake of notational convenience, skip the variables x, y, z below, and write for example $V(x, y, z) = V$.

Variational Calculus and Lagrangian Multipliers

The integral involving the kinetic energy can be written as, with the function ψ vanishing strongly for large values of x, y, z (given here by the limits a and b),

$$\int_a^b dx dy dz \psi^* \left(-\frac{1}{2} \nabla^2 \right) \psi dx dy dz = \psi^* \nabla \psi \Big|_a^b + \int_a^b dx dy dz \frac{1}{2} \nabla \psi^* \nabla \psi.$$

We will drop the limits a and b in the remaining discussion. Inserting this expression into the expectation value for the energy and taking the variational minimum we obtain

$$\delta E = \delta \left\{ \int dx dy dz \left(\frac{1}{2} \nabla \psi^* \nabla \psi + V \psi^* \psi \right) \right\} = 0.$$

Variational Calculus and Lagrangian Multipliers

The constraint appears in integral form as

$$\int dx dy dz \psi^* \psi = \text{constant},$$

and multiplying with a Lagrangian multiplier λ and taking the variational minimum we obtain the final variational equation

$$\delta \left\{ \int dx dy dz \left(\frac{1}{2} \nabla \psi^* \nabla \psi + V \psi^* \psi - \lambda \psi^* \psi \right) \right\} = 0.$$

We introduce the function f

$$f = \frac{1}{2} \nabla \psi^* \nabla \psi + V \psi^* \psi - \lambda \psi^* \psi = \frac{1}{2} (\psi_x^* \psi_x + \psi_y^* \psi_y + \psi_z^* \psi_z) + V \psi^* \psi - \lambda \psi^* \psi,$$

where we have skipped the dependence on x, y, z and introduced the shorthand ψ_x, ψ_y and ψ_z for the various derivatives.

Variational Calculus and Lagrangian Multipliers

For ψ^* the Euler-Lagrange equations yield

$$\frac{\partial f}{\partial \psi^*} - \frac{\partial}{\partial x} \frac{\partial f}{\partial \psi_x^*} - \frac{\partial}{\partial y} \frac{\partial f}{\partial \psi_y^*} - \frac{\partial}{\partial z} \frac{\partial f}{\partial \psi_z^*} = 0,$$

which results in

$$-\frac{1}{2} (\psi_{xx} + \psi_{yy} + \psi_{zz}) + V \psi = \lambda \psi.$$

We can then identify the Lagrangian multiplier as the energy of the system. The last equation is nothing but the standard Schrodinger equation and the variational approach discussed here provides a powerful method for obtaining approximate solutions of the wave function.

Definitions and notations

Before we proceed we need some definitions. We will assume that the interacting part of the Hamiltonian can be approximated by a two-body interaction. This means that our Hamiltonian is written as the sum of some onebody part and a twobody part

$$\hat{H} = \hat{H}_0 + \hat{H}_I = \sum_{i=1}^N \hat{h}_0(x_i) + \sum_{i<j}^N \hat{v}(r_{ij}), \quad (1)$$

with

$$H_0 = \sum_{i=1}^N \hat{h}_0(x_i). \quad (2)$$

The onebody part $u_{\text{ext}}(x_i)$ is normally approximated by a harmonic oscillator potential or the Coulomb interaction an electron feels from the nucleus. However, other potentials are fully possible, such as one derived from the self-consistent solution of the Hartree-Fock equations to be discussed here.

Definitions and notations

Our Hamiltonian is invariant under the permutation (interchange) of two particles. Since we deal with fermions however, the total wave function is antisymmetric. Let \hat{P} be an operator which interchanges two particles. Due to the symmetries we have ascribed to our Hamiltonian, this operator commutes with the total Hamiltonian,

$$[\hat{H}, \hat{P}] = 0,$$

meaning that $\Psi_\lambda(x_1, x_2, \dots, x_A)$ is an eigenfunction of \hat{P} as well, that is

$$\hat{P}_{ij} \Psi_\lambda(x_1, x_2, \dots, x_i, \dots, x_j, \dots, x_A) = \beta \Psi_\lambda(x_1, x_2, \dots, x_i, \dots, x_j, \dots, x_A),$$

where β is the eigenvalue of \hat{P} . We have introduced the suffix ij in order to indicate that we permute particles i and j . The Pauli principle tells us that the total wave function for a system of fermions has to be antisymmetric, resulting in the eigenvalue $\beta = -1$.

Definitions and notations

In our case we assume that we can approximate the exact eigenfunction with a Slater determinant

$$\Phi(x_1, x_2, \dots, x_A, \alpha, \beta, \dots, \sigma) = \frac{1}{\sqrt{N!}} \begin{vmatrix} \psi_\alpha(x_1) & \psi_\alpha(x_2) & \dots & \dots & \psi_\alpha(x_A) \\ \psi_\beta(x_1) & \psi_\beta(x_2) & \dots & \dots & \psi_\beta(x_A) \\ \dots & \dots & \dots & \dots & \dots \\ \psi_\sigma(x_1) & \psi_\sigma(x_2) & \dots & \dots & \psi_\sigma(x_A) \end{vmatrix} \quad (3)$$

where x_i stand for the coordinates and spin values of a particle i and $\alpha, \beta, \dots, \gamma$ are quantum numbers needed to describe remaining quantum numbers.

Definitions and notations

The single-particle function $\psi_\alpha(x_i)$ are eigenfunctions of the onebody Hamiltonian h_i , that is

$$\hat{h}_0(x_i) = \hat{t}(x_i) + \hat{u}_{\text{ext}}(x_i),$$

with eigenvalues

$$\hat{h}_0(x_i) \psi_\alpha(x_i) = (\hat{t}(x_i) + \hat{u}_{\text{ext}}(x_i)) \psi_\alpha(x_i) = \varepsilon_\alpha \psi_\alpha(x_i).$$

The energies ε_α are the so-called non-interacting single-particle energies, or unperturbed energies. The total energy is in this case the sum over all single-particle energies, if no two-body or more complicated many-body interactions are present.

Definitions and notations

Let us denote the ground state energy by E_0 . According to the variational principle we have

$$E_0 \leq E[\Phi] = \int \Phi^* \hat{H} \Phi d\tau$$

where Φ is a trial function which we assume to be normalized

$$\int \Phi^* \Phi d\tau = 1,$$

where we have used the shorthand $d\tau = dx_1 dx_2 \dots dx_A$.

Definitions and notations

In the Hartree-Fock method the trial function is the Slater determinant of Eq. (3) which can be rewritten as

$$\Phi(x_1, x_2, \dots, x_A, \alpha, \beta, \dots, \nu) = \frac{1}{\sqrt{N!}} \sum_p (-1)^p \hat{P} \psi_\alpha(x_1) \psi_\beta(x_2) \dots \psi_\nu(x_A)$$

where we have introduced the antisymmetrization operator \hat{A} defined by the summation over all possible permutations of two particles.

Definitions and notations

It is defined as

$$\hat{A} = \frac{1}{N!} \sum_p (-1)^p \hat{P}, \quad (4)$$

with p standing for the number of permutations. We have introduced for later use the so-called Hartree-function, defined by the simple product of all possible single-particle functions

$$\Phi_H(x_1, x_2, \dots, x_A, \alpha, \beta, \dots, \nu) = \psi_\alpha(x_1) \psi_\beta(x_2) \dots \psi_\nu(x_A).$$

Definitions and notations

Both \hat{H}_0 and \hat{H}_I are invariant under all possible permutations of any two particles and hence commute with \hat{A}

$$[H_0, \hat{A}] = [H_I, \hat{A}] = 0. \quad (5)$$

Furthermore, \hat{A} satisfies

$$\hat{A}^2 = \hat{A}, \quad (6)$$

since every permutation of the Slater determinant reproduces it.

Definitions and notations

The expectation value of \hat{H}_0

$$\int \Phi^* \hat{H}_0 \Phi d\tau = N! \int \Phi_H^* \hat{A} \hat{H}_0 \hat{A} \Phi_H d\tau$$

is readily reduced to

$$\int \Phi^* \hat{H}_0 \Phi d\tau = N! \int \Phi_H^* \hat{H}_0 \Phi_H d\tau,$$

where we have used Eqs. (5) and (6). The next step is to replace the antisymmetrization operator by its definition and to replace \hat{H}_0 with the sum of one-body operators

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{i=1}^N \sum_p (-1)^p \int \Phi_H^* \hat{h}_0 \hat{P} \Phi_H d\tau.$$

Definitions and notations

The integral vanishes if two or more particles are permuted in only one of the Hartree-functions Φ_H because the individual single-particle wave functions are orthogonal. We obtain then

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{i=1}^N \int \Phi_H^* \hat{h}_0 \Phi_H d\tau.$$

Orthogonality of the single-particle functions allows us to further simplify the integral, and we arrive at the following expression for the expectation values of the sum of one-body Hamiltonians

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{\mu=1}^N \int \psi_\mu^*(x) \hat{h}_0 \psi_\mu(x) dx d\tau. \quad (7)$$

Definitions and notations

We introduce the following shorthand for the above integral

$$\langle \mu | \hat{h}_0 | \mu \rangle = \int \psi_\mu^*(x) \hat{h}_0 \psi_\mu(x) dx,$$

and rewrite Eq. (7) as

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{\mu=1}^N \langle \mu | \hat{h}_0 | \mu \rangle. \quad (8)$$

Definitions and notations

The expectation value of the two-body part of the Hamiltonian is obtained in a similar manner. We have

$$\int \Phi^* \hat{H}_I \Phi d\tau = N! \int \Phi_H^* \hat{A} \hat{H}_I \hat{A} \Phi_H d\tau,$$

which reduces to

$$\int \Phi^* \hat{H}_I \Phi d\tau = \sum_{i \leq j=1}^N \sum_p (-)^p \int \Phi_H^* \hat{v}(r_{ij}) \hat{P} \Phi_H d\tau,$$

by following the same arguments as for the one-body Hamiltonian.

Definitions and notations

Because of the dependence on the inter-particle distance r_{ij} , permutations of any two particles no longer vanish, and we get

$$\int \Phi^* \hat{H}_I \Phi d\tau = \sum_{i < j=1}^N \int \Phi_H^* \hat{v}(r_{ij}) (1 - P_{ij}) \Phi_H d\tau.$$

where P_{ij} is the permutation operator that interchanges particle i and particle j . Again we use the assumption that the single-particle wave functions are orthogonal.

