

# Topic 23: Special Relativity

## Advanced Placement Physics

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Olympiads School

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# Introduction

The slides on **special relativity** is an expanded version of the slides used for Physics 12 (and with some calculus). For many of you, this is a review. There are 2 versions of the slides that are downloadable from the school website:

- The long version
  - More background information and derivations and integrations
  - 23a-relativity\_long.pdf
- The short version
  - More “to the point”
  - The version that I am using in this class
  - 23a-relativity\_short.pdf

There is also a handout on how to solve and interpret the time dilation example problem.

# Frame of Reference

A **frame of reference** is a hypothetical mobile “laboratory” an observer uses to make measurements (e.g. mass, lengths, time). At a minimum, it must include:

- A ruler to measure lengths
- A clock to measure the passage of time
- A scale to compare forces
- A balance to measure masses

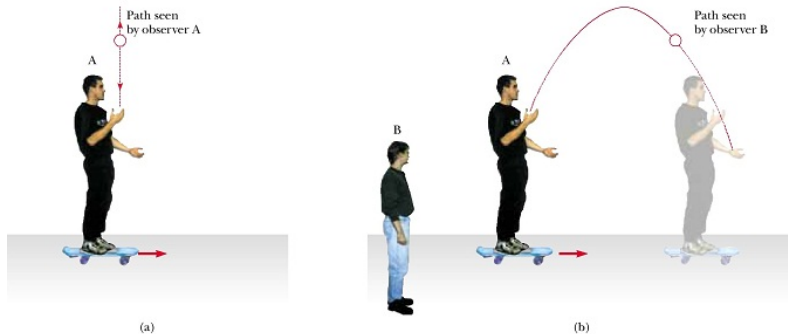
High-school textbooks often refer to the frame of reference as a “coordinate system”. While it certainly includes that, this definition often makes it difficult to understand special relativity.

# Frame of Reference

- We assume that the hypothetical laboratory is *perfect*
- What “instruments” are used are unimportant
- What matters is the *motion* (at rest, uniform motion, acceleration etc) of your laboratory, and how it affects the measurement that you make
- “From the point of view of. . .”

# Frame of Reference

Although both observers see different motion, but they agree on the equations that govern the motion (laws of motion).



# Newtonian (Classical) Relativity

In Newtonian physics, space and time are *absolute*:

- 1 m is 1 m no matter where you are in the universe
- 1 s is 1 s no matter where you are in the universe
- Measurements of space and time do not depend on motion

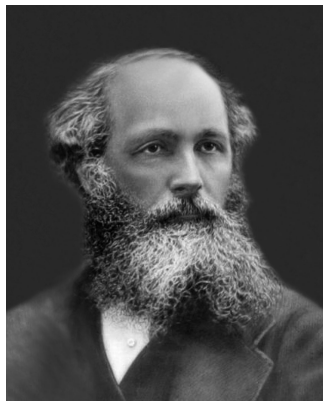
If space and time are absolute, then *all* velocities are relative

- Measured velocities depend on the motion of the observer
- An **inertial** frame of reference moving in uniform motion (constant velocity, without acceleration)

## The Principle of Relativity

All laws of physics must apply equally in all inertial frames of reference.

# New Physics: Maxwell's Equations



James Clerk Maxwell

- Classical laws of electrodynamics
- Published in 1861 and 1862
- Explains the relationship between
  - Electricity
  - Electric Circuits
  - Magnetism
  - Optics
- Previously these disciplines are thought to be separate and not related

# Maxwell's Equations in a Vacuum

Everything Comes Back to This

$$\nabla \cdot \mathbf{E} = 0$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

$$\nabla \times \mathbf{B} = \mu_0 \epsilon_0 \frac{\partial \mathbf{E}}{\partial t}$$

