

# Computational Physics 2: Definitions of the many-body problem

Morten Hjorth-Jensen Email [morten.hjorth-jensen@fys.uio.no](mailto:morten.hjorth-jensen@fys.uio.no)<sup>1,2</sup>

<sup>1</sup>Department of Physics, University of Oslo

<sup>2</sup>Department of Physics and Astronomy and National Superconducting Cyclotron Laboratory, Michigan State University

Spring 2016

## Quantum Many-particle Methods

- Large-scale diagonalization (Iterative methods, Lanczo's method, dimensionalities  $10^{10}$  states), discussed in [FYS-KJM4480](#)
- Coupled cluster theory, favoured method in quantum chemistry, molecular and atomic physics. Applications to ab initio calculations in nuclear physics as well for large nuclei, discussed in [FYS-KJM4480](#)
- Perturbative many-body methods, discussed in [FYS-KJM4480](#)
- Density functional theories/Mean-field theory and Hartree-Fock theory, also discussed in [FYS-KJM4480](#)
- Monte-Carlo methods (Only in FYS4411, Computational quantum mechanics)
- Green's function theories (depending on interest)
- and other. The physics of the system hints at which many-body methods to use.

## Selected Texts and Many-body theory

- Blaizot and Ripka, *Quantum Theory of Finite systems*, MIT press 1986
- Negele and Orland, *Quantum Many-Particle Systems*, Addison-Wesley, 1987.

- Fetter and Walecka, *Quantum Theory of Many-Particle Systems*, McGraw-Hill, 1971.
- Helgaker, Jorgensen and Olsen, *Molecular Electronic Structure Theory*, Wiley, 2001.
- Mattuck, *Guide to Feynman Diagrams in the Many-Body Problem*, Dover, 1971.
- Dickhoff and Van Neck, *Many-Body Theory Exposed*, World Scientific, 2006.

## Definitions

An operator is defined as  $\hat{O}$  throughout. Unless otherwise specified the number of particles is always  $N$  and  $d$  is the dimension of the system. In nuclear physics we normally define the total number of particles to be  $A = N + Z$ , where  $N$  is total number of neutrons and  $Z$  the total number of protons. In case of other baryons such as isobars  $\Delta$  or various hyperons such as  $\Lambda$  or  $\Sigma$ , one needs to add their definitions. Hereafter,  $N$  is reserved for the total number of particles, unless otherwise specified.

## Definitions

The quantum numbers of a single-particle state in coordinate space are defined by the variable

$$x = (\mathbf{r}, \sigma),$$

where

$$\mathbf{r} \in \mathbb{R}^d,$$

with  $d = 1, 2, 3$  represents the spatial coordinates and  $\sigma$  is the eigenspin of the particle. For fermions with eigenspin 1/2 this means that

$$x \in \mathbb{R}^d \oplus \left(\frac{1}{2}\right),$$

and the integral  $\int dx = \sum_{\sigma} \int d^d r = \sum_{\sigma} \int d\mathbf{r}$ , and

$$\int d^N x = \int dx_1 \int dx_2 \dots \int dx_N.$$

## Definitions

The quantum mechanical wave function of a given state with quantum numbers  $\lambda$  (encompassing all quantum numbers needed to specify the system), ignoring time, is

$$\Psi_{\lambda} = \Psi_{\lambda}(x_1, x_2, \dots, x_N),$$

with  $x_i = (\mathbf{r}_i, \sigma_i)$  and the projection of  $\sigma_i$  takes the values  $\{-1/2, +1/2\}$  for particles with spin  $1/2$ . We will hereafter always refer to  $\Psi_\lambda$  as the exact wave function, and if the ground state is not degenerate we label it as

$$\Psi_0 = \Psi_0(x_1, x_2, \dots, x_N).$$

## Definitions

Since the solution  $\Psi_\lambda$  seldomly can be found in closed form, approximations are sought. Here we define an approximative wave function or an ansatz to the exact wave function as

$$\Phi_\lambda = \Phi_\lambda(x_1, x_2, \dots, x_N),$$

with

$$\Phi_0 = \Phi_0(x_1, x_2, \dots, x_N),$$

being the ansatz to the ground state.

