

Topic 23: Special Relativity

Advanced Placement Physics

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Olympiads School

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Introduction

The slides on **special relativity** is a condensed version of the slides used for Physics 12 (and with some calculus). For many of you, this is a review. There are 2 versions of the slides that are downloadable from the school website:

- The long version
 - More background information and derivations and integrations
 - 23a-relativity_long.pdf
- The short version
 - More “to the point”
 - The version that I am using in this class
 - 23a-relativity_short.pdf

There is also a handout on how to solve and interpret the time dilation example problem.

Newtonian (Classical) Relativity

In Newtonian physics, space and time are *absolute*:

- 1 m is 1 m no matter where you are in the universe
- 1 s is 1 s no matter where you are in the universe
- Measurements of space and time do not depend on motion

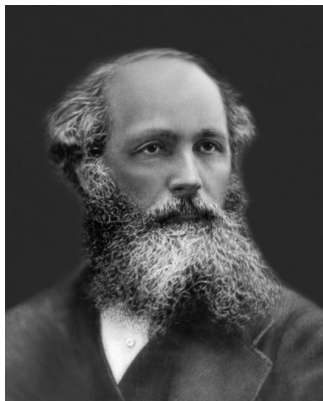
If space and time are absolute, then *all* velocities are relative

- Measured velocities depend on the motion of the observer

The Principle of Relativity

All laws of physics must apply equally in all inertial frames of reference.

New Physics: Maxwell's Equations



James Clerk Maxwell

- Classical laws of electrodynamics
- Published in 1861 and 1862
- Explains the relationship between
 - Electricity
 - Electric Circuits
 - Magnetism
 - Optics
- Previously these disciplines are thought to be separate and not related

Maxwell's Equations in a Vacuum

Everything Comes Back to This

$$\nabla \cdot \mathbf{E} = 0$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

$$\nabla \times \mathbf{B} = \mu_0 \epsilon_0 \frac{\partial \mathbf{E}}{\partial t}$$

Disturbances in \mathbf{E} and \mathbf{B} travel as an “electromagnetic wave”, with a speed:

$$c = \frac{1}{\sqrt{\epsilon_0 \mu_0}} = 299\,792\,458 \text{ m/s}$$

Peculiar features of Maxwell's equation

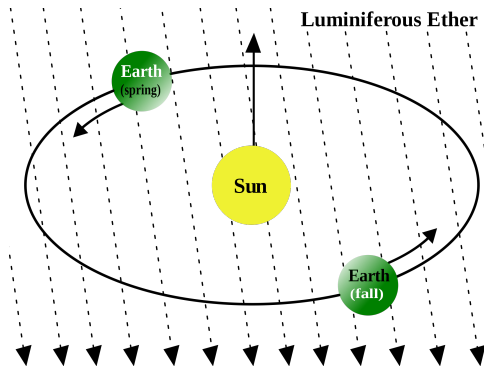
- Makes no mention of the *medium* in which EM waves travels
- When applying *Galilean transformation* (classical equation for calculating *relative velocity*) to Maxwell's equations, asymmetry is introduced
- Gauss's law for magnetism break down: magnetic field lines appear to have beginnings/ends
- In *some* inertial frames of reference, Maxwell's equations are simple and elegant, but in another inertial frame of reference, they are ugly and complex
- Physicists at the time theorized that—perhaps—there is/are actually *preferred* inertial frame(s) of references
- This violate the *principle of relativity*

The Illusive Aether

- Maxwell's hypothesis: the speed of light c_0 is relative to a hypothetical "luminiferous aether"
- In order for this "aether" (or "ether") to exist, it must have some fantastic (as in, a fantasy, too good to be true!) properties:
 - *All* space is filled with aether
 - Massless
 - Zero viscosity
 - Non-dispersive
 - Incompressible
 - Continuous at a very small (sub-atomic) scale