Definitions and notations

We obtain

$$\int \Phi^* \hat{H}_I \Phi d\tau = \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \left[\int \psi_\mu^*(x_i) \psi_\nu^*(x_j) \hat{v}(r_{ij}) \psi_\mu(x_i) \psi_\nu(x_j) dx_i dx_j \right. \\ \left. - \int \psi_\mu^*(x_i) \psi_\nu^*(x_j) \hat{v}(r_{ij}) \psi_\nu(x_i) \psi_\mu(x_j) dx_i dx_j \right]. \quad (9)$$

The first term is the so-called direct term. It is frequently also called the Hartree term, while the second is due to the Pauli principle and is called the exchange term or just the Fock term. The factor 1/2 is introduced because we now run over all pairs twice.

Definitions and notations

The last equation allows us to introduce some further definitions. The single-particle wave functions $\psi_\mu(x)$, defined by the quantum numbers μ and x are defined as the overlap

$$\psi_\alpha(x) = \langle x | \alpha \rangle.$$

Definitions and notations

We introduce the following shorthands for the above two integrals

$$\langle \mu\nu | \hat{v} | \mu\nu \rangle = \int \psi_\mu^*(x_i) \psi_\nu^*(x_j) \hat{v}(r_{ij}) \psi_\mu(x_i) \psi_\nu(x_j) dx_i dx_j,$$

and

$$\langle \mu\nu | \hat{v} | \nu\mu \rangle = \int \psi_\mu^*(x_i) \psi_\nu^*(x_j) \hat{v}(r_{ij}) \psi_\nu(x_i) \psi_\mu(x_j) dx_i dx_j.$$

Derivation of Hartree-Fock equations in coordinate space

Let us denote the ground state energy by E_0 . According to the variational principle we have

$$E_0 \leq E[\Phi] = \int \Phi^* \hat{H} \Phi d\tau$$

where Φ is a trial function which we assume to be normalized

$$\int \Phi^* \Phi d\tau = 1,$$

where we have used the shorthand $d\tau = dx_1 dx_2 \dots dx_A$.

Derivation of Hartree-Fock equations in coordinate space

In the Hartree-Fock method the trial function is a Slater determinant which can be rewritten as

$$\Psi(x_1, x_2, \dots, x_A, \alpha, \beta, \dots, \nu) = \frac{1}{\sqrt{N!}} \sum_p (-1)^p P \psi_\alpha(x_1) \psi_\beta(x_2) \dots \psi_\nu(x_A)$$

where we have introduced the anti-symmetrization operator \hat{A} defined by the summation over all possible permutations p of two fermions. It is defined as

$$\hat{A} = \frac{1}{N!} \sum_p (-1)^p \hat{P},$$

with the Hartree-function given by the simple product of all possible single-particle function

$$\Phi_H(x_1, x_2, \dots, x_A, \alpha, \beta, \dots, \nu) = \psi_\alpha(x_1) \psi_\beta(x_2) \dots \psi_\nu(x_A).$$

Derivation of Hartree-Fock equations in coordinate space

Our functional is written as

$$E[\Phi] = \sum_{\mu=1}^N \int \psi_\mu^*(x_i) \hat{h}_0(x_i) \psi_\mu(x_i) dx_i + \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \left[\int \psi_\mu^*(x_i) \psi_\nu^*(x_j) \hat{v}(r_{ij}) \psi_\mu(x_i) \psi_\nu(x_j) dx_i dx_j \right]$$

The more compact version reads

$$E[\Phi] = \sum_{\mu} \langle \mu | \hat{h}_0 | \mu \rangle + \frac{1}{2} \sum_{\mu\nu} [\langle \mu\nu | \hat{v} | \mu\nu \rangle - \langle \nu\mu | \hat{v} | \mu\nu \rangle].$$

Derivation of Hartree-Fock equations in coordinate space

Since the interaction is invariant under the interchange of two particles it means for example that we have

$$\langle \mu\nu | \hat{v} | \mu\nu \rangle = \langle \nu\mu | \hat{v} | \nu\mu \rangle,$$

or in the more general case

$$\langle \mu\nu | \hat{v} | \sigma\tau \rangle = \langle \nu\mu | \hat{v} | \tau\sigma \rangle.$$

Derivation of Hartree-Fock equations in coordinate space

The direct and exchange matrix elements can be brought together if we define the antisymmetrized matrix element

$$\langle \mu\nu | \hat{v} | \mu\nu \rangle_{AS} = \langle \mu\nu | \hat{v} | \mu\nu \rangle - \langle \mu\nu | \hat{v} | \nu\mu \rangle,$$

or for a general matrix element

$$\langle \mu\nu | \hat{v} | \sigma\tau \rangle_{AS} = \langle \mu\nu | \hat{v} | \sigma\tau \rangle - \langle \mu\nu | \hat{v} | \tau\sigma \rangle.$$

It has the symmetry property

$$\langle \mu\nu | \hat{v} | \sigma\tau \rangle_{AS} = -\langle \mu\nu | \hat{v} | \tau\sigma \rangle_{AS} = -\langle \nu\mu | \hat{v} | \sigma\tau \rangle_{AS}.$$

The antisymmetric matrix element is also hermitian, implying

$$\langle \mu\nu | \hat{v} | \sigma\tau \rangle_{AS} = \langle \sigma\tau | \hat{v} | \mu\nu \rangle_{AS}.$$

Derivation of Hartree-Fock equations in coordinate space

With these notations we rewrite the Hartree-Fock functional as

$$\int \Phi^* \hat{H}_I \Phi d\tau = \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \langle \mu\nu | \hat{v} | \mu\nu \rangle_{AS}. \quad (10)$$

Adding the contribution from the one-body operator \hat{H}_0 to (10) we obtain the energy functional

$$E[\Phi] = \sum_{\mu=1}^N \langle \mu | \hat{h}_0 | \mu \rangle + \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \langle \mu\nu | \hat{v} | \mu\nu \rangle_{AS}. \quad (11)$$

In our coordinate space derivations below we will spell out the Hartree-Fock equations in terms of their integrals.

Derivation of Hartree-Fock equations in coordinate space

If we generalize the Euler-Lagrange equations to more variables and introduce N^2 Lagrange multipliers which we denote by $\epsilon_{\mu\nu}$, we can write the variational equation for the functional of E

$$\delta E - \sum_{\mu\nu} \epsilon_{\mu\nu} \delta \int \psi_\mu^* \psi_\nu = 0.$$

For the orthogonal wave functions ψ_i this reduces to

$$\delta E - \sum_{\mu=1}^N \epsilon_{\mu\mu} \delta \int \psi_\mu^* \psi_\mu = 0.$$

Derivation of Hartree-Fock equations in coordinate space

Variation with respect to the single-particle wave functions ψ_μ yields then

$$\sum_{\mu=1}^N \int \delta \psi_\mu^* \hat{h}_0(x_i) \psi_\mu dx_i + \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \left[\int \delta \psi_\mu^* \psi_\nu^* \hat{v}(r_{ij}) \psi_\mu \psi_\nu dx_i dx_j - \int \delta \psi_\mu^* \psi_\nu^* \hat{v}(r_{ij}) \psi_\mu \psi_\nu dx_i dx_j \right] - \int \delta \psi_\mu^* \psi_\mu dx_i = 0$$

Derivation of Hartree-Fock equations in coordinate space

Although the variations $\delta\psi$ and $\delta\psi^*$ are not independent, they may in fact be treated as such, so that the terms dependent on either $\delta\psi$ and $\delta\psi^*$ individually may be set equal to zero. To see this, simply replace the arbitrary variation $\delta\psi$ by $i\delta\psi$, so that $\delta\psi^*$ is replaced by $-i\delta\psi^*$, and combine the two equations. We thus arrive at the Hartree-Fock equations

$$\left[-\frac{1}{2} \nabla_i^2 + \sum_{\nu=1}^N \int \psi_\nu^*(x_j) \hat{v}(r_{ij}) \psi_\nu(x_j) dx_j \right] \psi_\mu(x_i) - \left[\sum_{\nu=1}^N \int \psi_\nu^*(x_j) \hat{v}(r_{ij}) \psi_\mu(x_j) dx_j \right] \psi_\mu(x_i) = \epsilon_{\mu\mu} \psi_\mu(x_i) \quad (12)$$

Notice that the integration $\int dx_j$ implies an integration over the spatial coordinates \mathbf{r}_j and a summation over the spin-coordinate of fermion j . We note that the factor of $1/2$ in front of the sum involving the two-body interaction, has been removed. This is due to the fact that we need to vary both $\delta\psi_\mu^*$ and $\delta\psi_\nu^*$. Using the symmetry properties of the two-body interaction and interchanging μ and ν as summation indices, we obtain two identical terms.

Derivation of Hartree-Fock equations in coordinate space

The two first terms in the last equation are the one-body kinetic energy and the electron-nucleus potential. The third or *direct* term is the averaged electronic repulsion of the other electrons. As written, the term includes the *self-interaction* of electrons when $\mu = \nu$. The self-interaction is cancelled in the fourth term, or the *exchange* term. The exchange term results from our inclusion of the Pauli principle and the assumed determinantal form of the wave-function. Equation (12), in addition to the kinetic energy and the attraction from the atomic nucleus that confines the motion of a single electron, represents now the motion of a single-particle modified by the two-body interaction. The additional contribution to the Schrodinger equation due to the two-body interaction, represents a mean field set up by all the other bystander electrons, the latter given by the sum over all single-particle states occupied by N electrons.

The Hartree-Fock equation is an example of an integro-differential equation. These equations involve repeated calculations of integrals, in addition to the solution of a set of coupled differential

Analysis of Hartree-Fock equations in coordinate space

A theoretically convenient form of the Hartree-Fock equation is to regard the direct and exchange operator defined through

$$V_\mu^d(x_i) = \int \psi_\mu^*(x_j) \hat{v}(r_{ij}) \psi_\mu(x_j) dx_j$$

and

$$V_\mu^{ex}(x_i) g(x_i) = \left(\int \psi_\mu^*(x_j) \hat{v}(r_{ij}) g(x_j) dx_j \right) \psi_\mu(x_i),$$

respectively.