## Definitions

The wave function  $\Psi_\lambda$  is sought in the Hilbert space of either symmetric or anti-symmetric  $N$ -body functions, namely

$$\Psi_\lambda \in \mathcal{H}_N := \mathcal{H}_1 \oplus \mathcal{H}_1 \oplus \dots \oplus \mathcal{H}_1,$$

where the single-particle Hilbert space  $\hat{H}_1$  is the space of square integrable functions over  $\mathbb{R}^d \oplus (\sigma)$  resulting in

$$\mathcal{H}_1 := L^2(\mathbb{R}^d \oplus (\sigma)).$$

## Definitions

Our Hamiltonian is invariant under the permutation (interchange) of two particles. Since we deal with fermions however, the total wave function is antisymmetric. Let  $\hat{P}$  be an operator which interchanges two particles. Due to the symmetries we have ascribed to our Hamiltonian, this operator commutes with the total Hamiltonian,

$$[\hat{H}, \hat{P}] = 0,$$

meaning that  $\Psi_\lambda(x_1, x_2, \dots, x_N)$  is an eigenfunction of  $\hat{P}$  as well, that is

$$\hat{P}_{ij} \Psi_\lambda(x_1, x_2, \dots, x_i, \dots, x_j, \dots, x_N) = \beta \Psi_\lambda(x_1, x_2, \dots, x_j, \dots, x_i, \dots, x_N),$$

where  $\beta$  is the eigenvalue of  $\hat{P}$ . We have introduced the suffix  $ij$  in order to indicate that we permute particles  $i$  and  $j$ . The Pauli principle tells us that the total wave function for a system of fermions has to be antisymmetric, resulting in the eigenvalue  $\beta = -1$ .

## Definitions and notations

The Schrodinger equation reads

$$\hat{H}(x_1, x_2, \dots, x_N) \Psi_\lambda(x_1, x_2, \dots, x_N) = E_\lambda \Psi_\lambda(x_1, x_2, \dots, x_N), \quad (1)$$

where the vector  $x_i$  represents the coordinates (spatial and spin) of particle  $i$ ,  $\lambda$  stands for all the quantum numbers needed to classify a given  $N$ -particle state and  $\Psi_\lambda$  is the pertaining eigenfunction. Throughout this course,  $\Psi$  refers to the exact eigenfunction, unless otherwise stated.

## Definitions and notations

We write the Hamilton operator, or Hamiltonian, in a generic way

$$\hat{H} = \hat{T} + \hat{V}$$

where  $\hat{T}$  represents the kinetic energy of the system

$$\hat{T} = \sum_{i=1}^N \frac{\mathbf{p}_i^2}{2m_i} = \sum_{i=1}^N \left( -\frac{\hbar^2}{2m_i} \nabla_i^2 \right) = \sum_{i=1}^N t(x_i)$$

while the operator  $\hat{V}$  for the potential energy is given by

$$\hat{V} = \sum_{i=1}^N \hat{u}_{\text{ext}}(x_i) + \sum_{j < i=1}^N v(x_i, x_j) + \sum_{i < j < k=1}^N v(x_i, x_j, x_k) + \dots \quad (2)$$

Hereafter we use natural units, viz.  $\hbar = c = e = 1$ , with  $e$  the elementary charge and  $c$  the speed of light. This means that momenta and masses have dimension energy.

## Definitions and notations

If one does quantum chemistry, after having introduced the Born-Oppenheimer approximation which effectively freezes out the nucleonic degrees of freedom, the Hamiltonian for  $N = n_e$  electrons takes the following form

$$\hat{H} = \sum_{i=1}^{n_e} t(x_i) - \sum_{i=1}^{n_e} k \frac{Z}{r_i} + \sum_{i < j}^{n_e} \frac{k}{r_{ij}},$$

with  $k = 1.44$  eVnm

## Definitions and notations

We can rewrite this as

$$\hat{H} = \hat{H}_0 + \hat{H}_I = \sum_{i=1}^{n_e} \hat{h}_0(x_i) + \sum_{i < j}^{n_e} \frac{1}{r_{ij}}, \quad (3)$$

where we have defined

$$r_{ij} = |\mathbf{r}_i - \mathbf{r}_j|,$$

and

$$\hat{h}_0(x_i) = \hat{t}(x_i) - \frac{Z}{x_i}. \quad (4)$$

The first term of Eq. (3),  $H_0$ , is the sum of the  $N$  *one-body* Hamiltonians  $\hat{h}_0$ . Each individual Hamiltonian  $\hat{h}_0$  contains the kinetic energy operator of an electron and its potential energy due to the attraction of the nucleus. The second term,  $H_I$ , is the sum of the  $n_e(n_e - 1)/2$  two-body interactions between each pair of electrons. Note that the double sum carries a restriction  $i < j$ .