Disturbances in  $\mathbf{E}$  and  $\mathbf{B}$  travel as an “electromagnetic wave”, with a speed:

$$c = \frac{1}{\sqrt{\epsilon_0 \mu_0}} = 299\,792\,458 \text{ m/s}$$



# Peculiar features of Maxwell's equation

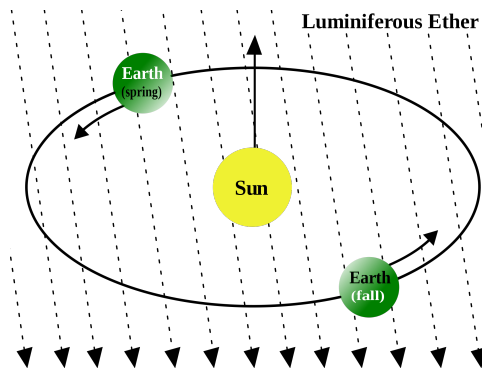
- Makes no mention of the *medium* in which EM waves travels
- When applying *Galilean transformation* (classical equation for calculating *relative velocity*) to Maxwell's equations, asymmetry is introduced
- Gauss's law for magnetism break down: magnetic field lines appear to have beginnings/ends
- In *some* inertial frames of reference, Maxwell's equations are simple and elegant, but in another inertial frame of reference, they are ugly and complex
- Physicists at the time theorized that—perhaps—there is/are actually *preferred* inertial frame(s) of references
- This violate the *principle of relativity*

# The Illusive Aether

- Maxwell's hypothesis: the speed of light  $c_0$  is relative to a hypothetical "luminiferous aether"
- In order for this "aether" (or "ether") to exist, it must have some fantastic (as in, a fantasy, too good to be true!) properties:
  - *All* space is filled with aether
  - Massless
  - Zero viscosity
  - Non-dispersive
  - Incompressible
  - Continuous at a very small (sub-atomic) scale

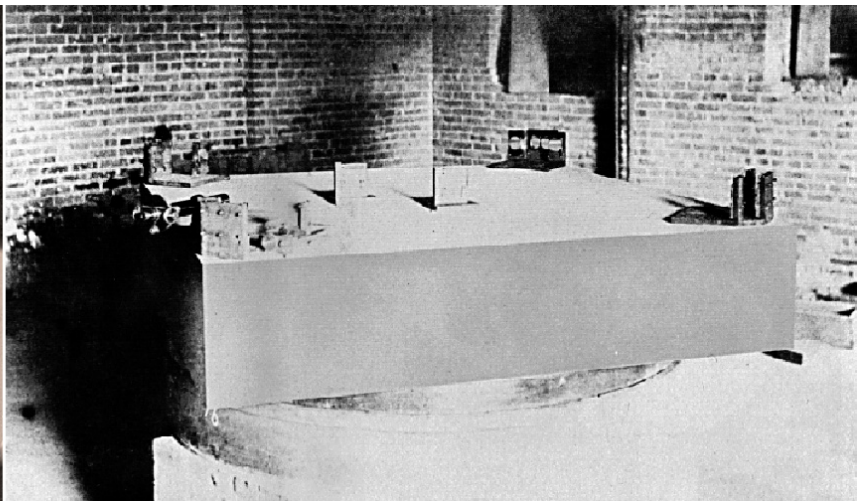
# The Michelson-Morley Experiment

If ether exists, then at different times of the year, the Earth will have a different relative velocity with respect to it:



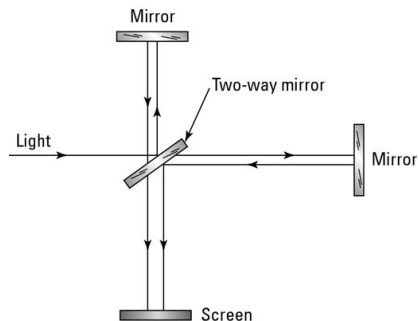
And it will cause light to either speed up, or slow down.

# The Michelson Interferometer



The experiment is ingenious but very difficult. . .