Analysis of Hartree-Fock equations in coordinate space

The function $g(x_i)$ is an arbitrary function, and by the substitution $g(x_i) = \psi_\nu(x_i)$ we get

$$V_\mu^{ex}(x_i) \psi_\nu(x_i) = \left(\int \psi_\mu^*(x_j) \hat{v}(r_{ij}) \psi_\nu(x_j) dx_j \right) \psi_\mu(x_i).$$

We may then rewrite the Hartree-Fock equations as

$$\hat{h}^{HF}(x_i) \psi_\nu(x_i) = \epsilon_\nu \psi_\nu(x_i),$$

with

$$\hat{h}^{HF}(x_i) = \hat{h}_0(x_i) + \sum_{\mu=1}^N V_\mu^d(x_i) - \sum_{\mu=1}^N V_\mu^{ex}(x_i),$$

and where $\hat{h}_0(i)$ is the one-body part. The latter is normally chosen as a part which yields solutions in closed form. The harmonic oscillator is a classical problem thereof. We normally rewrite the last equation as

$$\hat{h}^{HF}(x_i) = \hat{h}_0(x_i) + \hat{g}^{HF}(x_i).$$

Hartree-Fock by varying the coefficients of a wave function expansion

Another possibility is to expand the single-particle functions in a known basis and vary the coefficients, that is, the new single-particle wave function is written as a linear expansion in terms of a fixed chosen orthogonal basis (for example the well-known harmonic oscillator functions or the hydrogen-like functions etc). We define our new Hartree-Fock single-particle basis by performing a unitary transformation on our previous basis (labelled with greek indices) as

$$\psi_p^{HF} = \sum_{\lambda} C_{p\lambda} \phi_{\lambda}. \quad (13)$$

In this case we vary the coefficients $C_{p\lambda}$. If the basis has infinitely many solutions, we need to truncate the above sum. We assume that the basis ϕ_{λ} is orthogonal. A unitary transformation keeps the orthogonality, as discussed in exercise 1 below.

Hartree-Fock by varying the coefficients of a wave function expansion

It is normal to choose a single-particle basis defined as the eigenfunctions of parts of the full Hamiltonian. The typical situation consists of the solutions of the one-body part of the Hamiltonian, that is we have

$$\hat{h}_0 \phi_{\lambda} = \epsilon_{\lambda} \phi_{\lambda}.$$

The single-particle wave functions $\phi_{\lambda}(\mathbf{r})$, defined by the quantum numbers λ and \mathbf{r} are defined as the overlap

$$\phi_{\lambda}(\mathbf{r}) = \langle \mathbf{r} | \lambda \rangle.$$

Hartree-Fock by varying the coefficients of a wave function expansion

In our discussions hereafter we will use our definitions of single-particle states above and below the Fermi (F) level given by the labels $ijkl \dots \leq F$ for so-called single-hole states and $abcd \dots > F$ for so-called particle states. For general single-particle states we employ the labels $pqrs \dots$.

Hartree-Fock by varying the coefficients of a wave function expansion

In Eq. (11), restated here

$$E[\Phi] = \sum_{\mu=1}^N \langle \mu | h | \mu \rangle + \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \langle \mu \nu | \hat{v} | \mu \nu \rangle A S,$$

we found the expression for the energy functional in terms of the basis function $\phi_{\lambda}(\mathbf{r})$. We then varied the above energy functional with respect to the basis functions $|\mu\rangle$. Now we are interested in defining a new basis defined in terms of a chosen basis as defined in Eq. (13). We can then rewrite the energy functional as

$$E[\Phi^{HF}] = \sum_{i=1}^N \langle i | h | i \rangle + \frac{1}{2} \sum_{ij=1}^N \langle ij | \hat{v} | ij \rangle A S, \quad (14)$$

where Φ^{HF} is the new Slater determinant defined by the new basis of Eq. (13).

Hartree-Fock by varying the coefficients of a wave function expansion

Using Eq. (13) we can rewrite Eq. (14) as

$$E[\Psi] = \sum_{i=1}^N \sum_{\alpha\beta} C_{i\alpha}^* C_{i\beta} \langle \alpha | h | \beta \rangle + \frac{1}{2} \sum_{ij=1}^N \sum_{\alpha\beta\gamma\delta} C_{i\alpha}^* C_{j\beta}^* C_{i\gamma} C_{j\delta} \langle \alpha\beta | \hat{v} | \gamma\delta \rangle A S. \quad (15)$$

Hartree-Fock by varying the coefficients of a wave function expansion

We wish now to minimize the above functional. We introduce again a set of Lagrange multipliers, noting that since $\langle i | j \rangle = \delta_{i,j}$ and $\langle \alpha | \beta \rangle = \delta_{\alpha,\beta}$, the coefficients $C_{i\gamma}$ obey the relation

$$\langle i | j \rangle = \delta_{i,j} = \sum_{\alpha\beta} C_{i\alpha}^* C_{j\beta} \langle \alpha | \beta \rangle = \sum_{\alpha} C_{i\alpha}^* C_{j\alpha},$$

which allows us to define a functional to be minimized that reads

$$F[\Phi^{HF}] = E[\Phi^{HF}] - \sum_{i=1}^N \epsilon_i \sum_{\alpha} C_{i\alpha}^* C_{i\alpha}. \quad (16)$$

Hartree-Fock by varying the coefficients of a wave function expansion

Minimizing with respect to $C_{i\alpha}^*$, remembering that the equations for $C_{i\alpha}^*$ and $C_{i\alpha}$ can be written as two independent equations, we obtain

$$\frac{d}{dC_{i\alpha}^*} \left[E[\Phi^{HF}] - \sum_j \epsilon_j \sum_{\alpha} C_{j\alpha}^* C_{j\alpha} \right] = 0,$$

which yields for every single-particle state i and index α (recalling that the coefficients $C_{i\alpha}$ are matrix elements of a unitary (or orthogonal for a real symmetric matrix) matrix) the following Hartree-Fock equations

$$\sum_{\beta} C_{i\beta} \langle \alpha | h | \beta \rangle + \sum_{j=1}^N \sum_{\beta \gamma \delta} C_{j\beta}^* C_{j\delta} C_{i\gamma} \langle \alpha \beta | \hat{v} | \gamma \delta \rangle_{AS} = \epsilon_i^{HF} C_{i\alpha}.$$

Hartree-Fock by varying the coefficients of a wave function expansion

We can rewrite this equation as (changing dummy variables)

$$\sum_{\beta} \left\{ \langle \alpha | h | \beta \rangle + \sum_j \sum_{\gamma \delta} C_{j\gamma}^* C_{j\delta} \langle \alpha \gamma | \hat{v} | \beta \delta \rangle_{AS} \right\} C_{i\beta} = \epsilon_i^{HF} C_{i\alpha}.$$

Note that the sums over greek indices run over the number of basis set functions (in principle an infinite number).

Hartree-Fock by varying the coefficients of a wave function expansion

Defining

$$h_{\alpha\beta}^{HF} = \langle \alpha | h | \beta \rangle + \sum_{j=1}^N \sum_{\gamma \delta} C_{j\gamma}^* C_{j\delta} \langle \alpha \gamma | \hat{v} | \beta \delta \rangle_{AS},$$

we can rewrite the new equations as

$$\sum_{\gamma} h_{\alpha\beta}^{HF} C_{i\gamma} = \epsilon_i^{HF} C_{i\alpha}. \quad (17)$$

The latter is nothing but a standard eigenvalue problem. Compared with Eq. (12), we see that we do not need to compute any integrals in an iterative procedure for solving the equations. It suffices to tabulate the matrix elements $\langle \alpha | h | \beta \rangle$ and $\langle \alpha \gamma | \hat{v} | \beta \delta \rangle_{AS}$ once and for all. Successive iterations require thus only a look-up in tables over one-body and two-body matrix elements. These details will be discussed below when we solve the Hartree-Fock equations

Hartree-Fock algorithm

Our Hartree-Fock matrix is thus

$$\hat{h}_{\alpha\beta}^{HF} = \langle \alpha | \hat{h}_0 | \beta \rangle + \sum_{j=1}^N \sum_{\gamma \delta} C_{j\gamma}^* C_{j\delta} \langle \alpha \gamma | \hat{v} | \beta \delta \rangle_{AS}.$$

The Hartree-Fock equations are solved in an iterative way starting with a guess for the coefficients $C_{j\gamma} = \delta_{j,\gamma}$ and solving the equations by diagonalization till the new single-particle energies ϵ_i^{HF} do not change anymore by a prefixed quantity.

Hartree-Fock algorithm

Normally we assume that the single-particle basis $|\beta\rangle$ forms an eigenbasis for the operator \hat{h}_0 , meaning that the Hartree-Fock matrix becomes

$$\hat{h}_{\alpha\beta}^{HF} = \epsilon_{\alpha} \delta_{\alpha,\beta} + \sum_{j=1}^N \sum_{\gamma \delta} C_{j\gamma}^* C_{j\delta} \langle \alpha \gamma | \hat{v} | \beta \delta \rangle_{AS}.$$

The Hartree-Fock eigenvalue problem

$$\sum_{\beta} \hat{h}_{\alpha\beta}^{HF} C_{i\beta} = \epsilon_i^{HF} C_{i\alpha},$$

can be written out in a more compact form as

$$\hat{h}^{HF} \hat{C} = \epsilon^{HF} \hat{C}.$$

Hartree-Fock algorithm

The Hartree-Fock equations are, in their simplest form, solved in an iterative way, starting with a guess for the coefficients $C_{i\alpha}$. We label the coefficients as $C_{i\alpha}^{(n)}$, where the subscript n stands for iteration n . To set up the algorithm we can proceed as follows:

- We start with a guess $C_{i\alpha}^{(0)} = \delta_{i,\alpha}$. Alternatively, we could have used random starting values as long as the vectors are normalized. Another possibility is to give states below the Fermi level a larger weight.
- The Hartree-Fock matrix simplifies then to (assuming that the coefficients $C_{i\alpha}$ are real)

$$\hat{h}_{\alpha\beta}^{HF} = \epsilon_{\alpha} \delta_{\alpha,\beta} + \sum_{j=1}^N \sum_{\gamma \delta} C_{j\gamma}^{(0)} C_{j\delta}^{(0)} \langle \alpha \gamma | \hat{v} | \beta \delta \rangle_{AS}.$$

Hartree-Fock algorithm

Solving the Hartree-Fock eigenvalue problem yields then new eigenvectors $C_{i\alpha}^{(1)}$ and eigenvalues $\epsilon_i^{HF(1)}$.