## Definitions and notations

The potential energy term due to the attraction of the nucleus defines the onebody field  $u_i = u_{\text{ext}}(x_i)$  of Eq. (2). We have moved this term into the  $\hat{H}_0$  part of the Hamiltonian, instead of keeping it in  $\hat{V}$  as in Eq. (2). The reason is that we will hereafter treat  $\hat{H}_0$  as our non-interacting Hamiltonian. For a many-body wavefunction  $\Phi_\lambda$  defined by an appropriate single-particle basis, we may solve exactly the non-interacting eigenvalue problem

$$\hat{H}_0 \Phi_\lambda = w_\lambda \Phi_\lambda,$$

with  $w_\lambda$  being the non-interacting energy. This energy is defined by the sum over single-particle energies to be defined below. For atoms the single-particle energies could be the hydrogen-like single-particle energies corrected for the charge  $Z$ . For nuclei and quantum dots, these energies could be given by the harmonic oscillator in three and two dimensions, respectively.

## Definitions and notations

We will assume that the interacting part of the Hamiltonian can be approximated by a two-body interaction. This means that our Hamiltonian is written as

$$\hat{H} = \hat{H}_0 + \hat{H}_I = \sum_{i=1}^N \hat{h}_0(x_i) + \sum_{i<j}^N V(r_{ij}), \quad (5)$$

with

$$H_0 = \sum_{i=1}^N \hat{h}_0(x_i) = \sum_{i=1}^N (\hat{t}(x_i) + \hat{u}_{\text{ext}}(x_i)). \quad (6)$$

The onebody part  $u_{\text{ext}}(x_i)$  is normally approximated by a harmonic oscillator potential or the Coulomb interaction an electron feels from the nucleus. However, other potentials are fully possible, such as one derived from the self-consistent solution of the Hartree-Fock equations.

## Definitions and notations

Our Hamiltonian is invariant under the permutation (interchange) of two particles. Since we deal with fermions however, the total wave function is antisymmetric. Let  $\hat{P}$  be an operator which interchanges two particles. Due to the symmetries we have ascribed to our Hamiltonian, this operator commutes with the total Hamiltonian,

$$[\hat{H}, \hat{P}] = 0,$$

meaning that  $\Psi_\lambda(x_1, x_2, \dots, x_N)$  is an eigenfunction of  $\hat{P}$  as well, that is

$$\hat{P}_{ij}\Psi_\lambda(x_1, x_2, \dots, x_i, \dots, x_j, \dots, x_N) = \beta\Psi_\lambda(x_1, x_2, \dots, x_i, \dots, x_j, \dots, x_N),$$

where  $\beta$  is the eigenvalue of  $\hat{P}$ . We have introduced the suffix  $ij$  in order to indicate that we permute particles  $i$  and  $j$ . The Pauli principle tells us that the total wave function for a system of fermions has to be antisymmetric, resulting in the eigenvalue  $\beta = -1$ .

## Definitions and notations

In our case we assume that we can approximate the exact eigenfunction with a Slater determinant

$$\Phi(x_1, x_2, \dots, x_N, \alpha, \beta, \dots, \sigma) = \frac{1}{\sqrt{N!}} \begin{vmatrix} \psi_\alpha(x_1) & \psi_\alpha(x_2) & \dots & \dots & \psi_\alpha(x_N) \\ \psi_\beta(x_1) & \psi_\beta(x_2) & \dots & \dots & \psi_\beta(x_N) \\ \dots & \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots & \dots \\ \psi_\sigma(x_1) & \psi_\sigma(x_2) & \dots & \dots & \psi_\sigma(x_N) \end{vmatrix}, \quad (7)$$

where  $x_i$  stand for the coordinates and spin values of a particle  $i$  and  $\alpha, \beta, \dots, \gamma$  are quantum numbers needed to describe remaining quantum numbers.

## Definitions and notations

The single-particle function  $\psi_\alpha(x_i)$  are eigenfunctions of the onebody Hamiltonian  $h_i$ , that is

$$\hat{h}_0(x_i) = \hat{t}(x_i) + \hat{u}_{\text{ext}}(x_i),$$

with eigenvalues

$$\hat{h}_0(x_i)\psi_\alpha(x_i) = (\hat{t}(x_i) + \hat{u}_{\text{ext}}(x_i))\psi_\alpha(x_i) = \varepsilon_\alpha\psi_\alpha(x_i).$$

The energies  $\varepsilon_\alpha$  are the so-called non-interacting single-particle energies, or unperturbed energies. The total energy is in this case the sum over all single-particle energies, if no two-body or more complicated many-body interactions are present.