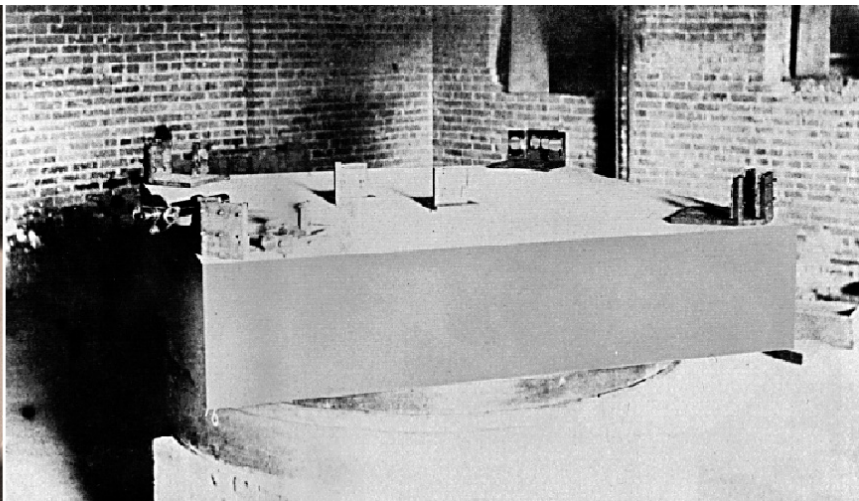
The Michelson-Morley Experiment

If ether exists, then at different times of the year, the Earth will have a different relative velocity with respect to it:



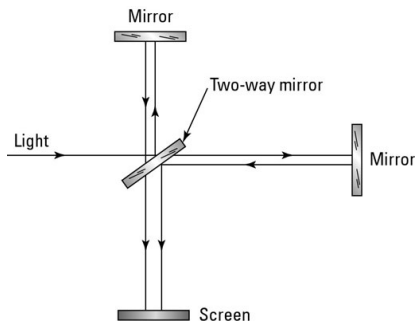
And it will cause light to either speed up, or slow down.

The Michelson Interferometer



The experiment is ingenious but very difficult. . .

The Michelson Interferometer



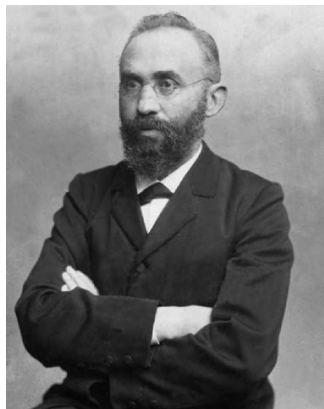
- A beam of light is split into two using a two-way (half-silvered) mirror
- The two beams are reflected off mirrors and finally arriving at the screen where interference patterns are observed
- The two paths are the same length, so if the *speed* of the light changes, we should see an interference pattern
- **Except none were ever found!**

What To Do with “Null Result”

The Michelson-Morley experiment failed to detect the illusive ether, even after many refinements. What does this mean?

- Majority view
 - **The experiment was flawed!**
 - Keep improving the experiment (or design a better experiment) and the ether will eventually be found
- Minority view:
 - **The hypothesis is wrong!**
 - The experiment showed it for what it is: ether cannot be found
- A few physicists: There must be **another explanation** that saves both experiment and theory

Hendrik Lorentz



Hendrik Antoon Lorentz

- Considered the Michelson-Morley experiment to be significant
- Objects travelling in the direction of ether contracts in length, nullifying the experimental results
- Lorentz Factor:

$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

- *No known physical phenomenon can cause anything to contract*
- Lorentz was on to something, but his thinking was wrong

Making Maxwell's Equations Work

Albert Einstein in 1905, Age 26



Albert Einstein

- Einstein believed in the principle of relativity, and therefore rejected the concept of a preferred frame of reference
- The failure of the Michelson-Morley experiment to find the flow of ether proves that it does not exist
- In order to make the equations to work again, Einstein revisited two most fundamental concepts in physics: space and time

Special Relativity

- Largely ignored by most physicists at first, until Max Planck took an interest in it
- Soon adopted by many physicists
- “Special” relativity because it describes a “special case” without effects of forces (e.g. gravity) & acceleration
- Later published theory of “general relativity” (much more complicated)

Postulates

The Principle of Relativity

All laws of physics must apply equally in all inertial frames of reference.

The Principle of Invariant Light Speed

As measured in any inertial frame of reference, light is always propagated in empty space with a definite velocity c that is independent of the state of motion of the emitting body.