- With the new eigenvalues we can set up a new Hartree-Fock potential

$$\sum_{j=1}^N \sum_{\gamma\delta} C_{j\gamma}^{(1)} C_{j\delta}^{(1)} \langle \alpha\gamma | \hat{v} | \beta\delta \rangle A_S.$$

The diagonalization with the new Hartree-Fock potential yields new eigenvectors and eigenvalues. This process is continued till for example

$$\frac{\sum_p |\epsilon_i^{(n)} - \epsilon_i^{(n-1)}|}{m} \leq \lambda,$$

where λ is a user prefixed quantity ($\lambda \sim 10^{-8}$ or smaller) and p runs over all calculated single-particle energies and m is the number of single-particle states.

Analysis of Hartree-Fock equations and Koopman's theorem

We can rewrite the ground state energy by adding and subtracting $\hat{u}^{HF}(x_i)$

$$E_0^{HF} = \langle \Phi_0 | \hat{H} | \Phi_0 \rangle = \sum_{i \leq F} \langle i | \hat{h}_0 + \hat{u}^{HF} | j \rangle + \frac{1}{2} \sum_{i \leq F} \sum_{j \leq F} [\langle ij | \hat{v} | ij \rangle - \langle ij | \hat{v} | ji \rangle] - \sum_{i \leq F} \langle i | \hat{u}^{HF} | i \rangle$$

which results in

$$E_0^{HF} = \sum_{i \leq F} \epsilon_i^{HF} + \frac{1}{2} \sum_{i \leq F} \sum_{j \leq F} [\langle ij | \hat{v} | ij \rangle - \langle ij | \hat{v} | ji \rangle] - \sum_{i \leq F} \langle i | \hat{u}^{HF} | i \rangle.$$

Our single-particle states $ijk \dots$ are now single-particle states obtained from the solution of the Hartree-Fock equations.

Analysis of Hartree-Fock equations and Koopman's theorem

Using our definition of the Hartree-Fock single-particle energies we obtain then the following expression for the total ground-state energy

$$E_0^{HF} = \sum_{i \leq F} \epsilon_i - \frac{1}{2} \sum_{i \leq F} \sum_{j \leq F} [\langle ij | \hat{v} | ij \rangle - \langle ij | \hat{v} | ji \rangle].$$

This form will be used in our discussion of Koopman's theorem.

Analysis of Hartree-Fock equations and Koopman's theorem

Atomic physics case

We have

$$E[\Phi^{HF}(N)] = \sum_{i=1}^H \langle i | \hat{h}_0 | i \rangle + \frac{1}{2} \sum_{ij=1}^N \langle ij | \hat{v} | ij \rangle A_S,$$

where $\Phi^{HF}(N)$ is the new Slater determinant defined by the new basis of Eq. (13) for N electrons (same Z). If we assume that the single-particle wave functions in the new basis do not change when we remove one electron or add one electron, we can then define the corresponding energy for the $N-1$ systems as

$$E[\Phi^{HF}(N-1)] = \sum_{i=1, i \neq k}^N \langle i | \hat{h}_0 | i \rangle + \frac{1}{2} \sum_{ij=1, i, j \neq k}^N \langle ij | \hat{v} | ij \rangle A_S,$$

where we have removed a single-particle state $k \leq F$, that is a state below the Fermi level.

Analysis of Hartree-Fock equations and Koopman's theorem

Calculating the difference

$$E[\Phi^{HF}(N)] - E[\Phi^{HF}(N-1)] = \langle k | \hat{h}_0 | k \rangle + \frac{1}{2} \sum_{i=1, i \neq k}^N \langle ik | \hat{v} | ik \rangle A_S - \frac{1}{2} \sum_{j=1, j \neq k}^N \langle kj | \hat{v} | kj \rangle A_S$$

we obtain

$$E[\Phi^{HF}(N)] - E[\Phi^{HF}(N-1)] = \langle k | \hat{h}_0 | k \rangle + \frac{1}{2} \sum_{j=1}^N \langle kj | \hat{v} | kj \rangle A_S$$

which is just our definition of the Hartree-Fock single-particle energy

$$E[\Phi^{HF}(N)] - E[\Phi^{HF}(N-1)] = \epsilon_k^{HF}$$

Analysis of Hartree-Fock equations and Koopman's theorem

Similarly, we can now compute the difference (we label the single-particle states above the Fermi level as $abcd > F$)

$$E[\Phi^{HF}(N+1)] - E[\Phi^{HF}(N)] = \epsilon_a^{HF}.$$

These two equations can thus be used to the electron affinity or ionization energies, respectively. Koopman's theorem states that for example the ionization energy of a closed-shell system is given by the energy of the highest occupied single-particle state. If we assume that changing the number of electrons from N to $N+1$ does not change the Hartree-Fock single-particle energies and eigenfunctions, then Koopman's theorem simply states that the ionization energy of an atom is given by the single-particle energy of the last bound state. In a similar way, we can also define the electron affinities.

Analysis of Hartree-Fock equations and Koopman's theorem

As an example, consider a simple model for atomic sodium, Na. Neutral sodium has eleven electrons, with the weakest bound one being confined the 3s single-particle quantum numbers. The energy needed to remove an electron from neutral sodium is rather small, 5.1391 eV, a feature which pertains to all alkali metals. Having performed a Hartree-Fock calculation for neutral sodium would then allows us to compute the ionization energy by using the single-particle energy for the 3s states, namely $\epsilon_{3s}^{\text{HF}}$. From these considerations, we see that Hartree-Fock theory allows us to make a connection between experimental observables (here ionization and affinity energies) and the underlying interactions between particles. In this sense, we are now linking the dynamics and structure of a many-body system with the laws of motion which govern the system. Our approach is a reductionistic one, meaning that we want to understand the laws of motion in terms of the particles or degrees of freedom which we believe are the fundamental ones. Our Slater determinant, being constructed as the product of various single-particle functions, follows this

Analysis of Hartree-Fock equations, Koopman's theorem

With similar arguments as in atomic physics, we can now use Hartree-Fock theory to make a link between nuclear forces and separation energies. Changing to nuclear system, we define

$$E[\Phi^{\text{HF}}(A)] = \sum_{i=1}^N \langle i | \hat{h}_0 | i \rangle + \frac{1}{2} \sum_{ij=1}^N \langle ij | \hat{v} | ij \rangle_{AS},$$

where $\Phi^{\text{HF}}(A)$ is the new Slater determinant defined by the new basis of Eq. (13) for A nucleons, where $A = N + Z$, with N now being the number of neutrons and Z the number of protons. If we assume again that the single-particle wave functions in the new basis do not change from a nucleus with A nucleons to a nucleus with $A - 1$ nucleons, we can then define the corresponding energy for the $A - 1$ systems as

$$E[\Phi^{\text{HF}}(A - 1)] = \sum_{i=1, i \neq k}^N \langle i | \hat{h}_0 | i \rangle + \frac{1}{2} \sum_{ij=1, i, j \neq k}^N \langle ij | \hat{v} | ij \rangle_{AS},$$

Analysis of Hartree-Fock equations and Koopman's theorem

Calculating the difference

$$E[\Phi^{\text{HF}}(A)] - E[\Phi^{\text{HF}}(A - 1)] = \langle k | \hat{h}_0 | k \rangle + \frac{1}{2} \sum_{i=1, i \neq k}^N \langle ik | \hat{v} | ik \rangle_{AS} - \frac{1}{2} \sum_{j=1, j \neq k}^N \langle kj | \hat{v} | kj \rangle_{AS}$$

which becomes

$$E[\Phi^{\text{HF}}(A)] - E[\Phi^{\text{HF}}(A - 1)] = \langle k | \hat{h}_0 | k \rangle + \frac{1}{2} \sum_{j=1}^N \langle kj | \hat{v} | kj \rangle_{AS}$$

which is just our definition of the Hartree-Fock single-particle energy

$$E[\Phi^{\text{HF}}(A)] - E[\Phi^{\text{HF}}(A - 1)] = \epsilon_k^{\text{HF}}$$

Analysis of Hartree-Fock equations and Koopman's theorem

Similarly, we can now compute the difference (recall that the single-particle states $abcd > F$)

$$E[\Phi^{\text{HF}}(A + 1)] - E[\Phi^{\text{HF}}(A)] = \epsilon_a^{\text{HF}}.$$

If we then recall that the binding energy differences

$$BE(A) - BE(A - 1) \quad \text{and} \quad BE(A + 1) - BE(A),$$

define the separation energies, we see that the Hartree-Fock single-particle energies can be used to define separation energies. We have thus our first link between nuclear forces (included in the potential energy term) and an observable quantity defined by differences in binding energies.

Analysis of Hartree-Fock equations and Koopman's theorem

We have thus the following interpretations (if the single-particle field do not change)

$$BE(A) - BE(A - 1) \approx E[\Phi^{\text{HF}}(A)] - E[\Phi^{\text{HF}}(A - 1)] = \epsilon_k^{\text{HF}},$$

and

$$BE(A + 1) - BE(A) \approx E[\Phi^{\text{HF}}(A + 1)] - E[\Phi^{\text{HF}}(A)] = \epsilon_a^{\text{HF}}.$$

If we use ^{16}O as our closed-shell nucleus, we could then interpret the separation energy

$$BE(^{16}\text{O}) - BE(^{15}\text{O}) \approx \epsilon_{0p_{1/2}'}^{\text{HF}},$$

and

$$BE(^{16}\text{O}) - BE(^{15}\text{N}) \approx \epsilon_{0p_{1/2}''}^{\text{HF}}.$$

Analysis of Hartree-Fock equations and Koopman's theorem

Similarly, we could interpret

$$BE(^{17}\text{O}) - BE(^{16}\text{O}) \approx \epsilon_{0d_{5/2}'}^{\text{HF}},$$

and

$$BE(^{17}\text{F}) - BE(^{16}\text{O}) \approx \epsilon_{0d_{5/2}''}^{\text{HF}}.$$

We can continue like this for all $A \pm 1$ nuclei where A is a good closed-shell (or subshell closure) nucleus. Examples are ^{22}O , ^{24}O , ^{40}Ca , ^{48}Ca , ^{52}Ca , ^{54}Ca , ^{56}Ni , ^{68}Ni , ^{78}Ni , ^{90}Zr , ^{88}Sr , ^{100}Sn , ^{132}Sn and ^{208}Pb , to mention some possible cases.