## Definitions and notations

Let us denote the ground state energy by  $E_0$ . According to the variational principle we have

$$E_0 \leq E[\Phi] = \int \Phi^* \hat{H} \Phi d\tau$$

where  $\Phi$  is a trial function which we assume to be normalized

$$\int \Phi^* \Phi d\tau = 1,$$

where we have used the shorthand  $d\tau = d\mathbf{r}_1 d\mathbf{r}_2 \dots d\mathbf{r}_N$ .

## Brief reminder on some linear algebra properties

Before we proceed with a more compact representation of a Slater determinant, we would like to repeat some linear algebra properties which will be useful for our derivations of the energy as function of a Slater determinant, Hartree-Fock theory and later the nuclear shell model.

The inverse of a matrix is defined by

$$\mathbf{A}^{-1} \cdot \mathbf{A} = I$$

A unitary matrix  $\mathbf{A}$  is one whose inverse is its adjoint

$$\mathbf{A}^{-1} = \mathbf{A}^\dagger$$

A real unitary matrix is called orthogonal and its inverse is equal to its transpose. A hermitian matrix is its own self-adjoint, that is

$$\mathbf{A} = \mathbf{A}^\dagger.$$

## Basic Matrix Features

### Matrix Properties Reminder.

Relations	Name	matrix elements
$A = A^T$	symmetric	$a_{ij} = a_{ji}$
$A = (A^T)^{-1}$	real orthogonal	$\sum_k a_{ik} a_{jk} = \sum_k a_{ki} a_{kj} = \delta_{ij}$
$A = A^*$	real matrix	$a_{ij} = a_{ij}^*$
$A = A^\dagger$	hermitian	$a_{ij} = a_{ji}^*$
$A = (A^\dagger)^{-1}$	unitary	$\sum_k a_{ik} a_{jk}^* = \sum_k a_{ki}^* a_{kj} = \delta_{ij}$

## Basic Matrix Features

Since we will deal with Fermions (identical and indistinguishable particles) we will form an ansatz for a given state in terms of so-called Slater determinants determined by a chosen basis of single-particle functions.

For a given  $n \times n$  matrix  $\mathbf{A}$  we can write its determinant

$$\det(\mathbf{A}) = |\mathbf{A}| = \begin{vmatrix} a_{11} & a_{12} & \dots & \dots & a_{1n} \\ a_{21} & a_{22} & \dots & \dots & a_{2n} \\ \dots & \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots & \dots \\ a_{n1} & a_{n2} & \dots & \dots & a_{nn} \end{vmatrix},$$

in a more compact form as

$$|\mathbf{A}| = \sum_{i=1}^{n!} (-1)^{p_i} \hat{P}_i a_{11} a_{22} \dots a_{nn},$$

where  $\hat{P}_i$  is a permutation operator which permutes the column indices  $1, 2, 3, \dots, n$  and the sum runs over all  $n!$  permutations. The quantity  $p_i$  represents the number of transpositions of column indices that are needed in order to bring a given permutation back to its initial ordering, in our case given by  $a_{11} a_{22} \dots a_{nn}$  here.

## Basic Matrix Features, simple $2 \times 2$ determinant

A simple  $2 \times 2$  determinant illustrates this. We have

$$\det(\mathbf{A}) = \begin{vmatrix} a_{11} & a_{12} \\ a_{21} & a_{22} \end{vmatrix} = (-1)^0 a_{11} a_{22} + (-1)^1 a_{12} a_{21},$$

where in the last term we have interchanged the column indices 1 and 2. The natural ordering we have chosen is  $a_{11} a_{22}$ .

## Definitions and notations

With the above we can rewrite our Slater determinant in a more compact form. In the Hartree-Fock method the trial function is the Slater determinant of Eq. (7) which can be rewritten as

$$\Phi(x_1, x_2, \dots, x_N, \alpha, \beta, \dots, \nu) = \frac{1}{\sqrt{N!}} \sum_P (-1)^P \hat{P} \psi_\alpha(x_1) \psi_\beta(x_2) \dots \psi_\nu(x_N) = \sqrt{N!} \hat{A} \Phi_H,$$

where we have introduced the antisymmetrization operator  $\hat{A}$  defined by the summation over all possible permutations of two particles.