# The Michelson Interferometer



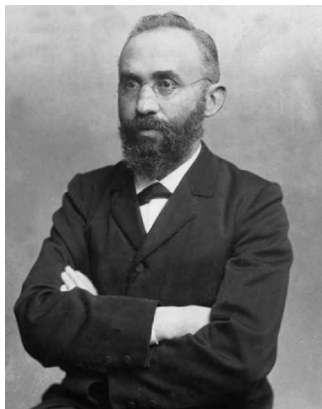
- A beam of light is split into two using a two-way (half-silvered) mirror
- The two beams are reflected off mirrors and finally arriving at the screen where interference patterns are observed
- The two paths are the same length, so if the *speed* of the light changes, we should see an interference pattern
- **Except none were ever found!**

# What To Do with “Null Result”

The Michelson-Morley experiment failed to detect the illusive ether, even after many refinements. What does this mean?

- Majority view
  - **The experiment was flawed!**
  - Keep improving the experiment (or design a better experiment) and the ether will eventually be found
- Minority view:
  - **The hypothesis is wrong!**
  - The experiment showed it for what it is: ether cannot be found
- A few physicists: There must be **another explanation** that saves both experiment and theory

# Hendrik Lorentz



Hendrik Antoon Lorentz

- Considered the Michelson-Morley experiment to be significant
- Objects travelling in the direction of ether contracts in length, nullifying the experimental results
- Lorentz Factor:

$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

- *No known physical phenomenon* can cause anything to contract
- Lorentz was on to something, but his thinking was wrong

# Making Maxwell's Equations Work

Albert Einstein in 1905, Age 26



Albert Einstein

- Einstein believed in the principle of relativity, and therefore rejected the concept of a preferred frame of reference
- The failure of the Michelson-Morley experiment to find the flow of ether proves that it does not exist
- In order to make the equations to work again, Einstein revisited two most fundamental concepts in physics: space and time



# Special Relativity

- Largely ignored by most physicists at first, until Max Planck took an interest in it
- Soon adopted by many physicists
- “Special” relativity because it describes a “special case” without effects of forces (e.g. gravity) & acceleration
- Later published theory of “general relativity” (much more complicated)

# Postulates

## The Principle of Relativity

All laws of physics must apply equally in all inertial frames of reference.

## The Principle of Invariant Light Speed

As measured in any inertial frame of reference, light is always propagated in empty space with a definite velocity  $c$  that is independent of the state of motion of the emitting body.

Published in the journal *Annalen der Physik* on September 26, 1905 in the article *On the Electrodynamics of Moving Bodies* when Einstein was 26 years old working as a patent clerk in Switzerland

# What's so Special About Special Relativity?

## Classical (Newtonian) relativity:

- Space and time are absolute, therefore
- The speed of light must be relative to the observer

## Einstein's special relativity:

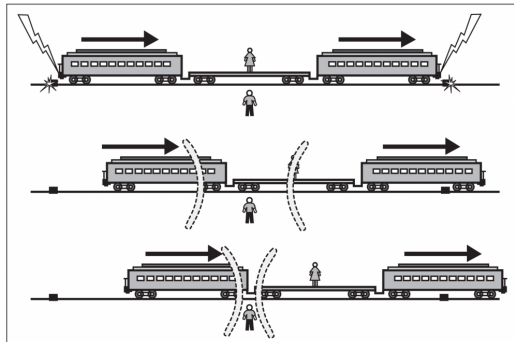
- The speed of light is absolute, therefore
- Space and time must be relative to the observer

We must modify our traditional concepts:

- Measurement of space (our ruler in the  $x$ -,  $y$ - and  $z$ -directions)
- Measurement of time (our clock)
- Concept of simultaneity (whether two events happens at the same time)

# Simultaneity: Thought Experiment

Lightning bolt strikes the ends of a moving train



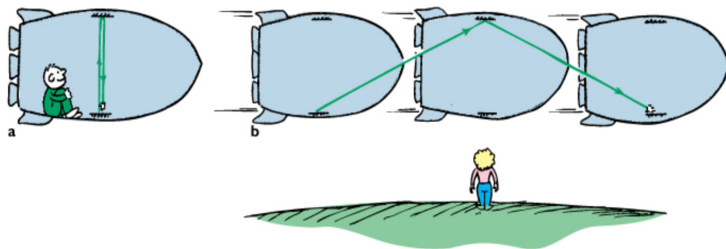
- The man on the ground sees the lightning bolt striking at the same time
- The woman on the moving train sees the lightning bolt on the right first