Analysis of Hartree-Fock equations and Koopman's theorem

We can thus make our first interpretation of the separation energies in terms of the simplest possible many-body theory. If we also recall that the so-called energy gap for neutrons (or protons) is defined as

$$\Delta S_n = 2BE(N, Z) - BE(N-1, Z) - BE(N+1, Z),$$

for neutrons and the corresponding gap for protons

$$\Delta S_p = 2BE(N, Z) - BE(N, Z-1) - BE(N, Z+1),$$

we can define the neutron and proton energy gaps for ^{16}O as

$$\Delta S_\nu = \epsilon_{0d_{5/2}}^{\text{HF}} - \epsilon_{0p_{1/2}}^{\text{HF}},$$

and

$$\Delta S_\pi = \epsilon_{0d_{5/2}}^{\text{HF}} - \epsilon_{0p_{1/2}}^{\text{HF}}.$$

Basic definitions and motivation

Our final aim is to develop a Hartree-Fock code which can be used to study properties of both atoms and molecules.

For a molecular system, the eigenfunctions of the Hartree-Fock equations are called *molecular orbitals* (MOs). It is important to distinguish these from the perhaps more familiar *atomic orbitals*, where we can think of (as we have done till now) that the electrons occupy some well-defined hydrogen-like states.

An example like the H_2 -molecule may suffice. In the ground state the electrons are not occupying the $1s$ -orbitals of atomic Hydrogen. The molecular system is entirely different from the atomic one, with an entirely different Hamiltonian, and the eigenstates of the Hartree-Fock equations will therefore also be different.

Basic definitions and motivation

In order to solve the Hartree-Fock equations, we need to expand the molecular orbitals in a known set of basis functions

$$\phi_k(\vec{r}) = \sum_{\mu=1}^M \chi_{\mu k}(\vec{r}). \quad (18)$$

The question we then need to answer is how to choose basis functions $\chi_{\mu k}$? The answer to this is primarily dictated by the following three criteria:

- The functions should make physically sense, i.e., they should have a large probability where the electrons are likely to be and small elsewhere.
- It should be possible to integrate the functions efficiently.
- The solution of the Hartree-Fock equations should converge towards the Hartree-Fock limit as the number of basis functions increases.

The first point suggests that we choose atomic orbitals as basis

Basic definitions, Slater orbitals

A much set of state functions are the so-called Slater type orbitals (STO), defined as

$$\chi^{STO}(r, \theta, \phi, n, l, m) = \frac{(2a)^{n+1/2}}{[(2n)!]^{1/2}} r^{n-1} \exp(-ar) Y_l^m(\theta, \phi), \quad (19)$$

where n is the principal quantum number, l and m are the orbital momentum quantum numbers, $Y_l^m(\theta, \phi)$ are the spherical harmonics familiar from the solution of the Schrödinger equation for the Hydrogen atom, and a is an exponent which determines the radial decay of the function. The main attractive features of the STOs are that they have the correct exponential decay with increasing r and that the orbital components are hydrogenic. Furthermore, they can be used to represent in a better way weakly bound states or even resonances. The hydrogen-like state functions represent only bound states. STOs are mainly used in atomic Hartree-Fock calculations. When doing molecular calculations, however, they have the disadvantage that the two-particle integrals $\langle \mu\lambda | g | \nu\lambda \rangle$ need to compute an anti-symmetrized matrix interaction

Basic definitions

A clever trick which makes multiple center integrals easier to handle is to replace the exponential term $\exp(-ar)$ with $\exp(-ar^2)$, that is, to functions proportional with Gaussians. This simplifies considerably the integrals to evaluate since the product of two Gaussians centered on nuclei with positions \vec{A} and \vec{B} is equal to one Gaussian centered on some point \vec{P} on the line between them:

$$\exp(-a|\vec{r}-\vec{A}|^2) \cdot \exp(-b|\vec{r}-\vec{B}|^2) = K_{AB} \exp(-p|\vec{r}-\vec{P}|^2), \quad (20)$$

where

$$K_{AB} = \exp\left(-\frac{ab}{a+b}|\vec{A}-\vec{B}|^2\right), \quad (21)$$

$$\vec{P} = \frac{a\vec{A} + b\vec{B}}{a+b}, \quad (22)$$

$$p = a + b. \quad (23)$$

This is the so-called *Gaussian product theorem*.

Basic definitions

The general functional form of a normalised Gaussian type orbital centered at \vec{A} is given by

$$G_{ijk}(a, \vec{r}_A) = \left(\frac{2a}{\pi}\right)^{3/4} \frac{[(8a)^{i+j+k} i! j! k!]}{[(2i)! (2j)! (2k)!]} x_A^i y_A^j z_A^k \exp(-ar_A^2), \quad (24)$$

where $\vec{r}_A = \vec{r} - \vec{A}$ and the integers i, j, k determine the orbital momentum quantum number $l = i + j + k$.

Disadvantages

The greatest drawback with GTOs is that they do not have the proper exponential radial decay. This can be fixed by forming linear combinations of GTOs to resemble the STOs

$$\chi^{CGTO}(\vec{r}_A, i, j, k) = \sum_{p=1}^L d_p G_{ijk}(a_p, \vec{r}_A). \quad (25)$$

These are called STO-LG basis functions, where L refers to the number of Gaussians used in the linear combination. The individual Gaussians are called *primitive* basis functions and the linear combinations are called *contracted* basis functions, hence the label CGTO (Contracted Gaussian Type Orbital). These are the functions we will employ in this project.

A very common choice for the STO-LG basis sets is $L = 3$. It is important to note that the parameters (a_p, d_p) are static and that the linear combination of Gaussians constitute *one single* basis function. With the wording "basis function", we will refer to a contracted basis function.

Basic definitions

The STO-LG basis sets belong to the family of *minimal basis sets*. It means that there is one and only one basis function per atomic orbital. The STO-LG basis sets for the Hydrogen and Helium atoms, for example, contain only one basis function for the 1s atomic orbital. This basis function is, as explained above, composed of a linear combination of L primitives. For the atoms Lithium through Neon the STO-LG basis sets contain 5 basis functions; one for each of the atomic orbitals 1s, 2s, 2p_x, 2p_y and 2p_z.

p	1	2	3
d _p	0.1543	0.5353	0.4446
a _p	3.4252	0.6239	0.1688

Cartesian Gaussians and how to compute integrals

Using GTOs as basis functions improves the speed of the integrations which must be done when setting up for example the Fock matrix. The Cartesian Gaussian functions centered at \vec{A} are given by

$$G_{ijk}(a, \vec{r}_A) = x_A^i y_A^j z_A^k \exp(-a r_A^2), \quad (26)$$

where $\vec{r}_A = \vec{r} - \vec{A}$. These will be our so-called primitive basis functions (see also the flow chart of Hartree-Fock code). They factorize in the Cartesian components

$$G_{ijk}(a, \vec{r}_A) = G_i(a, x_A) G_j(a, y_A) G_k(a, z_A), \quad (27)$$

where

$$G_i(a, x_A) = x_A^i \exp(-a x_A^2), \quad (28)$$

and the other factors are defined similarly. Each of the components obey the simple recurrence relation

Integral evaluations

We introduce shorthand notation

$$G_a(\vec{r}) = G_{ikm}(a, \vec{r}_A), \quad (30)$$

$$G_b(\vec{r}) = G_{jln}(b, \vec{r}_B), \quad (31)$$

and define the overlap distribution as

$$\Omega_{ab}(\vec{r}) = G_a(\vec{r}) G_b(\vec{r}). \quad (32)$$

Using the Gaussian product theorem (20) this can be written as

$$\Omega_{ab}(\vec{r}) = K_{AB} x_A^i x_B^j y_A^k y_B^l z_A^m z_B^n \exp(-p r_p^2), \quad (33)$$

where

$$K_{AB} = \exp\left(-\frac{ab}{a+b} R_{AB}^2\right)$$

$$\vec{R}_{AB} = \vec{A} - \vec{B}$$

Integral evaluations

Because the Gaussians G_a and G_b factorize in their Cartesian components, such a factorization applies to the overlap distribution as well, namely

$$\Omega_{ab}(\vec{r}) = \Omega_{ij}(x) \Omega_{kl}(y) \Omega_{mn}(z), \quad (35)$$

where

$$\Omega_{ij} = K_{AB}^x x_A^i x_B^j \exp(-p x_p^2) \quad (36)$$

and

$$K_{AB}^x = \exp\left(-\frac{ab}{a+b} X_{AB}^2\right)$$

$$X_{AB} = A_x - B_x. \quad (37)$$

The distributions $\Omega_{kl}(y)$ and $\Omega_{mn}(z)$ are defined similarly.

Integral evaluations

A very useful representation of the Cartesian GTOs is given by the so-called Hermite Gaussians. These functions simplify the integrations significantly. The Hermite Gaussians centered at \vec{P} are defined by

$$\Lambda_{tuv}(p, \vec{r}_p) = \left(\frac{\partial}{\partial P_x}\right)^t \left(\frac{\partial}{\partial P_y}\right)^u \left(\frac{\partial}{\partial P_z}\right)^v \exp(-p r_p^2), \quad (38)$$

where $\vec{r}_p = \vec{r} - \vec{P}$. They factorize as (like the cartesian GTOs also do)

$$\Lambda_{tuv}(p, \vec{r}_p) = \Lambda_t(p, x_p) \Lambda_u(p, y_p) \Lambda_v(p, z_p), \quad (39)$$

where

$$\Lambda_t(p, x_p) = \left(\frac{\partial}{\partial P_x}\right)^t \exp(-p x_p^2), \quad (40)$$

and the other factors are defined similarly.