## Definitions and notations

It is defined as

$$\hat{A} = \frac{1}{N!} \sum_p (-)^p \hat{P}, \quad (8)$$

with  $p$  standing for the number of permutations. We have introduced for later use the so-called Hartree-function, defined by the simple product of all possible single-particle functions

$$\Phi_H(x_1, x_2, \dots, x_N, \alpha, \beta, \dots, \nu) = \psi_\alpha(x_1) \psi_\beta(x_2) \dots \psi_\nu(x_N).$$

## Definitions and notations

Both  $\hat{H}_0$  and  $\hat{H}_I$  are invariant under all possible permutations of any two particles and hence commute with  $\hat{A}$

$$[H_0, \hat{A}] = [H_I, \hat{A}] = 0. \quad (9)$$

Furthermore,  $\hat{A}$  satisfies

$$\hat{A}^2 = \hat{A}, \quad (10)$$

since every permutation of the Slater determinant reproduces it.

## Definitions and notations

The expectation value of  $\hat{H}_0$

$$\int \Phi^* \hat{H}_0 \Phi d\tau = N! \int \Phi_H^* \hat{A} \hat{H}_0 \hat{A} \Phi_H d\tau$$

is readily reduced to

$$\int \Phi^* \hat{H}_0 \Phi d\tau = N! \int \Phi_H^* \hat{H}_0 \hat{A} \Phi_H d\tau,$$

where we have used Eqs. (9) and (10). The next step is to replace the anti-symmetrization operator by its definition and to replace  $\hat{H}_0$  with the sum of one-body operators

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{i=1}^N \sum_p (-)^p \int \Phi_H^* \hat{h}_0 \hat{P} \Phi_H d\tau.$$

## Definitions and notations

The integral vanishes if two or more particles are permuted in only one of the Hartree-functions  $\Phi_H$  because the individual single-particle wave functions are orthogonal. We obtain then

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{i=1}^N \int \Phi_H^* \hat{h}_0 \Phi_H d\tau.$$

Orthogonality of the single-particle functions allows us to further simplify the integral, and we arrive at the following expression for the expectation values of the sum of one-body Hamiltonians

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{\mu=1}^N \int \psi_{\mu}^*(\mathbf{r}) \hat{h}_0 \psi_{\mu}(\mathbf{r}) d\mathbf{r}. \quad (11)$$

### Definitions and notations

We introduce the following shorthand for the above integral

$$\langle \mu | \hat{h}_0 | \mu \rangle = \int \psi_{\mu}^*(\mathbf{r}) \hat{h}_0 \psi_{\mu}(\mathbf{r}) d\mathbf{r},$$

and rewrite Eq. (11) as

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{\mu=1}^N \langle \mu | \hat{h}_0 | \mu \rangle. \quad (12)$$

### Definitions and notations

The expectation value of the two-body part of the Hamiltonian is obtained in a similar manner. We have

$$\int \Phi^* \hat{H}_I \Phi d\tau = N! \int \Phi_H^* \hat{A} \hat{H}_I \hat{A} \Phi_H d\tau,$$

which reduces to

$$\int \Phi^* \hat{H}_I \Phi d\tau = \sum_{i \leq j=1}^N \sum_p (-)^p \int \Phi_H^* V(r_{ij}) \hat{P} \Phi_H d\tau,$$

by following the same arguments as for the one-body Hamiltonian.

### Definitions and notations

Because of the dependence on the inter-particle distance  $r_{ij}$ , permutations of any two particles no longer vanish, and we get

$$\int \Phi^* \hat{H}_I \Phi d\tau = \sum_{i < j=1}^N \int \Phi_H^* V(r_{ij}) (1 - P_{ij}) \Phi_H d\tau.$$

where  $P_{ij}$  is the permutation operator that interchanges particle  $i$  and particle  $j$ . Again we use the assumption that the single-particle wave functions are orthogonal.

## Definitions and notations

We obtain

$$\int \Phi^* \hat{H}_I \Phi d\tau = \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \left[ \int \psi_{\mu}^*(x_i) \psi_{\nu}^*(x_j) V(r_{ij}) \psi_{\mu}(x_i) \psi_{\nu}(x_j) dx_i dx_j \right. \quad (13)$$

$$\left. - \int \psi_{\mu}^*(x_i) \psi_{\nu}^*(x_j) V(r_{ij}) \psi_{\nu}(x_i) \psi_{\mu}(x_j) dx_i dx_j \right]. \quad (14)$$

The first term is the so-called direct term. It is frequently also called the Hartree term, while the second is due to the Pauli principle and is called the exchange term or just the Fock term. The factor 1/2 is introduced because we now run over all pairs twice.