# Simultaneity: Thought Experiment

- The two observers disagree on the result, but
  - Neither person is wrong
  - Neither person is misinformed
- Both observers are valid *inertial* frames of reference
- This means that simultaneity depends on your motion

**Events that are simultaneous in one inertial frame of reference are not simultaneous in another.**

# Time Dilation: A Thought Experiment

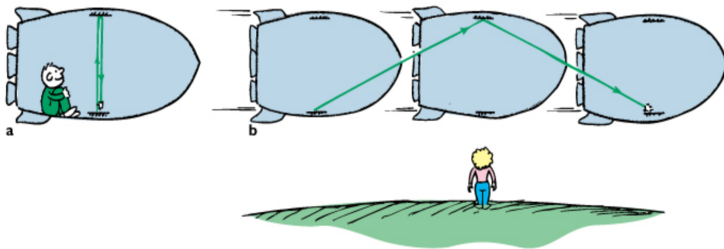


I'm on a spaceship travelling in deep space, and I shine a light from  $A$  to  $B$ . The distance between  $A$  and  $B$  is really just:

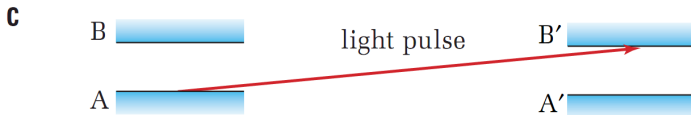
$$|AB| = ct$$

I know the speed of light  $c$ , and I know how long it took for the light pulse to reach  $B$ .

# Time Dilation: A Thought Experiment

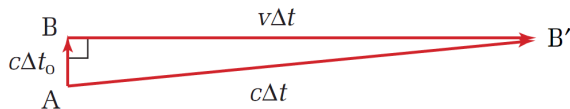


You are in space station watching my spaceship go past you at speed  $v$ . You would see that same beam of light travel from  $A$  to  $B'$  instead.



# Time Dilation: A Thought Experiment

D



$$c^2\Delta t^2 = v^2\Delta t^2 + c^2\Delta t_0^2$$

$$(c^2 - v^2) \Delta t^2 = c^2\Delta t_0^2$$

$$\left(1 - \frac{v^2}{c^2}\right) \Delta t^2 = \Delta t_0^2$$

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$



# Time Dilation: A Thought Experiment

$$t' = \frac{t}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

- $t$  is called the **proper time**. It is the time measured by a person at rest relative to the object or event.
- $t'$  is called the expanded time or **dilated time**. It is the time measured by a moving observer in another inertial frame of reference. Since  $\sqrt{1 - \left(\frac{v}{c}\right)^2}$  is always smaller than 1,  $t'$  is always greater than  $t$ .

## Example Problem

**Example 1a:** A rocket speeds past an asteroid at  $0.800c$ . If an observer in the rocket sees  $10.0\text{ s}$  pass on her watch, how long would that time interval be as seen by an observer on the asteroid?

## Example Problem

**Example 1a:** A rocket speeds past an asteroid at  $0.800c$ . If an observer in the rocket sees  $10.0\text{ s}$  pass on her watch, how long would that time interval be as seen by an observer on the asteroid?

**Example 1b:** A rocket speeds past an asteroid at  $0.800c$ . If an observer in the *asteroid* sees  $10.0\text{ s}$  pass on his watch, how long would that time interval be as seen by an observer on the *rocket*?

## Example Problem

**Example 1a:** A rocket speeds past an asteroid at  $0.800c$ . If an observer in the rocket sees  $10.0\text{ s}$  pass on her watch, how long would that time interval be as seen by an observer on the asteroid?