Integral evaluations

However, their recurrence relation is quite different from that of the Cartesian Gaussians:

$$\begin{aligned}\Lambda_{t+1}(p, x_P) &= \left(\frac{\partial}{\partial P_x}\right)^t \frac{\partial}{\partial P_x} \exp(-p x_P^2) \\ &= \left(\frac{\partial}{\partial P_x}\right)^t 2p x_P \exp(-p x_P^2) \\ &= 2p \left[-t \left(\frac{\partial}{\partial P_x}\right)^{t-1} + x_P \left(\frac{\partial}{\partial P_x}\right)^t \right] \exp(-p x_P^2) \\ &= 2p [-t \Lambda_{t-1} + x_P \Lambda_t],\end{aligned}\quad (41)$$

where we have used that

$$\left(\frac{\partial}{\partial x}\right)^t x f(x) = t \left(\frac{\partial}{\partial x}\right)^{t-1} f(x) + x \left(\frac{\partial}{\partial x}\right)^t f(x). \quad (42)$$

The recurrence relation reads

$$x_P \Lambda_t = \frac{1}{2p} \Lambda_{t+1} + t \Lambda_{t-1}. \quad (43)$$

Integral evaluations

Our first goal is to compute the overlap integral

$$S_{ab} = \langle G_a | G_b \rangle = \int d\vec{r} \Omega_{ab}(\vec{r}) \quad (44)$$

between two Gaussians centered at the points \vec{A} and \vec{B} . Since the overlap distribution Ω_{ab} factorizes in the Cartesian components, the integrals over x , y and z can be calculated independently of each other

$$\begin{aligned}S_{ab} &= \langle G_i | G_j \rangle \langle G_k | G_l \rangle \langle G_m | G_n \rangle \\ &= S_{ij} S_{kl} S_{mn}.\end{aligned}\quad (45)$$

The x component of the overlap integral, for example, is given by

$$S_{ij} = \int dx \Omega_{ij}(x) \quad (46)$$

Integral evaluations

In equation (46) the two-center GTOs have been reduced to a one-center Gaussian. However, the integral is still not straightforward to calculate because of the powers x_A^i and x_B^j . A smart way to deal with this is to express the Cartesian Gaussian in terms of the Hermite Gaussians. Note that (40) is a polynomial of order t in x multiplied by the exponential function. In equation (46) the polynomial is of order $i+j$. This means that we can express the overlap distribution $\Omega_{ij}(x)$ in equation (46) in terms of the Hermite Gaussians in (40) in the following way:

$$\Omega_{ij}(x) = \sum_{t=0}^{i+j} E_t^{ij} \Lambda_t(p, x_P), \quad (47)$$

where E_t^{ij} are constants. Note that the sum is over t only. The indices i and j are static and are determined from the powers of x in G_i and G_j . We use them as labels on the coefficients E_t^{ij} because different sets of indices will lead to different sets of coefficients.

Integral evaluations

To get the overlap integral in the x -direction we integrate (47) over \mathbb{R} , which now turns out to be extremely easy; the only term that survives the integration is the term for $t=0$:

$$\int dx \Lambda_t(p, x_P) = \int dx \left(\frac{\partial}{\partial P_x}\right)^t \exp(-p x_P^2), \quad (48)$$

$$= \left(\frac{\partial}{\partial P_x}\right)^t \int dx \exp(-p x_P^2), \quad (49)$$

$$= \sqrt{\frac{\pi}{p}} \delta_{t0}. \quad (50)$$

We have used Leibniz' rule, which says that the differentiation of an integrand with respect to a variable which is not an integration variable can be moved outside the integral. Thus the integral in (46) is simply

$$S_{ij} = E_0^{ij} \sqrt{\frac{\pi}{p}}. \quad (51)$$

Integral evaluations

The same procedure can be used for the integrals with respect to y and z , which means that the total overlap integral is

$$S_{ab} = E_0^{ij} E_0^{kl} E_0^{mn} \left(\frac{\pi}{p}\right)^{3/2}. \quad (52)$$

We need thereafter to determine the coefficients E_t^{ij} . When $i=j=0$ in equation (47) we have

$$E_0^{0,0} = K_{AB}^x. \quad (53)$$

The other coefficients are found via the following recurrence relations

$$\begin{aligned}E_t^{i+1,j} &= \frac{1}{2p} E_{t-1}^{ij} + X_{PA} E_t^{ij} + (t+1) E_{t+1}^{ij} \\ E_t^{i,j+1} &= \frac{1}{2p} E_{t-1}^{ij} + X_{PB} E_t^{ij} + (t+1) E_{t+1}^{ij}.\end{aligned}\quad (54)$$

Integral evaluations

Analogous expressions hold for the coefficients E_u^{kl} and E_v^{mn} . The first equation in (54) can be derived by comparing two equivalent ways of expanding the product $G_{i+1} G_j$ in Hermite Gaussians. The first way is

$$G_{i+1} G_j = \sum_{t=0}^{i+j+1} E_t^{i+1,j} \Lambda_t, \quad (55)$$

and the second way is

$$\begin{aligned}G_{i+1} G_j &= x_A G_i G_j \\ &= [(x - P_x) + (P_x - A_x)] \sum_{t=0}^{i+j} E_t^{ij} \Lambda_t \\ &= \sum_{t=0}^{i+j} [x_P + X_{PA}] E_t^{ij} \Lambda_t \\ &\quad + \sum_{t=0}^{i+j-1} E_t^{ij} \Lambda_t.\end{aligned}\quad (56)$$

Integral evaluations, kinetic energy

Note that the change in summation indices in equation (56) implies that we must define

$$E_t^{ij} = 0, \quad \text{if } t < 0 \text{ or } t > i + j. \quad (57)$$

Next we turn to the evaluation of the kinetic integral:

$$\begin{aligned} T_{ab} &= -\frac{1}{2} \langle G_a | \nabla^2 | G_b \rangle \\ &= -\frac{1}{2} \langle G_{ikm}(a, \vec{r}_A) | \nabla^2 | G_{jln}(b, \vec{r}_B) \rangle \\ &= -\frac{1}{2} (T_{ij} S_{kl} S_{mn} + S_{ij} T_{kl} S_{mn} + S_{ij} S_{kl} T_{mn}), \end{aligned} \quad (58)$$

where

$$T_{ij} = \int dx G_i(a, x_A) \frac{\partial^2}{\partial x^2} G_j(b, x_B), \quad (59)$$

and the other factors are defined in the same way. Performing the

Integral evaluations, one-body Coulomb interaction

We now turn to the Coulomb integral due to the interaction between the electrons and the atomic nuclei

$$V_{ab} = \langle G_a | \frac{1}{r_C} | G_b \rangle, \quad (61)$$

where $r_C = |\vec{r} - \vec{C}|$. As before, the overlap distribution is expanded in terms of Hermite Gaussians

$$\begin{aligned} V_{ab} &= \int d\vec{r} \frac{\Omega_{ab}(\vec{r})}{r_C} \\ &= \sum_{tuv} E_t^{ij} E_u^{kl} E_v^{mn} \int d\vec{r} \frac{\Lambda_{tuv}(p, \vec{r}_P)}{r_C} \\ &= \sum_{tuv} E_{tuv}^{ab} \int d\vec{r} \frac{\Lambda_{tuv}(p, \vec{r}_P)}{r_C}. \end{aligned} \quad (62)$$

Here we have used the shorthand notation

Integral evaluations, one-body Coulomb interaction

We will show that this three-dimensional integral can actually be converted to a one-dimensional one. The trick is to observe that the factor $1/r_C$ can be replaced by the integral

$$\frac{1}{r_C} = \frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} dt \exp(-r_C^2 t^2). \quad (65)$$

Inserting this into V_p and using the Gaussian product theorem gives

$$\begin{aligned} V_p &= \int \exp(-p r_P^2) \left(\frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} \exp(-r_C^2 t^2) dt \right) d\vec{r} \\ &= \frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} \int \exp\left(-\frac{pt^2}{p+t^2} R_{PC}^2\right) \exp[-(p+t^2)r_S^2] d\vec{r} dt, \end{aligned} \quad (66)$$

where $\vec{R}_{PC} = \vec{P} - \vec{C}$ and $\vec{r}_S = \vec{r} - \vec{S}$ for some point \vec{S} . Doing the integral over the spatial coordinates reveals that the specific value of \vec{S} is not relevant

Integral evaluations, one-body Coulomb interaction

Next we change integration variable from t to u by defining

$$u^2 = \frac{t^2}{p+t^2}. \quad (70)$$

This will change the range of integration from $[0, \infty)$ to $[0, 1]$. This is beneficial because the final integral at which we arrive will be calculated numerically. The change of variables leads to

$$V_p = \frac{2\pi}{p} \int_0^1 \exp(-p R_{PC}^2 u^2) du \quad (71)$$

$$= \frac{2\pi}{p} F_0(p R_{PC}^2), \quad (72)$$

where $F_0(x)$ is a special instance of the Boys function $F_n(x)$ which is defined as

$$F_n(x) = \int_0^1 \exp(-xt^2) t^{2n} dt. \quad (73)$$

Integral evaluations, one-body Coulomb interaction

We have now a simplified way of calculating the integral of Λ_{000}/r_C . However, we need to integrate Λ_{tuv}/r_C for general values of t , u and v . These integrals are actually not that hard to do once the Boys function is calculated:

$$V_{ab} = \frac{2\pi}{p} \sum_{tuv} E_{tuv}^{ab} \int d\vec{r} \frac{\Lambda_{tuv}(p, \vec{r}_P)}{r_C} \quad (74)$$

$$= \frac{2\pi}{p} \sum_{tuv} E_{tuv}^{ab} \frac{\partial^{t+u+v} F_0(p R_{PC}^2)}{\partial P_x^t \partial P_y^u \partial P_z^v} \quad (75)$$

$$= \frac{2\pi}{p} \sum_{tuv} E_{tuv}^{ab} R_{tuv}(p, \vec{R}_{PC}), \quad (76)$$

where we have defined

$$R_{tuv}(p, \vec{R}_{PC}) = \frac{\partial^{t+u+v} F_0(p R_{PC}^2)}{\partial P_x^t \partial P_y^u \partial P_z^v}. \quad (77)$$