## Definitions and notations

The last equation allows us to introduce some further definitions. The single-particle wave functions  $\psi_{\mu}(x)$ , defined by the quantum numbers  $\mu$  and  $x$  are defined as the overlap

$$\psi_{\alpha}(x) = \langle x | \alpha \rangle.$$

## Definitions and notations

We introduce the following shorthands for the above two integrals

$$\langle \mu\nu | \hat{v} | \mu\nu \rangle = \int \psi_{\mu}^*(x_i) \psi_{\nu}^*(x_j) V(r_{ij}) \psi_{\mu}(x_i) \psi_{\nu}(x_j) dx_i dx_j,$$

and

$$\langle \mu\nu | \hat{v} | \nu\mu \rangle = \int \psi_{\mu}^*(x_i) \psi_{\nu}^*(x_j) V(r_{ij}) \psi_{\nu}(x_i) \psi_{\mu}(x_j) dx_i dx_j.$$

## Definitions and notations

The direct and exchange matrix elements can be brought together if we define the antisymmetrized matrix element

$$\langle \mu\nu | \hat{v} | \mu\nu \rangle_{\text{AS}} = \langle \mu\nu | \hat{v} | \mu\nu \rangle - \langle \mu\nu | \hat{v} | \nu\mu \rangle,$$

or for a general matrix element

$$\langle \mu\nu | \hat{v} | \sigma\tau \rangle_{\text{AS}} = \langle \mu\nu | \hat{v} | \sigma\tau \rangle - \langle \mu\nu | \hat{v} | \tau\sigma \rangle.$$

It has the symmetry property

$$\langle \mu\nu | \hat{v} | \sigma\tau \rangle_{\text{AS}} = -\langle \mu\nu | \hat{v} | \tau\sigma \rangle_{\text{AS}} = -\langle \nu\mu | \hat{v} | \sigma\tau \rangle_{\text{AS}}.$$

## Definitions and notations

The antisymmetric matrix element is also hermitian, implying

$$\langle \mu\nu | \hat{v} | \sigma\tau \rangle_{\text{AS}} = \langle \sigma\tau | \hat{v} | \mu\nu \rangle_{\text{AS}}.$$

With these notations we rewrite Eq. (14) as

$$\int \Phi^* \hat{H}_I \Phi d\tau = \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \langle \mu\nu | \hat{v} | \mu\nu \rangle_{\text{AS}}. \quad (15)$$

## Definitions and notations

Combining Eqs. (12) and (15) we obtain the energy functional

$$E[\Phi] = \sum_{\mu=1}^N \langle \mu | \hat{h}_0 | \mu \rangle + \frac{1}{2} \sum_{\mu=1}^N \sum_{\nu=1}^N \langle \mu\nu | \hat{v} | \mu\nu \rangle_{\text{AS}}. \quad (16)$$

which we will use as our starting point for the Hartree-Fock calculations later in this course.

## Our system, quantum dots

We consider a system of electrons confined in a pure isotropic harmonic oscillator potential  $V(\mathbf{r}) = m^* \omega_0^2 r^2 / 2$ , where  $m^*$  is the effective mass of the electrons in the host semiconductor,  $\omega_0$  is the oscillator frequency of the confining potential, and  $\mathbf{r} = (x, y, z)$  denotes the position of the particle.

The Hamiltonian of a single particle trapped in this harmonic oscillator potential simply reads

$$\hat{H} = \frac{\mathbf{p}^2}{2m^*} + \frac{1}{2} m^* \omega_0^2 \mathbf{r}^2$$

where  $\mathbf{p}$  is the canonical momentum of the particle.

## Our system, quantum dots

When considering several particles trapped in the same quantum dot, the Coulomb repulsion between those electrons has to be added to the single particle Hamiltonian which gives

$$\hat{H} = \sum_{i=1}^{N_e} \left( \frac{\mathbf{p}_i^2}{2m^*} + \frac{1}{2} m^* \omega_0^2 \mathbf{r}_i^2 \right) + \frac{e^2}{4\pi\epsilon_0\epsilon_r} \sum_{i < j} \frac{1}{|\mathbf{r}_i - \mathbf{r}_j|},$$

where  $N_e$  is the number of electrons,  $-e$  ( $e > 0$ ) is the charge of the electron,  $\epsilon_0$  and  $\epsilon_r$  are respectively the free space permittivity and the relative permittivity of the host material (also called dielectric constant), and the index  $i$  labels the electrons.