Published in the journal *Annalen der Physik* on September 26, 1905 in the article *On the Electrodynamics of Moving Bodies* when Einstein was 26 years old working as a patent clerk in Switzerland

What's so Special About Special Relativity?

Classical (Newtonian) relativity:

- Space and time are absolute, therefore
- Speed of light must be relative to the observer

Einstein's special relativity:

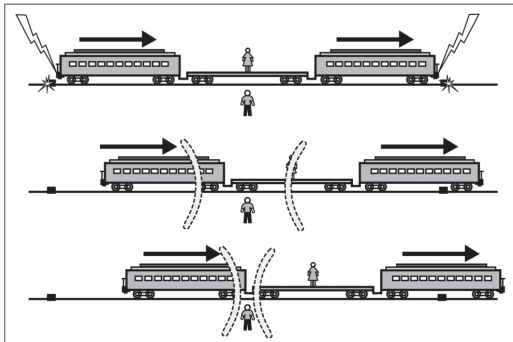
- Speed of light is absolute, therefore
- space and time must be relative to the observer

We must modify our traditional concepts:

- Measurement of space (our ruler in the x -, y - and z -directions)
- Measurement of time (our clock)
- Concept of simultaneity (whether or not two events happens at the same time)

Simultaneity: Thought Experiment

Lightning bolt strikes the ends of a moving train



- The man on the ground sees the lightning bolt striking at the same time
- The woman on the moving train sees the lightning bolt on the right first

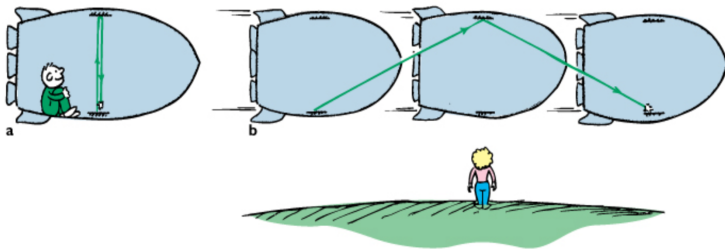
Simultaneity: Thought Experiment

- The two observers disagree on the result, but
 - Neither person is wrong
 - Neither person is misinformed
- Both observers are valid *inertial* frames of reference
- This means that simultaneity depends on your motion

Events that are simultaneous in one inertial frame of reference are not simultaneous in another.

Abandoning Concept of Absolute Time: Time Dilation

A “thought experiment”

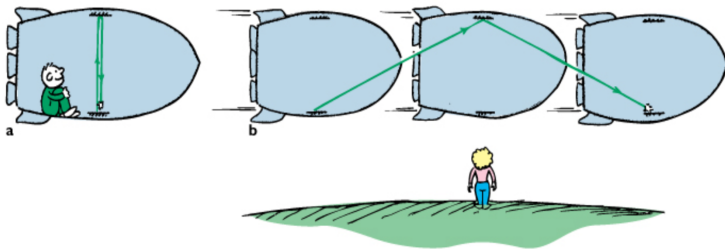


I'm on a spaceship travelling in deep space, and I shine a light from A to B . The distance between A and B is really just:

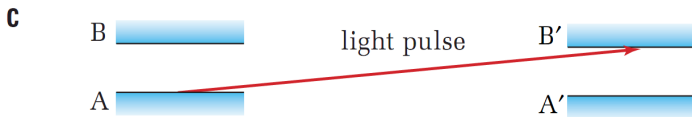
$$|AB| = c\Delta t_0$$

I know the speed of light c , and I know how long it took for the light pulse to reach B . (The reason I used Δt_0 will be obvious later.)

Time Dilation: A Thought Experiment

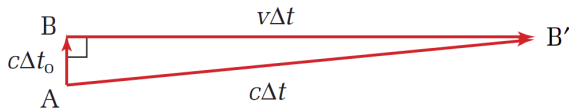


You are in space station watching my spaceship go past you at speed v . You would see that same beam of light travel from A to B' instead.