Integral evaluations, one-body Coulomb interaction

We need to calculate derivatives of the function F_0 . Note first that

$$\frac{d}{dx} F_n(x) = -F_{n+1}(x). \quad (78)$$

This means that it is possible to derive analytical expressions for the Coulomb term V_{ab} . However, in practice they are calculated recursively in a manner similar to the way we calculate the coefficients E_t^{ij} . Before presenting the recursion relations, we introduce the so-called auxiliary Hermite integrals

$$R_{tuv}^n(p, \vec{R}_{PC}) = (-2p)^n \frac{\partial^{t+u+v} F_n(p R_{PC}^2)}{\partial P_x^t \partial P_y^u \partial P_z^v}. \quad (79)$$

Integral evaluations, one-body Coulomb interaction

By starting with the source terms $R_{000}^n(p, \vec{R}_{PC}) = (-2p)^n F_n(pR_{PC}^2)$ we can reach the targets $R_{tuv}^n(p, \vec{R}_{PC}) = R_{tuv}(p, \vec{R}_{PC})$ through the following recurrence relations

$$\begin{aligned} R_{t+1,u,v}^n &= tR_{t-1,u,v}^{n+1} + X_{PC}R_{tuv}^{n+1} \\ R_{t,u+1,v}^n &= uR_{t,u-1,v}^{n+1} + Y_{PC}R_{tuv}^{n+1} \\ R_{t,u,v+1}^n &= vR_{t,u,v-1}^{n+1} + Z_{PC}R_{tuv}^{n+1}. \end{aligned} \quad (80)$$

The first of these are derived as follows

$$R_{t+1,u,v}^n = (-2p)^n \frac{\partial^{t+u+v}}{\partial P_x^t \partial P_y^u \partial P_z^v} [2p X_{PC} F_n(pR_{PC}^2)] \quad (81)$$

$$= (-2p)^{n+1} \frac{\partial^{t+u+v}}{\partial P_x^t \partial P_y^u \partial P_z^v} [X_{PC} F_{n+1}(pR_{PC}^2)] \quad (82)$$

$$= (-2p)^{n+1} \frac{\partial^{u+v}}{\partial P_y^u \partial P_z^v} \left[t \frac{\partial^{t-1}}{\partial P_x^{t-1}} + X_{PC} \frac{\partial^t}{\partial P_x^t} \right] F_{n+1}(pR_{PC}^2)$$

Integral evaluations, two-body Coulomb interaction

Finally we show how to calculate the Coulomb integral due to the interaction between the electrons. It is given by

$$\begin{aligned} g_{acbd} &= \langle G_a G_c | \frac{1}{r_{12}} | G_b G_d \rangle \\ &= \int \int \frac{\Omega_{ab}(\vec{r}_1) \Omega_{cd}(\vec{r}_2)}{r_{12}} d\vec{r}_1 d\vec{r}_2 \\ &= \sum_{tuv} \sum_{\tau\nu\phi} E_{tuv}^{ab} E_{\tau\nu\phi}^{cd} \int \int \frac{\Lambda_{tuv}(p, \vec{r}_1) \Lambda_{\tau\nu\phi}(q, \vec{r}_2)}{r_{12}} d\vec{r}_1 d\vec{r}_2 \\ &= \sum_{tuv} \sum_{\tau\nu\phi} E_{tuv}^{ab} E_{\tau\nu\phi}^{cd} \frac{\partial^{t+u+v}}{\partial P_x^t \partial P_y^u \partial P_z^v} \frac{\partial^{\tau+\nu+\phi}}{\partial Q_x^\tau \partial Q_y^\nu \partial Q_z^\phi} \\ &\quad \int \int \frac{\exp(-p\vec{r}_1^2) \exp(-q\vec{r}_2^2)}{r_{12}} d\vec{r}_1 d\vec{r}_2, \end{aligned} \quad (85)$$

where, similar to p and \vec{r}_1 , we have defined

Integral evaluations, two-body Coulomb interaction

Thus we need to evaluate the integral

$$V_{pq} = \int \int \frac{\exp(-p\vec{r}_1^2) \exp(-q\vec{r}_2^2)}{r_{12}} d\vec{r}_1 d\vec{r}_2. \quad (87)$$

By first integrating over \vec{r}_1 and using equation (71) this can be written as

$$V_{pq} = \int \left(\frac{2\pi}{p} \int_0^1 \exp(-p r_{2p}^2 u^2) du \right) \exp(-q\vec{r}_2^2) d\vec{r}_2. \quad (88)$$

Next we change the order of integration and use the Gaussian product theorem to get

$$\begin{aligned} V_{pq} &= \frac{2\pi}{p} \int_0^1 \int \exp\left(-\frac{pq u^2}{pu^2 + q} R_{PQ}^2\right) \exp[-(pu^2 + q)r_{2S}^2] d\vec{r}_2 du \\ &= \frac{2\pi}{p} \int_0^1 \exp\left(-\frac{pq u^2}{pu^2 + q} R_{PQ}^2\right) \left(\frac{\pi}{pu^2 + q}\right)^{3/2} du. \end{aligned}$$

Integral evaluations, two-body Coulomb interaction

Again, the value of \vec{S} in $\vec{r}_{2S} = \vec{r}_2 - \vec{S}$ is not relevant. If we now make the change of variables

$$\frac{v^2}{p+q} = \frac{u^2}{pu^2 + q}, \quad (90)$$

we get the result

$$V_{pq} = \frac{2\pi^{5/2}}{pq\sqrt{p+q}} F_0\left(\frac{pq}{p+q} R_{PQ}^2\right). \quad (91)$$

From this we get the final answer

$$\begin{aligned} g_{acbd} &= \frac{2\pi^{5/2}}{pq\sqrt{p+q}} \sum_{tuv} \sum_{\tau\nu\phi} (-1)^{\tau+\nu+\phi} E_{tuv}^{ab} E_{\tau\nu\phi}^{cd} \\ &\quad \frac{\partial^{t+u+v+\tau+\nu+\phi}}{\partial P_x^{t+\tau} \partial P_y^{u+\nu} \partial P_z^{v+\phi}} F_0\left(\frac{pq}{p+q} R_{PQ}^2\right) \end{aligned}$$

Integral evaluations, two-body Coulomb interaction and Boys function

Calculating the Coulomb integrals boils down to evaluating the Boys function

$$F_n(x) = \int_0^1 \exp(-xt^2) t^{2n} dt. \quad (94)$$

Doing this by standard numerical procedures is computationally expensive and should therefore be avoided. Here we describe a possible way to calculate the Boys function efficiently.

First note that if x is very large, the function value will hardly be affected by changing the upper limit of the integral from 1 to ∞ . Doing this is useful since the integral can be calculated exactly.

Thus, we have the following approximation for the Boys function for large x :

$$F_n(x) \approx \frac{(2n-1)!!}{2^{n+1}} \sqrt{\frac{\pi}{x^{2n+1}}}. \quad (x \text{ large}) \quad (95)$$

For small values of x there seems to be no escape from numerical

Integral evaluations, two-body Coulomb interaction and Boys function

The computational cost can be reduced even further by calculating the Boys function according to the description above only for the highest values of n needed; for lower values of n the function can be found via the recursion relation

$$F_n(x) = \frac{2xF_{n+1}(x) + e^{-x}}{2n+1}, \quad (97)$$

which can be shown by integrating the function by parts.

Integral evaluations, summarizing

The Gaussian functions are given by

$$\begin{aligned} G_a(\vec{r}) &= G_{ikm}(a, \vec{r}_A) = x_A^i y_A^j z_A^k \exp(-a r_A^2), \\ G_b(\vec{r}) &= G_{jln}(b, \vec{r}_B) = x_B^j y_B^l z_B^n \exp(-b r_B^2), \end{aligned} \quad (98)$$

where $\vec{r}_A = \vec{r} - \vec{A}$ and $\vec{r}_B = \vec{r} - \vec{B}$. We further define

$$\begin{aligned} p &= a + b, \\ \vec{P} &= \frac{a\vec{A} + b\vec{B}}{a + b}. \end{aligned} \quad (99)$$

Integral evaluations, summarizing

The overlap integral

$$S_{ab} = \langle G_a | G_b \rangle \quad (100)$$

is calculated as

$$S_{ab} = E_0^{ij} E_0^{kl} E_0^{mn} \left(\frac{\pi}{p} \right)^{3/2}, \quad (101)$$

where

$$E_0^{i=0,j=0} = \exp\left(-\frac{ab}{a+b} X_{AB}^2\right), \quad (102)$$

and the desired coefficients are found via

$$\begin{aligned} E_t^{i+1,j} &= \frac{1}{2p} E_{t-1}^{ij} + X_{PA} E_t^{ij} + (t+1) E_{t+1}^{ij}, \\ E_t^{i,j+1} &= \frac{1}{2p} E_{t-1}^{ij} + X_{PB} E_t^{ij} + (t+1) E_{t+1}^{ij}. \end{aligned} \quad (103)$$

Integral evaluations, summarizing

The kinetic integral is calculated as

$$T_{ab} = -\frac{1}{2} (T_{ij} S_{kl} S_{mn} + S_{ij} T_{kl} S_{mn} + S_{ij} S_{kl} T_{mn}), \quad (104)$$

where

$$T_{ij} = 4b^2 S_{i,j+2} - 2b(2j+1) S_{i,j} + j(j-1) S_{i,j-2}. \quad (105)$$

Integral evaluations, summarizing

The Coulomb integral

$$V_{ab} = \langle G_a | \frac{1}{r_C} | G_b \rangle \quad (106)$$

is calculated as

$$V_{ab} = \frac{2\pi}{p} \sum_{tuv} E_{tuv}^{ab} R_{tuv}(p, \vec{R}_{PC}), \quad (107)$$

where

$$E_{tuv}^{ab} = E_t^{ij} E_u^{kl} E_v^{mn}, \quad (108)$$

and $R_{tuv}(p, \vec{R}_{PC})$ is found by first calculating the source term

$$R_{000}^0(p, \vec{R}_{PC}) = (-2p)^n F_n(p R_{PC}^2) \quad (109)$$

and then iterating towards the target $R_{tuv}^0(p, \vec{R}_{PC}) = R_{tuv}(p, \vec{R}_{PC})$ via the recurrence relations

Integral evaluations, summarizing

The Coulomb integral

$$g_{abcd} = \langle G_a G_c | \frac{1}{r_{12}} | G_b G_d \rangle \quad (111)$$

is calculated as

$$g_{abcd} = \frac{2\pi^{5/2}}{pq\sqrt{p+q}} \sum_{tuv} \sum_{\tau\nu\phi} (-1)^{\tau+\nu+\phi} E_{tuv}^{ab} E_{\tau\nu\phi}^{cd} R_{t+\tau, u+\nu, v+\phi}(\alpha, \vec{R}_{PQ}), \quad (112)$$

where

$$\alpha = \frac{pq}{p+q}. \quad (113)$$

Hartree-Fock equations with spin degrees of freedom

Till now we have mainly dealt with the Hartree-Fock equations using anti-symmetrized matrix elements. Since the Hamiltonian does not contain operators acting on the spin states, it is common to write out the spin-degrees of freedom in an explicit way. We wrote the Hartree-Fock equations as

$$\hat{h}^{HF}(x_i) \psi_p(x_i) = \epsilon_p \psi_p(x_i),$$

with

$$\hat{h}^{HF}(x_i) = \hat{h}_0(x_i) + \sum_{j=1}^N V_j^d(x_i) - \sum_{j=1}^N V_j^{ex}(x_i),$$

and where $\hat{h}_0(i)$ is the one-body part. In this equations we include both spin and spatial degrees of freedom.