**Example 1b:** A rocket speeds past an asteroid at  $0.800c$ . If an observer in the *asteroid* sees  $10.0\text{ s}$  pass on his watch, how long would that time interval be as seen by an observer on the *rocket*?

How can that be?!

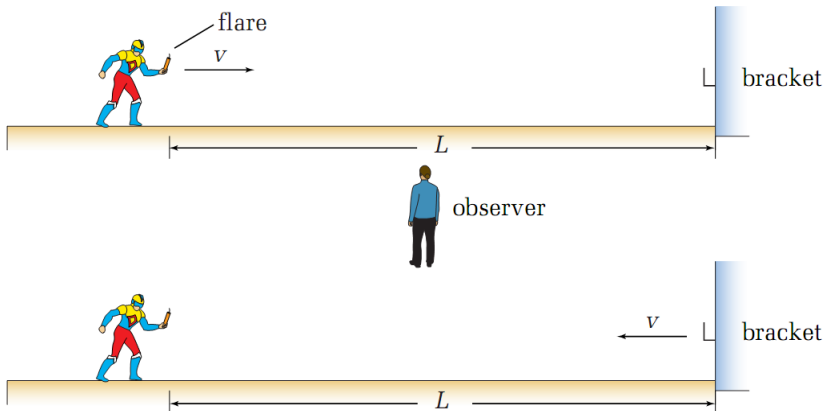
# Abandoning Concept of Absolute Space: Length Contraction

## Another Example

Captain Quick is a comic book hero who can run at nearly the speed of light. In his hand, he is carrying a flare with a lit fuse set to explode in  $1.5\text{ }\mu\text{s}$ . The flare must be placed into its bracket before this happens. The distance ( $L$ ) between the flare and the bracket is 402 m.

# Abandoning Concept of Absolute Space: Length Contraction

## Another Example



# Abandoning Concept of Absolute Space: Length Contraction

## Another Example

If Captain Quick runs at  $2.00 \times 10^8$  m/s, according to classical mechanics, he will not make it in time:

$$t = \frac{L}{v} = \frac{402 \text{ m}}{2.00 \times 10^8 \text{ m/s}} = 2.01 \times 10^{-6} \text{ s} = 2.01 \mu\text{s}$$

But according to relativistic mechanics, he makes it just in time. . .

# Abandoning Concept of Absolute Space: Length Contraction

## Another Example

To a stationary observer, the time on the flare is slowed:

$$t' = \frac{t}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} = \frac{1.5 \times 10^{-6} \text{ s}}{\sqrt{1 - \left(\frac{2}{3}\right)^2}} = \frac{1.5 \times 10^{-6} \text{ s}}{0.7454} = 2.01 \times 10^{-6} \text{ s}$$

The stationary observer sees a passage of time of  $t' = 2.01 \mu\text{s}$ , but Captain Quick, who is in the same reference frame as the flare, experiences a passage of time of  $t = 1.50 \mu\text{s}$ , precisely the time for the flare to explode.



# Abandoning Concept of Absolute Space: Length Contraction

## Another Example

- So, if Captain Quick sees only  $t = 1.50 \mu\text{s}$ , then how far did he travel?
- He isn't travelling any faster, so the only other possibility is that **the distance actually got shorter** (in his frame of reference).
- How much did the distance contract?

$$L' = L \sqrt{1 - \left(\frac{v}{c}\right)^2}$$

For this example:

$$L' = L \sqrt{1 - \left(\frac{v}{c}\right)^2} = (402 \text{ m}) \cdot \sqrt{1 - \left(\frac{2}{3}\right)^2} = 300 \text{ m}$$

# Lorentz Factor

The **Lorentz factor**  $\gamma$  is a short-hand for writing length contraction, time dilation and relativistic mass:

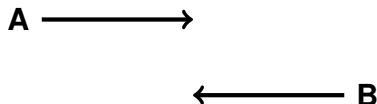
$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