Time Dilation: A Thought Experiment

D



$$c^2 \Delta t^2 = v^2 \Delta t^2 + c^2 \Delta t_0^2$$

$$(c^2 - v^2) \Delta t^2 = c^2 \Delta t_0^2$$

$$\left(1 - \frac{v^2}{c^2}\right) \Delta t^2 = \Delta t_0^2$$

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

Time Dilation: A Thought Experiment

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

- Δt_0 is called the **proper time**. It is the time measured by a person at rest relative to the object or event.
- Δt is called the expanded time or **dilated time**. It is the time measured by a moving observer in another inertial frame of reference. Since $\sqrt{1 - \left(\frac{v}{c}\right)^2}$ is always smaller than 1, Δt is always greater than Δt_0 .

Example Problem

Example 1a: A rocket speeds past an asteroid at $0.800c$. If an observer in the rocket sees 10.0 s pass on her watch, how long would that time interval be as seen by an observer on the asteroid?

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Example 1b: A rocket speeds past an asteroid at $0.800c$. If an observer in the *asteroid* sees 10.0 s pass on his watch, how long would that time interval be as seen by an observer on the *rocket*?

Example Problem

Example 1a: A rocket speeds past an asteroid at $0.800c$. If an observer in the rocket sees 10.0 s pass on her watch, how long would that time interval be as seen by an observer on the asteroid?

Example 1b: A rocket speeds past an asteroid at $0.800c$. If an observer in the *asteroid* sees 10.0 s pass on his watch, how long would that time interval be as seen by an observer on the *rocket*?

How can that be?!

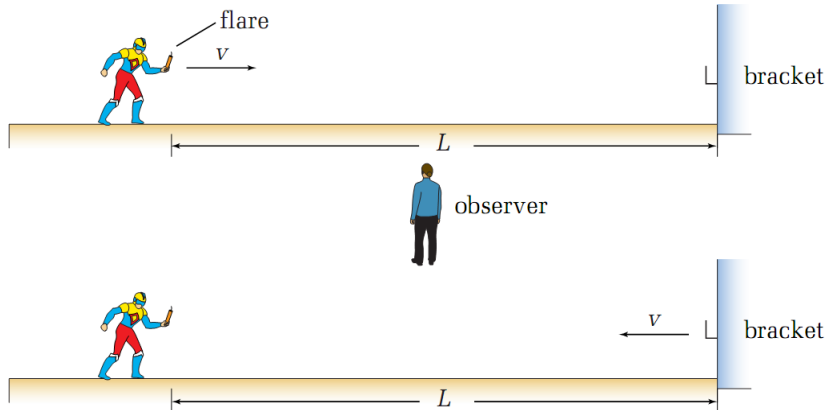
Abandoning Concept of Absolute Space: Length Contraction

Another Example

Captain Quick is a comic book hero who can run at nearly the speed of light. In his hand, he is carrying a flare with a lit fuse set to explode in $1.5\text{ }\mu\text{s}$. The flare must be placed into its bracket before this happens. The distance (L) between the flare and the bracket is 402 m.

Abandoning Concept of Absolute Space: Length Contraction

Another Example



Abandoning Concept of Absolute Space: Length Contraction

Another Example

If Captain Quick runs at $2.00 \times 10^8 \text{ m/s}$, according to classical mechanics, he will not make it in time:

$$\Delta t = \frac{L}{v} = \frac{402 \text{ m}}{2.00 \times 10^8 \text{ m/s}} = 2.01 \times 10^{-6} \text{ s} = 2.01 \mu\text{s}$$

But according to relativistic mechanics, he makes it just in time...

Abandoning Concept of Absolute Space: Length Contraction

Another Example

To a stationary observer, the time on the flare is slowed:

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} = \frac{1.5 \times 10^{-6} \text{ s}}{\sqrt{1 - \left(\frac{2}{3}\right)^2}} = \frac{1.5 \times 10^{-6} \text{ s}}{0.7454} = 2.01 \times 10^{-6} \text{ s}$$

The stationary observer sees a passage of time of $\Delta t = 2.01 \mu\text{s}$, but Captain Quick, who is in the same reference frame as the flare, experiences a passage of time of $\Delta t_0 = 1.50 \mu\text{s}$, precisely the time for the flare to explode.