## Our system, quantum dots

We assume that the magnetic field  $\mathbf{B}$  is static and along the  $z$  axis. At first we ignore the spin-dependent terms. The Hamiltonian of these electrons in a magnetic field now reads

$$\hat{H} = \sum_{i=1}^{N_e} \left( \frac{(\mathbf{p}_i + e\mathbf{A})^2}{2m^*} + \frac{1}{2}m^*\omega_0^2\mathbf{r}_i^2 \right) + \frac{e^2}{4\pi\epsilon_0\epsilon_r} \sum_{i<j} \frac{1}{|\mathbf{r}_i - \mathbf{r}_j|}, \quad (17)$$

$$= \sum_{i=1}^{N_e} \left( \frac{\mathbf{p}_i^2}{2m^*} + \frac{e}{2m^*}(\mathbf{A} \cdot \mathbf{p}_i + \mathbf{p}_i \cdot \mathbf{A}) + \frac{e^2}{2m^*}\mathbf{A}^2 + \frac{1}{2}m^*\omega_0^2\mathbf{r}_i^2 \right) \quad (18)$$

$$+ \frac{e^2}{4\pi\epsilon_0\epsilon_r} \sum_{i<j} \frac{1}{|\mathbf{r}_i - \mathbf{r}_j|}, \quad (19)$$

where  $\mathbf{A}$  is the vector potential defined by  $\mathbf{B} = \nabla \times \mathbf{A}$ .

## Our system, quantum dots

In coordinate space,  $\mathbf{p}_i$  is the operator  $-i\hbar\nabla_i$  and by applying the Hamiltonian on the total wave function  $\Psi(\mathbf{r})$  in the Schroedinger equation, we obtain the following operator acting on  $\Psi(\mathbf{r})$

$$\mathbf{A} \cdot \mathbf{p}_i + \mathbf{p}_i \cdot \mathbf{A} = -i\hbar(\mathbf{A} \cdot \nabla_i + \nabla_i \cdot \mathbf{A})\Psi \quad (20)$$

$$= -i\hbar(\mathbf{A} \cdot (\nabla_i\Psi) + \nabla_i \cdot (\mathbf{A}\Psi)) \quad (21)$$

## Our system, quantum dots

We note that if we use the product rule and the Coulomb gauge  $\nabla \cdot \mathbf{A} = 0$  (by choosing the vector potential as  $\mathbf{A} = \frac{1}{2}\mathbf{B} \times \mathbf{r}$ ),  $\mathbf{p}_i$  and  $\nabla_i$  commute and we obtain

$$\nabla_i \cdot (\mathbf{A}\Psi) = \mathbf{A} \cdot (\nabla_i\Psi) + \underbrace{(\nabla_i \cdot \mathbf{A})\Psi}_0 = \mathbf{A} \cdot (\nabla_i\Psi)$$

## Our system, quantum dots

This leads us to the following Hamiltonian:

$$\hat{H} = \sum_{i=1}^{N_e} \left( -\frac{\hbar^2}{2m^*}\nabla_i^2 - i\hbar\frac{e}{m^*}\mathbf{A} \cdot \nabla_i + \frac{e^2}{2m^*}\mathbf{A}^2 + \frac{1}{2}m^*\omega_0^2\mathbf{r}_i^2 \right) + \frac{e^2}{4\pi\epsilon_0\epsilon_r} \sum_{i<j} \frac{1}{|\mathbf{r}_i - \mathbf{r}_j|},$$

## Our system, quantum dots

The linear term in  $\mathbf{A}$  becomes, in terms of  $\mathbf{B}$ :

$$\frac{-i\hbar e}{m^*} \mathbf{A} \cdot \nabla_i = -\frac{i\hbar e}{2m^*} (\mathbf{B} \times \mathbf{r}_i) \cdot \nabla_i \quad (22)$$

$$= \frac{-i\hbar e}{2m^*} \mathbf{B} \cdot (\mathbf{r}_i \times \nabla_i) \quad (23)$$

$$= \frac{e}{2m^*} \mathbf{B} \cdot \mathbf{L} \quad (24)$$

where  $\mathbf{L} = -i\hbar(\mathbf{r}_i \times \nabla_i)$  is the orbital angular momentum operator of the electron  $i$ .

If we assume that the electrons are confined in the  $xy$ -plane, the quadratic term in  $\mathbf{A}$  can be written as

$$\frac{e^2}{2m^*} \mathbf{A}^2 = \frac{e^2}{8m^*} (\mathbf{B} \times \mathbf{r})^2 = \frac{e^2}{8m^*} B^2 r_i^2$$

## Our system, quantum dots

Until this point we have neglected the intrinsic magnetic moment of the electrons which is due to the electron spin in the host material. We will now add its effect to the Hamiltonian. This intrinsic magnetic moment is given by

$$M_s = -g_s(e\mathbf{S})/(2m),$$

where  $g_s$  is the effective spin gyromagnetic ratio.