Then time dilation and length contraction can be written simply as:

$$t' = \gamma t$$

$$L' = \frac{L}{\gamma}$$

## Let's Summarize



If Person A and Person B are moving at constant speed with respect to one another (doesn't matter if they're moving towards, or away from each other)

- They cannot agree whether any events happens at the same time or not
- Each sees the other's clock running slow
- Each sees the other “contracted” in length along the direction of motion

## Example Problem

**Example 2:** A spacecraft passes Earth at a speed of  $2.00 \times 10^8$  m/s. If observers on Earth measure the length of the spacecraft to be 554 m, how long would it be according to its passengers?

## Relativistic Momentum

In Physics 12, you were taught that momentum is mass times velocity. And back in Physics 11, you were taught that velocity is displacement over time. **These definitions have not changed.**

$$\mathbf{p} = m \frac{d\mathbf{x}}{dt}$$

But now that you know  $d\mathbf{x}$  and  $dt$  depend on motion, we can find the “relativistic momentum”:

$$\mathbf{p} = m \frac{d\mathbf{x}}{dt} = \frac{m d\mathbf{x}}{\sqrt{1 - \left(\frac{v}{c}\right)^2} dt} = \frac{m \mathbf{v}}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

# Relativistic Mass

From the relativistic momentum expression, we can see that there is a relativistic aspect to mass as well. The apparent mass  $m'$  as measured by a moving observer is related to its rest mass by the Lorentz factor:

$$m' = \frac{m}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} = \gamma m$$

The real (intrinsic) mass has not increased, but a moving observer will note that the object behaves as if it is more massive. As  $v \rightarrow c$ ,  $m \rightarrow \infty$ !

## Force and Work

In Physics 12, you were taught that force is the rate of change of momentum with respect to time. **This definition has not changed.**

$$\mathbf{F} = \frac{d\mathbf{p}}{dt}$$

and that work is the integral of the dot product between force and displacement vectors:

$$W = \int \mathbf{F} \cdot d\mathbf{x} = \int \frac{d\mathbf{p}}{dt} \cdot d\mathbf{x}$$

Since we now have a relativistic expression for momentum, we substitute that new expression into the expression for force, and then integrate.

## Work and Energy

For 1D motion, we can rearrange the terms in the integral:

$$W = \int F dx = \int \frac{dp}{dt} dx = \int v dp$$

Assuming that both velocity and momentum are continuous in time. Since momentum is a function of both  $\gamma$  and  $v$ , we must apply the chain rule to find the infinitesimal change in momentum ( $dp$ ) with respect to  $d\gamma$  and  $dv$ :

$$p = \gamma m v \quad \rightarrow \quad dp = \gamma dv + v d\gamma$$

Substituting that into the integral, we have:

$$W = \int v dp = \int m v (\gamma dv + v d\gamma) = \int m (\gamma v dv + v^2 d\gamma)$$



# Work and Energy

We want to integrate with respect to  $\gamma$ , so we need to express  $v$  and  $dv$  in terms of  $\gamma$  using its definition:

$$v^2 = c^2 \left[ 1 - \left( \frac{1}{\gamma} \right)^2 \right] \quad dv = \frac{c^2}{\gamma^3 v} d\gamma$$

# Work and Energy

Putting everything together, we have

$$W = \int m(\gamma v dv + v^2 d\gamma) = \int m \left[ \frac{c^2}{\gamma^2} + c^2 \left( 1 - \frac{1}{\gamma^2} \right) \right] d\gamma$$

We end up with a surprisingly simple integral:

$$W = \int_1^\gamma mc^2 d\gamma$$

The limit of the integral is from 1 because at  $v = 0$ ,  $\gamma = 1$

## Work and Kinetic Energy

The integral gives us this expression:

$$W = \gamma mc^2 - mc^2 = K$$

We know from the work-kinetic energy theorem that the work  $W$  done is equal to the change in kinetic energy  $K$ , therefore

$$K = m'c^2 - mc^2$$

Variable	Symbol	SI Unit
Kinetic energy of an object	$K$	J
Relativistic mass (measured in moving frame)	$m'$	kg
Rest mass (measured in stationary frame)	$m$	kg
Speed of light	$c_0$	m/s