Abandoning Concept of Absolute Space: Length Contraction

Another Example

- So, if Captain Quick sees only $\Delta t_0 = 1.50 \mu\text{s}$, then how far did he travel?
- He isn't travelling any faster, so the only other possibility is that **the distance actually got shorter** (in his frame of reference).
- How much did the distance contract?

$$L = L_0 \sqrt{1 - \left(\frac{v}{c}\right)^2}$$

For this example:

$$L = L_0 \sqrt{1 - \left(\frac{v}{c}\right)^2} = (402 \text{ m}) \cdot \sqrt{1 - \left(\frac{2}{3}\right)^2} = 300 \text{ m}$$

Lorentz Factor

The **Lorentz factor** γ is a short-hand for writing length contraction, time dilation and relativistic mass:

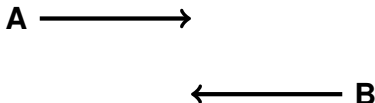
$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

Then time dilation and length contraction can be written simply as:

$$\Delta t = \gamma \Delta t_o$$

$$L = \frac{L_o}{\gamma}$$

Let's Summarize



If Person A and Person B are moving at constant speed with respect to one another (doesn't matter if they're moving towards, or away from each other)

- They cannot agree whether any events happens at the same time or not
- Each sees the other's clock running slow
- Each sees the other “contracted” in length along the direction of motion

Example Problem

Example 2: A spacecraft passes Earth at a speed of 2.00×10^8 m/s. If observers on Earth measure the length of the spacecraft to be 554 m, how long would it be according to its passengers?

Relativistic Momentum

In Physics 12, you were taught that momentum is mass times velocity. And back in Physics 11, you were taught that velocity is displacement over time:

$$\mathbf{p} = m \frac{d\mathbf{x}}{dt}$$

Now that you know $d\mathbf{x}$ and dt depend on motion, we can find the “relativistic momentum”:

$$\mathbf{p} = m \frac{d\mathbf{x}}{dt} = \frac{m d\mathbf{x}}{\sqrt{1 - \left(\frac{v}{c}\right)^2} dt} = \frac{m \mathbf{v}}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

Relativistic Mass

From the relativistic momentum expression, we can see that there is a relativistic aspect to mass as well. The apparent mass m' as measured by a moving observer is related to its rest mass by the Lorentz factor:

$$m' = \frac{m}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} = \gamma m$$

The real (intrinsic) mass has not increased, but a moving observer will note that the object behaves as if it is more massive. As $v \rightarrow c$, $m \rightarrow \infty$!

Force and Work

In Physics 12, you were taught that force is the rate of change of momentum with respect to time. This has not changed

$$\mathbf{F} = \frac{d\mathbf{p}}{dt}$$

and that work is the integral of the dot product between force and displacement vectors:

$$W = \int \mathbf{F} \cdot d\mathbf{x} = \int \frac{d\mathbf{p}}{dt} \cdot d\mathbf{x}$$

We can substitute the relativistic expression for momentum into the expression for force, and then integrate

Work and Energy

The integral is not particularly difficult. For a 1D motion with no acceleration, we can rearrange the terms in the integral:

$$W = \int F dx = \int \frac{dp}{dt} dx = \int v dp$$

Assuming that both velocity and momentum are continuous in time. We need to take the derivative of the relativistic momentum term:

$$p = \gamma m v \quad \rightarrow \quad dp = \gamma dv + v d\gamma$$

Substituting that into the integral, we have:

$$W = \int v dp = \int m v (\gamma dv + v d\gamma)$$

Work and Energy

The $d\gamma$ term requires some careful (but not very difficult) derivation:

$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} \quad \rightarrow \quad d\gamma = \frac{v/c^2}{(1 - v^2/c^2)^{3/2}} dv$$

We want to integrate with respect to γ , so we need to express v and dv in terms of γ using its definition:

$$v^2 = c^2 \left[1 - \left(\frac{1}{\gamma} \right)^2 \right] \quad dv = \frac{c^2}{\gamma^3 v} d\gamma$$

Work and Energy

Putting everything together, we have

$$W = \int mv(\gamma dv + v d\gamma) = \int m \left[\frac{c^2}{\gamma^2} + c^2 \left(1 - \frac{1}{\gamma^2} \right) \right] d\gamma$$