We see that the spin magnetic moment  $M_s$  gives rise to an additional interaction energy linear in the magnetic field,

$$\hat{H}_s = -M_s \cdot \mathbf{B} = g_s \frac{e}{2m} B \hat{S}_z = g_s \frac{\omega_c}{2} \hat{S}_z$$

## Our system, quantum dots

In order to simplify the computation, the Hamiltonian can be rewritten on dimensionless form. For this purpose, we introduce the following constants:

- The oscillator frequency  $\omega = \omega_0 \sqrt{1 + \omega_c^2/(4\omega_0^2)}$ ,
- A new energy unit  $\hbar\omega$ ,
- A new length unit, the oscillator length defined by  $l = \sqrt{\hbar/(m^*\omega)}$ , also called the characteristic length unit.

We rewrite the Hamiltonian in dimensionless units using

$$\mathbf{r} \longrightarrow \frac{\mathbf{r}}{l}, \quad \nabla \longrightarrow l \nabla \quad \text{and} \quad \hat{L}_z \longrightarrow \hat{L}_z$$

## Our system, quantum dots

A new dimensionless parameter  $\lambda = l/a_0^*$  (where  $a_0^* = 4\pi\epsilon_0\epsilon_r\hbar^2/(e^2m^*)$  is the effective Bohr radius) describes the strength of the electron-electron interaction. Large  $\lambda$  implies strong interaction and/or large quantum dot.

Since both  $\hat{L}_z$  and  $\hat{S}_z$  commute with the Hamiltonian we can perform the calculations separately in subspaces of given quantum numbers  $L_z$  and  $S_z$ .

The simplified dimensionless Hamiltonian becomes

$$\hat{H} = \sum_{i=1}^{N_e} \left[ -\frac{1}{2} \nabla_i^2 + \frac{1}{2} r_i^2 \right] + \lambda \sum_{i<j} \frac{1}{r_{ij}} + \sum_{i=1}^{N_e} \left( \frac{1}{2} \frac{\omega_c}{\hbar\omega} L_z^{(i)} + \frac{1}{2} g_s^* \frac{\omega_c}{\hbar\omega} S_z^{(i)} \right),$$

## Our system, quantum dots

The last sum which is proportional to the magnetic field involves only the quantum numbers  $L_z$  and  $S_z$  and not the operators themselves. Therefore these terms can be put aside during the resolution, the squizzing effect of the magnetic field being included simply in the parameter  $\lambda$ . The contribution of these terms will be added when the other part has been solved. This brings us to the simple and general form of the Hamiltonian

$$\hat{H} = \sum_{i=1}^{N_e} \left( -\frac{1}{2} \nabla_i^2 + \frac{1}{2} r_i^2 \right) + \lambda \sum_{i<j} \frac{1}{r_{ij}}.$$

The form

$$\hat{H} = \sum_{i=1}^{N_e} \left( -\frac{1}{2} \nabla_i^2 + \frac{1}{2} r_i^2 \right) + \lambda \sum_{i<j} \frac{1}{r_{ij}},$$

is however not so practical since the interaction carries a strength  $\lambda$ . Why?

## Our system, quantum dots

We rewrite it as a one-body part

$$\hat{H}_0 = \sum_{i=1}^{N_e} \left( -\frac{1}{2} \nabla_i^2 + \frac{\omega^2}{2} r_i^2 \right),$$

and interacting part

$$\hat{V} = \sum_{i<j}^{N_e} \frac{1}{|\mathbf{r}_i - \mathbf{r}_j|}.$$

The unperturbed part of the Hamiltonian yields the single-particle energies

$$\epsilon_i = \omega (2n + |m| + 1),$$

where  $n = 0, 1, 2, 3, \dots$  and  $m = 0, \pm 1, \pm 2, \dots$ . The index  $i$  runs from  $0, 1, 2, \dots$ .

## Small tasks

Set up the harmonic oscillator wave function in cartesian coordinates for an electron with  $n_x = n_y = 0$  and find the oscillator energy. Use this result to find the unperturbed energy

$$\int \Phi^* \hat{H}_0 \Phi d\tau = \sum_{\mu=1}^N \langle \mu | h | \mu \rangle.$$

for two electrons with the same quantum numbers.

- Repeat for six electrons (find the relevant harmonic oscillator quantum numbers)
- Repeat for 12 and 20 electrons