# Relativistic Energy

## What This All Means

$$K = m'c^2 - mc^2$$

The minimal energy that any object has, regardless of it's motion (or lack of) is its **rest energy**:

$$E_0 = mc^2$$

The **total energy** of an object has is

$$E_T = m'c^2 = \gamma mc^2$$

The difference between total energy and rest energy is the kinetic energy:

$$K = E_T - E_0$$

# Relativistic Energy

## What This All Means

$$E = mc^2$$

### Mass-energy equivalence:

- Whenever there is a change of energy, there is also a change of mass
- “Conservation of mass” and “conservation of energy” must be combined into a single concept of **conservation of mass-energy**
- Mass-energy equivalence doesn't merely mean that mass can be converted into energy, and vice versa (although this is true), but rather, one can be converted into the other **because they are fundamentally the same thing**

## Example Problem

**Example 3:** An electron has a rest mass of  $9.11 \times 10^{-31}$  kg. In a detector, it behaves as if it has a mass of  $12.55 \times 10^{-31}$  kg. How fast is that electron moving relative to the detector?

# Kinetic Energy—Classical vs. Relativistic

**Relativistic:**

$$K = \frac{mc^2}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} - mc^2$$

**Newtonian:**

$$K = \frac{1}{2}mv^2$$

But are they really that different?

- If space and time are indeed relative quantities, then the relativistic equation for  $K$  must apply to all velocities
- But we know that when  $v \ll c$ , the Newtonian expression works perfectly
- i.e. The Newtonian expression for  $K$  must be a very good approximation for the relativistic expression for  $K$  for  $v \ll c$

# Binomial Series Expansion

The **binomial series** is the Maclaurin series for the function  $f(x) = (1 + x)^\alpha$ , given by:

$$(1 + x)^\alpha = \sum_{k=0}^{\infty} \binom{\alpha}{k} x^k = 1 + \alpha x + \frac{\alpha(\alpha - 1)}{2!} x^2 + \dots$$

In the case of relativistic kinetic energy, we use:

$$x = - \left( \frac{v}{c} \right)^2 \quad \text{and} \quad \alpha = -\frac{1}{2}$$



## Binomial Series Expansion

Substituting these terms into the equation:

$$K = mc^2 \left( 1 + \frac{1}{2} \frac{v^2}{c^2} + \frac{3}{8} \frac{v^4}{c^4} + \dots \right) - mc^2$$
$$\approx \frac{1}{2} mv^2 + \frac{3}{4} m \frac{v^4}{c^2} + \dots$$

For  $v \ll c$ , we can ignore the high-order terms. The leading term reduces to the Newtonian expression

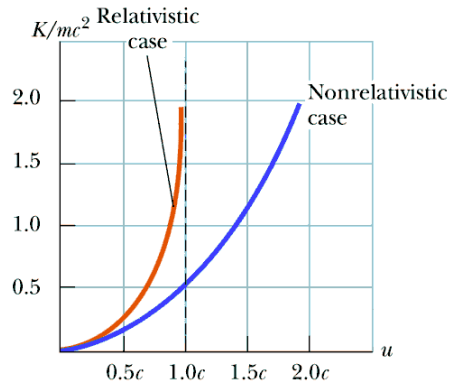
# Comparing Classical and Relativistic Energy

In classical mechanics:

$$K = \frac{1}{2}mv^2$$

In relativistic mechanics:

$$K = \gamma mc^2 - mc^2$$



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The classical expression is accurate for speeds up to  $v \approx 0.3c$ .

## Example Problem

**Example 4:** A rocket car with a mass of  $2.00 \times 10^3$  kg is accelerated from rest to  $1.00 \times 10^8$  m/s. Calculate its kinetic energy:

1. Using the classical equation
2. Using the relativistic equation