What we end up with is surprisingly simple:

$$W = \int_{\gamma=1}^{\gamma} mc^2 d\gamma$$

The limit of the integral is from 1 because at $v = 0$, $\gamma = 1$

Work and Kinetic Energy

The integral gives us this expression:

$$W = \gamma mc^2 - mc^2 = K$$

We know from the work-kinetic energy theorem that the work W done is equal to the change in kinetic energy K , therefore

$$K = m'c^2 - mc^2$$

Variable	Symbol	SI Unit
Kinetic energy of an object	K	J
Relativistic mass (measured in moving frame)	m'	kg
Rest mass (measured in stationary frame)	m	kg
Speed of light	c_0	m/s

Relativistic Energy

What This All Means

$$K = m'c^2 - mc^2$$

The minimal energy that any object has, regardless of it's motion (or lack of) is its **rest energy**:

$$E_0 = mc^2$$

The **total energy** of an object has is

$$E_T = m'c^2 = \gamma mc^2$$

The difference between total energy and rest energy is the kinetic energy:

$$K = E_T - E_0$$

Relativistic Energy

What This All Means

$$E = mc^2$$

Mass-energy equivalence:

- Whenever there is a change of energy, there is also a change of mass
- “Conservation of mass” and “conservation of energy” must be combined into a single concept of **conservation of mass-energy**
- Mass-energy equivalence doesn't merely mean that mass can be converted into energy, and vice versa (although this is true), but rather, one can be converted into the other **because they are fundamentally the same thing**

Example Problem

Example 3: An electron has a rest mass of 9.11×10^{-31} kg. In a detector, it behaves as if it has a mass of 12.55×10^{-31} kg. How fast is that electron moving relative to the detector?

Kinetic Energy–Classical vs. Relativistic

Relativistic:

$$K = \frac{mc^2}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} - mc^2$$

Newtonian:

$$K = \frac{1}{2}mv^2$$

But are they really that different?

- If space and time are indeed relative quantities, then the relativistic equation for K must apply to all velocities
- But we know that when $v \ll c$, the Newtonian expression works perfectly
- i.e. The Newtonian expression for K must be a very good approximation for the relativistic expression for K for $v \ll c$

Binomial Series Expansion

The **binomial series** is the Maclaurin series for the function $f(x) = (1 + x)^\alpha$, given by:

$$(1 + x)^\alpha = \sum_{k=0}^{\infty} \binom{\alpha}{k} x^k = 1 + \alpha x + \frac{\alpha(\alpha - 1)}{2!} x^2 + \dots$$

In the case of relativistic kinetic energy, we use:

$$x = - \left(\frac{v}{c} \right)^2 \quad \text{and} \quad \alpha = -\frac{1}{2}$$

Binomial Series Expansion

Substituting these terms into the equation:

$$\begin{aligned} K &= mc^2 \left(1 + \frac{1}{2} \frac{v^2}{c^2} + \frac{3}{8} \frac{v^4}{c^4} + \dots \right) - mc^2 \\ &\approx \frac{1}{2} mv^2 + \frac{3}{4} m \frac{v^4}{c^2} + \dots \end{aligned}$$

For $v \ll c$, we can ignore the high-order terms. The leading term reduces to the Newtonian expression

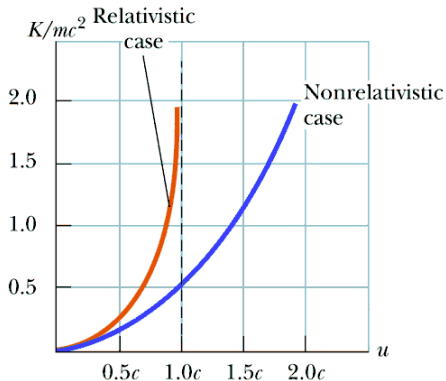
Comparing Classical and Relativistic Energy

In classical mechanics:

$$K = \frac{1}{2}mv^2$$

In relativistic mechanics:

$$K = \gamma mc^2 - mc^2$$



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The classical expression is accurate for speeds up to $v \approx 0.3c$.

Example Problem

Example 4: A rocket car with a mass of 2.00×10^3 kg is accelerated from rest to 1.00×10^8 m/s. Calculate its kinetic energy:

1. Using the classical equation
2. Using the relativistic equation