The direct part is defined as

$$V_p^d(x_i) = \int \psi_p^*(x_j) \psi_p(x_j) \hat{v}(r_{ij}) dx_j$$

Hartree-Fock equations with spin degrees of freedom

If now deal explicitly with the spin degrees of freedom, we can write the single-particle state

$$\psi_p(\mathbf{r}_i) = \phi_p(\mathbf{r}_i)\xi_{\sigma_i},$$

where ξ_{σ_i} is a standard Pauli spinor. Since we have only two possible spin values, the direct terms reduces then to (recall that we have defined $\int d\mathbf{r}_j = \sum_{\sigma_j} \int d\mathbf{r}_j$)

$$V_p^d(\mathbf{r}_i)\phi_q(\mathbf{r}_i) = 2 \int \phi_p^*(\mathbf{r}_j)\phi_p(\mathbf{r}_i)\hat{v}(\mathbf{r}_{ij})d\mathbf{r}_j\phi_q(\mathbf{r}_i)$$

while the exchange operator (or Fock operator) is

$$V_p^{ex}(\mathbf{r}_i)\phi_q(\mathbf{r}_i) = \left(\int \phi_p^*(\mathbf{r}_j)\hat{v}(\mathbf{r}_{ij})\phi_q(\mathbf{r}_j)d\mathbf{r}_j \right) \phi_p(\mathbf{r}_i).$$

Hartree-Fock equations with spin degrees of freedom

In our Hartree-Fock calculation we expand the single-particle functions in terms of known basis functions (hydrogen-like one, STOs, GTOs etc), namely

$$\phi_p(\mathbf{r}_i) = \sum_{k=1}^d C_{pk}\chi_k(\mathbf{r}_i),$$

d is the number of basis functions $\chi_k(\mathbf{r}_i)$. The Hartree-Fock equations become then

$$\hat{h}^{HF}\hat{C}_p = \epsilon^{HF}\hat{S}\hat{C}_p,$$

where \hat{S} is the overlap matrix needed in case the basis functions $\chi_k(\mathbf{r}_i)$ are not normalized (typically, GTOs are not). Our Hartree-Fock Hamiltonian leads then to matrix elements (in a bra-ket notation)

$$\langle p|\hat{h}^{HF}|q\rangle = \langle p|\hat{h}_0|q\rangle + \sum_{kr} \sum_{ks} C_{kr}^* C_{ks} (2\langle pr|\hat{v}|qs\rangle - \langle pr|\hat{v}|sq\rangle),$$

Hartree-Fock equations with spin degrees of freedom

The other integrals are

$$\langle pr|\hat{v}|qs\rangle = \int \int \chi_p^*(\mathbf{r}_i)\chi_q^*(\mathbf{r}_j)\hat{v}(\mathbf{r}_{ij})\chi_r(\mathbf{r}_i)\chi_s(\mathbf{r}_j)d\mathbf{r}_i d\mathbf{r}_j.$$

If we then introduce the density matrix defined as

$$D_{pq} = \sum_{k \leq F} C_{kp} C_{kq}^*,$$

we can rewrite the Hartree-Fock matrix elements as

$$\langle p|\hat{h}^{HF}|q\rangle = \langle p|\hat{h}_0|q\rangle + \sum_{rs} D_{rs} (2\langle pr|\hat{v}|qs\rangle - \langle pr|\hat{v}|sq\rangle),$$

meaning that the only quantity we need to calculate at every iteration is the density matrix, which is evaluated using the eigenvector obtained from the diagonalization of the Hartree-Fock matrix.

Hartree-Fock equations with spin degrees of freedom

From an algorithmic point of view, we see now that we need, with our GTO basis, to evaluate

- The overlap matrix \hat{S} .
- The kinetic energy and one-body interaction matrix elements $\langle p|\hat{h}_0|q\rangle$ and finally the
- two-body interaction matrix elements $\langle pr|\hat{v}|qs\rangle$.

All these elements can be computed once and for all and stored in memory. The overlap matrix S and the one-body matrix elements $\langle p|\hat{h}_0|q\rangle$ can be stored as simple one-dimensional arrays, or alternatively as matrices of small dimensions. The time-consuming part in the Hartree-Fock calculations involves the calculation of the two-body matrix. Furthermore, the storage of these matrix elements plays also an important role, in particular we wish to access the table of matrix elements as fast as possible.

Hartree-Fock equations with spin degrees of freedom

In a brute force algorithm for storing the matrix elements, if we have d basis functions, we end up with the need of storing d^4 matrix elements. We can reduce this considerably by the following considerations. In the calculation of the two-body matrix elements $\langle pr|\hat{v}|qs\rangle$ we have the following symmetries

- Invariance under permutations, that is
- $$\langle pr|\hat{v}|qs\rangle = \langle rp|\hat{v}|sq\rangle.$$
- The functions entering the evaluation of the integrals are all real, meaning that if we interchange $p \leftrightarrow q$ or $r \leftrightarrow s$, we end up with the same matrix element.

This reduces by a factor of eight the total number of matrix elements to be stored.

Hartree-Fock equations with spin degrees of freedom

Furthermore, in setting up a table for the two-body matrix elements we can convert the need of using four indices $pqrs$ of $\langle pr|\hat{v}|qs\rangle$, which in a brute forces way could be coded as a four-dimensional array, to a two-dimensional array V_{lm} , where l and m stand for all possible two-body configurations pq .

Each number l and m in V_{lm} should then point to a set of single-particle states (p, q) and (r, s) .

In our case, since we have symmetries which allow us to set $p \leq q$, we have, with d single particle states a total of $d(d+1)/2$ two-body configurations. How do we store such a matrix? The simplest thing to do is to convert it into a one-dimensional array. How do we achieve that?

Hartree-Fock equations with spin degrees of freedom

We now have a matrix V of dimension $n \times n$ and we want to store the elements V_{lm} as a one-dimensional array A using $0 \leq l \leq m \leq n-1$. For

- $l = 0$ we have n elements
- $l = 1$ we have $n-1$ elements
- ...
- $l = \nu$ we have $n-\nu$ elements
- ...
- $l = n-1$ we have 1 element,

and the total number is

$$\sum_{\nu=0}^{n-1} (n-\nu) = \frac{n(n+1)}{2}.$$

Hartree-Fock equations with spin degrees of freedom

To find the number $\text{number}(l, m)$ in a one-dimensional array A which corresponds to a matrix element V_{lm} , we note that

$$\text{number}(l, m) = \sum_{\nu=0}^{l-1} (n-\nu) + m - l = \frac{l(2n-l-1)}{2} + m.$$

The first matrix element $V(0, 0)$ is obviously given by the element $A(0)$.

We have thus reduced a four dimensional array to a one-dimensional array, where the given pairs (p, q) and (r, s) point to the matrix indices l and m , respectively. The latter are used to find the explicit number $\text{number}(l, m)$ which points to the desired matrix element stored in a one-dimensional array.

Exercises: Derivation of Hartree-Fock equations

Exercise 1

Consider a Slater determinant built up of single-particle orbitals ψ_λ , with $\lambda = 1, 2, \dots, N$.

The unitary transformation

$$\psi_a = \sum_{\lambda} C_{a\lambda} \psi_\lambda,$$

brings us into the new basis. The new basis has quantum numbers $a = 1, 2, \dots, N$. Show that the new basis is orthonormal. Show that the new Slater determinant constructed from the new single-particle wave functions can be written as the determinant based on the previous basis and the determinant of the matrix C . Show that the old and the new Slater determinants are equal up to a complex constant with absolute value unity. (Hint: C is a unitary matrix).

Exercises: Derivation of Hartree-Fock equations

Exercise 2

Consider the Slater determinant

$$\Phi_0 = \frac{1}{\sqrt{n!}} \sum_p (-)^p P \prod_{i=1}^n \psi_{\alpha_i}(x_i).$$

A small variation in this function is given by

$$\delta\Phi_0 = \frac{1}{\sqrt{n!}} \sum_p (-)^p P \psi_{\alpha_1}(x_1) \psi_{\alpha_2}(x_2) \dots \psi_{\alpha_{i-1}}(x_{i-1}) (\delta\psi_{\alpha_i}(x_i)) \psi_{\alpha_{i+1}}(x_{i+1}) \dots \psi_{\alpha_n}(x_n)$$

Show that

$$\langle \delta\Phi_0 | \sum_{i=1}^n \{t(x_i) + u(x_i)\} + \frac{1}{2} \sum_{i \neq j=1}^n v(x_i, x_j) | \Phi_0 \rangle = \sum_{i=1}^n \langle \delta\psi_{\alpha_i} | \hat{t} + \hat{u} | \psi_{\alpha_i} \rangle + \sum_{i \neq j=1}^n \langle \delta\psi_{\alpha_i} | v | \psi_{\alpha_j} \rangle$$